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ABSTRACT

We show that the Green's function solution of the Anderson model calculated by the functional derivative method developed earlier yields the density of states having the three-peak structure proposed by Grüner and Zawadowski and that the central peak indeed introduces the $\ln T$ dependence.

INTRODUCTION

It is generally accepted that the problem of localized moments in metals may be discussed more naturally by the Anderson model¹ than by the s-d model, although the Kondo effect² has been formulated solely by the s-d model. The Schrieffer-Wolff transformation³ can convert the Anderson model to the s-d model, but this is possible only in the limit where the density of states for d electrons is negligible at the Fermi surface. To circumvent the difficulties, Grüner and Zawadowski⁴ have proposed the density of states exhibiting a three-peak structure. At high temperatures where many body corrections are not strong enough to modify the Hartree-Fock results, two broad but well-separated peaks corresponding to the majority- and minority-spin electrons appears. The conduction electron scattering is then described with two phase shifts. At zero temperature, on the other hand, the system is in the singlet ground state with a well-defined single phase shift, introducing a narrow third peak. They have assumed that this narrow resonance is responsible for the Kondo effect. Zlatić et al⁵ have constructed a scattering amplitude by adding this narrow resonance but did not find the $\ln T$ dependence.

THE CALCULATION OF THE GREEN'S FUNCTION

In this paper, we shall show that the Green's function solution of the nondegenerate Anderson model calculated by the functional derivative method developed earlier⁶ yields the density of states having the three-peak structure and that the central peak indeed introduce the $\ln T$ dependence. Let us consider the equations of motion for the two types of Green's

functions $\Gamma_{k\lambda\sigma}^{(+)} \equiv \langle\langle C_{d\sigma}(t) N_{d\bar{\sigma}}^{(+)}(t) C_{\lambda\sigma}^{\dagger}(t') \rangle\rangle$, where $N_{d\sigma}^{(+)}(t) \equiv N_{d\sigma}(t) = C_{d\sigma}^{\dagger}(t) C_{d\sigma}(t)$, $N_{d\bar{\sigma}}^{(-)}(t) = 1 -$

$N_{d\bar{\sigma}}(t)$, and $\lambda = d$ or k . Here, in line with the Hubbard treatment⁷ of narrow bands, the projection operators $N_{d\sigma}^{(\pm)}$ are inserted so that the intra-atomic interaction U is correctly calculated whenever an electron with opposite spin $\bar{\sigma}$ appears at the impurity site. The one-electron Green's function is then given by $G = \Gamma^{(+)} + \Gamma^{(-)}$.

The equations of motion involve the following two types of terms

$$g_1 = \sum_k V_{dk} \langle\langle C_{k\sigma}(t) [N_{d\bar{\sigma}}^{(-)}(t) - \langle N_{d\bar{\sigma}} \rangle] C_{\lambda\sigma}^{\dagger}(t') \rangle\rangle, \quad (1)$$

$$g_2 = \sum_k V_{dk} \langle\langle C_{d\sigma}(t) [C_{d\bar{\sigma}}^{\dagger}(t) C_{k\bar{\sigma}}(t) - \langle C_{d\bar{\sigma}}^{\dagger} C_{k\bar{\sigma}} \rangle] C_{\lambda\sigma}^{\dagger}(t') \rangle\rangle - V_{kd} \langle\langle C_{d\sigma}(t) [C_{k\bar{\sigma}}^{\dagger}(t) C_{d\bar{\sigma}}(t) - \langle C_{k\bar{\sigma}}^{\dagger} C_{d\bar{\sigma}} \rangle] C_{\lambda\sigma}^{\dagger}(t') \rangle\rangle, \quad (2)$$

where V_{kd} is the hopping matrix element between conduction electron k and localized electron d . If one neglects g_1 and g_2 , the equations can be solved immediately, yielding G_0 . Since g_1 and g_2 can be reduced to functional derivatives of G , they can be calculated iteratively by inserting G_0 . The reason why, instead of the Hartree-Fock solution, G_0 is used as the zeroth order approximation is that the intra-atomic interaction U is included rigorously in each step of perturbation calculation, making our result distinctly different from any other existing calculation. The result is

$$2\pi G_{dd\sigma}^{-1}(\omega) = \frac{(\omega - \epsilon_d)(\omega - \epsilon_d - U) - iG_{dd\sigma}(\omega)Y(\omega)}{\omega - \epsilon_d - (1 - n_{d\bar{\sigma}})U} - \sum_k \frac{|V_{dk}|^2}{\omega - \epsilon_k} \frac{[\omega - \epsilon_d - (\frac{3}{4} - \frac{n_{d\bar{\sigma}}}{2})U]^2}{(\omega - \omega_a)(\omega - \omega_b) - iG_{dd\sigma}(\omega)Y(\omega)}, \quad (3)$$

where ϵ_k and ϵ_d are the energies of electrons k and d ; $n_{d\bar{\sigma}} = \langle n_{d\bar{\sigma}} \rangle$, and

$$Y(\omega) = (U/2)^2 M_{\bar{\sigma}}^{-1} [\omega - \epsilon_d - (1 - n_{d\bar{\sigma}})U] (\omega - \omega_x)^{-1} (\omega - \omega_x)^{-1},$$

$$\omega_{a,b} = \epsilon_d + (\frac{3}{4} - \frac{n_{d\bar{\sigma}}}{2})U \mp \sqrt{n_{d\bar{\sigma}}(1 - n_{d\bar{\sigma}})U},$$

$$\omega_{\pm x} = \epsilon_d + (\frac{3}{4} - \frac{n_{d\bar{\sigma}}}{2})U \pm [iG_{dd\sigma}(\omega)](U/4), \quad (4)$$

$$M_{\bar{\sigma}} = \frac{1}{2} \sum_k [V_{dk} \langle C_{d\bar{\sigma}}^{\dagger} C_{k\bar{\sigma}} \rangle + V_{kd} \langle C_{k\bar{\sigma}}^{\dagger} C_{d\bar{\sigma}} \rangle].$$

If $\sum_k |v_{dk}|^2 / (\omega - \epsilon_k)$ is replaced by ϵ_k , the structure of above expression becomes exactly the same as that of the Hubbard model obtained in Ref. 8. The value of $\sum_k |v_{dk}|^2 / (\omega - \epsilon_k)$ may be evaluated as $\Gamma - i\Delta$, where $\Gamma (< 0)$ and Δ are the effective energy shift and the width parameter of the d states, respectively. However, the term involving $iG_{dd\sigma}(\omega)Y(\omega)$ in the summation $\sum_k \dots$ should be handled with care, since $iG_{dd\sigma}(\omega) = [1 - f(\omega)]A(\omega)$ where $f(\omega)$ is the Fermi-Dirac distribution function and $A(\omega)$ the spectral function. In this particular term, $\sum_k |v_{dk}|^2 / (\omega - \epsilon_k)$ cannot be reduced to $\Gamma - i\Delta$, and instead the relation

$$\sum_k |v_{dk}|^2 f(\omega) / (\omega - \epsilon_k) = \kappa(\omega) - i\mathcal{K}(\omega) \quad , \quad (5)$$

should be used where $\kappa(\omega)$ is the Kondo integral responsible for the $\ln T$ dependence of the resistivity and $\mathcal{K}(\omega) = F(\omega)\Delta$. Then

$$2\pi G_{dd\sigma}^{-1}(\omega) = \frac{(\omega - \epsilon_d)(\epsilon - \epsilon_d - U) - [1 - f(\omega)]A(\omega)Y(\omega)}{\omega - \epsilon_d - (1 - n_{d\bar{\sigma}})U} - \frac{(\Gamma - i\Delta)[\omega - \epsilon_d - (\frac{3}{4} - \frac{n_{d\bar{\sigma}}}{2})U]^2}{(\omega - \omega_a)(\omega - \omega_b) - (1 - k + i\ell)A(\omega)Y(\omega) + \dots} \quad , \quad (6)$$

where $k = (\kappa\Gamma + \mathcal{K}\Delta) / (\Gamma^2 + \Delta^2)$ and $\ell = (\mathcal{K}\Gamma - \kappa\Delta) / (\Gamma^2 + \Delta^2)$.

THE RESULTS AND DISCUSSIONS

Since M_{σ} , Γ and Δ are all proportional to $|V^2|$, the Kondo term $(k - i\ell)$ is of order $|V^2|$ in $G_{dd\sigma}(\omega)$ and of order $|V^6|$ in the conduction electron Green's function $G_{k\sigma}(\omega)$ in agreement with the result of the Schrieffer-Wolff transformation. However, the $\ln T$ dependence will appear only if the density of d electrons, $A(\omega)$, is finite at the Fermi level. We shall show this by solving the polynomial equation $G_{dd\sigma}^{-1}(\omega) = 0$ graphically. Since the method is completely parallel to that used in the Hubbard model,⁸ we shall only outline the procedure and the results.

In the limit of $\Gamma = \Delta = 0$, the poles of G is calculated by solving the quartic equation in ω :

$$(\omega - \epsilon_d)(\omega - \epsilon_d - U) - [1 - f(\omega)]Y(\omega) = 0 \quad , \quad (7)$$

where $f(\omega) \approx 0$ since we are calculating the scattering amplitude of conduction electron k traveling right above the Fermi level; $G_{dd\sigma}(\omega)$ is a part of $G_{kk\sigma}(\omega)$ for this electron. The term $Y(\omega)$ is generated by the correction term g_2 and tends to introduce magnetism.

In fact, the solution calculated by Eq. (7) satisfies the condition for magnetism $(\delta n_{d\sigma} / \delta n_{d\sigma}) / \delta n_{d\sigma} < -1$. As long as k_d is small, however, the poles ω_1' and ω_2' will not deviate drastically from the isolated atomic levels ϵ_d and $\epsilon_d + U$ while the additional poles $\omega_{\pm x}'$ are physically unimportant since they are nearly equal to $\omega_{\pm x}$ and the spectral weights are vanishingly small.

The solutions of

$$(\omega - \omega_a)(\omega - \omega_b) - (1 - k + i\ell)A(\omega)Y(\omega) = 0 \quad (8)$$

can be calculated similarly. At high temperature where $\kappa(\omega)$ is small, $k \approx f(\omega)$ and $\ell \approx 0$. In the nonmagnetic limit $n_{d\sigma} = n_{d\sigma} \approx 1/2$, $\omega_a = \epsilon_d$ and $\omega_b = \epsilon_d + U$ and hence Eq. (8) becomes identical to Eq. (7), yielding $\omega_1' = \omega_1^a$ and $\omega_2' = \omega_2^a$. In the fully magnetic limit $n_{d\sigma} \approx 1$ and $n_{d\sigma} = 0$, on the other hand, $\omega_a = \omega_b = \epsilon_d + (3/4)U$. As the magnetization $n_{d\sigma} - n_{d\sigma}^a$ decreases, the value of ω decreases from $\epsilon_d + (3/4)U$ to ϵ_d , while the value ω_b increases from $\epsilon_d + (3/4)U$ to $\epsilon_d + U$. Note that the difference between ω_1' and ω_2' remains large as long as the difference $n_{d\sigma} - n_{d\sigma}^a$ is appreciable.

The complete solutions $\bar{\omega}_1$, $\bar{\omega}_2$ etc. of Eq. (6) can now be obtained graphically by assuming $\Delta = \ell = 0$. The contribution from the imaginary parts of the equation can then be estimated perturbationally. For the nonmagnetic case, however, the calculation is trivial since $\omega_1' = \omega_1^a$ and $\omega_2' = \omega_2^a$. $\bar{\omega}_1$ and $\bar{\omega}_2$ (or $\bar{\omega}_1$ and $\bar{\omega}_2$) will split into two and the separation, being proportional to $\sqrt{|\Gamma|}$, would be small as compared with the width parameter Δ of ω and the two peaks $\bar{\omega}_1$ and $\bar{\omega}_2$ (or $\bar{\omega}_1$ and $\bar{\omega}_2$) will look like a broad single peak. As the magnetization $n_{d\sigma} - n_{d\sigma}^a$ increases from zero to one, the difference $\omega_1' - \omega_2'$ increases, yielding two distinct solutions $\bar{\omega}_1$ and $\bar{\omega}_2$. Although $\bar{\omega}_1$ remains more or less constant, $\bar{\omega}_2$ increases from ϵ_d to $\epsilon_d + (3/4)U$. $\bar{\omega}_1$ and $\bar{\omega}_2$ will remain exhibiting a single peak.

As temperature is lowered, $\kappa(\omega)$ and hence $(1 - k)$ and ℓ increase, making ω_1' and ω_2' larger. The calculated value of $n_{d\sigma}$ will then decrease and, to maintain the self-consistent requirement $\delta n_{d\sigma} / \delta n_{d\sigma} < -1$, $n_{d\sigma}$ has to increase, resulting in the reduction of the magnetization $n_{d\sigma} - n_{d\sigma}^a$ as well as ω .

Let us now summarize the foregoing results. At high temperature where the Kondo integral is negligible, magnetic solution with the three-peak structure appears as long as the term $Y(\omega)$ which enhances magnetism dominates over the term which is proportional to $\Gamma - i\Delta$ and generated from the correction term g , and which tends to suppress magnetism. As long as

there exists a small density of d states at the Fermi level ϵ_F , the value of $(1-k)$ increases as T decreases, tending to make $\bar{\omega}$ larger. If the third peak $\bar{\omega}$ lies below ϵ_F , the peak $\bar{\omega}$ will be shifted ϵ_F , increasing the density of d states and enhancing the $\ln T$ dependence. As T is lowered further, the magnetization $n_{d\uparrow} - n_{d\downarrow}$ will decrease, making $\bar{\omega}$ smaller. Consequently $\bar{\omega}$ will not increase and instead remain constant near ϵ_F . If, on the other hand, the third peak $\bar{\omega}$ appears above the Fermi level, the tendency to increase $\bar{\omega}$ again introduces two competing procedures. One is the tendency to reduce the density of d states at ϵ_F and hence the tendency to increase $\bar{\omega}$ will be suppressed greatly. The second effect is to reduce the magnetization $n_{d\uparrow} - n_{d\downarrow}$, making $\bar{\omega}$ and hence $\bar{\omega}$ smaller. If $\bar{\omega}$ becomes smaller and approaches ϵ_F , the density of d states at ϵ_F increases, enhancing the $\ln T$ dependence in $(1-k)$. Hence the third peak $\bar{\omega}$ will be shifted towards ϵ_F and remains there.

As the magnetization $n_{d\uparrow} - n_{d\downarrow}$ becomes small, however, $\bar{\omega}$ becomes nearly equal to ω_1 and will not decrease further. Then the mechanism to keep $\bar{\omega}$ around ϵ_F will disappear and $\bar{\omega}$ starts to increase, thus reducing the density of d states at ϵ_F and hence reducing the $\ln T$ dependence. The system will then behave like spin fluctuations. This is possible because the imaginary parts of Eq. (6) are proportional to $[\omega - \epsilon_d - (3/4 - n_{d\uparrow}/2)U]^2$ or $[\omega - \epsilon_d - (1 - n_{d\downarrow})U]$, and hence the width of the third peak remains narrow.

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