

RE-ENTRY FLIGHT DEMONSTRATION NO. 1 (RFD-1): OPTICAL DATA AND FUEL-ELEMENT EXPERIMENT

Department 7410

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ABSTRACT

This report on the RFD-1 optical data and external fuel-element experiment includes a description of the instruments and test components used, a presentation of the data obtained, an explanation of the methods of data reduction employed, and a statement of the conclusions derived. It covers the theory, design, quaifiction tests, flight-test data, and results of the external fuel-element expertment. Also presented is a theoretical analysis of observed versus predicted abletion times and altitudes for the external fuel elements. In addition, this report presents recommendations for improvements to data acquisition and reduction methods in future, similar flight tests.

## ACKNOWLEDGMENT

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## Department Number

1110
7220
7240
7620

Department Name
Materials and Process Department I
Test Range Department
Test Support Department
Programming Department

Outside agencies who have contributed to the work presented in this report are acknowledged by specific references to their support throughout the text.

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Nuclear power supplies designed for use in space vehicles must incorporate safety features to preclude significant radiation hazards to the earth's population during orbital-decay re-entries or in the event of aborted missions. Sandia Corporation, a prime nuclear-weapon contractor to the AEC, has been authorized by the AEC Division of Reactor Development to act as the prime contractor for the independent safety evaluation of aerospace nuclear power systems. The Aerospace Nuclear Safety Program at Sandia includes research and development studies, ground testing, flight testing, system analysis, and independent safety assessment.

Re-entry Flight Demonstration Number One (RFD-1) was the first re-entry flight test conducted under this program since its inception in March 1962. The SNAP-10A (systems for nuclear auxiliary power) reactor, designed and constructed by Atomics International (AI), was selected for RFD-1 because of its proposed early use as a nuclear auxiliary power supply for earth satellites. An inert version of the SNAP-10A was flown on RFD-1 to determine the effectiveness of the safety design. The simulated reactor was mounted on a re-entry vehicle (RV) which was placed into the required trajectory by a four-stage Scout missile launched from the National Aeronautics and Space Administration (NASA) Wallops Station, Wallops, Island, Va.

Sandia carried out its assignment in the Aerospace Nuclear Safety Program by performing the following tasks:

1. Design of the flight-test experiment and the configuration of the simulated test reactor (STR), in cooperation with AI.
2. Study of the capabilities of the Scout launch vehicle, and recommendation of a trajectory which would assure that the desired information would be obtained.
3. Design of the RV and the telemetry (TM) system, and coordination of interface problems with AI, NASA, and Ling-Temco-Vought Corporation.
4. Theoretical predictions of flight-test outcomes.
5. Preparation of documents on support requirements for the Atlantic Missile Range (AMR) and the NASA Wallops and Bermuda stations.
6. Provision of complementary downrange instrumentation for the collection of TM and optical data.
7. Management of flight-implementation activities.
8. Data reduction and analysis.
9. Comparison of flight-test results with theoretical calculations.

The following Sandia Corporation reports, together with the present volume, comprise the final documentation of RFD-1:

| Report No. | Title |
| :---: | :---: |
| SC-RR-64-501 | Re-entry Flight Demonstration Number One (RFD-1): Final Flight-Test Plan |
| SC-RR-64-502 | Re-entry Flight Demonstration Number One (RFD-1): Data Book |
| SC-RR-64-510 | Re-entry Flight Demonstration Number One (RFD-1): Comparison of the Preflight and observed Trajectories |
| SC-RR-64-511 | Re-entry Flight Demonstration Number One (RFD-1): Design, Development, and Performance of the Reentry Vehicle |
| SC-RR-64-515 | Re-entry Flight Demonstration Number One (RFD-1): Preflight Disassembly Analysis and Observed Disassembly of the Simulated SNAP-10A Reactor |
| SC-RR-64-517 | Re-entry Flight Demonstration Number One (RFD-1): Atmospheric Sciences Support |

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## LIST OF ABBREVIATIONS

RFD-1 Re-entry Flight Demonstration Number One
AI Atomics International
SNAP Systems for Nuclear Auxiliary Power
RV Re-entry Vehicle
NASA National Aeronautics and Space Administration
STR Simulated Test Reactor
TM
Telemetry
AMR Atlantic Missile Range
UZrH Uranium-Zirconium Hydride
AFSWC Air Force Special Weapons Command
TACODA Target Coordinate Data
TRSS Time-Resolved Streak Spectrograph
BRT Bermuda Range ..... Time
$W / C_{D} A \quad$ Ballistic Coefficient
RS Re-entry System

# RE-ENTRY FLIGHT DEMONSTRATION NO. 1 (RFD-1): <br> OPTICAL DATA AND FUEL-ELEMENT EXPERIMENT 

## SECTION I -- INTRODUCTION AND SUMMARY

## Introduction

SNAP-10A is a 500 -watt electrical power supply designed for use in space vehicles. Heat to operate the power supply is provided by a $30-\mathrm{kw}$ nuclear reactor. After normal operation of the SNAP-10A in space, the reactor fuel elements will contain long-lived fission products. These products could pose a radiation hazard if a space vehicle carrying a SNAP-10A reactor were to re-enter the earth's atmosphere before they had decayed to a relatively harmless level. Since the orbital life of a space vehicle cannot be predicted with surety before launch, the possibility of re-entry at any time after launch must be considered. It is therefore necessary to ensure that the reactor fuel elements will be disposed of in such a manner that no radiation hazard will result. One method proposed to ensure safe disposal is to use re-entry heating to disassemble the reactor and burn up the fuel elements thus exposed; atmospheric diffusion and high-altitude winds would then disperse the resulting debris widely enough so that no radiation hazard would be posed.

In May 1963, Sandia conducted the RFD-1 flight test to investigate the effectiveness of this method. A flight test was necessary because it is not possible to exactly simulate the orbital-decay re-entry environment in laboratory experiments. The main objectives of RFD-1 were to evaluate reactor disassembly and fuel-element burnup (particle size and dispersion are being investigated separately). Other objectives included the confirmation of analytical techniques for predicting re-entry phenomena, and the development of optical and electronic instrumentation for use in future flight tests.

In the fuel-element experiment, four groups of simulated fuel elements, or rods, were simultaneously ejected from the RFD-1 RV early in re-entry. Each rod consisted essentially of a core of tracer material surrounded by a uranium-zirconium-hydride (UZrH) wall, with the UZrH in turn covered by a 15-mil cladding of Hastelloy N. During re-entry, the times and altitudes of rod burnup were optically recorded, using spectrographic analysis to identify each group of rods, and the results were compared with the computed heat inputs for the RFD-1 trajectory.

This report presents the conclusions drawn concerning the data recorded and the probable fate of the groups of fuel rods. It also presents recommendations, drawn on the basis of these conclusions, for modifications and refinements to existing instrumentation and to present methods of data acquisition and reduction for application in future flight tests.

## Summary

Re-entry of the $R V$, the reactor, and the simulated tracer-loaded fuel rods was photographed with 61 cameras, including plate, spectral, and framing types. Twenty cameras were located at High Point, Bermuda, and forty-one were airborne in aircraft provided by NASA and the Air Force Special Weapons Command (AFSWC).

Many of the films were underexposed because of the distances involved and the effects of atmospheric attenuation. However, much information was extracted from the films and other records obtained, permitting valid conclusions to be drawn about the re-entry events and the degree to which the test objectives were achieved. Data from the plate, spectral, and framing cameras were reduced by individual type, and all the results were then correlated to confirm agreement of the conclusions.

The re-entry was first visible 325.40 seconds after launch, when the RV was at an altitude of 266,000 feet. Melting and disassembly of the reactor were evident on all the films which viewed its re-entry. Information on these observations was used in the analysis of reactor disassembly (see SC-RR-64-515).

A11 of the rods for the fuel-element experiment were 1.25 inches in diameter and 12.25 inches long. The tracer materials and UZrH wall thicknesses for each group of rods were as follows:

| Group | UZrH Wa11 <br> Thickness (in.) |  |
| :---: | :---: | :--- |
|  | 0.101 | Tracer |
| 2 | 0.184 | Strontium |
| 3 | 0.306 | Barium |
| 4 | 0.406 | Silver |
| 4 |  | Gold |

The strontium-loaded rods were visible from 342.27 seconds after launch ( 220,100 feet) to 363.32 seconds ( 161,400 feet). It is believed that they were completely consumed at this altitude. Theoretical predictions based on laboratory data indicated that, for the heat inputs from the RFD-1 trajectory, complete burnup should have occurred at 364.30 seconds ( 158,000 feet). This close agreement was within the accuracy expected of the RFD-1 test.

The barium-loaded rods were visible from 348.00 seconds (205, 200 feet) to 371.50 seconds ( 142,000 feet), at which time they are believed to have been completely consumed. Again, theoretical predictions agreed closely with the flighttest results; predicted burnup was at 374.60 seconds ( 135,000 feet).

The trails of the silver- and gold-loaded rods were identified, but no evidence of the silver or gold tracers was found on any of the films. Presumably, neither of these rod groups ablated enough to expose the tracers. This was as predicted in both the preflight and postflight analyses of the test.

SC-RR-64-501 describes the complete flight-test plan as it was drawn up before the test. The discussion in this section will also cover the flight-test plan, but it will be more general in the areas of the RV, the reactor, and the trajectories, and more specific in the areas of the optical instrumentation, the optical data, and the fuel-element experiment.

## SNAP-10A System

Figure 1 shows the SNAP-10A. A comprehensive discussion of the system appears in SC-RR-64-515; the principal system components are:

1. A core vessel housing 37 UZrH fuel elements which provide heat to operate the system.
2. Four beryllium reflectors to control the nuclear reaction.
3. A conical radiator which houses the thermoelectric conversion elements and also houses the tubes which carry a eutectic mixture of sodium and potassium ( $\mathrm{NaK}-78$ ) used to transfer heat from the reactor core to the thermoelectric junctions.
4. A pump to circulate the NaK-78 through the system.


Figure 1. SNAP-10A reactor

The proposed design of the SNAP-10A includes features to promote disassembly of the reactor under re-entry heating, and consequent exposure of the rods. The steps in disassembly (Figure 2) are:

1. Melting of fusible links in a band which holds the reflectors in place on the reactor, and subsequent separation of the band.
2. Spring ejection of the reflectors.
3. Burn-off of the NaK pump.
4. Melting of the lip weld which holds the lid of the reactor core vessel in place, and subsequent separation of the lid.
5. Expulsion of the fuel elements from the core vessel through the effects of aerodynamic drag forces on the reactor.


Figure 2. SNAP-10A desired orbital-decay re-entry sequence

If this sequence of events results in exposure of the rods at an altitude sufficiently high, the rods will be subjected to considerable aerodynamic heating. Assuming the requisite heat input, the rods will melt, ablate into small particles, and disperse in the atmosphere over a wide area. If adequate dilution of the particles containing fission products is achieved, their potential for offering a radiation hazard will be eliminated.

Test Objectives

## Reactor Disassembly

One objective of RFD-1 was to provide information on the disassembly sequence of the SNAP-10A so that the current design could be evaluated. Because of the payload limitation of the Scout launch vehicle, some deviations from the actual SNAP-10A design were required in the STR to reduce its weight; these included the substitution of lighter, aluminum reflectors for the beryllium reflectors actually used in the SNAP-10A, and omission of the fuel elements from inside the reactor core vessel, as
well as other, minor changes. It is not expected that these deviations affected the results of the experiment appreciably. Reactor disassembly is covered comprehensively in SC-RR-64-515.

## Fuel-Element Ablation

Knowledge of the response of UZrH to an environment of re-entry into the atmosphere is very limited. Calculations indicate that, in an orbital-decay reentry, aerodynamic heating alone may be insufficient to assure complete ablation of the fuel elements. However, the available heat input could be augmented by the exothermic chemical reactions of UZrH with the re-entry atmosphere. Conversely, if a hard oxide layer is formed on the surface of UZrH fuel elements, as it frequently does on specimens heat-tested in the laboratory, this phenomenon may retard ablation. Since exact simulation of the re-entry environment is not possible in laboratory experiments, it was decided to investigate the ablation of UZrH fuel elements on RFD-1.

The objectives of the fuel-element experiment were:

1. To obtain data, for UZrH in general and the SNAP-10A fuel rods in particular, on ablation rate and on the volume consumed by burnup during re-entry.
2. To confirm theoretical calculations and laboratory data pertaining to re-entry ablation.
3. To evaluate optical instrumentation and data-reduction techniques for use in future flight tests.

Complete ablation of the re-entering fuel elements is not absolute proof that radiation hazards will be eliminated, since adequate dispersion of the resulting debris in the atmosphere is also necessary. However, the collection of ablation particles was not compatible with the other objectives of RFD-1, so future flight tests may be required for this purpose. Both analytical and laboratory investigations of particle sizes and dispersion are currently under way.

## Fuel-Element Experiment

## Concept

The experimental concept was to simultaneously eject four groups of three tracer-loaded, simulated fuel rods each from the RFD-1 RV early in re-entry, and record their burnup with spectral, plate, and motion-picture cameras. In normal operation, of course, the fuel rods would be carried within the reactor core vessel, and released as a consequence of reactor disassembly. For RFD-1, however, the rods were ejected from an external mounting, since calculations indicated that melting of the lip weld securing the core-vessel cover, which must separate from the vessel before the rods can be released, would occur too late in the trajectory for the rods to be exposed to the desired maximum heating.

The construction of the tracer-loaded rods, with varying thicknesses of UZrH surrounding the tracer material, has already been noted. Since all the rods should follow the same trajectory and therefore experience essentially the same heating, the UZrH should ablate away first from the group with the thinnest UZrH wall, exposing the tracer material contained in that group. The tracer element should then flare, emitting radiation characteristic of that particular material. Because of the different wall thickness of UZrH in the four groups, each tracer element should flare at a different time and altitude. Comparison between the volume of UZrH consumed and the calculated heating for the trajectory should then provide indications of the rate and volume of ablation versus re-entry heating.

## Fuel-Rod Design

Figure 3 shows a cutaway view of a normal SNAP-10A fuel element. Figures 4 through 7 are cutaways of the tracer-loaded, simulated fuel rods used for the RFD-1 fuel-element experiment. Configurations of the tracer-loaded rods were dictated largely by the physical and mechanical properties of the materials included. Since UZrH is extremely brittle, the minimum wall thickness that would be machined without fracturing the rod was 0.100 inch. Conversely, it was surmised that a minimum of 0.300 cubic inch of tracer material would be necessary to permit spectral observation at ranges of 100 to 200 miles. The void required to accommodate this amount of tracer material and its insulation limited the maximum wall thickness of UZrH to 0.406 inch. One group of rods was made with the minimum thickness, and one group with the maximum; wall thicknesses of the remaining two groups were spaced about equally between these extremes. Wall thicknesses for the four groups were 0.101 , $0.184,0.306$, and 0.406 inch, resulting in volume ratios of $1.00,1.65,2.45$, and 2.92 , respectively. Since predictions indicated that only a small volume of fuel element would be consumed during the RFD-1 re-entry, these ratios were considered to be near optimum. Table I lists the weights and volumes of principal materials in each group of rods, and Figure 8 shows the 12 tracer-loaded rods flown on RFD-1.


Figure 3. SNAP-10A fuel rod


Figure 4. SNAP-10A fuel rod with strontium tracer


Figure 5. SNAP-10A fuel rod with barium tracer


Figure 6. SNAP-10A fuel rod with silver tracer


Figure 7. SNAP-10A fuel rod with gold tracer

TABLE I
RFD-1 Tracer-Loaded Fuel Rods

| $\begin{gathered} \text { Group } \\ \text { No. } \\ \hline \end{gathered}$ | Tracer |  |  | UZrH |  | Cladding |  | Lead Ballast |  | Total Weight (1b) | $\begin{aligned} & \text { Ballistic Coefficient, } \\ & \text { W/C } C_{D}^{A}\left(1 \mathrm{~b} / \mathrm{ft}^{2}\right) \\ & \hline \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Material | Weight (1b) | Volume $\left(\ln .^{3}\right)$ | Weight (1b) | Volume $\left(\operatorname{in} .^{3}\right)$ | Weight $\qquad$ | Volume $\text { (in. }{ }^{3} \text { ) }$ | Weight (1b) | Volume $\left(\ln .^{3}\right)$ |  |  |
| 1 | Strontium | 0.205 | 1.965 | 0.955 | 4.30 | 0.336 | 0.982 | 1.040 | 2.550 | *2.812 | 43.70 |
| 2 | Barium | 0.217 | 1.526 | 1.543 | 7.10 | 0.325 | 0.982 | 0.577 | 1.394 | *2.864 | 44.40 |
| 3 | Silver | 0.426 | 1.132 | 2.311 | 10.59 | 0.326 | 0.982 | -- | -- | *3.160 | 49.10 |
| 4 | Gold | 0.203 | 0.287 | 2.731 | 12.25 | 0.327 | 0.982 | -- | -- | *3.294 | 51.10 |
| SNAP-10A | -- | -- | -- | 3.140 | 14.17 | 0.329 | 0.982 | -- | -- | 3.469 | 53.90 |

[^0]

Figure 8. Full complement of tracer-loaded fuel rods for RFD-1 external fuel-element experiment

The four tracer elements, strontium, barium, silver, and gold, were selected because of their intense and easily identified spectral characteristics. (Figures 35, 36,37 , and 38 show the wave lengths and intensities of the various tracers.)

Calculations indicated that insulation would be necessary to delay melting of the tracers as long as possible during ablation of the fuel elements. Insulation would also help to maintain the pressure of the tracer vapor at a low level. (A high pressure would have ruptured the rod walls and exposed the tracers, giving a premature indication of ablation.) Fiberfrax insulation ( $51 \% \mathrm{AlO}_{3}, 47 \% \mathrm{SiO}_{\text {, }}$ ) was used because of its high melting point $\left(3300^{\circ} \mathrm{F}\right)$ and relatively good insulating properties. In addition to the insulation, a small vent hole was provided in each end of the rod to ensure against pressure build-up.

To ensure that all of the fuel rods received essentially the same heat input, it was necessary that they follow essentially the same trajectory. This meant that all 12 rods had to have a nearly identical W/CDA. Since the density of strontium and barium is much lower than that of silver, gold, and UZrH, lead ballast was inserted into the centers of the rods containing strontium and barium tracers.

The brittleness of UZrH made it impossible to machine screw threads in the fuel elements for end plugs to contain the tracers. Accordingly, after the tracers and insulation had been inserted into the rods, alumina ceramic was sprayed into the ends to form a plug.

Hastelloy $N$ cladding with a thin, internal coating of ceramic barrier material, as used on the actual SNAP-10A fuel rods, was included on the experimental fue 1 rods. Preflight calculations indicated that, since this cladding is only 0.015 inch thick, it would melt away early in the re-entry.

## Fuel-Rod Mounting and Ejection

The 12 tracer-loaded fuel-rods were mounted externally on the RV in four groups of 3 rods each, with each group 90 degrees from the next (Figure 9). The mounting brackets were designed to release all 12 rods simultaneously 282 seconds after launch, at an altitude of 350,000 feet. Release was effected by detonating explosive boits, which were initiated by signals programmed through a timer in the RFD-1 TM system. Immediately after rod release, the wires which held the mounting brackets in place were severed by an explosively actuated cutter, permitting all the bracket hardware to fall free of the RV.


Figure 9. RFD-1 re-entry vehicle

## Structural and Functional Tests

Before the RFD-1 flight test, a series of qualification tests was conducted on both the RV and the fuel rods. Tests concerning the fuel rods are described briefly below.

Shock, Vibration, Acceleration, and Balance -- The RV was subjected to sinusoidal and random vibration, linear acceleration, shock, and a test for spin balance. In each of these tests, a set of the experimental fuel rods was mounted on the RV. The rods, which were radiographed before and after testing, suffered no observable detrimental effects as a result of any of the tests.

Ejection -- During development of the fuel-rod brackets, a series of tests was conducted to monitor their performance. Dummy fuel rods were mounted in the brackets, which were in turn mounted on a simulated RV cone (Figure 10A). This assembly was spun at the predicted RV spin rate of 150 rpm . The explosive bolts were then detonated, releasing the entire bracket mechanism and the dummy fuel rods. As a result of centrifugal force, the fuel rods and bracket components accelerated away from the simulated RV cone. Cameras mounted above the assembly photographed the entire sequence of events (Figure 10B). On the basis of comparison between distance traveled and time (both measured from the film), a history of ejection velocity and of the direction in which the fuel rods and brackets traveled was determined. This information was used in trajectory predictions, both before and after the flight test.

(A) Before ejection

(B) At ejection

Figure 10. Ejection test of fuel rods for RFD-1 external fuel-element experiment

## Qualification and Ablation Tests

The tracer-loaded fuel rods were extensively modified from the normal SNAP-10A fuel elements, giving the rods physical properties quite different from those of the normal fuel elements. Machining the cavities for the tracer materials reduced the strength of the rods, and inclusion of the tracer materials and insulation disrupted the normal heat-transfer pattern of the UZrH fuel element. In addition, an internal pressure build-up caused by tracer vapor was possible. Since the possible consequences of these deviations were not obvious, a series of tests was conducted in a l-megawatt, electric-arc tunnel. Two-inch-iong test specimens simulating the experimental designs and materials except for length and deletion of the external cladding were subjected to the predicted re-entry heat input pressure and enthalpy. Figure 11 is a cutaway view of a typical test specimen. (Specimens of actual SNAP-10A fuel elements were also tested for comparison.) The support rod was inserted in a rotating mechanism and rotated about its longitudinal axis at 2.4 rps . This procedure simulated the tumbling re-entry of a cylinder as closely as was possible, and distributed the heating around the cylinder circumference.


Figure 11. Tracer-loaded fuel-rod specimen for tests in the electric-arc tunnel

The objectives of these tests were:

1. To ascertain the structural integrity of the tracer-loaded fuel rods under the RFD-1 re-entry environment.
2. To observe the mode of fuel-rod ablation and tracer exposure.
3. To establish factors to be used in interpreting the flight-test data, including extrapolation from experimental to actual (SNAP-10A) fuel-element ablation, and extrapolation from the RFD-1 trajectory to an orbital-decay trajectory.
4. To verify and record the spectral characteristics of the tracer materials.
5. To test the operation of the spectral and event cameras to be used in the test flight.

Figures 12A and 12B show typical tests. Figure 12A shows fuel-rod ablation and the flare of the strontium tracer. Figure 12 B shows the flare from the barium tracer. The marked difference in color of the two flares may be seen in color photoprints on file in Sandia Division 7411. This phenomenon was used to aid in identifying the fuel elements during data evaluation. Results of the tests and the conclusions derived were:

1. Design of the specimens was acceptable, although minor modifications to the end caps were shown to be necessary. No serious structural defects were observed.
2. The mode of fuel-rod ablation and tracer exposure was not uniform. In all cases, the UZrH developed "hot spots" and ablated at these hot areas. This was especially true where flaws such as scratches or cracks were present in the UZrH; eventually a hole (or holes) would ablate through the UZrH in the flawed area and expose the tracer material. As ablation continued, the tracer would flare through the holes in a pulsating pattern (depending on its orientation relative to the jet).

Some of this problem of nonuniform ablation was caused by attenuation of the jet temperature from the center to the outer circumference. It was surmised that during re-entry, heating would be more evenly distributed over the entire surface of the fuel rod, and that these effects would therefore be somewhat reduced. Ideally, all of the UZrH would have ablated away before the tracers were exposed. Since this was not possible, the observed mode of tracer exposure was considered in interpretation of the RFD-1 data.
3. Little difference was noted between the $B T U / 1 b$ required to ablate the test specimens and that required to ablate solid SNAP-10A fuel elements. Inclusion of the tracer materials and insulation appeared to have a negiligible effect on the total ablation of the rods.
4. The record of spectral characteristics of the tracers was satisfactory. These films were used for comparison with the flight-test films to evaluate results.

The heat inputs and pressure required to ablate the fuel elements were used for both preflight predictions and postflight analysis. (See Section IV of this report, and also SCDR 124-63, RFD-1 Fuel Rod Qualification Tests - Phase I, which presents details of the tests described above, including methods of analysis and results.)


Figure 12. RFD-1 fuel-rod ablation and qualification tests in electric-arc tunnel

## Preflight Predictions

Data from fuel-rod qualification tests conducted in the hyperthermal tunnel were used as the basis for calculating predictions of burnup in the flight test. Since the tunnel can produce only a constant-flow condition (as opposed to the pulse profile of heating generated in an actual re-entry), the predictions were made from considerations of total heat input. Justification for this type of analysis is the fact that present computer programs are not able to describe adequately the real effects of surface oxidation, combustion, and hydrogen diffusion on rod heating and temperature response. Work to develop such a program is in progress, but in the meantime only gross effects can be considered.

The SNAP-000439 theoretical trajectory and heating curve was used as the basis for the preflight predictions. This theoretical trajectory assumed a higher re-entry velocity than that achieved in the flight test, and therefore produced higher heat-input conditions than the actual RFD-1 trajectory. By means of the analysis described in detail later in this report (Section IV, Pp. 114-126), the following predictions of burnup altitudes were made:

| Tracer |  | Predicted Burnup Altitude |
| :--- | :--- | :--- |
| Strontium |  | 166,000 to 177,000 feet |
| Barium |  | 126,000 to 145,000 feet |
| Silver |  | Tracer not exposed |
| Gold |  | Tracer not exposed |

## SECTION III -- OPTICAL DATA ACQUISITION

Seven aircraft, three ships, and two ground stations adjacent to the flight path and in the impact area were used for data acquisition during the test. Figure 13 shows their locations relative to the re-entry flight path.


Figure 13. RFD-1 re-entry flight path and locations of downrange instrument stations

The intended function of the ships was to record TM data and to recover the RV and the remains of the STR after impact. Four of the aircraft were equipped to receive TM data from the RV (a presentation of the TM data is included in SC-RR-64502, SC-RR-64-511, and SC-RR-64-515). Four of the aircraft were also equipped to receive the SARAH beacon transmitted from the RV after impact to assist in recovery. Three aircraft were employed as airborne platforms for the optical instrumentation.

The ground station at Cooper's Island was headquarters for RFD-1 operations at Bermuda. This was also the location of the NASA FPS-16 radar, which tracked the RV C-band beacon and recorded its trajectory. High Point, Bermuda, was the only ground-based camera station.

## Optical Instrumentation

Twenty ground-based and 41 airborne cameras photographed the re-entry. The reasons for using both ground-based and airborne cameras were:

1. In the event of any major breakdown, such as electrical power failure, at least partial coverage of the re-entry was still possible.
2. The closeness of the aircraft to the re-entry flight path was not limited. Since re-entry closer to Bermuda than 90 nautical miles was not allowed, ground-camera slant ranges were restricted to at least that distance.
3. The effects of atmospheric attenuation were reduced by the 10,000 -foot altitude of the aircraft cameras.
4. Because of their large light-gathering systems, many of the groundbased cameras could not be installed in aircraft.
5. Ground-based cameras are more stable and are not subject to aircraft vibrations. These two factors are important in reduction of the data.

## Ground-Based Cameras

A11 ground-based cameras were stationed at High Point (Figures 14 and 15), on the south coast of Bermuda. High Point is 125 feet above sea level, with an unobstructed view of the re-entry flight path. The facilities there were essentially self-sustaining. Auxiliary equipment included:

1. Diesel-electric generators, 65 and 45 KVA .
2. A power-distribution system capable of converting to generator or commercial power.
3. An operations building.
4. A photo trailer for loading cameras and developing film.
5. A transportainer to house recording and TACODA (a parallaxcompensating computer) equipment.
6. Two transportainers for spare parts.
7. A storage building.

Table II lists the films and the mechanical features of the cameras used at High Point to record the RFD-1 re-entry. The TACODA (target coordinate data) system used to control the tracking of certain of the cameras during part of the trajectory is described briefly below, followed by descriptions of the optical instrumentation at High Point.


Figure 14. Bermuda ground stations


Figure 15. Camera station at High Point, Bermuda

TABLE II
High Point Cameras

| $\begin{gathered} \text { Camera } \\ \text { No. } \end{gathered}$ | Description | Grating or Filter | Lens |  |  | Film | Sampling Rate | Illustration |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Focal <br> Length (in.) | Aperture | Field of View |  |  |  |
| ME-16-1 | Photosonics 70-mm motion-picture camera | None | 117.5 | f/7.0 | $1.1 \times 1.1^{\circ}$ | Super Anscochrome | 90 frames/sec | Figure 16 |
| ME-16-2 | Mitchell 35-mm motion-picture camera | None | 18 | f/3.0 | $3.4 \times 2.3^{\circ}$ | Super Anscochrome | 96 frames/sec | Figure 16 |
| ME-16-3 | Mitchell 35-mm motion-picture camera | None | 2 | f/2.0 | $27.0 \times 20.0^{\circ}$ | Super Anscochrome | 96 frames/sec | Figure 16 |
| LA-24-1 | K-37 time-resolved streak spectrograph | Bausch and Lomb 600-lines/mm grating | 117.5 | f/5.0 | $0^{\circ} 27^{\prime}$ wide | Tri-X Aerecon | 3/16 in./sec | Figure 17 |
| LA-24-2 | Photometer | Filter wheel with 24 special interference filters | $\begin{aligned} & \text { 12-in. dia. } \\ & \text { cassigrainian } \end{aligned}$ | -- | $0^{\circ} 24^{\prime}$ wide | None | Filter wheel turning <br> 600 rpm | Figure 17 |
| LA-24-3 | Mitche11 35-mm motion-picture camera | None | 40 | f/5.7 | $1.5 \times 1.0^{\circ}$ | Super Anscochrome | 96 frames/sec | Figure 17 |
| $\begin{aligned} & P-1, P-2, \\ & \text { and } P-3 \end{aligned}$ | K-37 plate cameras | None | 12 | f/2.5 | $53.0 \times 45.2^{\circ}$ | ```10 x 12-in. glass plates with Kodak 103-F emulsion``` | Continuous | Not shown |
| $\begin{aligned} & T-1, T-2, \\ & \text { and } T-3 \end{aligned}$ | K-37 plate trajectory cameras | None | 12 | f/2.5 | $53.0 \times 45.2^{\circ}$ | ```10 x 12-in. glass plates with Kodak 103-F emulsion``` | Chopped: <br> open 7 sec <br> closed 2 sec ; <br> repeat | Figure 19 |
| S-1 through $s-6$ | K-37 plate spectrographs | Bausch and Lomb 600-1ines/mim grating | 12 | f/2.5 | $53.0 \times 45.2^{\circ}$ | ```10 x 12-in. glass plates with Kodak 103-F emulsion``` | Continuous | Figure 18 |
| C-1 | Photosonics 70-mm cinespectrograph | Bausch and Lomb 600-1ines/mm grating | 4 | f/1.9 | $30 \times 30^{\circ}$ | Kodak Linagraph Shellburst | 10 frames/sec | Figure 20 |
| $\begin{aligned} & \mathrm{H}-1 \\ & \text { (hand held) } \end{aligned}$ | Be 11 and Howe $1116-\mathrm{mm}$ motion-picture camera | None | 6 | f/2.5 | $4.0 \times 2.8^{\circ}$ | Super Anscochrome | 126 frames/sec | Not shown |

TACODA -- TACODA is a parallax-compensating computer system through which the tracking mounts for High Point optical instruments were slaved to the FPS-16 radar at Cooper's Island. Operation is as follows:

1. Well before re-entry began, the FPS-16 made the initial Bermuda acquisition of the RV, picking up and automatically tracking the C -band beacon.
2. Position data from the FPS-16 were fed to the TACODA, which was located at High Point, through a land telephone-1ine link from Cooper's Island. From these data, the TACODA computed the changing azimuths and elevations for the tracking mounts, compensating for the parallax introduced by the 12 -mile separation between the FPS-16 and the mounts.
3. Signals reflecting these azimuths and elevations were continually transmitted from the TACODA to the drive systems for the mounts, which then automatically tracked the not-yet visible RV.
4. When the tracking-mount operators at High Point sighted the RV, first through 5.6 -power and then through 32 -power telescopes, they switched to manual control. This was done because objects other than the RV required tracking, and because visual tracking is more accurate than radar once an object is sighted.

During the RFD-1 flight, the TACODA system functioned perfectly, guiding the tracking mounts to the RV before it was visible to the operators through their tracking scopes.

ME-16 Tracking Mount -- The entire ME-16 mount and protective dome (Figure 16) is motorized to rotate 360 degrees in azimuth and 90 degrees in elevation. Three cameras were included on the mount for RFD-1 coverage. The main optics, contained in the large center tube, were connected to a $70-\mathrm{mm}$ Photosonics $10-\mathrm{B}$ high-speed motion-picture camera. Light from the re-entry was collected by a Newtonian telescope 16 inches in diameter and 117.5 inches in focal length, and focused on a parabolic mirror at the rear of the tube. This mirror reflected the light via two flat-angled mirrors to a focus on the film plane of the camera. A rotating prism and shutter system in the camera directed the image onto the film. Each rotation of the prism recorded one frame on the film. With a film frame rate of 90 frames/ sec, this camera is capable of pulling 1700 feet in 100 seconds.

The remaining two cameras were $35-\mathrm{mm}$ Mitchell shutter-operated framing cameras with focal lengths of 18 and 2 inches, respectively. They were not connected to the Newtonian telescope in the main optics. A timing code received from the Bermuda Range Time (BRT) transmitter was recorded on the film edges by a neon lamp.

LA-24 Tracking Telescope -- This tracking mount operates similarly to the ME-16. However it mounted different instrumentation.

The large center tube shown in Figure 17 housed a parabolic mirror 24 inches in diameter with a focal length of 117.5 inches. The light from the re-entry was reflected from this mirror and brought to a focus at the focal point of an Aeroektar 12 -inch f.1. (focal length) lens. On leaving this lens, the light rays (which were parallel) passed through a $600-1$ ines/mm grating. After diffraction by the grating, the light was brought to another focus on the film plane of the camera by a second 12 -inch-f.1., f/2.5 lens. The 9 -inch wide film was pulled perpendicular to the dispersion of the grating at $3 / 16$ inch per second. From the resulting spectral record and the time code recorded on the edge of the film, it was possible to determine when spectrally distinct events occurred.

The smaller tube on the left in Figure 17 housed the photometer, consisting of a 12 -inch cassigrainian system for gathering light which was then passed through 24 narrow band-pass filters. These filters, approximately 10 A at half width, were centered on predetermined spectral lines of the tracer materials: strontium, barium, silver, and gold. Two spectral lines per tracer were chosen as the lines which, for that material, had the best chance of being detected. The probable relative intensities of ilines and interfering lines were considered in this choice.


Figure 16. ME-16 tracking telescope


Figure 17. LA-24 tracking telescope

To distinguish between spectral line radiation and overall intensity fluctuations, background filters were used to monitor the background radiation $50 \&$ on each side of the spectral line of interest. The filters were mounted around the periphery of a disc which was spun at 600 rpm to provide $0.1-s e c o n d$ time resolution for each wavelength monitored.

The radiation transmitted through each filter as it passed the opening in the lens system was detected by a dry-ice-cooled RCA-C-7265 fourteen-stage photo-
 range of from $5^{\prime} \times 10^{-13}$ to $2 \times 10^{-8}$ lumens. This tube transmitted an electrical pulse of measured intensity to a C.E.C. oscillograph. Time was recorded on the oscillograph simultaneously with the photometer pulses, making it possible to determine the elements flaring during the re-entry and the times of their appearances.

The shelf on the right in Figure 17 supported a Sandia-designed 40 -inch-f. 1. , $\mathrm{f} / 6.8$ lens mounted with a high-speed $35-\pi m$ Mitchell motion-picture camera. This camera provided high-magnification, slow-motion movies of objects and events recorded by the streak spectrograph and photometer.

Plate Cameras -- The plate cameras were essentially continuous-exposure cameras. The optics were war-surplus Kodak Aeroektar 12-inch-f.1., f/2.5 aerialmapping lenses; the camera bodies were cast by Sandia. The film, located at the rear of the housing, consisted of a $10 \times 12$-inch glass film plate coated with Kodak 103-F film emulsion. Images of the re-entry objects were focused on the film plate by the lenses. The lens shutter was an iris-type diaphragm which was kept open so that the film was continuously exposed during re-entry. The re-entry was recorded as a series of streaks across the film. Three cameras oriented at different angles were required to record the entire re-entry.

Spectrographic Plate Cameras -These cameras (Figure 18) were similar to the plate cameras. However, a Bausch and Lomb 600-1ines/mm trans-mission-diffraction grating blazed at $5500 \AA$ was attached to the front of each lens. Light from the re-entry objects was separated, as it passed through the grating, into spectral groupings characteristic of the radiation from the elements comprising the different materials in the RS. This light was then focused onto the film plane by the lens. Two arrays of three cameras each were used, One group was oriented at a high elevation angle and the other at a lower elevation angle to assure coverage of the entire re-entry envelope.

Chopped Trajectory Plate Cameras -- These cameras (Figure 19) used lens systems and glass film plates similar to those in the plate and spectrographic plate cameras. The shutter system, located in the cylindrical housing, was operated by an electrically actuated spring. A pre-set timing mechanism opened the shutter for 7 seconds and then closed it for 2 seconds. This sequence continued during the entire re-entry, causing the re-entry trails on the film to be interrupted at these intervals. The times of openings and closing were also recorded on the C.E.C. oscillograph. Comparison of the times with


Figure 18. Spectrographic plate cameras
the measured and theoretical trajectories furnished a history of altitude, range, and velocity for the various objects (this is discussed in Section IV, pp. 41-134).

Cinespectrograph -- This instrument (Figure 20) was hand-tracked and operated. It was equipped with a 4-inch-f.1., f/1.9 lens and a Bausch and Lomb 600-1ines/mm :ransmission grating, and provided ten $70-\mathrm{mm}$ spectral pictures per second.


Figure 19. Chopped trajectory plate cameras

## Airborne Cameras

Three aircraft, one DC-4 and two C-54's, supported RFD-1 camera operations. Their locations during the reentry were as shown in Figure 13.

The DC-4 was furnished and outfitted by NASA's Langley Research Center, Hampton, Virginia. It was designated NASA-238. Although this aircraft was not included in the formal RFD-1 plan, films from its cameras furnished much of the data used for evaluation of the re-entry events. Table III lists the cameras used on this plane to photograph the RFD-1 re-entry.

The two C-54's and their flight crews were furnished by AFSWC, Kirtland AFB, Albuquerque, New Mexico. Sandia instrumented these planes with cameras and TMreceiving equipment. AFSWC installed additional cameras for experimentation, documentary films, and re-entry analysis. These aircraft were designated AFSWC-46l and

TABLE III
NASA Airborne Cameras

|  |  |  | Lens |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Camera No. | Description | Grating | Type | Focal Length (in.) | Aperture | Field of View | Film | Sampling Rate |
| AC-37 | KG-24 spectral camera | 600 lines/mm | Aeroektar | 7 | f/2.5 | $40.0 \times 40.0^{\circ}$ | Royal-X Pan <br> black-and-white | Continuous |
| AC-84 | KG-24 spectral camera | 75 lines/mm | Aeroektar | 7 | f/2.5 | $40.0 \times 40.0^{\circ}$ | Royal-X Pan <br> black-and-white | Continuous |
| AC-85 | KG-24 spectral camera | 400 lines/mm | Aeroektar | 7 | f/2.5 | $40.0 \times 40.0^{\circ}$ | Royal-X Pan black-and-white | Continuous |
| $\begin{aligned} & \mathrm{AC}-36, \mathrm{AC}-38 \\ & \mathrm{AC}-80, \mathrm{AC}-86 \end{aligned}$ | KG-24 plate cameras | None | Aeroektar | 7 | f/2.5 | $40.0 \times 40.0^{\circ}$ | Royal-X Pan black-and-white | Continuous |
| AC-1 | RC-5 plate trajectory camera | None | Avigon | 6 | f/2.6 | $73.0 \times 73.0^{\circ}$ | Royal-X Pan <br> black-and-white | Chopped: <br> open 0.25 sec , <br> closed 0.25 <br> sec; repeat |
| AC-101 | Flight Research 35-mm motionpicture camera | None | Bell and <br> Howe 11 <br> Eyemax | 10 | f/4.5 | $5.6 \times 4.2^{\circ}$ | Royal-X Pan <br> black-and-white | 10 frames/sec |
| 113 | ```Cine Kodak 16-mm motion- picture camera``` | None | Kodak Anastigmat | 2 | f/2.0 | $10.4 \times 8.3^{\circ}$ | Ektachrome ER color | 16 frames/sec |

AFSWC-521. Figure 21 shows the bank of Sandia cameras carried in AFSWC-52l. Included were a streak spectrograph, a photometer, a $16-\mathrm{mm}$ motion-picture camera, and three K-24 spectral cameras; the optical instrumentation in AFSWC-461 was essentially the same, although the photometer was omitted. Figure 22 shows the locations of cameras carried in the AFSWC aircraft, and Table IV lists these cameras. The major cameras in these planes are described briefly below.


Figure 21. Sandia airborne cameras


Figure 22. Optical instrumentation in AFSWC Aircraft 461 and 521 (refer to Table III)

TABLE IV
Airborne Cameras, AFSWC Aircraft 461 and 521

| Camera No, | Aircraft | Description | Grating or Filter | Lens |  |  |  | Type | $\begin{aligned} & \text { ASA } \\ & \text { Speed } \end{aligned}$ | $\qquad$ | Sampling Rate |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | Type | Focal Length (in.) | Aperture | Field of View |  |  |  |  |
| 13 | 461, 521 | Photometer | Filter wheel with 24 special interference filters | $\underset{\text { mirror }}{\text { 12-in.-dia. }}$ | -- | -- | $0^{\circ} 24^{\prime}$ | -- | -- | -- | $\begin{aligned} & \text { Filter whee } 1 \\ & \text { turning } \\ & 600 \mathrm{rpm} \end{aligned}$ |
| 14 | 461, 521 | Milliken DBM-5 16-mm motionpicture camera | None | -- | 6 | f/2.5 | $4.0 \times 2.8^{\circ}$ | Super Anscochrome | 160 | 1200 | 400 frames/sec |
| 15 | 521, 521 | General Electric <br> aircraft-type <br> K-37 time- <br> resolved streak <br> spectrograph | Bausch and Lomb 600-1ines/mm grating | Aeroektar | 12 | f/2.5 | $3 \times 42^{\circ}$ | Tri-X Aerecon | 200 | -- | 3/16-in./sec |
| $\begin{aligned} & 2-13 \\ & 2-15 \end{aligned}$ | $\begin{aligned} & 461 \\ & 521 \end{aligned}$ | $\begin{aligned} & \text { Hulcher Mode } 1 \\ & 10070-\mathrm{mm} \\ & \text { motion-picture } \\ & \text { camera } \end{aligned}$ | Bausch and Lomb 600-1ines/mm grating | Aeroektar | 7 | f/2.5 | $18 \times 18^{\circ}$ | DuPont 140 | 320 | 400 | 20 frames/sec |
| 2-9 | 461 | Milliken DBM-5 16-mm motionpicture camera | None | Raptar | 2 | f/2.5 | $11.7 \times 8.4{ }^{\circ}$ | $\begin{aligned} & \text { Ansco Ultraspeed } \\ & \text { color } \end{aligned}$ | 500 | 400 | 128 frames/sec |
| 2-11 | 521 | Milliken DBM-5 $16-\mathrm{mm}$ motionpicture camera | None | Angenieux | 3 | ¢/2.5 | $7.5 \times 5.4^{\circ}$ | $\begin{aligned} & \text { Ansco Ultraspeed } \\ & \text { color } \end{aligned}$ | 500 | 400 | 128 frames/sec |
| 2-20 | 461, 521 | Aerosonics <br> ballistic camera | None | Metrogon | 6 | f/6.3 | $67 \times 79^{\circ}$ | RXP $8 \times 10 \mathrm{in}$. | 1250 | One plate | Random chop: <br> 0.04-0.065 sec |
| z-18 | 461, 521 | Mitchell Model B $35-\mathrm{mm}$ motionpicture camera | None | Baltar | 6 | f/2.5 | $6.9 \times 9.2^{\circ}$ | RXP | 1250 | 1000 | 96 frames/sec |
| $\begin{aligned} & \text { Z-11 } \\ & \text { z-13 } \end{aligned}$ | $\begin{aligned} & 461 \\ & 521 \end{aligned}$ | General Electric aircraft-type K-37 | None | Aeroektar | 12 | f/2.5 | $42 \times 42^{\circ}$ | RXP | 650 | 75 | $\begin{aligned} & 1.25 \mathrm{sec}(461) \\ & 0.5 \mathrm{sec}(521) \end{aligned}$ |
| Z-17 | 461, 521 | Avco-Everett <br> ballistic camera | None | Metrogon | 6 | f/6.3 | $67 \times 79^{\circ}$ | RXP $8 \times 10 \mathrm{in}$. | 1250 | One plate | Random chop: $0.04-0.065 \mathrm{sec}$ |
| z-16 | 461, 521 | Aerosonics <br> ballistic camera | None | Metrogon | 6 | f/6.3 | $67 \times 79^{\circ}$ | RXP $8 \times 10 \mathrm{in}$. | 1250 | One plate | Random chop: <br> 0.04-0.065 sec |
| 2-12 | 461, 521 | $\begin{aligned} & \text { Eastman Kodak } \\ & \text { K-24 } \end{aligned}$ | None | Aeroektar | 7 | f/2.5 | $39 \times 39^{\circ}$ | RXP | 650 | 56 | $\begin{aligned} & 0.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |
| 2-14 | 461, 521 | Eastman Kodak $\mathrm{K}-24$ | None | Aeroektar | 7 | £/2.5 | $39 \times 39^{\circ}$ | RXP | 650 | 56 | $\begin{aligned} & 0.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |
| $\begin{aligned} & z-7 \\ & 2-8 \end{aligned}$ | $\begin{aligned} & 461 \\ & 521 \end{aligned}$ | $\begin{aligned} & \text { Eastman Kodak } \\ & \text { K-24 } \end{aligned}$ | None | Aeroektar | 7 | f/2.5 | $39 \times 39^{\circ}$ | Tri-X | 200 | 56 | $\begin{aligned} & 0.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |
| 2-10 | 461, 521 | Eastman Kodak K-24 | Bausch and Lomb 600-lines/mm grating | Aeroektar | 7 | f/4.0 | $39 \times 39^{\circ}$ | RXP | 650 | 56 | $\begin{aligned} & 2.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |
| z-2 | 461, 521 | Eastman Kodak K-24 | Bausch and Lomb 600-1ines/mm | Aeroektar | 7 | f/5.6 | $39 \times 39^{\circ}$ | RXP | 650 | 56 | $\begin{aligned} & 2.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |
| 2-1 | 461, 521 | $\underset{\mathrm{K}-24}{\text { Eastman Kodak }}$ $K-24$ | Bausch and Lomb 600-1ines/mun grating | Aeroektar | 7 | f/2.5 | $39 \times 39^{\circ}$ | RXP | 650 | 56 | $\begin{aligned} & 2.5 \mathrm{sec}, 4-\mathrm{sec} \\ & \text { cycle } \end{aligned}$ |

Time-Resolved Streak Spectrograph -- This instrument was similar to the TRSS in the LA-24, but employed a refracting system for light gathering in lieu of a reflecting telescope.

Sampling Photometer -- The photometer was the same as the one mounted on the LA-24.

K-24 Spectral Cameras -- These cameras were fitted with 7-inch-f.1., f/2.5 Aeroektar lenses. A $300-1$ ines $/ \mathrm{mm}$ diffraction grating was included to produce spectral films. These cameras furnished one 5-1/2-inch-wide photo per second.

## Description of the Re-entry

Since optical coverage was one of the primary methods employed to record RFD-l events, it was imperative that no clouds or haze be present between High Point and the expected flight path at re-entry. After numerous delays caused by such adverse weather conditions, the Scout missile bearing RFD-1 was launched at 04-38-14 Zulu time on May 22, 1963. Impact of the re-entry vehicle and reactor into the ocean occurred approximately 500 seconds later.

## Trajectory

Before the flight, trajectories were computed for the RV, the fourth-stage motor, the simulated fuel rods, and the reflectors. They were computed for the nominal, 2- $\sigma$ low steep, and $2-\sigma$ high shallow re-entry conditions. These are discussed in SC-RR-64-510. Figure 23 shows the envelope of computed azimuths and elevations for the RV and fuel-rod trajectories, referenced to High Point; the actual RFD-1 RV trajectory is superimposed for comparison. The actual trajectory was somewhat high and shallow compared with the predicted nominal trajectory. This was actually better from the standpoint of optical data acquisition, since the re-entry path remained above the ocean haze for a longer time.


Figure 23. Envelope of RFD-1 re-entry trajectories

## Events

In the nominal trajectory, re-entry was predicted to be visible 325 seconds after launch, at an altitude of 240,000 feet. For the possible steep or shallow re-entry trajectories, the times at which the RV would reach this altitude were 309 and 343 seconds, respectively. It was decided to uncap the plate cameras at 240 seconds, since they were used essentially for time exposures. However, because of limitations on film footage, the motion-picture cameras required a more
exact starting time, and were therefore started upon command from the NASA FPS -16 radar plot board when the re-entry vehicle reached an altitude of 270,000 feet. This occurred at 324 seconds in the actual flight.

The actual re-entry flight path was as shown in Figure 13. The times of reentry, of separation of the fourth-stage motor and the simulated fuel rods, and of all events and sequences in general were very close to those which had been predicted.

Data Obtained

Almost all of the optical instrumentation functioned satisfactorily during the re-entry. Performance was less than satisfactory in three instances: (1) one ground-based and two aircraft cameras did not operate, (2) timing did not record on several films, and (3) many of the films were underexposed to varying degrees as a result of the distances involved and of atmospheric attenuation. However, the data obtained were ample to permit satisfactory reduction for investigation of the test objectives. Table $V$ is a listing of all optical instruments included in the test, showing the amount and quality of data obtained from each.

TABLE V
RFD-1 Photographic Data

## High Point Plate Cameras

| Camera No. | Objects Photographed | Remarks |
| :---: | :---: | :---: |
| P-1 | 4th-stage motor and early portion of bracket trajectories | Good data; clear. |
| P-2 | A11 objects; center portion of re-entry | Good data; slightly out of focus and somewhat underexposed. |
| P-3 | Reactor burnup during final portion of reentry | Good data; slightly underexposed. |
| T-1 | Same as P-1 but time chopped | Badly fogged due to light leak; furnished valuable information on 4 th-stage timing. |
| T-2 | Same as P-2 but time chopped | Excellent exposure; good resolution; used entensively in data analysis. |
| T-3 | Same as P-3 but time chopped | Excellent exposure; good resolution; used extensively in data analysis. |
| High Point Spectral Cameras |  |  |
| S-1 | 4th-stage motor and early portion of bracket trajectories | Spectra off plate; not usable. |
| S-2 | A11 objects; center portion of re-entry | Faint but good data. |
| S-3 | Reactor burnup during final portion of reentry | Zero order not sharp, but first-order spectra clear. |
| S-4 | Same as S-1 | Zero order first-order spectra not sharp due to poor focus and underexposure. |


| Camera No. | Objects Photographed | Remarks |
| :---: | :---: | :---: |
| S-5 | Same as S-2 | Fainter than S-2; valuable information on strontium flares obtained. |
| S-6 | Same as S-3 | Good focus; zero order sharp; spectra less distinct. |
| LA-24 TRSS | Mistracked (tracked 4th-stage motor) | Very good spectral data; elements of 4 th stage clear. |
| LA-24 <br> Photometer | Mistracked (tracked 4th-stage motor) | Good record; confirmed LA-24 TRSS. |
| LA-24-3 | Mistracked (tracked 4th-stage motor) | Camera malfunction; no film record. |
| Cinespectrograph | Re-entry vehicle | Badly underexposed; not usable. |
| High Point Framing Cameras |  |  |
| ME-16-1 | Experimental fuel rods | Underexposed; few faint objects visible; not usable. |
| ME-16-2 | Experimental fuel rods | Good film; RV, fuel-rod flares, and brackets visible; color of flares used to identify tracers. |
| ME-16-3 | Experimental fuel rods | Badly underexposed; not used. |
| H-1 | Re-entry vehicle | Faint record of RV; not used. |
| AFSWC Aircraft 521 Plate Cameras |  |  |
| $\begin{aligned} & Z-20, Z-17, \\ & Z-16, Z-14, \\ & Z-13, Z-12, \\ & \text { and } Z-8 \end{aligned}$ | Entire re-entry | Severe vibration; poor resolution; not used in data analysis. |
| AFSWC Aircraft 521 Spectral Cameras |  |  |
| Z-10 | Entire re-entry | Fair focus; underexposed; RV, fuel-rod, and bracket spectra faintly visible. |
| Z-2 | Re-entry vehicle | Faint spectra. |
| Z-1 | Re-entry vehicle | Sharp focus; vibration apparent; clear spectra of RV. |
| Z-15 | Experimental fuel rods | Underexposed; not usable. |
| 13 | Experimental fuel rods | No data; mistrack due to narrow field of view. |
| 15 | RV and fuel rods | RV, fuel-rod, and bracket spectra identified; good timing. |
| AFSWC Aircraft 521 Framing Cameras |  |  |
| 2-18 | Re-entry vehicle | Underexposed; no data. |
| 14 | Experimental fuel rods | Faint objects visible; not used in data analysis. |

Camera No. Objects Photographed
Z-11 Re-entry vehicle

AFSWC Aircraft 461 Plate Cameras

```
Z-20, Z-17, Entire re-entry
Z-16, z-12,
Z-7, Z-14,
and Z-11
```


## AFSWC Aircraft 461 Spectral Cameras

| $Z-10, Z-2$, <br> and $Z-1$ | Entire re-entry |
| :--- | :--- |
| $Z-13$ | Experimental fuel rods |
| 15 | $R V$ and fuel rods |

No data; possibly due to camera orientation.

Underexposed; not usable.
RV, fuel-rod, and bracket spectra visible; no timing.

Underexposed; no data.
Mistrack; no data.
Underexposed; no data.

NASA Aircraft 238 Plate Cameras

```
AC-38, AC-36, Entire re-entry
AC-80, and
AC-86
```

AC-1 Entire re-entry

## NASA Aircraft 238 Spectral Cameras

| AC-85 | Entire re-entry |
| :--- | :--- |
| AC-37 | Entire re-entry |
| AC-84 | Entire re-entry |

NASA Aircraft 238 Framing Cameras

| $A C-101$ | Entire re-entry |
| :--- | :--- |
| $A C-113$ | Entire re-entry |

Good clear film; use limited due to lack of timing and information on aircraft orientation.

Good timing; no detail resolution; used for time only.

Excellent film; RV, fuel-rod flares, and 4th-stage motor very clear.

Slightly out of focus; RV spectra visible.
Only second-order spectra visible; RV, brackets, and 4th-stage motor identified.

[^1]
## SECTION IV -- DATA EVALUATION

Reduction of the films from RFD-1 was not straightfoward. Reasons for the difficulties encountered were:

1. The majority of the motion-picture and spectrographic films were underexposed. This was caused by atmospheric attenuation, a toofast film-frame rate, and the relatively great distance of the cameras from the re-entry.
2. Superposition of the 16 groups of re-entry objects on the plate films made it difficult to distinguish individual objects with certainty.
3. Timing chops on the plate-camera films were too long to allow accurate connection of re-entry streaks across the chops. In addition, the time between chops was too long to permit precise event/time correlation.

In spite of these problems, much usable information was extracted from the films and records. It will be presented in this section.

## Methods of Analysis

Because a number of different kinds of films and records were obtained from RFD-1, it was necessary to reduce the data from each type of camera (plate, spectral, and framing) separately and then to correlate the results. The procedures included in reduction were:

1. Calibration of the Sandia plate-camera films from the star background, and establishment of lines of sight for each re-entry object.
2. Replotting, on an expanded scale, the trajectories from the plate cameras. This allowed separation of the individual streaks for identification as described in Item 3, below.
3. Construction of a re-entry picture of all the objects, based on the theoretical trajectory program described by J. A. Allensworth in The TTA Generalized Rigid Body Theoretical Trajectory Program for Digital Computer, SC-TM-64-526. This picture was then compared with the plates replotted as described in 2, above. Comparison between the theoretical and observed re-entry patterns confirmed identification of many of the objects.
4. Investigation of the spectrographic films. Spectral characteristics of the various re-entry objects were compared with the spectra recorded on the films. Times of occurrence of events were recorded for comparison with the other data.
5. Tabulation of motion-picture film events versus time. Timing marks on the film allowed a precise record of event time. Information from these films was used to confirm data from the preceding steps and to provide exact timing for them.
6. Correlation of all records. Throughout the process of data reduction, the data from each individual film and record were continuously compared with data from the others to confirm tentative findings. When all the separate analyses were completed, a compilation of all the data was made to assure agreement and to allow a synthesis of the entire re-entry sequence.

Time
Since time was the only initially known quantity recorded on each film and record, it was used as the common base for the identification and comparison of events. WWV time, broadcast from Station WWV, Washington, D.C., was recorded on the C.E.C. oscillograph. BRT, generated at Cooper's Island, was synchronized with the WWV time. BRT was transmitted via cable to all ground cameras and was radioed to the aircraft cameras. This time was recorded on the film edges by neon lamps as a serial decimal code. The NASA aircraft generated its own binary time-code, which was also synchronized with WWV time. All times on the films were measurable to 0.01 second.

Detailed discussion of the data reduction procedure and results follows.

Plate-Camera Data

Sixteen groups of objects were re-entering the atmosphere during the reentry phase of RFD-1. Table VI lists their physical characteristics.

TABLE VI
Physical Characteristics of RFD-1

| Item | Weight (lb) | $W / C_{D}{ }^{\text {( }}\left(1 \mathrm{~b} / \mathrm{ft}^{2}\right)$ |
| :---: | :---: | :---: |
| RV-reactor | *532.9 to 330.0 | *372.00 to 248.35 |
| 4th-stage motor | 67.0 | 14.0 |
| Strontium-loaded fuel rods (ea) | 2.812 | 43.7 |
| Barium-loaded fuel rods (ea) | 2.864 | 44.4 |
| Silver-loaded fuel rods (ea) | 3.160 | 49.1 |
| Gold-loaded fuel rods (ea) | 3.294 | 51.1 |
| Brackets 1 and 2 (ea) | 4.23 | 6.43 |
| Rod holders 1 and 2 (ea) | 0.25 | 2.0 |
| Reflectors 1 and 2 (ea) | 7.80 | 8.69 |
| Reflector springs 1-4 (ea) | 0.50 | 12.0 |

Variation is due to ejection of fourth-stage motor, fuel rods, brackets,
holders, reflectors, and springs, and to disassembly of the reactor.

In spite of a large variation in $W / C_{D} A$, all the objects (except the fourth-stage rocket motor, which was retro-rocketed away from the flight path) followed very nearly the same trajectory. They were, however, displaced in time. On the platecamera films, the continuous procession of re-entry objects appeared as a series of superimposed streaks (see Figures 24 and 25). Except for the RV and the fourthstage motor, which were obvious, the various items could not be identified from


Figure 24. RFD-1 re-entry (from Sandia Plate Camera P-3) (0.1 inch * 7800 feet)


Figure 25. RFD-1 re-entry (trom Sandia Plate Camera T-2) (0.1 inch $\approx 7800$ feet)
mere visual inspection. Comparison of the calculated trajectories with those measured from the plate-camera films was made in an effort to:

1. Identify individual re-entry objects, especially the four groups of tracer-loaded fuel rods.
2. Furnish a history of altitudes and velocities versus time. This information was needed for the computation of heat inputs to the RV and fuel rods, and for the location of observed events (disassembly of the RV and exposure of the fue1-rod tracers) in space.
3. Correlate the FPS-16 radar and the plate-camera data.

## Camera Calibration

On each plate film, at least 11 star images distributed over the region of interest were selected and measured on a Mann comparator. The directions corresponding to those images were established from data given in Boss', General Catalogue of 33,342 Stars for the Epoch 1950, and from WWV timing recorded on the C.E.C. oscillograph during the re-entry flight. Apparent lines of sight to the selected stars were obtained by applying the atmospheric-refraction corrections in Manual of Geodetic Astronomy. Figure 26 illustrates the phenomena of atmospheric refraction and apparent lines of sight. The refractions were adjusted to High Point surface temperature and pressure.


Figure 26. Atmospheric refraction

For each camera, (1) direction of the optical axis, (2) plate coordinates of the intersection between the optical axis and the film emulsion, (3) image distance, (4) orientation of the plate coordinate system, and (5) radial-distortion coefficients were determined by minimizing the sum of the squares of the distances between predicted and measured star images (Figure 27). Those distances appeared to be randomly distributed about zero, with a standard deviation of less than 0.00020 inch in each coordinate. Since 12 -inch lenses were used, a probable relative error of no greater than 3.4 seconds of arc is indicated. The effects of plate-surface irregularities and tangential distortion are quite small and were therefore disregarded.


Figure 27. Calibration for Plate Camera T-2, predicted versus measured star images

Since the points on the plate-film re-entry trails were less distinct than the star images, and since photographic emulsion distorts images that are very close together, the data points may be in error by a significantly larger amount. Comparison of two lines of sight to the same object, as determined by two different cameras, indicates that a bias error as large as 30 seconds of arc may be present. No estimate of the error in atmospheric-refraction correction can be made because no independent data are available for comparison. However, it is reasonable to assume that no line of sight is in error by as much as 2 minutes of arc. Also, the relative error between two lines determined from the same camera plate is less than $1 / 2$ minute of arc. The $1 / 2$ minute of arc represents a distance of 100 feet in the trajectory and 0.0018 inch on the plate films (Figures 24 and 25).

## Plate-Film Measurements

To compare the plate-film data with the theoretical trajectories, it was necessary to convert plate-film data and trajectories to a common coordinate system. The quantities selected for comparison were azimuth and elevation of the reentry objects as viewed from the camera stations at High Point. These quantities were chosen to facilitate the use of existing computer programs. (After the data from RFD-1 had been reduced, a computer program was formulated to predict the plate coordinates of successive images of a moving object, given position and camera calibration parameters. That program will be used on future flight tests.)

Measurements of the plate-camera films to allow conversion of the re-entry trail's apparent lines of sight to azimuths and elevations were made with a Mann comparator. This instrument includes a 40 -power microscope for measuring coordinates on two perpendicular axes. Scale resolution is $1 / 2$ micron. Measurement
limitation due to the plate-film resolution was approximately 5 microns. This represents 12 feet in the trajectory and 0.000197 inch on the plate-camera films.

The flares and other points of special interest on the image of each reentry trail were measured and recorded. The nondescript portions of the image trails were described by a series of points at 0.040 -inch intervals. Each point was converted to an apparent line of sight and a large-scale graph (10 inches per degree) of apparent elevation angle versus azimuth was constructed.

## Theoretical Trajectories

Theoretical trajectories were computed on the Control Data 1604 computer, using the program described in SC-TM-64-526, The TTA Generalized Rigid Body Theoretical Trajectory Program for Digital Computer.

## RV Theoretical Versus Radar Trajectories

Differences in measured trajectories of the 16 re-entry objects were small. The maximum vertical separation of the highest and lowest object was 3600 feet. To compute trajectories accurate enough for comparison with the plate-film data it was essential first to establish agreement with a known trajectory. The NASA FPS-16 radar had provided this trajectory. During re-entry, the FPS 16 tracked the RV C-band beacon from 167 seconds ( 450,000 feet) to 360 seconds ( 168,800 feet). Raw data from this track were corrected for the effects of atmospheric refraction. Resulting geometric azimuths, elevations, and slant ranges were used to calculate the geometric coordinates of the RV relative to the radar station. No smoothing was used in the process. An estimate of the initial position, velocity, and reentry angle of the RV was made from corrected radar data. Using these initial conditions, a trajectory for the RV was computed to obtain theoretical coordinates relative to the radar station. Aerodynamic drag on the RV was calculated on the basis of TM from the flight test and the characteristics predicted for the RV during re-entry. Drag characteristics of the RV changed during re-entry as a consequence of generator disassembly (Figure 28).


Figure 28. Ballistic coefficient as a function of time for the postflight theoretical calculation of re-entry trajectory

The standard 1959 ARDC atmosphere was used to define density, temperature. and pressure. This agreed closely with the measurements discussed in SC-RR-64-517.

The difference between the radar and theoretical values was determined for each coordinate. The resulting differences were then plotted as a function of time, and a straight line was fitted to the plot. The slope of the straight line was taken to be the correction in initial value of the corresponding velocity, while the intercept at the initial time was applied as the correction to the coordinate. The trajectory was then repeated, using the corrected values. This iterative procedure was continued until coordinate deviations were randomly dispersed about zero. The matching was based upon the points every 5 seconds from 259.4 to 349.4 seconds. In this range, only five points ( $259.4,264.4,269.4$, 329.4 , and 344.4 seconds) exceeded 250 feet (Figure 29). The disagreement between the radar and theoretical RV trajectories is within the expected accuracy of the radar.


Figure 29. Coordinate distances between radar-observed and theoretical trajectories

With agreement between the radar data and the theoretical RV trajectory established, a comparison of this trajectory with the plate-film trajectory was made. Coordinates from the theoretical RV trajectory were used to compute apparent azimuths and elevations relative to the camera locations at High Point. Those lines agreed with the corresponding lines computed from the plate-camera data to within 1.5 minutes of arc ( 300 feet) from 338 to 392 seconds. Deviation in time, measured from plate-camera chops, from the theoretically computed time at common azimuths and elevations did not exceed 0.25 second at any point. Since the radar lost contact with the RV C-band beacon at 360.00 seconds ( 168,000 feet), the theoretical trajectory from this point to impact was used to define that portion of the RV flight path.

The close agreement of the RV theoretical trajectory with both radar and plate-camera films indicated that this method was adequate for identification of the remaining re-entry objects. Theoretical trajectories were then computed for all the re-entry objects. The initial conditions for the various re-entry items were taken from the RV trajectory at the time of their ejection from the RV. Times of ejection were recorded during the flight from telemetered switches. These ejection times are shown in Table VII. The direction of ejection of each item was calculated from the telemetered orientation of the RV as measured by a roll-stabilized, free-gyro system. Orientation of the RV and components at launch is shown in Figure 30A. Orientation and ejection directions and velocities for the fuel rods, brackets, and holders is shown in Figure 30B. Ejection velocities and relative directions were measured during the preflight tests discussed in Section II.

TABLE VII
RFD-1 Ejection Times

| Time from Launch (sec) | Items Ejected |
| :---: | :---: |
| 282.09 | 12 experimental fuel rods |
|  | 2 fuel-rod brackets |
|  | 2 fuel-rod holders |
| 322.94 | 2 reflectors |
|  | 4 reflector springs |
|  | 1 reflector band |


A. Orientation at launch
B. Orientation at ejection

Figure 30. Ejection velocities for components ejected in the external fuel-rod experiment

In each case, the resulting ejection velocity and direction were applied as a perturbation of the initial conditions taken from the RV. Theoretical trajectories for each re-entry item are presented in Tables VIII through XX. A large scale graph similar to the graph of the measured trajectories described on page 46 was then constructed. However, an overlay combining both measured and computed trajectories proved too cumbersome for practical presentation. Figures 31 and 32 present similar information at a greatly distorted scale. The abscissa represents the azimuth of each object. The ordinate scale, however, represents the variation in elevation of each object from the RV. The advantages of this method of presentation are:

1. It allows expansion of the vertical scale, which provides more separation between individual objects.
2. The differential effect of drag and ejection conditions can be seen more readily, since, in differential plotting of the vertical scale, only the relative curvature of the various trajectories is apparent.
3. The entire trajectory is brought within the span of vision as a result of contracting the horizontal scale. This has a twofold effect: (a) it accentuates the differential behavior of the object, and (b) it facilitates reconstruction of the curves from their disconnected line segments.
4. The combination of differential plotting and contraction of the horizontal scale makes possible comparison between the behavior of the objects relative to one another as recorded in the optical data and their behavior relative to one another as predicted by the theoretical trajectory. Consequently, correlation of the two types of data can be established by both the position and shape of the resulting curve. This eliminates the confusion which would be created by a slight translation of the entire system of streaks in the vertical direction.

Comparison of the measured and theoretical trajectories shown in Figures 31 and 32 revealed the following:

1. The actual re-entry patterns of 13 of the 15 re-entry objects agreed with the patterns predicted for them. The plotted theoretical trajectories paralleled the plotted trails taken from Sandia Plate Cameras T-2 and P-2 for the major portion of the re-entry. During the latter part of the trajectories, most measured objects fell slightly lower than the theoretical trajectories, as a result of ablation and the associated reduction in $W / C D A$. Each group of three fuel rods was considered to be visible as one object. Also, considering their angles of ejection, it was assumed that two reflector springs appeared as one object. The two fuel-rod holders did not follow any of the predicted trajectories, but rather re-entered at a much steeper angle.
2. The average deviation in elevation between the measured and theoretical trajectories was within 2 minutes of arc ( 400 feet) for all the objects except the fuel-rod holders. This is within the overall accuracy of the optical system and measurements. The extremely low $W / C_{D} A$ of the fuel-rod holders caused them to decrease in velocity before a high temperature was reached. It is possible that they impacted intact.

The correlation between actual and predicted trajectories and the identification of objects is considered valid in spite of the deviation in elevation. Small inaccuracies in the $W / C_{D A}$, or perturbations in initial conditions, could cause such deviations; in addition, the limitations in accuracy of the optical system (mentioned above) contributed.

RFD-1 RV Re-entry Trajectory










VT
-64
GMMA
$-644$ DATA-D=

| RVE | 3?.3467=LAT |  | -64.6538=LUNG |  | - $\quad 0.0000=0 R N T$ |  |  | $64=E L E V$ |  | -64484 $=x$ |  | -36074 $=$ Y. |  | $-0=2$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $W^{\top}$ | CD AREA |  | VI GA | MMAI | U1 |  | LAT | NG | $V T$ | M | A TAU |  |  |  |
| 51.1 | 1.001 .00 | 11 | \% |  | 4391 |  | 276 | 22 | 1. | 10. | .778 |  |  |  |
| TIME | HITE | LAT | LONG | $\checkmark T$ | gamma | TAU | 0 | ac | DOC | ${ }^{1}$ | RaNGE | SR | A ZM | ELOBS |
| 382.00 | 117640 | 29.91 | -65.22 | 10521 | -10.3 | -45.8 | 786.644 | 6671.2 | 79.9 | -15. | 303.437 | 867076. | 72.27 | 6.580 |
| 384.00 | 113997 | 29.87 | -65.18 | 9555. | -10.6 | -45.8 | 763.672 | 6814.9 | 64.0 | -15. | 306.666 | 878763. | 173.29 | 6.226 |
| 386.00 | 110602 | 29.84 | -65.14 | 8627. | -10.9 | -45.8 | 726.344 | 6928.6 | 50.1 | -14. | 309.587 | 889576. | 174.18 | 5.905 |
| 388.00 | 107443 | 29.81 | -65.10 | 7751. | 11.3 | -45.9 | 682.380 | 7016.8 | 38.6 | -13. | 312.215 | 899488. | -174.97 | 5.614 |
| 390.00 | 10450 A | 29.78 | -65.07 | 6940. | -11.8 | -45.9 | 625.191 | 7084.1 | 29.1 | -12. | 314.569 | 908508. | -175.66 | 5.353 |
| 392.00 | 101778 | 29.75 | -65.04 | 6199. | -12.3 | -45.9 | 570.402 | 7134.7 | 21.8 | -11. | 316.671 | 916667. | 176.26 | 5.116 |
| 94.00 | 99235. | 29.73 | -65.02 | 5532. | 12.8 | -45.9 | 512.200 | 7172.3 | 16.2 | -9.8 | 318.544 | 924018. | -176.79 | 4.900 |
| 96.00 | 96858. | 29.71 | -65.00 | 4935. | -13 | -45. | 458.845 | 7200.3 | 12.0 | -8.7 | 320.211 | 930621. | 177.25 | 4.704 |
| 398.00 | 94630. | 29.69 | -64.98 | 4406. | -14.2 | -45 | 405.960 | 7220.9 | 8.82 | -7.7 | 321.694 | 936540. | 177.65 | 4.524 |
| 400.00 | 92531. | 29.68 | -64.96 | 3938. | 15.0 | -46.0 | 362.040 | 7236.1 | 0.54 | -6.8 | 325.015 | 941843. | 178.01 | 4.359 |
| 402.00 | 90544. | 29.66 | -64.94 | 3527. | 15 | -46.0 | 318.903 | 7247.4 | 4.84 | -6.0 | 324.192 | 946590. | 178.32 | 4.206 |
| 404.00 | 88654. | 29.65 | -64.93 | 3167. | -16.9 | -46.0 | 283.832 | 7255.8 | 3.62 | -5.3 | 325.243 | 950842. | 178.60 | 4.064 |
| 406.00 | 86848. | 29.64 | -64.92 | 2849. | -18.1 | -46.0 | 251.952 | 7262.1 | 2.72 | -4.6 | 326.182 | 954652. | 178.84 | 3.932 |
| 408.00 | 85114. | 29.63 | -64.91 | 2571. | -19.3 | -46.0 | 222.567 | 7266.8 | 2.05 | -4.0 | 327.022 | 958066. | 179.06 | 3.307 |
| 410.00 | 83440. | 29.62 | -64.90 | 2327. | -20.7 | -46.0 | 199.523 | 7270.4 | 1.57 | -3.6 | 327.775 | 961129. | -179.25 | 3.688 |
| 412.00 | 81818. | 29.61 | -64.89 | 2113. | -22.2 | -46.0 | 178.302 | 7273.2 | 1.20 | -3.1 | 328.452 | 963880. | 179.43 | 3.576 |
| 414.00 | 80240 . | 29.60 | -64.88 | 1926. | -23.9 | -46.0 | 159.273 | 7275.3 | . 931 | -2.1 | 329.060 | 966352. | -179.58 | 3.469 |
| 416.00 | 78698. | 29.50 | -64.87 | 1762. | -25.7 | -46.0 | 144.182 | 7276.9 | . 731 | -2.4 | 329.609 | 968576. | 179.72 | 3.366 |
| 418.00 | 77187. | 29.59 | -64.86 | 1617. | -27.6 | -46.0 | 130.761 | 7278.2 | . 579 | -2.1 | 330.104 | 970578. | 179.84 | 3.266 |
| 420.00 | 75702. | 29.59 | -64.86 | 1491. | 29.6 | -46.1 | 118.833 | 7279.3 | . 464 | -1.8 | 330.551 | 972381. | 179.96 | 3.169 |
| 422.00 | 74238. | 29.58 | -64.85 | 1380. | 31.8 | -46 | 109.257 | 7280.1 | . 376 | -1.6 | 330.955 | 974005. | 179.94 | 3.074 |
| 424.00 | 72792. | 29.58 | -64.85 | 1282. | 34 | -46 | 101.397 | 7280.8 | . 310 | -1.4 | 331.322 | 975468. | 179.85 | 2.981 |
| 426.00 | 71362. | 29.57 | 64.84 | 1196. |  | -46. | 94.2903 | 7281.4 | . 257 | 1.3 | 331.653 | 976786. | 179.77 | 2.989 |
| 8.00 | 69947. | 29.57 | -64.84 | 112 | 39 | -46 | 88.0607 | 7281.8 | . 216 | -1.1 | 331.954 | 977972. | 179.70 | 2.799 |
| 430.00 | 68544. | 29.5 | -64.84 | 1054. | -41 | -46 | 83.7586 | 7282.2 | . 184 | -.98 | 332.226 | 979038. | 179.63 | 2.711 |
| 432.00 | 67153. | 29.56 | -64.83 | 994.is | -44.1 | -46.1 | 79.7462 | $72^{82} 2.6$ | . 159 | 87 | 332.473 | 979995. | 179.57 | 2.625 |
| 434.00 | 65775. | 29.56 | -64.83 | 941.4 | -46.8 | -46.1 | 76.1382 | 7282.9 | .138 | 77 | 332.695 | 980852. | 179.51 | 2.539 |
| 436.00 | 64409. | 29.56 | -64.83 | 894.9 | -49.4 | -46.1 | 73.4033 | 7283.1 | .121 |  | 332.897 | 981618. | 179.46 | 2.456 |
| 438.00 | 63057. | 29.56 | -64.82 | 853.2 | -52.1 | -46.1 | 71.4504 | 7283.4 | . 108 | 61 | 333.079 | 982301. | 179.42 | 2.373 |
| 440.00 | 61718. | 29.55 | -64.82 | 815.7 | 54.7 | -46.1 | 69.5767 | 7283.6 | .097 | . 55 | 333.242 | 982907. | 179.38 | 2.292 |
| 442.00 | 60395. | 29.55 | -64.82 | 782.1 | -57.2 | -46.0 | 67.8392 | 7283.7 | . 087 | 49 | 333.389 | 983444. | 179.34 | 2.212 |
| 444.00 | 59088. | 29.55 | -64.82 | 751.0 | -59.7 | -46.0 | 66.8887 | 7283.9 | . 080 | . 45 | 333.521 | 983917. | 179.31 | 2.133 |
| 446.00 | 57799. | 29.55 | -64.82 | 723.9 | -62.1 | -46.0 | 66.1413 | 7284.1 | . 073 | .41 | 333.639 | 984332. | 179.28 | 2.056 |
| 448.00 | 56530. | 29.55 | -64.82 | 698.2 | -64.4 | -46.0 | 65.2968 | 7284.2 | . 067 | 38 | 333.745 | 984693. | 179.26 | 1.980 |
| 450.00 | 552 Al . | 29.55 | -64.81 | 674.7 | -66.6 | -46.0 | 64.4313 | 7284.3 | . 062 | 35 | 333.838 | $98500{ }^{7}$ | 179.23 | 1.906 |
| 452.00 | 54053. | 29.55 | -64.81 | 652.9 | -68.7 | -46.0 | 64.1822 | 7284.5 | . 058 | 33 | 333.921 | $985277^{\circ}$ | 179.21 | 1.833 |
| 454.00 | 52848. | 29.55 | -64.81 | 632.2 | -70.7 | -46.0 | 63.9178 | 7284.6 | . 054 | . 31 | 333.995 | $98550{ }^{\circ}$. | 179.19 | 1.762 |
| 456.00 | 51667. | 29.54 | -64.81 | 612.6 | -72.5 | -45.9 | 63.4514 | 7284.7 | . 050 | . 29 | 334.060 | 985704. | 179.18 | 1.692 |
| 458.00 | 50511. | 29.54 | -64.81 | 594.3 | -74.2 | -45. | 62.8793 | 7284.8 | .047 | 27 | 334.117 | 985869. | 179.16 | 1.624 |
| 460.00 | 49380. | 29.54 | -64.81 | 577.1 | -75.8 | -45.9 | 62.6355 | 7284.9 | . 044 | . 26 | 334.167 | 986005. | 179.15 | 1.558 |
| 462.00 | 48273. | 29.54 | -64.81 | 560.6 | -77.3 | -45.9 | 62.5176 | 7284.9 | . 041 | -. 25 | 334.211 | 986116. | 179.14 | 1.493 |
| 464.00 | 47192. | 29.54 | -64.81 | 544.6 | -78.7 | -45.8 | 62.1643 | 7285.0 | . 038 | -. 24 | 334.249 | 986205. | 179.13 | 1.430 |
| 466.00 | 46137. | 29.54 | -64.81 | 529.5 | -79.9 | -45.8 | 61.6636 | 7285.1 | . 036 | 23 | 334.282 | 986275. | 179.12 | 1.368 |
| 468.00 | 45106. | 29.54 | -64.81 | 515.3 | -A1. 1 | -45.7 | 61.0921 | 7285.2 | . 034 | -. 21 | 334.310 | 986329. | 179.12 | 1.308 |
| 470.00 | 44100. | 29.54 | -64.81 | 501.8 | -R2.1 | -45.6 | 61.0220 | 7285.2 | . 032 | -. 21 | 334.335 | 986367. | 179.11 | 1.249 |
| 472.00 | 43118. | 29.54 | -64.81 | 488.5 | -R3. 1 | -45.6 | 60.7693 | 7285.3 | . 030 | -. 20 | 334.356 | 986393. | 179.11 | 1.192 |
| 474.00 | 42150. | 29.54 | -64.81 | 475.8 | -83.9 | -45.5 | 60.3630 | 7285.4 | . 028 | . 19 | 334.374 | 986408. | 179.10 | 1.136 |
| 476.00 | 41225. | 29.54 | -64.81 | 463.8 | -84.7 | -45.4 | 59.8640 | 7285.4 | . 027 | -. 18 | 334.390 | 986414. | 179.10 | 1.08 ? |
| 478.00 | 40313. | 29.54 | -64.81 | 452.5 | -85. 4 | -45.3 | 59.3233 | 7285.5 | . 025 | -. 17 | 334.403 | 986413. | 179.09 | 1.029 |
| 490.00 | 39421. | 29.54 | -64.81 | 441.0 | -96.0 | -45.2 | 59.0249 | 7285.5 | . 024 | -. 16 | 334.415 | 986404. | 179.09 | . 9767 |


| ITA7622 S CONTRULJI | $\begin{array}{lll} 56-64 & 51 \\ \text { JIM AUG } \end{array}$ |  | 000631 | 041203 | BR1 | DATA－B＝RFDA |  | a refraction |  | НАTA-C: | AIMOS= | theoretl | CAL ARDC UATA－D＝ | 1959 ATMOSPHERE |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| UBSERVER | 32．5407＝LAT |  | －64．65S8mLUNG |  |  | O．ODOD＝URNT |  | 643ELEV |  | －04484＝x |  | －36074EY． |  | $-0=2$ |
| WI | CD AHEA |  | Vi GiA | MHAI |  | aul | LAT | LUNG | VT | gamma | TAU |  |  |  |
| 6.451 | 1.001 .00 | 209 | 3． | 6380 | 0.740 | 500 33 | 7000－69 | －69．45 | 4 | ．2513 | －43．183 |  |  |  |
| TIME | HItt | LAI | LONG | v 1 | GAMMA | TAU |  | 0 UC | DOC | A！ | hange | SK | A2M | ELOBS |
| 282．00 | 370528 | 33.47 | －69．45 | 19974 | －6． 25 | 5－43．2 | ． 014192 | 2.00000 | 1.37 | ．100 | ，UU0000 | 1535275 | －71．203 | 11.86 |
| 284．00 | 366157 | 33． 39 | －69．36 | 19981 | －6：31 | －43．2 | ． 018185 | 5 2．9148 | 1.55 | .14 | 6.41953 | 1500152 | －71．906 | 12.07 |
| 286．00 | 361743 | 33． 32 | －69．20 | 1998） | －6．37 | －43．3 | ． 022470 | 0 6．2047 | 1.73 | ．14 | 12．8418 | 1465246 | －72．645 | 12.28 |
| 288．00 | 357280 | 33． 25 | －69．1／ | 19994 | －6．43 | －43－4 | －029890 | 9．8846 | 2.99 | ．14 | 19．2067 | 1430574 | －73．421 | 12.49 |
| 290．00 | 352780 | 33．17 | －69．08 | cuoui | －6．49 | －43：4 | ． 039395 | 514.167 | 2.29 | .10 | 23．6944 | 1396135 | －74．237 | 12.71 |
| 292．00 | 346244 | 33．10 | －68．99 | cuovi | －6．55 | －43．5 | ． 050628 | 19．052 | 2.00 | ．14 | 32.1249 | 1362009 | －75．095 | 12.93 |
| 294：00 | 343658 | 33.02 | －68．89 | 20014 | －6． 61 | －43－5 | .1164585 | 24．518 | 2.93 | .10 | 30－5580 | 1328164 | －76．000 | 13.15 |
| 296．00 | 339030 | 32．75 | －68．80 | cu0zu | －6． 6.67 | －43．6 | －041040 | 0 50．798 | 3.29 | .14 | 44.9938 | 129463s | －76．953 | 13.38 |
| 298．00 | 334359 | 32．58 | －68．72 | 20021 | －6．73 | －43．7 | $\underline{-146623}$ | 3 57．836 | 3.77 | ． 14 | 51.4324 | 1261454 | －77．959 | 13.61 |
| 300.00 | 329646 | 32－80 | －68．62 | 20035 | －6．79 | －43．7 | －134136 | 6 45．846 | 4.23 | ． 0 y | 57.8735 | 122863s | －79．021 | 13.84 |
| 302.00 | 324889 | 32．13 | －68．52 | 20039 | －6．85 | －43． | $\underline{.182622}$ | 254.918 | 4.94 | ． 0 ¢ | 64.3172 | 1196265 | －80．144 | 14．04 |
| 304：00 | 320090 | 32．63 | －68．43 | 20044 | －6－91 | 1－43．8 | － 251588 | 865.496 | 3.56 | ． 0 | 700．7634 | 1164324 | －81．331 | 14.32 |
| 306．00 | 315249 | 32． | －08．34 | 20045 | －6．97 | －43．9 | ． 326276 | ¢ 77.597 | 0.61 | ．0\％ | 77.2118 | 1132812 | －82．587 | 14.56 |
| 308.00 | 310366 | 32：${ }^{\text {2 }}$ | －68．23 | cuoss | －7－02 | －43．9 | $\because 42<711$ | 191.743 | 7.52 | ． 05 | 35．0623 | 1101952 | －83．918 | 14.79 |
| 310.00 | 305440 | 32．45 | －68．16 | くu0bs | －7：08 | －44．0 | ． 012152 | 108．21 | 9.05 | ． 0.8 | 90.1145 | 1071614 | －85．327 | 15.03 |
| 312.00 | 300475 | 32．35 | －68．07 | 20050 | －7－14 | －44－1 | .810589 | 9127.09 | 10.4 |  | 96.5679 | 1041914 | －86．821 | 15.27 |
| 314.00 | 295465 | 32．28 | －67．97 | cuus4 | －7－20 | －44．1 | 1.14846 | 6150.38 | 12.4 | －． 00 | 103.022 | 1012904 | －88．405 | 15.50 |
| 316.00 | 290417 | 32.20 | －67．88 | 20040 | －7－26 | － 44.2 | 1．20292 | 2176.85 | 14.2 | －．11 | 109.475 | 984659． | －90．083 | 15.73 |
| 318.00 | 285329 | 32.13 | －67．79 | cooso | －7－32 | －44．2 | 2.13378 | 8 207．18 | 16.8 | －． 2 i | 117.927 | 957240 ． | －91．863 | 15.95 |
| 320.00 | 280205 | 32． 05 | －67．70 | ＜U0＜＜ | －7－38 | －44．3 | 2.19095 | 5 243．88 | 19.2 | －． 31 | $12<.375$ | 930746. | －93．747 | 16.16 |
| 322.00 | 275141 | 31．98 | －07．61 | 19990 | －7．44 | －44．5 | 5.98746 | 285．17 | 22.9 | －． 4 4 | 120.817 | 905244. | －95．742 | 16.36 |
| 324.00 | 269045 | 31－90 | －67．32 | 19934 | －7－50 | －44 | 5.22935 | 5334.85 | 26.1 | －． 6 y | 135.249 | 890835. | －97．849 | 16.54 |
| 326．00 | 264619 | 31－03 | －67．4s | 19904 | －7：56 | －44 | 7.36817 | 7391.49 | 30.8 | －1．0 | 141.668 | 857620. | －100．07 | 16.70 |
| 324：00 | 259507 | 31： 5 | －67．34 | 19828 | －7：63 | －44 | 9.99440 | － 457.22 | 54.8 | －1．9 | 144．066 | 835705. | －102．41 | 16.83 |
| 330.00 | 254190 | 31－08 | －07．25 | 19720 | －7．69 | －44 | 12.6351 | 1531.46 | 39.5 | －1．00 | 154．437 | 815202. | －104．86 | 16.94 |
| 332．00 | 248812 | 31．00 | －67．17 | 19591 | －7．76 | －44 | 15．9226 | ＋ 614.16 | 43.7 | －2．j | 10 u .770 | 796219. | －107．42 | 17.01 |
| 334.00 | 243525 | 31－93 | －07．08 | 19420 | －7．82 | －44 | 20.2480 | －706．11 | 48.3 | －3．0］ | 16\％．054 | 778861. | －110．08 | 17.04 |
| 336.00 | 238243 | 31．46 | －66．99 | $1920<$ | －7－89 | －44．7 | 24.9926 | 6 807．53 | 52.4 | －3．80 | 173．277 | 763220. | －112．82 | 17.03 |
| 338.00 | 232981 | 31－38 | －06．91 | 1895S | －7－97 | 7－44．8 | 30.6905 | $5916 .<4$ | 56.3 | －4．0 | 179.422 | 749394. | －115．64 | 16.97 |
| 340.00 | 2271b1 | 31． 31 | －66．82 | i86U4 | －8－05 | －44．8 | 30.9002 | 21032.1 | 59.4 | －5．0 | 185.471 | 737424. | －118．50 | 16.86 |
| 342.00 | 222569 | 31－く4 | －66．74 | 18214 | －8．13 | －44．9 | 43.6943 | 31153.4 | 01.7 | －6．？ | 191.403 | 727348. | －121．39 | 16.70 |
| 344：00 | 217454 | 31．10 | －06．00 | 17746 | －8．22 | －44．9 | 90．8717 | 71278.3 | 05.0 | －7．0 | 191．198 | 719164 ： | －124．28 | 16.49 |
| 346.00 | 212425 | 31－11 | －66．34 | 17212 | －8．32 | －45．0 | 57.9921 | 12404.5 | 62.9 | －8，y | 204.832 | 71282 C ． | －127．13 | 16.24 |
| 348：00 | 207494 | 31－04 | －66．51 | 16004 | －8：43 | －45．0 | 04.9909 | 1529．3 | 01.7 | －10： | 208． 282 | 708259． | －129．93 | 15.94 |
| 350.00 | 202047 | 30－78 | －60．44 | 1ヶ931 | －8．56 | －45．1 | 71．2519 | 9 1650.2 | 59.0 | －11－ | 213．525 | 705337. | －132．64 | 15.61 |
| 352．00 | 198019 | 30－92 | －66．37 | 1520＜ | －8．69 | －45．1 | 76.0699 | 9 1764．6 | 53.1 | －12－ | 218.542 | 703909 ． | －135．24 | 15.25 |
| 354：00 | 193503 | 30．87 | －06．3U | 14430 | －8．84 | －45．1 | 80.1517 | 71870.5 | 50.6 | －12－ | 223.315 | 703798. | －137：70 | 14.87 |
| 356．00 | 189150 | 30－31 | －66．24 | is62s | －9．01 | －45．2 | 82.7602 | 21966.6 | 45．4 | －13－ | 221.833 | 704811． | －140：03 | 14.48 |
| 358.00 | 184964 | 30－76 | －66．18 | $12790^{\circ}$ | －9．20 | －45．2 | 83.8080 | 2052．0 | 59.9 | －13： | 232．087 | 706751 ． | －142：19 | 14.09 |
| 360.00 | 180904 | 36． 72 | －00．13 | 11964 | －9．41 | －45．2 | 83.0643 | 3126.4 | 34.5 | －13： | 230.072 | 709427． | －144：21 | 13.69 |
| 362.00 | 177138 | 30.67 | －66．08 | 11144 | －9．65 | －45．3 | 42．2071 | 12190.3 | 29.4 | －13－ | 2Sy． 790 | 712662． | －146．06 | 13.50 |
| 364：00 | 173488 | 30.03 | －06．us | 10341 | －9．91 | －45．3 | $80: 1402$ | 2 2244．3 | 24.7 | －12－ | 243．245 | 716294． | －147：76 | 12.93 |
| 366.00 | 170011 | 30－59 | －65．99 | 9569. | －1u． 2 | －45．3 | 77.4720 | － 2289.5 | 20.5 | －12－ | 240.444 | 720183. | －149．31 | 12.56 |
| 368.00 | 160702 | 30－50 | －65．93 | 8819： | －10．5 | －45．4 | 14．4706 | 2326．8 | 16.9 | －11： | 24y． 394 | 724209. | －150．72 | 12.21 |
| 370．00 | 103550 | 30：52 | －05．91 | 8105 ． | －10．9 | －45．4 | 70．4898 | 2357．3 | 13.7 | －11： | 254.147 | 728273. | －152．00 | 11.88 |
| 372．00 | 164566 | 30－50 | －65．48 | 1434． | －11．3 | －-45.4 | 65.9658 | 83381.9 | 11.0 | －10－ | 254.595 | 732290. | －153．16 | 11.56 |
| 374．00 | 157121 | 30－47 | －65．85 | 6809 ？ | －11．8 | －45．4 | 61．5438 | 82401.6 | 8.79 | －9．9 | 250．872 | 736219. | －154．21 | 11.26 |
| 376.00 | 155u10 | 30－44 | －65．82 | －230． | －12．3 | －45．b | 20．65s7 | 2417.3 | 6.97 | －8．0． | 258.953 | 739999. | －155：15 | 10.98 |
| 378.00 | 252424 | 30．42 | －65．79 | 2697． | －12．8 | －45．3 | 52．7716 | 62429.8 | 3.55 | －8．0 | 204.853 | 743607. | －156．00 | 10.72 |
| 380.00 | 149953 | 30－40 | －65．71 | s2us？ | －13．4 | －45．5 | 48.3857 | 72439.7 | 4.37 | －7．3 | 264．586 | 747022. | －156．77 | 10.47 |





| TTAT622 3 B CONTROLJI | $\begin{aligned} & 3 B-64 \\ & \text { JIM AUG } \end{aligned}$ | SNAP | $\begin{aligned} & 000633 \\ & \text { DATA-AE } \end{aligned}$ | 091263 | H1 | DATA-R=RFDA |  | REFRACTION |  | DATA-C $=$ | ATMOS = | THFGRET | AL ARDC DATA-D= | 1959 ATMOSPHERE |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| OBSERVER | 32.34 | 7=LAT | 4 | $38=10 N$ |  | . 0000 | RNT | $64=E$ |  |  | 4484=x | -36074 | $4=Y$ 。 | $-0=2$ |
| WT | CD AREA |  | $v$ I | MMAI | Ui |  | LAT | ONG | $v T$ | GAMMA | tau |  |  |  |
| 2.00 | 1.001 .00 | O 209 | 33. | 4880 | 74000 |  | 76000 | 9.45 | 74. | . 2355 | -43.183 |  |  |  |
| TIME | HITE | LAT | LONG | $v i$ | gamma | tal | 0 | ac | DOC | A 1 | Range | SH | AZM | Elobs |
| 282.00 | 370528 | 33.47 | -69.45 | 19974 | -6.24 | -43.2 | .014192 | . 00000 | 1.37 | . 10 | . 000000 | 1535275 | -71.203 | 11.86 |
| 284.00 | 366168 | 33.39 | -69.36 | 19980 | -6.30 | -43.2 | .018174 | 2.9193 | 1.55 | .10 | 6.41965 | 1500154 | -71.906 | 12.07 |
| 286.00 | 361765 | 33.32 | -69.26 | 19986 | -6.36 | -43.3 | . 022446 | 6.2026 | 1.73 | .09 | 12.8419 | 1465251 | -72.645 | 12.28 |
| 288.00 | 357319 | 33.25 | -69.17 | 19992 | -6.41 | -43.4 | .029816 | 9.8822 | 1.99 | . 09 | 19.2667 | 1430583 | -73.421 | 12.49 |
| 290.00 | 352831 | 33.17 | -69.08 | 19998 | -6.47 | -43.4 | . 039293 | 14.154 | 2.29 | .09 | 25.6940 | 1396169 | -74.237 | 12.71 |
| 292.00 | 348300 | 33.10 | -68.99 | 20004 | - 0.53 | -43.5 | . 050442 | 19.031 | 2.59 | . 08 | 32.1237 | 1367031 | -75.095 | 12.93 |
| 294.00 | 343726 | 33.02 | -68.89 | 20009 | -6.59 | -43.5 | . 064351 | 24.544 | 2.93 | . 08 | 38.5558 | 1328191 | -75.999 | 13.15 |
| 296.00 | 339110 | 32.95 | -68.80 | 20014 | -6.65 | -43.6 | .080556 | 30.745 | 3.28 | .07 | 44.9901 | 1294674 | -76.953 | 13.38 |
| 298.00 | 334451 | 32.88 | -68.71 | 20018 | -6.71 | -43.7 | .106034 | 37.758 | 3.76 | . 06 | 51.4265 | 1261509 | -77.958 | 13.61 |
| 300.00 | 329750 | 32.90 | -68.62 | 20022 | -6.77 | -43.7 | . 132926 | 45.731 | 4.21 | . 05 | 57.8648 | 1228726 | -79.020 | 13.85 |
| 302.00 | 325008 | 32.73 | -68.52 | 20024 | -6.83 | -43.8 | .181159 | 54.810 | 4.91 | . 02 | 54.3046 | 1196360 | -80.142 | 14.08 |
| 304.00 | 320223 | 32.65 | -68.43 | 20025 | -6.89 | -43.8 | . 229805 | 65.269 | 5.53 |  | 70.7456 | 1164447 | -81.328 | 14.32 |
| 306.00 | 315397 | 32.58 | -68.34 | 20024 | -6.95 | -43.9 | . 322556 | 77.285 | 6.55 | -. 04 | 77.1872 | 1133030 | -82.582 | 14.56 |
| 308.00 | 310531 | 32.50 | -68.25 | 20014 | -7.01 | -43.9 | . 418076 | 91.306 | 7.45 | -. 09 | 33.6286 | 1102154 | -83.911 | 14.80 |
| 310.00 | 305624 | 32.43 | -68.16 | 20011 | -7.07 | -44.0 | . 602142 | 107.59 | 8.93 | -. 18 | 70.0686 | 1071873 | -85.317 | 15.04 |
| 312.00 | 300679 | 32.35 | -68.07 | 19996 | -7.13 | -44.1 | . 797607 | 126.81 | 10.3 | -. $2^{8}$ | 96.5055 | 1042244 | -86.806 | 15.28 |
| 314.00 | 295697 | 32.28 | -67.98 | 19973 | -7.19 | -44.1 | 1.1230 A | 149.11 | 12.1 | . 44 | 102.937 | 1013332 | -88.383 | 15.51 |
| 316.00 | 290680 | 32.21 | -67.89 | 19940 | -7.25 | -44.2 | 1.46834 | 175.11 | 13.8 | -. 61 | 109.360 | 985208. | -90.053 | 15.74 |
| 318.00 | 285631 | 32.13 | -67.80 | 19891 | -7.31 | -44.2 | 2.06414 | 205.10 | 16.3 | -. 91 | 115.771 | 957853. | -91.818 | 15.96 |
| 320.00 | 280554 | 32.06 | -67.71 | 19823 | -7.38 | -44.3 | 2.69140 | 239.92 | 18.5 | -1.2 | 122.164 | 931656. | -93.084 | 16.17 |
| 322.00 | 275455 | 31.98 | -67.62 | 19728 | -7.44 | -44.3 | 3.78538 | 279.89 | 21.7 | -1.8 | 128.531 | 906413. | -95.051 | 16.36 |
| 324.00 | 270341 | 31.91 | -67.53 | 19597 | -7.50 | -44.4 | 4.94123 | 325.95 | 24.3 | -2.3 | 134.863 | 682330. | -97.719 | 16.54 |
| 326.00 | 265223 | 31.83 | -67.44 | 19419 | -7.57 | -44.4 | 6.77534 | 378.06 | 28.0 | -3.3 | 141.146 | 859523. | -99.887 | 16.70 |
| 328.00 | 260112 | 31.76 | -67.35 | 19180 | -7.64 | -44.5 | 8.57796 | 436.90 | 30.7 | -4.2 | 147.362 | 838109. | -102.15 | 16.83 |
| 330.00 | 255026 | 31.69 | -67.27 | 18874 | -7.72 | -44.5 | 11.0609 | 501.15 | 33.6 | -5.4 | 153.492 | 818206. | -104.49 | 16.94 |
| 332.00 | 249982 | 31.62 | -67.18 | 18491 | -7.80 | -44.6 | 13.2608 | 570.22 | $35 . ?$ | -6.5 | 159.511 | 799916. | -106.90 | 17.01 |
| 334.00 | 244999 | 31.55 | -67.10 | 18026 | -7.88 | -44.6 | 16.3390 | 642.40 | 37.0 | -8.0 | 165.394 | 783326. | -109.37 | 17.04 |
| 336.00 | 240100 | 31.48 | -67.02 | 17467 | -7.98 | -44.7 | 18.7931 | 716.75 | 37.1 | -9.3 | 171.112 | 768493. | -111.86 | 17.04 |
| 338.00 | 235308 | 31.42 | -66.95 | 16822 | -8.08 | -44.7 | 21.9955 | 790.81 | 37.0 | -11. | 176.636 | 755445. | -114.35 | 17.00 |
| 340.00 | 230643 | 31.36 | -66.87 | 16085 | -R. 20 | -44.6 | 24.1979 | 863.24 | 35.2 | -12. | 181.937 | 744166. | -116.32 | 16.92 |
| 342.00 | 226135 | 31.30 | -66.80 | 15274 | -R. 34 | -44.8 | 26.705 | 931.65 | 33.1 | -13. | 186.989 | 734601. | -119.34 | 16.80 |
| 344.00 | 221795 | 31.24 | -66.74 | 14399 | -R. 49 | -44.9 | 28.0959 | 994.15 | 29.9 | -14. | 191.767 | 726653. | -121.57 | 16.65 |
| 346.00 | 217639 | 31.19 | -66.67 | 13490 | -R. 66 | -44.9 | 29.1683 | 1051.0 | 26.4 | -14. | 196.256 | 720190. | -123.81 | 16.48 |
| 348.00 | 213677 | 31.14 | -66.62 | 12552 | -8.85 | -45.0 | 29.5012 | 1100.3 | 22.8 | -15. | 200.447 | 715055. | -125.92 | 16.28 |
| 350.00 | 209914 | 31.09 | -66.56 | 11622 | -9.07 | -45.0 | 28.7801 | 1142.1 | 19.1 | -14. | 204.335 | 711078. | -127.90 | 16.07 |
| 352.00 | 206347 | 31.05 | -66.51 | 10712 | -9.32 | -45.0 | 28.2817 | 1177.0 | 15.9 | -14. | 207.925 | 708087. | -129.75 | 15.85 |
| 354.00 | 202972 | 31.01 | -66.47 | 9833. | -9.60 | -45.0 | 26.8853 | 1205.6 | 12.9 | -13. | 211.225 | 705918. | -131.45 | 15.62 |
| 356.00 | 199782 | 31.97 | -66.43 | 9008. | -9.92 | -45.1 | 25.0291 | 1228.7 | 10.3 | -12. | 214.249 | 704417. | -133.01 | 15.39 |
| 358.00 | 196761 | 30.94 | -66.39 | 8240. | -10.3 | -45.1 | 23.3508 | 1247.0 | 8.19 | -12. | 217.014 | 703451. | -134.44 | 15.15 |
| 360.00 | 193897 | 30.91 | -66.36 | 7530. | -10.7 | -45.1 | 21.5235 | 1261.6 | 6.48 | -11. | 219.540 | 702905. | -135.75 | 14.93 |
| 362.00 | 191176 | 30.88 | -66.32 | 6878. | -11.1 | -45.2 | 19.7101 | 1273.2 | 5.10 | -9.7 | 221.845 | 702682. | -136.94 | 14.70 |
| 364.00 | 188586 | 30.86 | -66.30 | 6285. | -11.6 | -45.2 | 17.9651 | 1282.2 | 4.01 | -8.8 | 223.947 | 702704. | -138.03 | 14.48 |
| 366.00 | 186111 | 30.84 | -66.27 | 5747. | -12.2 | -45.2 | 16.3042 | 1289.4 | 3.15 | -7.9 | 225.865 | 702907. | -139.02 | 14.27 |
| 368.00 | 183738 | 30.82 | -66.25 | 5261. | -12.8 | -45.2 | 14.7766 | 1295.0 | 2.48 | -7.2 | 227.616 | 703240. | -139.91 | 14.06 |
| 370.00 | 141457 | 30.80 | -66.22 | 4823. | -13.4 | -45.2 | 13.3759 | 1299.4 | 1.96 | -6.b | 229.217 | 703662. | -140.73 | 13.86 |
| 372.00 | 179755 | 30.78 | -66.20 | 4428. | -14.2 | -45.2 | 12.0995 | 1302.9 | 1.55 | -5.8 | 230.681 | 704144. | -141.48 | 13.66 |
| 374.00 | 177122 | 30.76 | -66.18 | 4073. | -14.9 | -45.3 | 10.9786 | 1305.6 | 1.23 | -5.2 | 232.022 | 704660. | -142.16 | 13.47 |
| 376.00 | 175050 | 30.75 | -66.17 | 3754. | -15.8 | -45.3 | 9.93584 | 1307.8 | . 984 | -4.7 | 233.253 | 705192. | -142.78 | 13.29 |
| 378.00 | 173029 | 30.74 | -66.15 | 3465. | -16.7 | -45.3 | 9.14360 | 1309.6 | . 794 | -4.3 | 234.383 | 705726. | -143.35 | 13.11 |
| 380.00 | 171055 | 30.72 | -66.14 | 3202. | 17. | -45.3 | 8.37613 | 1311.0 | . 641 | -3.9 | 235.421 | 706251. | -143.88 | 12.93 |



| TTA7622 3B-64 S CONTROLJIM AUG ? |  | SNAP | $000634$ | 091263 | H2 | DATA-B=RFDA |  | REFRACTION |  | DATA-C= | ATMOS = | THEORETICAL ARDC |  | 1959 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| CONTROLJI OBSERVER | $\begin{aligned} & \text { IN AUG } 28 \\ & 32.346 \end{aligned}$ | 7=LAT | $\begin{array}{r} \text { DATA-A }= \\ -64.6 \end{array}$ | $538=$ LON6 |  | . 00000 | A-B=RFDA RNT | REFRACT $64=E$ | ON |  | 4484=X . | -3607 | $\text { DATA-D }=$ |  |
| WT | CD area |  | VI GA | MMAI | taul |  | Lat | ONG | $V T$ | GAMMA | tau |  |  |  |
| 2.001 | 1.001 .00 |  |  | 5760 | 74000 | 033. | 67000 | . 45 | 74. | . 24 | -43.183 |  |  |  |
| TIME | HIte | Lat | LONG | $V^{7}$ | GAMHA | tau | 0 | OC | DOC | $4 T$ | RANGE | SR | A2M | Elobs |
| 282.00 | 370528 | 33.47 | -69.45 | 19974 | -6.24 | -43.2 | . 014192 | . 00000 | 1.37 | . 10 | . 000000 | 1535275 | -71.203 | 11.86 |
| 284.00 | 366161 | 33.39 | -69.36 | 19980 | -6.30 | -43.2 | .018180 | 2.9196 | 1.55 | .10 | 6.41955 | 1500153 | -71.906 | 12.07 |
| 286.00 | 361752 | 33.32 | -69.26 | 19986 | -6.36 | -43.3 | . 022459 | 6.2036 | 1.73 | . 09 | 12.8417 | 1465249 | -72.645 | 12.28 |
| 288.00 | 357300 | 33.25 | -69.17 | 19995 | -6. 42 | -43.4 | . 029856 | 9.8854 | 1.99 | . 09 | 19.2664 | 1430579 | -73.421 | 12.49 |
| 290.00 | 352805 | 33.17 | -69.08 | 19998 | -6.48 | -43.4 | . 039347 | 14.161 | 2.29 | . 09 | 25.6936 | 1396164 | -74.237 | 12.71 |
| 292.00 | 348268 | 33.10 | -68.99 | 20004 | -6.54 | -43.5 | . 050540 | 19.041 | 2.59 | . 08 | 32.123? | 1362024 | -75.095 | 12.93 |
| 294.00 | 343687 | 33.02 | -68.89 | 20009 | -6.60 | -43.5 | . 064468 | 24.560 | 2.93 | . 08 | 36.5552 | 1329183 | -75.999 | 13.15 |
| 296.00 | 339065 | 32.95 | -68.80 | 20014 | -6.66 | -43.6 | .080801 | 30.768 | 3.28 | .07 | 44.9894 | 1294665 | -76.952 | 13.38 |
| 298.00 | 334400 | 32.88 | -68.71 | 20016 | -6. 72 | -43.7 | .106314 | 37.791 | 3.76 | . 06 | 51.4258 | 1261498 | -77.958 | 13.61 |
| 300.00 | 329693 | 32.80 | -68.62 | 20022 | -6.78 | -43.7 | . 133512 | 45.778 | $4.2 ?$ | . 05 | 57.8640 | 1228714 | -79.020 | 13.84 |
| 302.00 | 324944 | 32.73 | -68.52 | 20024 | -6.84 | -43.8 | -181811 | 54.876 | 4.92 | . 02 | 64.3037 | 1196345 | -80.141 | 14.08 |
| 304.00 | 320153 | 32.65 | -68.43 | 20025 | -6.90 | -43.8 | . 23052 ? | 65.352 | 5.54 |  | 70.7447 | 1164431 | -81.327 | 14.32 |
| 306.00 | 315320 | 32.58 | -68.34 | 20024 | -6.96 | -43.9 | . 324067 | 77.390 | 6.57 | -. 05 | 77.186? | 1133011 | -82.582 | 14.56 |
| 308.00 | 310448 | 32.50 | -68.25 | 20014 | -7.02 | -43.9 | . 419711 | 91.44 .3 | 7.47 | -. 09 | 83.0274 | 1102134 | -83.910 | 14.80 |
| 310.00 | 305535 | 32.43 | -68.16 | 20011 | -7.08 | -44.0 | . 605699 | 107.77 | 8.98 | -. 18 | 90.0673 | 1071851 | -85.317 | 15.03 |
| 312.00 | 300543 | 32.35 | -68.07 | 19996 | -7.14 | -44.1 | . 801400 | 127.04 | 10.3 | -. 28 | 96.5041 | 1042220 | -86.806 | 15.27 |
| 314.00 | 295595 | 32.28 | -67.98 | 19973 | -7.20 | -44.1 | 1.13018 | 149.40 | $12 . ?$ | -. 44 | 102.936 | 1013307 | -88.383 | 15.50 |
| 316.00 | 290572 | 32.21 | -67.89 | 19934 | -7.26 | -44.2 | 1.47580 | 175.47 | 13.9 | -. 62 | 109.359 | 985182. | -90.052 | 15.73 |
| 318.00 | 285516 | 32.13 | -67.80 | 19891 | -7.32 | -44.2 | 2.07853 | 205.55 | 16.4 | -. 92 | 115.769 | 957926. | -91.818 | 15.95 |
| 320.00 | 280434 | 32.06 | -67.71 | 19822 | -7.38 | -44.3 | 2.70627 | 240.48 | 18.5 | -1.2 | 122.161 | 931628. | -93.683 | 16.16 |
| 322.00 | 275329 | 31.98 | -67.62 | 19726 | -7.45 | -44.3 | 3.91385 | 280.56 | 21.7 | -1.8 | 128.528 | 906384. | -95.650 | 16.36 |
| 324.00 | 270209 | 31.91 | -67.53 | 19594 | -7.51 | -44.4 | 4.92989 | 326.76 | 24.4 | -2.3 | 1.34.859 | 882302. | -97.718 | 16.53 |
| 326.00 | 265085 | 31.83 | -67.44 | 19415 | -7.58 | -44.4 | $6.8267 ?$ | 379.03 | 28.1 | -3.3 | 141.140 | 859496. | -99.885 | 16.69 |
| 328.00 | 259970 | 31.76 | -67.35 | 19174 | -7.65 | -44.5 | 8.63289 | 438.95 | 30.7 | -4.2 | 147.355 | 838085. | -102.14 | 16.87 |
| 330.00 | 254879 | 31.69 | -67.27 | 18866 | -7.73 | -44.5 | 11.1317 | 502.44 | 33.7 | -5.4 | 153.482 | 818185. | -104.49 | 16.92 |
| 332.00 | 249831 | 31.62 | -67.18 | 18480 | $-7.81$ | -44.6 | 13.3643 | 571.69 | 35.3 | -6.6 | 159.498 | 799901. | -104.90 | 16.99 |
| 334.00 | 244846 | 31.55 | -67.10 | 18012 | -7.89 | -44.6 | 16.4288 | 644.01 | 37.0 | -8.1 | 165.377 | 783.316. | -109.36 | 17.03 |
| 336.00 | 239945 | 31.48 | -67.02 | 17451 | -7.99 | -44.7 | 18.8844 | 718.40 | 37.1 | -9.3 | 171.090 | 768490. | -111.85 | 17.02 |
| 338.00 | 235152 | 31.42 | -66.95 | 15802 | -R.10 | -44.7 | 22.0934 | 792.47 | 37.0 | -11. | 176.609 | 755448. | -114.34 | 16.98 |
| 340.00 | 230490 | 31.36 | -66.87 | 16062 | -8.21 | -44.8 | 24.2671 | 864.82 | 35.1 | -12. | 181.903 | 744175. | -116.80 | 16.90 |
| 342.00 | 225980 | 31.30 | -66.80 | 15249 | -8. 35 | -44.8 | 26.7922 | 933.13 | 33.0 | -13. | 186.946 | 734614. | -119.22 | 16.79 |
| 344.00 | 221641 | 31. ? 4 | -66.74 | 14371 | -8. 50 | -44.9 | 28.1423 | 996.07 | 29.8 | -14. | 191.715 | 726669. | -121.55 | 16.64 |
| 346.00 | 217488 | 31.19 | -66.6A | 13460 | -8. 67 | -44.9 | 29.2250 | 1052.1 | 26.4 | -14. | 196.196 | 720205. | -123.78 | 16.47 |
| 348.00 | 213529 | 31.14 | -66.62 | 12521 | - -8.86 | -45.0 | 29.5137 | 1101.? | 22.7 | -15. | 200.377 | 715067. | -125.89 | 16.27 |
| 350.00 | 209770 | 31.09 | -66.56 | 11591 | -9.09 | -45.0 | 28.8074 | $1142 . \mathrm{B}$ | 19.0 | -14. | 204.255 | 711085. | -127.86 | 16.06 |
| 352.00 | 206207 | 31.05 | -66.52 | 10681 | -9.34 | -45.0 | 28.2662 | 1177.4 | 15.7 | -14. | 207.835 | 708085. | -129.70 | 15.84 |
| 354.00 | 202838 | 31.01 | -66.47 | 9803. | -9.62 | -45.0 | 26.8418 | 1205.9 | 12.8 | -13. | 211.125 | 705904. | -131.40 | 15.61 |
| 356.00 | 199651 | 30.98 | -66.43 | 8979. | -9.94 | -45.1 | 24.9936 | 1228.8 | 10.2 | -12. | 214.139 | 704391. | -132.96 | 15.38 |
| 358.00 | 196635 | 30.94 | -66.39 | 4?13. | -10.3 | -45.1 | 23.2969 | 1247.0 | 8.12 | -11. | 216.895 | 703410. | -134.38 | 15.14 |
| 360.00 | 193775 | 30.91 | -66.36 | 7505. | -10.7 | -45.1 | 21.4726 | 1261.5 | 6.42 | -11. | 219.41? | 702847. | -135.69 | 14.92 |
| 362.00 | 191058 | 30.89 | -66.33 | 6455. | -11.1 | -45.1 | 19.6509 | 1272.9 | 5.05 | -9.6 | 221.709 | 702608. | -133.87 | 14.69 |
| 364.00 | 188471 | 30.86 | -66.30 | 6263. | -11.6 | -45.2 | 17.9120 | 1281.9 | 3.97 | -8.8 | 223.804 | 702613. | -137.95 | 14.47 |
| 366.00 | 185999 | 30.94 | -66.27 | 5727. | -12.2 | -45.2 | 16.2488 | 1289.0 | 3.12 | -7.9 | 225.715 | 702799. | -138.94 | 14.26 |
| 388.00 | 183630 | 30.82 | -66.25 | 5243. | -12.8 | -45.2 | 14.7284 | 1294.5 | 2.46 | -7.1 | 227.460 | 703116. | -139.83 | 14.06 |
| 370.00 | 181351 | 30.30 | -66.23 | 4806. | -13.5 | -45.2 | 13.3279 | 1298.9 | 1.94 | -6.4 | 229.055 | 703523. | -140.65 | 13.85 |
| 372.00 | 179152 | 30.78 | -66.21 | 4413. | -14.2 | -45.2 | 12.0589 | 1302.3 | 1.53 | -5.8 | 230.514 | 703989. | -141.39 | 13.66 |
| 374.00 | 177021 | 30.77 | -66.19 | 4059. | -15.0 | -45.3 | 10.9390 | 1305.1 | 1.2 ? | -5.2 | 231.851 | 704490. | -142.07 | 13.47 |
| 376.00 | 174951 | 30.75 | -66.17 | 3741. | -15.9 | -45.3 | 9.90305 | 1307.3 | . 975 | -4.7 | 233.077 | $705100^{8 .}$ | -142.69 | 13.28 |
| 378.00 | 172932 | 30.74 | -66.16 | 3454. | -16.R | -45.3 | 9.11501 | 1309.0 | . 787 | -4.3 | 234.202 | 705528. | -143.26 | 13.10 |
| 380.00 | 170950 | 30.73 | -66.14 | 3192. | -17.8 | -45.3 | 9.3478? | 1310.4 | .0.63 | -3.9 | 235.237 | 706040 . | -143.78 | 12.93 |



| TTA7622 3 B CONTROLJI | 3B-64 SNA IM AUG 28 |  | $\begin{aligned} & 000650 \\ & \text { DATA-A }= \end{aligned}$ | 092463 | SPR1 |  | TA-B=RFDA | REFRAC | ON | DATA -Ca |  | EORETIC | CAL ARDC DATA-D= | 959 ATMOSPHERE $-0=2$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| OSSERVER | 32.346 | 77=LAT | -64.6 | $538=10 N$ |  | 0.0000 | ORNT | $64=$ |  |  | 4484=x | -3607 | = Y 。 | $-0=2$ |
| WT | CD AREA |  | $V I$ Game | MMAI | TAU |  | LAT | LONG | VT | GAMMA | TAU |  |  |  |
| 12.01 | 1.001 .00 | 2107 | 74. -7.10 | 0000 | 1.8340 |  | 40000 | 7.56 | 3. | -7.4375 | 4.349 |  |  |  |
| TIME | HITE | LAT | LONG | $v 1$ | GAMMA | tau | 0 | 0 C | DOC | 4 | RaNGE | SH | ZM | ELOBS |
| 323.00 | 272248 | 31.94 | -67.56 | 20123 | -7.44 | -44.3 | 4.70947 | . 00000 | 25.2 | -. 27 | . 000000 | 890857. | -96.818 | 16.48 |
| 324.00 | 269639 | 31.90 | -67.51 | 20113 | -7.47 | -44.4 | 5.39804 | 26.032 | 27.0 | -. 32 | 3.23916 | 878715. | -97.897 | 16.57 |
| 326.00 | 264394 | 31.83 | -67.42 | 20087 | -7.53 | -44.4 | 7.59915 | 85.036 | 31.9 | -. 51 | 9.71247 | 855335. | -100.15 | 16.74 |
| 328.00 | 259117 | 31.75 | -67.33 | 20048 | -7.59 | -44.5 | 9.96233 | 153.20 | 36.3 | -. 70 | 16.1760 | 833248. | -102.52 | 16.87 |
| 330.00 | 253810 | 31.68 | -67.24 | 19994 | -7.65 | -44.5 | 13.1565 | 231.14 | 41.5 | -. 97 | 22.6255 | 812563. | -105.01 | 16.98 |
| 332.00 | 248479 | 31.60 | -67.15 | 19923 | -7.71 | -44.6 | 16.7709 | 319.14 | 46.5 | -1.3 | 29.0550 | 793393. | -107.62 | 17.05 |
| 334.00 | 243126 | 31.52 | -67.07 | 19829 | -7.77 | -44.7 | 21.4717 | 417.58 | 52.0 | -1.7 | 35.4592 | 775844. | -110.35 | 17.09 |
| 336.00 | 237760 | 31.45 | -66.98 | 19708 | -7.83 | -44.7 | 26.9629 | 527.18 | 57.5 | -2.1 | 41.8295 | 760024. | -113.17 | 17.07 |
| 338.00 | 232386 | 31.38 | -66.89 | 19555 | -7.90 | -44 | 33.5103 | 647.78 | 63.1 | -2.7 | 48.1565 | 746028. | -116.09 | 17.01 |
| 340.00 | 227014 | 31.30 | -66.80 | 19363 | -7.97 | -44.8 | 41.3209 | 779.37 | 68.5 | -3.3 | 54.4287 | 733938. | -119.08 | 16.89 |
| 342.00 | 221654 | 31.23 | -66.72 | 19128 | -8.04 | -44.9 | 49.8360 | 921.29 | 73.3 | -4.0 | 60.6328 | 723813. | -122.13 | 16.72 |
| 344.00 | 216317 | 31.16 | -66.63 | 18843 | -8. 12 | -44 | 60.1055 | 1072.5 | 77.9 | -4.9 | 66.7536 | 715686. | -125.20 | 16.49 |
| 346.00 | 211019 | 31.09 | -66.55 | 18502 | -R. 20 | -45.0 | 70.2940 | 1231.6 | 81.0 | -5.7 | 72.7740 | 709557. | -128.27 | 16.20 |
| 348.00 | 205772 | 31.02 | -66.47 | 18102 | -R. 28 | -45.0 | 82.5330 | 1396.3 | 83.7 | -6.7 | 78.6754 | 705390. | -131.32 | 15.87 |
| 350.00 | 200593 | 30.95 | -66.39 | 17639 | -9.38 | -45.1 | 93.4492 | 1564.4 | 84.2 | -7.6 | 34.4380 | 703108. | -134.31 | 15.49 |
| 352.00 | 195499 | 30.88 | -66.31 | 17116 | -R.48 | -45.1 | 105.096 | 1732.4 | 83.6 | -8.6 | 90.0414 | 702598. | -137.21 | 15.07 |
| 354.00 | 190505 | 30.82 | -66.24 | 16532 | -A. 59 | -45.2 | 116.351 | 1897.9 | 81.6 | -9.6 | 95.4662 | 703712. | -140.01 | 14.62 |
| 356.00 | 185627 | 30.75 | -66.17 | 15890 | -8.71 | -45.2 | 126.556 | 2058.0 | 78.? | -10. | 100.693 | 706271. | -142.68 | 14.15 |
| 358.00 | 180881 | 30.70 | -66.10 | 15197 | -R. 85 | -45.2 | 135.217 | 2209.7 | 73.4 | -11. | 105.703 | 710078. | -145.20 | 13.67 |
| 360.00 | 176279 | 30.64 | -66.03 | 14462 | -9.00 | -45.3 | 142.093 | 2350.8 | 67.6 | -12. | 110.483 | 714921. | -147.57 | 13.19 |
| 362.00 | 171834 | 30.58 | -65.97 | 13692 | -9.17 | -45.3 | 149.031 | 2480.0 | 61.5 | -12. | 115.019 | 720588. | -149.77 | 12.70 |
| 364.00 | 167556 | 30.53 | -65.91 | 12888 | -9.36 | -45.4 | 154.254 | 2596.5 | 54.9 | -13. | 119.300 | 726869. | -151.81 | 12.23 |
| 366.00 | 163455 | 30.49 | -65.86 | 12063 | -9.57 | -45.4 | 156.734 | 2699.5 | 48.0 | -13. | 123.317 | 733561. | -153.68 | 11.77 |
| 368.00 | 159537 | 30.44 | -65.81 | 11233 | -9.80 | -45 | 156.344 | 2788.? | 41.1 | -13. | 127.066 | 740483. | -155.39 | 11.33 |
| 370.00 | 155806 | 30.40 | -65.76 | 10411 | -10.1 | -45 | 154.008 | 2864.4 | 34.7 | -13. | 130.546 | 747475. | -156.94 | 10.92 |
| 372.00 | 152260 | 30.36 | -65.72 | 9604. | -10.4 | -45. | 150.967 | 2927.8 | 28.8 | -12. | 133.762 | 754405. | -158.34 | 10.52 |
| 374.00 | 148899 | 30.33 | -65.68 | 8820. | -10.7 | -45. | 145.499 | 2980.1 | 23.6 | -12. | 136.720 | 761161. | -159.60 | 10.16 |
| 376.00 | 145717 | 30.29 | -65.64 | 8070. | -11.1 | -45.5 | 138.399 | 3022.6 | 19.0 | -11. | 139.428 | 767659. | -160.73 | 9.812 |
| 378.00 | 142708 | 30.26 | -65.61 | 7361. | -11.5 | -45.6 | 130.657 | 3056.6 | 15.1 | -11. | 141.900 | 773837. | -161.75 | 9.491 |
| 380.00 | 139862 | 30.24 | -65.58 | 6698. | -11.9 | -45.6 | 121.089 | 3083.5 | 11.9 | -9.9 | 144.148 | 779656. | -162.65 | 9.194 |
| 382.00 | 137169 | 30.21 | -65.55 | 6084. | -12.4 | -45.6 | 112.406 | 3104.6 | 9.32 | -9.2 | 146.188 | 785096. | -163.46 | 8.918 |
| 384.00 | 134617 | 30.19 | -65.53 | 5522. | -13.0 | -45.6 | 102.696 | 3121.1 | 7.23 | -8.3 | 148.037 | 790147. | -164.18 | 8.661 |
| 386.00 | 132194 | 30.17 | -65.50 | 5009. | -13.6 | -45.6 | 94.0632 | 3133.9 | 5.61 | -7.6 | 149.711 | 794817. | -164.83 | 8.421 |
| 388.00 | 129889 | 30.15 | -65.48 | 4544. | -14.3 | -45.6 | 84.9967 | 3143.7 | 4.32 | -6.8 | 151.225 | 799115. | -165.40 | 8.198 |
| 390.00 | 127688 | 30.14 | -65.47 | 4125. | -15.1 | -45.7 | 77.4943 | 3151.4 | 3.35 | -6.2 | 152.595 | 803059. | -165.92 | 7.990 |
| 392.00 | 125580 | 30.12 | -65.45 | 3747. | -16.0 | -45.7 | 69.8188 | 3157.3 | 2.59 | -5.5 | 153.834 | 806570. | -166.38 | 7.795 |
| 394.00 | 123556 | 30.11 | -65.43 | 3408. | -16.9 | -45.7 | 63.3409 | 3161.8 | 2.01 | -5.0 | 154.955 | 809968. | -166.79 | 7.611 |
| 396.00 | 121607 | 30.10 | -65.42 | 3104. | -18.0 | -45.7 | 57.2745 | 3165.4 | 1.56 | -4.5 | 155.971 | 812978. | -167.16 | 7.437 |
| 398.00 | 119722 | 30.09 | -65.41 | 2833. | -19.1 | -45.7 | 51.6536 | 3168.2 | 1.22 | -4.0 | 156.891 | 815720. | -167.50 | 7.272 |
| 400.00 | 117894 | 30.08 | -65.40 | 2590. | -70.4 | -45.7 | 47.1105 | 3170.3 | . 960 | -3.6 | 157.725 | 818216. | -167.80 | 7.115 |
| 402.00 | 116117 | 30.07 | -65.39 | 2372. | -21.7 | -45.7 | 42.7217 | 3172.0 | . 757 | -3.2 | 158.483 | 820487. | -168.07 | 6.965 |
| 404.00 | 114384 | 30.06 | -65.38 | 2177. | -23.2 | -45.7 | 38.9066 | 3173.4 | . 601 | -2.9 | 159.170 | 822550. | -168.31 | 6.821 |
| 406.00 | 112689 | 30.05 | -65.37 | 2003. | -24.7 | -45.7 | 35.7278 | 3174.4 | . 481 | -2.6 | 159.795 | 824423. | -168.53 | 6.683 |
| 408.00 | 111029 | 30.04 | -65.36 | 1847. | -26.4 | -45.8 | 32.7079 | 3175.3 | . 387 | -2.3 | 160.363 | 826122. | -168.73 | 6.550 |
| 410.00 | 109399 | 30.04 | -65.35 | 1708. | -28.2 | -45.8 | 30.1271 | 3176.0 | . 314 | -2.0 | 160.880 | 827661. | -168.91 | 6.421 |
| 412.00 | 107796 | 30.03 | -65.35 | 1584. | -30.1 | -45.8 | 28.0124 | 3176.6 | . 257 | -1.8 | 161.351 | 829055. | -169.08 | 6.296 |
| 414.00 | 106218 | 30.03 | -65.34 | 1472. | -32.2 | -45.8 | 26.0000 | 3177.0 | . 211 | -1.6 | 161.779 | 830314. | -169.23 | 6.174 |
| 416.00 | 104661 | 30.02 | -65.34 | 1372. | -34.3 | -45.8 | 24.2401 | 3177.4 | . 176 | -1.b | 162.168 | 831450. | -169.36 | 6.056 |
| 418.00 | 103124 | 30.02 | -65.33 | 1283. | - 36.5 | -45.8 | 22.9087 | 3177.8 | . 148 | -1.3 | 102.522 | 832473. | -169.48 | 5.940 |
| 420.00 | 101007 | 30.01 | -65.33 | 1202. | -38.8 | -45.8 | 21.6202 | 3178.0 | . 125 | -1.2 | 162.845 | 833392. | -169.59 | 5.827 |


| TTA7622 CONTROLJ | $\begin{array}{ll} 3 B-64 & S N \\ \text { JIM AUG } 28 \end{array}$ |  | $\begin{aligned} & 000650 \\ & \text { DATA-AE } \end{aligned}$ | 092463 | SPR1 | DATA-B=RFDA |  | REFRACTION |  | DATA-C= |  | THEORETICAL ARDC DATA-D= |  | 1959 ATMOSPHERE |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| OBSERVER | 32.346 | 7=LAT | -64.6 | $538=L O N$ |  | . 0000 | ORNT | $64=$ | EV |  | - | -36074 | =Y. | -0 2 |
| WT | CD AREA |  | VI | MMAI | taul |  | Lat | LONG | $v T$ | GAMM | TAU |  |  |  |
| 12.0 | 1.001 .00 | 214 | . 3 -20 | 2765 | 59413 |  | 10759 | 5.32 | . 4 | 41.21 | -45.782 |  |  |  |
| TIme | HITE | LAT | LONG | $\checkmark T$ | GAMMA | tau | 0 | OC | DaC | ${ }^{1} 1$ | RANGE | SR | AZM | ELOBS |
| 422.00 | 100109 | 30.01 | -65.32 | 1130. | -41.2 | -45.8 | 20.4218 | 3178.3 | . 106 | -1.0 | 163.137 | 834216. | -169.69 | 5.716 |
| 424.00 | 98628. | 30.01 | -65.32 | 1066. | -43.7 | -45.8 | 19.6319 | 3178.5 | . 092 | 95 | 163.403 | 834952. | -169.78 | 5.607 |
| 426.00 | 97165. | 30.00 | -65.32 | 1008. | -46.1 | -45.8 | 18.8602 | 3178.6 | . 080 | -. 86 | 203.644 | 835606. | -169.87 | 5.501 |
| 428.00 | 95721. | 30.00 | -65.31 | 955.6 | -48.6 | -45.8 | 18.1219 | 3178.8 | . 070 | -. 77 | 163.862 | 836186. | -169.94 | 5.398 |
| 430.00 | 94297. | 30.00 | -65.31 | 908.6 | -51.1 | -45.8 | 17.5827 | 3178.9 | . 062 | -. 69 | 164.059 | 836698. | -170.01 | 5.296 |
| 432.00 | 92891. | 30.00 | -65.31 | 865.9 | -53.6 | -45.8 | 17.1880 | 3179.0 | . 055 | -. 63 | 164.236 | 837147. | -170.07 | 5.197 |
| 434.00 | 91508. | 30.00 | -65.31 | 826.9 | -56.1 | -45.8 | 16.7710 | 3179.1 | . 049 | -. 57 | 164.396 | 837538. | -170.12 | 5.099 |
| 436.00 | 90146. | 29.99 | -65.30 | 791.7 | -58.5 | -45.8 | 16.3587 | 3179.2 | . 044 | -. 52 | 104.540 | 837876. | -170.17 | 5.003 |
| 438.00 | 88807. | 29.99 | -65.30 | 759.6 | -60.9 | -45.8 | 16.1991 | 3179.3 | . 040 | -. 48 | 164.668 | 838166. | -170-21 | 4.909 |
| 440.00 | 87492. | 29.99 | -65.30 | 729.5 | -63.2 | -45.7 | 16.0032 | 3179.4 | . 037 | -. 45 | 164.783 | 838413. | -170.25 | 4.818 |
| 442.00 | 86203. | 29.99 | -65.30 | 701.7 | -65.4 | -45.7 | 15.7707 | 3179.4 | . 033 | -. 41 | 164.885 | 838619 | -170.29 | 4.728 |
| 444.00 | 84940. | 29.99 | -65.30 | 676.2 | -67.5 | -45.7 | 15.5317 | 3179.5 | . 031 | -. 38 | 164.976 | 838790. | -170.32 | 4.641 |
| 446.00 | 83704. | 29.99 | -65.30 | 652.5 | -69.5 | -45.7 | 15.4674 | 3179.6 | . 028 | -. 36 | 105.056 | 838928. | -170.34 | 4.556 |
| 448.00 | 82496. | 29.99 | -65.30 | 629.9 | -71.4 | -45.7 | 15.3335 | 3179.6 | . 026 | -. 34 | 165.127 | 839038. | -170.37 | 4.473 |
| 450.00 | 81316. | 29.99 | -65.30 | 608.9 | -73.2 | -45.7 | 15.1624 | 3179.7 | . 024 | -. 31 | 165.189 | 839121. | -170.39 | 4.392 |
| 452.00 | 80165. | 29.99 | -65.30 | 589.4 | -74.8 | -45.6 | 14.9717 | 3179.7 | . 022 | -. 29 | 165.243 | 839182. | -170.41 | 4.313 |
| 454.00 | 79041. | 29.99 | -65.29 | 571.1 | -76 | -45.6 | 14.8919 | 3179.8 | . 021 | -. 28 | 165.291 | 839223. | -170.42 | 4.236 |
| 456.00 | 77944. | 29.98 | -65.29 | 553.7 | -77.8 | -45.6 | 14.7858 | 3179.8 | . 019 | -. 26 | 165.332 | 839247. | -170.44 | 4.162 |
| 458.00 | 76875. | 29.98 | -65.29 | 537.3 | -79.2 | -45.5 | 14.0442 | 3179.8 | . 018 | -. 24 | 165.368 | 839256. | -170.45 | 4.089 |
| 460.00 | 75833. | 29.98 | -65.29 | 522.0 | -80.4 | -45.5 | 14.4843 | 3179.9 | . 017 | . 23 | 165.399 | 839251. | -170.46 | 4.018 |
| 462.00 | 74816. | 29.98 | -65.29 | 507.7 | -81.5 | -45.4 | 14.3447 | 3179.9 | . 016 | -. 21 | 165.426 | 839235. | -170.47 | 3.949 |
| 464.00 | 73825. | 29.98 | -65.29 | 494.1 | -R2.5 | -45.3 | 14.3055 | 3179.9 | . 015 | -. 21 | 165.449 | 839210. | -170.48 | 3.882 |
| 466.00 | 72857. | 29.98 | -65.29 | 480.9 | -83.4 | -45.3 | 14.2189 | 3180.0 | . 014 | -. 20 | 165.469 | 839176. | -170.48 | 3.817 |
| 468.00 | 71913. | 29.98 | -65.29 | 468.4 | -84.2 | -45.2 | 14.1041 | 3180.0 | . 013 | -. 19 | 165.486 | 839136. | -170.49 | 3.753 |
| 470.00 | 70992. | 29.98 | -65.29 | 456.7 | -84.9 | -45.1 | 13.9756 | 3180.0 | . 012 | -. 17 | 165.501 | 839091. | -170.50 | 3.691 |
| 472.00 | 70093. | 29.98 | -65.29 | 445.7 | - 25.6 | -45.0 | 13.8426 | 3180.0 | . 012 | -. 16 | 165.514 | 839041. | -170.50 | 3.630 |
| 474.00 | 69215. | 29.98 | -65.29 | 435.2 | -86. 2 | -44.8 | 13.8161 | 3180.1 | . 011 | -. 16 | 165.524 | 838987. | -170.50 | 3.571 |
| 476.00 | 68356. | 29.98 | -65.29 | 425.0 | -86.7 | -44.7 | 13.7587 | 3180.1 | . 011 | -. 15 | 165.533 | 838931. | -170.51 | 3.513 |
| 478.00 | 67517. | 29.98 | -65.29 | 415.2 | -87.2 | -44.5 | 13.6765 | 3180.1 | . 010 | -. 15 | 165.541 | 838872. | -170.51 | 3.457 |
| 480.00 | 66697. | 29.98 | -65.29 | 405.4 | -87.6 | -44.3 | 13.5801 | 3180.1 | . 010 | -. 14 | 165.547 | 838812. | -170.51 | 3.402 |
| 482.00 | 65895. | 29.98 | -65.29 | 397.1 | -87.9 | -44.0 | 13.4786 | 3180.1 | . 009 | -. 13 | 165.553 | 838751. | -170.51 | 3.348 |
| 484.00 | 65109. | 29.98 | -65.29 | 389.0 | - 88.2 | -43.7 | 13.3774 | 3180.2 | . 009 | -. 12 | 165.557 | 838689. | -170.52 | 3. 295 |
| 486.00 | 64340 . | 29.98 | -65.29 | 381.1 | -88.5 | -43.4 | 13.3625 | 3180.2 | . 008 | -. 12 | 165.561 | 838627. | -170.52 | 3.244 |
| 488.00 | 63585. | 29.98 | -65.29 | 373.4 | -88.7 | -43.0 | 13.3320 | 3180.2 | . 008 | -. 12 | 165.565 | 838565. | -170.52 | 3.193 |
| 490.00 | 62846. | 29.98 | -65.29 | 365.9 | -R8. 9 | -42.5 | 13.2773 | 3180.2 | . 008 | -. 11 | 165.567 | 838503. | -170.52 | 3.144 |
| 492.00 | 62122. | 29.98 | -65.29 | 358.8 | -R9.1 | -42.0 | 13.2099 | 3180.2 | . 007 | -. 11 | 165.570 | 838442. | -170.52 | 3.094 |
| 494.00 | 61411. | 29.98 | -65.29 | 352.0 | -89.2 | -41.4 | 13.1377 | 3180.2 | . 007 | -. 10 | 165.572 | 838381. | -170.52 | 3.046 |
| 496.00 | 60713. | 29.98 | -65.29 | 345.6 | -89.3 | -40.6 | 13.0656 | 3180.3 | . 007 | -. 09 | 165.573 | 838321. | -170.52 | 2.998 |
| 498.00 | 60028. | 29.98 | -65.29 | 339.0 | -89.4 | -39.7 | 12.9957 | 3180.3 | . 006 | -. 09 | 165.575 | 838262 | -170.52 | 2.952 |
| 500.00 | 59355. | 29.98 | -65.29 | 333.8 | -89.5 | -38.7 | 13.0066 | 3180.3 | . 006 | -. 09 | 165.576 | 838204 | -170.52 | 2.906 |
| 502.00 | 58693. | 29.98 | -65.29 | 328.0 | -89. 6 | -37.6 | 12.9927 | 3180.3 | . 006 | -. 09 | 165.577 | 838147. | -170.52 | 2.861 |
| 504.00 | 58043. | 29.98 | -65.29 | 322.3 | -89.7 | -36.2 | 12.9579 | 3180.3 | . 006 | -.06 | 165.578 | 838091. | -170.52 | 2.817 |
| 506.00 | 57404. | 29.98 | -65.29 | 316.9 | -89.7 | -34.7 | 12.9132 | 3180.3 | . 005 | -. 08 | 165.579 | 838036. | -170.52 | 2.773 |
| 508.00 | 56775. | 29.98 | -65.29 | 311.7 | -89.8 | -33.0 | 12.8642 | 3180.3 | . 005 | -. 08 | 165.579 | 837982. | -170.52 | 2.730 |
| 510.00 | 56157. | 29.98 | -65.29 | 306.7 | -89.8 | -31.1 | 12.8143 | 3180.3 | . 005 | -. 07 | 165.580 | 837929. | -170.52 | 2.688 |
| 512.00 | 55548. | 29.98 | -65.29 | 302.1 | - 99.8 | -28.9 | 12.7655 | 3180.3 | . 005 | -. 07 | 165.581 | 837877. | -170.52 | 2.647 |
| 514.00 | 54949. | 29.98 | -65.29 | 297.0 | -99.8 | -26.7 | 12.7254 | 3180.4 | . 005 | -. 06 | 165.581 | 837826. | -170.52 | 2.606 |
| 516.00 | 54358. | 29.98 | -65.29 | 293.3 | -89.9 | -24.3 | 12.7473 | 3180.4 | . 005 | . 01 | 165.582 | 837776. | -170.52 | 2.565 |
| 518.00 | 53776. | 29.98 | -65.29 | 288.8 | -89.9 | -21.8 | 12.7417 | 3180.4 | . 004 | -. 01 | 165.582 | 837728. | -170.52 | 2.526 |
| 520.00 | 53202. | 29.98 | -65.29 | 284.5 | -89.9 | -19.3 | 12.7206 | 3180.4 | . 004 | 06 | 165.582 | 837680. | -170.52 | 2.487 |



| TTA7622 CONTROLJIM OBSERVER | 3B-64 SNA <br> JIM AUG 28 <br> 32.346 |  | $\begin{aligned} & 000651 \\ & \text { DATA-A }= \\ & -64.6 \end{aligned}$ | 092463 | SPR2 | . 0000 | TA-Gz ORNT | REFRAC |  | DATA-C | ATMOS $=$ $484=x$. | Theoretica $-3607$ | $\begin{aligned} & A L \\ & \text { DARDC } \\ & \text { ARA-D } \end{aligned}$ $=Y_{0}$ | 1959 ATMOSPHERE $-0=\mathrm{Z}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| OBSERVER | 32.34 | 7=LAT |  | ( | ONG 0.00 | .0000 |  |  |  |  |  | 36074 |  | $-0=2$ |
| WT | CD AREA |  | $\mathrm{VI} \mathrm{GA}^{\text {a }}$ | Mmal | Taul |  | LAT | ONG | $v T$ | GAMMA | TAU |  |  |  |
| 12.0 | 1.001 .00 | 2138 | 8.9-20 | 3291 | 7.43351 |  | 12396 | 5.33 |  | 41. 5 | -45.778 |  |  |  |
| TIME | HITE | Lat | LONG | $v$ T | t gamma | tau | 0 | 0 C | DOC | AT | Range | SR | A2M | ELOBS |
| 422.00 | 99882. | 30.01 | -65.33 | 1120. | . -41.5 | -45.8 | 20.2856 | 3172.5 | .104 | -1.0 | 162.615 | 833974. | -169.54 | 5.702 |
| 424.00 | 98404. | 30.01 | -65.33 | 1057. | . -44.0 | -45.8 | 19.5197 | 3172.8 | . 090 | -. 94 | 162.877 | 834694. | -169.63 | 5.594 |
| 426.00 | 96945. | 30.01 | -65.32 | 999.0 | - -46.5 | -45.8 | 18.7482 | 3172.9 | . 078 | -. 84 | 163.115 | 835334. | -169.71 | 5.489 |
| 428.00 | 95505. | 30.00 | -65.32 | 948.2 | 2-49.0 | -45.8 | 18.0153 | 3173.1 | . 068 | -. 75 | 163.330 | 835901. | -169.78 | 5.385 |
| 430.00 | 94084. | 30.00 | -65.32 | 902.0 | $0-51.5$ | -45.8 | 17.5254 | 3173.2 | . 061 | -. 68 | 163.524 | 836401. | -169.85 | 5.284 |
| 432.00 | 92682. | 30.00 | -65.32 | 859.8 | $8-54.0$ | -45.8 | 17.1250 | 3173.3 | . 054 | -. 62 | 163.699 | 836838. | -169.91 | 5.185 |
| 434.00 | 91302. | 30.00 | -65.31 | 821.4 | $4-56.5$ | -45.8 | 16.7058 | 3173.4 | . 048 | -. 57 | 163.856 | 837219. | -169.96 | 5.087 |
| 436.00 | 89944. | 30.00 | -65.31 | 786.7 | 7-58.9 | -45.8 | 16.3077 | 3173.5 | . 043 | -. 51 | 163.997 | 837548. | -170.01 | 4.992 |
| 438.00 | 88609. | 29.99 | -65.31 | 754.9 | 9-61.2 | -45.7 | 16.1727 | 3173.6 | . 040 | -. 48 | 164.124 | 837830. | -170.05 | 4.898 |
| 440.00 | 87298. | 29.99 | -65.31 | 725.2 | $2-63.5$ | -45.7 | 15.9686 | 3173.7 | . 036 | -. 44 | 164.237 | 838069. | -170.09 | 4.807 |
| 442.00 | 86013. | 29.99 | -65.31 | 697.7 | $7-65.7$ | -45.7 | 15.7322 | 3173.7 | . 033 | -. 41 | 164.337 | 838269. | $-170.13$ | 4.718 |
| 444.00 | 84755. | 29.99 | -65.31 | 672.5 | 5-67.8 | -45.7 | 15.5222 | 3173.8 | . 030 | -. 37 | 164.426 | 838433. | -170.16 | 4.631 |
| 446.00 | 83523. | 29.99 | -65.31 | 649.0 | $0-69.8$ | -45.7 | 15.4453 | 3173.9 | . 028 | . 36 | 164.505 | 838566. | -170.18 | 4.546 |
| 448.00 | 82319. | 29.99 | -65.30 | 626.7 | 7-71.7 | -45.7 | 15.3069 | 3173.9 | . 026 | -. 33 | 164.574 | 838670. | $-170.21$ | 4.463 |
| 450.00 | 81144. | 29.99 | -65.30 | 605.8 | 8-73.4 | -45.6 | 15.1323 | 3174.0 | . 024 | -. 31 | 164.635 | 838750. | $-170.23$ | 4.383 |
| 452.00 | 79996. | 29.99 | -65.30 | 586.6 | - -75.1 | -45.6 | 14.9406 | 3174.0 | . 022 | -. 29 | 164.689 | 838807. | -170.24 | 4.304 |
| 454.00 | 78877. | 29.99 | -65.30 | 568.5 | $5-76.6$ | -45.6 | 14.8816 | 3174.1 | . 021 | -. 27 | 164.735 | 838845. | -170.26 | 4.228 |
| 456.00 | 77784. | 29.99 | -65.30 | 551.2 | 2-78.0 | -45.5 | 14.7659 | 3174.1 | . 019 | -. 26 | 164.776 | 838865. | $-170.27$ | 4.153 |
| 458.00 | 76720. | 29.99 | -65.30 | 535.0 | 0-79.3 | -45.5 | 14.6206 | 3174.1 | . 018 | -. 24 | 164.811 | 838871. | $-170.29$ | 4.081 |
| 460.00 | 75681. | 29.99 | -65.30 | 519.8 | $8-80.5$ | -45.5 | 14.4599 | 3174.2 | . 017 | -. 22 | 164.842 | 838864. | $-170.30$ | 4.010 |
| 462.00 | 74668. | 29.99 | -65.30 | 505.6 | 6-81.6 | -45.4 | 14.3406 | 3174.2 | . 016 | -. 21 | 164.868 | 838847. | -170.31 | 3.942 |
| 464.00 | 73680. | 29.99 | -65.30 | 492.1 | $1-82.6$ | -45.3 | 14.2942 | 3174.2 | . 015 | -. 21 | 164.891 | 838820. | -170.31 | 3.875 |
| 466.00 | 72716. | 29.99 | -65.30 | 479.0 | $0-83.5$ | -45.3 | 14.2020 | 3174.3 | . 014 | -. 20 | 164.910 | 838785. | -170.32 | 3.809 |
| 468.00 | 71776. | 29.98 | -65.30 | 466.6 | $6-84.3$ | -45.2 | 14.0846 | 3174.3 | . 013 | -. 18 | 164.927 | 838744. | $-170.33$ | 3.746 |
| 470.00 | 70858. | 29.98 | -65.30 | 455.0 | 0 - 25.0 | -45.1 | 13.9553 | 3174.3 | . 012 | -. 17 | 164.941 | 838697. | -170.33 | 3.684 |
| 472.00 | 69962. | 29.98 | -65.30 | 444.1 | $1-85.7$ | -44.9 | 13.8275 | 3174.3 | . 012 | -. 16 | 164.953 | 838646. | -170.33 | 3.623 |
| 474.00 | 69087. | 29.98 | -65.30 | 433.7 | 7-26.3 | -44.8 | 13.8111 | 3174.4 | . 011 | -. 16 | 164.964 | 838592. | -170.34 | 3.564 |
| 476.00 | 68231. | 29.98 | -65.30 | 423.5 | $5-86.8$ | -44.6 | 13.7482 | 3174.4 | . 011 | -. 15 | 164.973 | 838535. | -170.34 | 3.507 |
| 478.00 | 67395. | 29.98 | -65.30 | 413.8 | $8-87.2$ | -44.5 | 13.6635 | 3174.4 | . 010 | -. 15 | 164.980 | 838476. | -170.34 | 3.451 |
| 480.00 | 66578. | 29.98 | -65.30 | 404.6 | - - 77.6 | -44.2 | 13.5655 | 3174.4 | . 009 | -. 14 | 164.986 | 838416. | -170.35 | 3.396 |
| 482.00 | 65778. | 29.98 | -65.30 | 395.9 | $9-88.0$ | -44.0 | 13.4637 | 3174.4 | . 009 | -. 13 | 164.992 | 838355. | -170.35 | 3.342 |
| 484.00 | 64995. | 29.98 | -65.30 | 387.8 | $8-88.3$ | -43.7 | 13.3636 | 3174.5 | . 009 | -. 12 | 164.996 | 838293. | -170.35 | 3.290 |
| 486.00 | 64227. | 29.98 | -65.30 | 380.0 | $0-88.5$ | -43.3 | 13.3647 | 3174.5 | . 008 | -. 12 | 165.000 | 838231. | -170.35 | 3.238 |
| 488.00 | 63475. | 29.98 | -65.30 | 372.3 | $3-88.7$ | -42.9 | 13.3258 | 3174.5 | . 008 | -. 12 | 165.003 | 838169. | $-170.35$ | 3.188 |
| 490.00 | 62736. | 29.98 | -65.30 | 364.8 | $8-88.9$ | -42.5 | 13.2682 | 3174.5 | .008 | -. 11 | 165.006 | 838107. | -170.35 | 3.138 |
| 492.00 | 62016. | 29.98 | -65.30 | 357.7 | 7-89.1 | -41.9 | 13.1999 | 3174.5 | . 007 | -. 10 | 165.008 | 838046. | -170.35 | 3.089 |
| 494.00 | 61307. | 29.98 | -65.30 | 351.0 | $0-89.2$ | -41.3 | 13.1274 | 3174.5 | . 007 | -. 10 | 165.010 | 837985. | -170.35 | 3.041 |
| 496.00 | 60611. | 29.98 | -65.30 | 344.7 | 7-R9.3 | -40.5 | 13.0553 | 3174.5 | . 007 | -. 09 | 165.012 | 837925. | -170.36 | 2.993 |
| 498.00 | 59928. | 29.98 | -65.30 | 338.8 | $8-89.4$ | -39.6 | 12.9947 | 3174.6 | . 006 | -. 09 | 165.013 | 837866. | -170.36 | 2.947 |
| 500.00 | 59256. | 29.98 | -65.30 | 333.0 | $0-79.5$ | -38.6 | 13.0084 | 3174.6 | . 006 | -. 09 | 165.014 | 837808. | -170.36 | 2.901 |
| 502.00 | 58596. | 29.98 | -65.30 | 327.1 | $1-89.6$ | -37.4 | 12.9878 | 3174.6 | . 006 | -. 09 | 165.015 | 837751. | -170.36 | 2.856 |
| 504.00 | 57948. | 29.98 | -65.30 | 321.5 | $5-89.7$ | -36.0 | 12.9513 | 3174.6 | . 006 | -. $0^{8}$ | 165.016 | 837695. | -170.36 | 2.812 |
| 506.00 | 57310. | 29.98 | -65.30 | 316.1 | 1-89.7 | -34.5 | 12.9058 | 3174.6 | . 005 | -. $0^{8}$ | 165.017 | 837640. | -170.36 | 2.768 |
| 508.00 | 56683. | 29.98 | -65.30 | 310.9 | $9-79.8$ | -32.7 | 12.8566 | 3174.6 | . 005 | -. 08 | 165.018 | 837586. | -170.36 | 2.726 |
| 510.00 | 56067. | 29.98 | -65.30 | 306.0 | - -89.8 | -30.8 | 12.8068 | 3174.6 | . 005 | -. 07 | 165.018 | 837534. | -170.36 | 2.684 |
| 512.00 | 55459. | 29.98 | -65.30 | 301.4 | 4 - 19.8 | -28.6 | 12.7584 | 3174.6 | . 005 | -. 07 | 165.019 | 837482. | -170.36 | 2.642 |
| 514.00 | 54861. | 29.98 | -65.30 | 297.0 | $0-89.8$ | -26.3 | 12.7290 | 3174.6 | . 005 | -. 06 | 265.019 | 837431. | -170.36 | 2.601 |
| 516.00 | 54271. | 29.98 | -65.30 | 292.6 | - -19.9 | -23.9 | 12.7494 | 3174.7 | .005 | -. 07 | 165.020 | 837382. | -170.36 | 2.561 |
| 518.00 | 53690. | 29.98 | -65.30 | 288.2 | $2-19.9$ | -21.4 | 12.7400 | 3174.7 | .004 | -. 07 | 165.020 | 837333. | $-170.36$ | 2.522 |
| 520.00 | 53118. | 29.98 | -65.30 | 283.9 | $9-89.9-1$ | -19.0 | 12.7171 | 3174.7 | .004 | -. 06 | 165.021 | 837285. | -170.36 | 2.483 |







Figure 31. Plot of theoretical re-entry trajectories for RFD-1 components


Figure 32. Plot of observed re-entry trajectories for RFD-1 components

Visual inspection of the plate-camera films shows the timing chops and distinctive events (see Figures 48A and 48B, p. 107). The two re-entry trails identified as the strontium-loaded and barium-loaded rods pulsate and flare very brightly in comparison with the other trails. The flares are caused by the tracers. The RV appears as a steady, fairly intense streak, as do the brackets. The remainder of the re-entry items are dimmer and have much less distinctive trails.

Confirmation of the plate-camera film evaluation by other portions of the data reduction is discussed on pages 104 to 114.

NASA Plate-Camera Films
Figures 33 and 34 were taken from NASA aircraft plate cameras AC-36 and AC-38. The films from the four NASA plate cameras were used to confirm and supplement data from the Sandia plate cameras. Many of the re-entry events (such as flares) were evident on both films. Also, the angles of view, due to the different locations of the cameras (see Figure 13), afforded essentially two different pictures of the same events. Several re-entry trails and events not clear on one film were obvious on the other. However, reduction of the NASA plates by the same procedure employed to reduce the Sandia plates was impossible for the following reasons:

1. The timing chopper on the NASA camera failed to operate. Consequently, the camera plates could not be correlated with the data from the other cameras except by means of distinctive events.
2. The position, turn rate, and bank angle of the aircraft could not be accurately determined (this aircraft was not tracked by the verlort radar). As a result, the plates could not be measured, nor could theoretical calculations of azimuth and elevation be made.

Spectral Data

The spectrographic-photometric data on RFD-1 were collected in a variety of ways. Unfortunately, most of the data suffered limitations of one kind or another. The most severe limitations were imposed by the re-entry conditions themselves: that is, the distances between the re-entering objects and the optical instrumentation. The distance to the objects, plus the attenuation of the intervening atmosphere, resulted in drastic reduction in the light intensity at the instrumentation. Since all the spectral instruments were light detectors of some description, the low light level resulted in no data collection in some instances, and in borderline data in others. However, low light levels at the instruments had been expected, and the instruments had been designed to cope with the problem, within the time and money restrictions of the experiment. An unexpected limitation making data reduction difficult was the loss of timing on several instruments and the relatively poor time resolution on others. In addition, the LA-24 tracking telescope was not tracking the proper objects for part of the time, so the data it might have obtained were lost.

Reduction of the spectrographic data required much cross-referencing between data in order to correlate events. Most of the optical data were recorded photographically. The best timing seemed to be on framing-camera film, so events that occurred on a spectrographic plate or film (nonframing) were associated with events on framing camera records whenever possible. Approximate timing was often determined by comparing chopped plate-camera data with spectral-camera data. Spectral films and plates were often underexposed; if an event could be located on one film, it was sometimes possible, knowing what to look for, to locate the same event on the less exposed films and plates. The less exposed film then could be used as corroborative evidence for the data first found. Also, the less exposed films sometimes had time, while the better exposed did not, and a comparison of the two made it possible to determine the time of an event.

Before reduction of the spectral data began, an effort was made to collect material specifications on all re-entering objects. This information was needed as an aid in identifying the spectra recorded and in associating spectra with specific objects.


Figure 33. Re-entry trajectories for RFD-1 components (from Camera AC-80)

Alt 123,000
Range 695.12

Re-entry Vehicle

Time 364.6
Alf 155,100 Range 661.53

Time 359.4 Alt 170,600 Range 644.8

Time 356.3 Alt 179800 Range 641.23

$$
4^{\text {th }} \text { Stage Motor }
$$

Figure 34 . RFD-1 re-entry (from Camera AC-1)

An early accumulation of materials and their specifications proved to be incomplete, and in many instances the specifications provided by a manufacturer proved to be incomplete and/or in error. Extensive laboratory spectral analyses were carried out during the data reduction. Many times such analyses were of a disjunctive nature: that is, either a material contained a particular element or it did not. In other instances, quantitative or semiquantitative analyses were requested and obtained. Some of the analyses and the results obtained are listed in Table XXI. Sodium, which was associated with several re-entry objects, is a common impurity and one easily detected spectrally because of its electronic configuration. If any significant amount were present during the re-entry, it should be easily detectable. Sodium investigations were conducted when some discrepancies arose between specifications and laboratory analyses. The fiberglass-filled covering used on the brackets and on the re-entry vehicle case were supposed to contain only trace amounts of sodium or a few parts per million, but analysis of this material indicated that it contained from 0.5 to 5 percent sodium.

TABLE XXI

## Material Analyses

| Item |  | Percentage of Element Detected |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Ca | Na | Li | Mg | A1 | Si | K | Pb |
| 1. | Large plastic sheet ${ }^{1}$ | 1-10 | 0.5-5 | ND* | >10 | >10 | >10 | ND | ND |
|  | Rubber sample ${ }^{2}$ | 0.1-1 | ND | ND | >10 | <1 | $<1$ | ND | ND |
| 3. | Heat paper ${ }^{3}$ | 0.1-1 | ND | 0.1 | <1 | 1-10 | $\cdot 1$ | ND | ND |
| 4. | Potting material ${ }^{3}$ | < 0.1 | ND | $<0.1$ | >10 | $>10$ | >10 | 1-10 | < 1 |
| 5. | Plastic inner liner (1) ${ }^{3}$ | 0.1-1 | 1-10 | $<0.1$ | $>10$ | $<1$ | $<1$ | 1-10 | $<1$ |
| 6. | $\begin{aligned} & \text { Plastic inner } \\ & \text { liner }(2)^{3} \end{aligned}$ | 0.1-1 | 1-10 | 0.1-1 | 1-10 | $-1$ | $<1$ | 1-10 | $<1$ |
| 7. | Electrolyte pad ${ }^{3}$ | 1-10 | 0.1-1 | > 10 | $>10$ | >10 | $-10$ | $>10$ | 1-10 |
| 8. | Fiberglass-filled material ${ }^{4}$ | 0.1-1 | ND | ND | 1-10 | $\because 1$ | $>10$ | ND | ND |
| 9. | Fiberglass-filled material ${ }^{5}$ | 1-10 | 0.1-1 | ND | 1-10 | >10 | $>10$ | ND | ND |
| 10. | Fiberglass-filled material | 1-10 | ND | ND | $>10$ | $\therefore 10$ | $>10$ | ND | ND |
| 11. | Fiberglass-filled material ${ }^{7}$ | 1-10 | ND | ND | 1-10 | $\therefore 10$ | $>10$ | ND | ND |

NOTES: 1. The fiberglass-filled ablating material covering the fuel-rod brackets and parts of the RV.
2. A rubber cushion between the fuel-rod brackets and the RV.
3. Components of the MC1192 battery.
4. Fiberglass-filled material from the forward flare of the RV.
5. Fiberglass-filled material from the access door of the RV.
6. Fiberglass-filled material from the front flange of the RV, to which the reactor was bolted.
7. Fiberglass-filled material from the aft flare of the RV.

Table XXII lists materials exposed to re-entry heating during the RFD-1 flight test.

TABLE XXII

## Materials Exposed to RFD-1 Re-entry Heating

Item
Reactor

## Reflectors

Reflector springs
Reflector band
Fins
Core vessel and tubing
Pump
Re-entry Vehicle
Ablation-material (heat shield)
Antenna window (heat shield)
Collar (ballast)
Fixed fuel rods
Standoffs
Experimental Fuel Elements
Strontium-loaded rod
Barium-loaded rod
Silver-loaded rod
Gold-1oaded rod
Brackets
Rod holders
Guillotine
Fourth-Stage Motor
Case
Adapter
Retro-rockets
Rocket brackets
Battery
Propellant gases

Materials Included

Aluminum (6061)
Rene 41
Stainless steel (316)
Aluminum (6061)
Stainless steel (316)
Stainless steel (316)

Fiberglass with phenolic resin
Teflon
Tungsten, nickel, copper
$\mathrm{Al}_{2} \mathrm{O}_{3}$, graphite
Titanium

UZrH, strontium, lead, $\mathrm{Al}_{2} \mathrm{O}_{3}$, Fiberfrax insulation, Hastelloy N
UZrH , barium, lead, $\mathrm{Al}_{2} \mathrm{O}_{3}$, Fiberfrax insulation, Hastelloy N
UZrH, silver, Fiberfrax insulation, Hastelloy N
UZrH, gold, Fiberfrax insulation, Hastelloy $N$
Aluminum (6061), fiberglass, Almag 35, MC1192 battery (see Table XXI)
A1mag 35
Stainless steel

Fiberglass
Magnesium
Stainless steel
Aluminum
MC1192 (similar to bracket battery)
Arcite 362

[^2]Figures $35,36,37$, and 38 show diagrams of the most intense lines of the spectrum of each of the four tracers. The intensities are those listed in Wallace Brode, Chemical Spectroscopy, 2nd ed. (published by John Wiley and Sons). The intensity values listed in this book have the same value as those listed in the MIT tables (compiled by George Harrison, published by John Wiley and Sons), but they are more convenient to use for the present problem. The intensities listed are almost certainly at variance with those under re-entry conditions. Thermal excitation is believed to be the primary mode of excitation during reentry, with electron excitation and collisions of the second kind being secondary. If thermal excitation is predominant and some sort of thermal equilibrium exists, the number of atoms in an excited state, $E_{n}$, is proportional to the Boltzman factor $\operatorname{e}\left[\exp \left(-E_{n} / k T\right)\right]$ where $k$ is the Boltzman factor, $T$ is the temperature in degrees Kelvin, and $\mathrm{En}_{\mathrm{n}}$ is the energy of the excited state. The intensity of a line radiated from this excited level can be written as $I_{n m}=A_{n m} N_{n} h \nu_{n m}$, where $A_{n m}$ is the Einstein transition probability, $N_{n}$ is the number of atoms in the nth state, $h$ is Planck's Constant, and $\nu_{n m}$ is the frequency of the emitted radiation. $N_{n}$ can be written as

$$
\left(\frac{N g_{n}}{Z}\right)\left(e^{\left(-E_{n} / k T\right)}\right),
$$

where N is the total number of a species (atom or molecule in a particular state of ionization) present, $g n$ is the statistical weight of the levels, and $Z$ is the partition function.

$$
Z=\sum_{i} g_{i}\left(e^{\left(-E_{i} / k T\right)}\right)
$$

for the species under investigation.
One can see from even this simplified description how dependent the number of atoms or molecules in a particular state of excitation is upon the energy value of the state, through the exponential function. Higher excitation energies result in lower intensities. To calculate absolute intensities under re-entry conditions would be virtually impossible. Even under laboratory conditions, and with only one or two species involved, calculations are difficult. In some cases, wave mechanics can be used to calculate realistic transition probabilities; in others, the transition probabilities can be determined experimentally to varying degrees of accuracy. In many instances, partition function can only be approximated. If electron collision excitation becomes significant, the exponential dependence falls out and thermal equilibrium no longer exists (see Atomic Spectra and Atomic Structure (C. Herzberg), page 160, Dover Publications). Many of the intensities listed in the literature are probably derived from combinations of thermally excited and electron-excited spectra, and so should be used only as rough guide lines.

Collected from four stations, three airborne and one on the ground, the data will be discussed by site category. The ground station was High point, and the airborne stations were the NASA DC-4 (238) and the two AFSWC C-54's ( 461 and 521) . The instrumentation on the NASA plane was provided by NASA, which gracioussly gave Sandia the original data acquired. The instrumentation on the AFSWC C-54's was provided partially by Sandia Corporation and partially by AFSWC. AFSWC also very graciously made its original data available. In addition, Mr. David Mick of AFSWC made a special effort to help interpret the timing records and to provide other helpful information.

Figure 35
Most intense spectral lines for strontium



Figure 36
Most intense spectral lines for barium


Figure 37. Most intense spectral lines for silver


Figure 38. Most intense spectral lines for gold

Spectral instrumentation at High Point consisted of a cinespectrograph, six plate spectrographs, an LA-24 TRSS mounted on and utilizing the $24-i n c h$ optics of the LA- 24 tracking telescope, and a photometer, also mounted on the LA-24.

Cinespectrograph -- The cinespectrograph was a hand-tracked Photosonics 10A $70-\mathrm{mm}$ framing camera operated at 10 frames/sec. The camera had a 4-inch-f.1., f/l.9 lens and a 600-lines/mm grating. The film used was Kodak Linagraph Shellburst. BRT was clearly visible on this film. The only exposures of the re-entry sequences recorded were a few underexposed frames at 364.75 seconds, which coincides with a large flare up of the RV, which was supposedly being tracked by this instrument. No usable data were acquired, since the optical aperture on this instrument was too small to collect enough light to record, and exposures were too weak to determine much about the quality of the focus.

Plate Spectrographs -- The plate spectrographs were all Sandia-built instruments with 12-inch-f.1., f/2.5 Kodak Aeroektar lenses and 600 -lines $/ \mathrm{mm}$ transmission gratings. The gratings were placed in front of the lenses, with their dispersion perpendicular to the predicted trajectory and their blaze oriented to enhance the first-order spectra. The photographic plates on which the data were recorded were $10 \times 12$ inches, with Kodak $103-F$ emulsion. This emulsion proved to have much better sensitivity in the red region from 6300 to $6900 \AA$ than did the Royal-X Pan used by NASA and AFSWC. Spectrographs S-1, S-2, and $S-3$ were tilted up so as to put their zero orders at the upper edge of the plate; $S-4, S-5$, and $S-6$ were tilted down to place their zero orders at the lower edge of the plate. These spectrograph positions were chosen so that at least one set of instruments would cover the trajectory even if it went either higher or lower than predicted. Cameras S-1, S-2, and S-3 were rotated in azimuth with respect to one another so as to cover the trajectory in azimuth. Cameras S-4, S-5, and S-6 were also rotated with respect to one another to cover the trajectory. Some of these instruments were not in as sharp a focus as would have been desirable. Also, when used off axis and at full aperture, these lenses are subject to distortions which create fuzzy off-axis images. The plate spectrographs all suffered from lack of light intensity, but usable data were collected.

The data from plate spectrographs, $S-1$ through $S-6$, will be discussed, instrument by instrument. Events referred to can be seen in Figures 48A and 48B, which were reproduced from the trajectory-camera data.

The spectra will be discussed in terms of the zero-order, first-order, and sometimes second-order spectra. The zero order is the undiffracted light that passes through the grating and the lens and then is imaged on the film or plate just as though the grating were removed (except that its intensity is reduced). The first-order spectrum is the one displayed with the greatest intensity because of the gratings chosen, and the second-order spectrum is the one recorded with less intensity and twice the dispersion of the first-order spectrum. The second order is further displaced from the zero order than the first is. Events will be referred to in terms of their spectral characteristics. For instance, RV flareup or burnup will signify that the materials believed to make up the reactor on the RV were undergoing re-entry heating, either on or off the RV, which was sufficient to vaporize them. For chromium or chromium-oxide spectra or other atomic or molecular spectra to exist under re-entry conditions means that the atoms or molecules were free at the moment they radiated their characteristic radiation.

S-1 Camera: This camera recorded only about 2 inches of the zero order of the radiation from the fourth stage of the launch vehicle. The image is fuzzy, because of an out-of-focus condition or because of lens aberrations. The image is close to the edge of the plate. The fourthstage image broadens as it progresses in time, which is consistent with other data indicating that the fourth stage broke into two objects. There are no spectra on this plate. Either the light level was too low to expose the plate, or, more likely, the spectra were off the photosensitive surface as a consequence of the camera orientation. The fourth stage was considerably below the actual and predicted trajectory of the re-entry objects of primary interest.

S-2 Camera: The data on this camera were the best of any of the plate spectrographs. Focus was sharp over the entire format, including star trails, zero-order images, and spectral lines, The zero order of the fourth stage and of the several objects in the "flare" or "fuelrod" region (i.e., that interval between about 336 and 361 seconds) are recorded, as are the spectra of some of these objects. The object with the lowest trajectory, with the exception of the fourth stage, has a strong sodium spectrum associated with it, as is attested by correspondence between the sodium spectrum and intensity variations in the zero order of that object. The sodium lines at 5890 and $5896 \AA$ are present from about 341 until about 349 seconds. This same object has magnesium lines at 5167, 5172, and 5183 A associated with it from 342.5 until 345 seconds. The magnesium lines have intensity fluctuations corresponding to those of the zero order and sodium spectrum, and they end with an intensity burst like that associated with a sodium intensity burst. Four intensity flareups quite definitely identified as strontium are visible in the zero order. The spectral bands associated with these flares are just detectable on this plate. The times that have been assigned to these flares by correlating the framing-camera data and chopped trajectory plate-camera data with the spectral data are 350.72 , $352.00,353.88$, and 355.98 seconds. A number of other flares were visible in the zero order of the "flare" region, but they were not intense enough to produce identifiable spectra. The fourth stage was quite intensely exposed and a spectrum was recorded. The zero order of the fourth stage clearly separates into two objects. Strong components in the fourth-stage spectrum on this plate are the sodium lines at 5890 and $5896 \AA$ and the lithium line at $6707 \AA$.

S-3 Camera: This camera covered the latter part of the trajectory. The zero order attributed to the RV is visible from 359 to 383.5 seconds. The zero order is not a sharp image, but the spectra and star trails in the center of the plate are in fairly good focus. The only measurable and identifiable spectra are associated with an intensity flareup at 365 seconds. In this region, sodium is present in low amounts and lithium is barely detectable. Chromium lines at 5204 , $5206,5208,5296,5328,5345$, and $5348 \AA$, plus strong chromium oxide bands with heads' at 5794,6051 , and $6394 \AA$ clearly indicate that chromium was undergoing re-entry burnup at this time.

S-4 Camera: This camera covered the early part of the re-entry. The only visible record is of the fourth stage. The zero order and spectra from this source are not sharp. The lithium line at $6707 \AA$, and the sodium lines at 5890 and $5896 \AA$ (associated with the fourth stage), are visible. The doubling of the zero-order image is not obvious. This may be a result of underexposure or of relatively poor image quality--again caused, perhaps, by focus or by lens aberration.

S-5 Camera: This camera was aimed to cover approximately the same portion of the trajectory as the S-2. However, the focus (or chromatic lens aberration) was worse on this camera than it was on the S-2, and the spectral lines are not as well exposed or resolved. More of the re-entry vehicle flareup is recorded on this plate than on the one from the S-2. The spectra early in the flare region have the sodium 5890 and $5896 \AA$ lines associated with the bottom object in the "flare" region. However, the zero order for this sodium doublet is off the plate. The spectra are present because of the angle at which the grating was oriented. The RV flareup at 365 seconds has the same general features as it had in the plate from the $S-3$, except that it is fainter. For some reason, perhaps the angle of the camera axis, the band structures associated with the strontium flares on the S-2 at $350.72,352.00$, 353.88 , and 355.98 seconds are more distinct on this plate than on the S-2.

S-6 Camera: This camera covered the latter part of the trajectory, as did the S-3. Its focus seems to have been better than the focus of the S-3 (the zero order of the re-entry vehicle is sharper), but the spectra are less distinct. The chrome-oxide bands associated with the re-entry vehicle flareup at 365 seconds are visible for about 1 or 2 seconds.

None of the pertinent data from the preceding films was distinct enough for reproduction and presentation in this report.

LA-24 Time-Resolved Streak Spectrograph -- The best-exposed film was that recorded with the LA- 24 TRSS. The basic reason for this was that the LA- 24 primary optics have a 24 -inch-aperture light-gathering system. A secondary reason for the good record was that the focus on the film plane was sharp. Unfortunately, the LA- 24 was tracking the fourth stage at the time this record was obtained; in addition, the time system failed on this spectrograph. However, the photometer was aligned so that its optical axis was essentially parallel with the optical axis of the LA- 24 primary, and they were pointed at the same point in space. By comparing the C.E.C. record of the photometer output with the LA- 24 TRSS spectrogram, it was possible to determine approximate timing for the TRSS spectrogram. The conclusion that the LA-24 was tracking the fourth stage was based on a number of data. Even before chopped trajectory plate-camera data were received from NASA, the object being tracked was identified as the fourth stage from the spectral characteristics of the radiation. Intense aluminum-oxide molecular bands were present in the fourth-stage spectra and were only weak or not present elsewhere. In addition, the intensity fluctuations in the radiation from the fourth stage as recorded on other instruments corresponds with that recorded on the LA-24 instruments. Also, an intense lithium line at $6707 \AA$ was present in the fourthstage data and not (or only faintly) elsewhere. Finally, the fourth stage was clearly seen to separate into two objects on the LA-24 TRSS data, as it was on several of the other plates and films. The time of the fourth-stage radiation and its intensity variation in time are shown in Figure 34.

Before Sandia received these data (shown in Figure 34), it was believed that the fourth-stage intensity maximum occurred much earlier, so it had been hard to reconcile the data. (As it turned out, both the RV and the fourth stage had their maximum intensity at about the same time, 365 seconds.) The reason so much effort was put into the LA-24 TRSS data was the possibility that the data were of re-entry objects of primary interest. But the fourth-stage data recorded on the photometer have the same duration and character as on all other data, including the data from NASA. Since these data have been reduced and are of peripheral interest, the spectrogram is shown in Figure 39, with the lines and bands identified. Note how the sodium and lithium lines split in two. The zero order also shows this splitting. The waviness of the spectral lines was caused by variations in tracking. Another reason for showing these data is that they illustrate the effectiveness of this method, even when the spectral source is over 100 miles away.

LA-24 Photometer -- This instrument operated very well, and if it had been tracking the proper objects it probably would have provided valuable data. The technique of filling a chamber around the photomultiplier with $\mathrm{CO}_{2}$ from a fire extinguisher, and so cooling the photomultiplier, reduced the noise level of the tube to a very satisfactory level. The filter wheel (operating at 600 rpm ) was so well balanced and designed that vibration effects were negligible.

The band pass of the filters used was to have been $10-\AA$ wide and were to monitor two intense lines per flare material and two regions about $50 \AA$ away from each of the spectral lines of interest. This was so that background radiation could be subtracted. The spectral line filters and background filters were chosen so as not to encompass other strong line radiation in their band pass. As it turned out, however, some strong radiation that had not been anticipated occurred during the re-entry, partly a consequence of the fact that the photometer was tracking the fourth stage. The photometer detected radiation from about 346 to 373 seconds, with relatively high-to-high intensity occurring from 358.5 to 371 seconds. One of the filters had been chosen to monitor the silver line at $5209 \AA$. This filter's band pass covered the chromium lines at 5204, 5206, and $5208 \AA$, thus becoming a "chromium filter." The chromium radiation received by the photometer was successfully discriminated against the background radiations.

In some cases, the filters proved to have wider band passes than the specified 10 A. After the RFD-1 flight test, Dr. Fred Roesler of the University of Wisconsin evaluated some of the filters with instrumentation he had designed for mapping the band passes of filters. Time limitations had not permitted earlier evaluation. The filters appeared to have $10-\AA$ band passes, but the peak transmission varied up to $30 \AA$ across a filter, creating an effective band pass of 30 to $40 \AA$. This cut down the transmitted light at any chosen wavelength and passed undesirable wavelengths. The peak transmission wavelengths of the 24 filters follow: gold lines 6278 and $5837 \AA$ with gold lines background $6228,6328,5787$, and $5877 \AA$; silver lines 5209 and $5465 \AA$ with silver lines background 5159, 5259, 5415,


Figure 39. $\underset{\binom{\text { Spectrogram of }}{\text { (from Camera } \mathrm{LA}-24-1 \text { ) }} \text { re-entry }}{ }$
and $5515 \AA$; strontium lines 4607 and $6878 \AA$ with strontium lines background 4555 , 4657,6828 , and $6928 \AA$; barium lines 5535 and $6142 \AA$ with barium lines background 5486, 5586, 6142, and $6228 \AA$. The lines chosen and their relative intensities were shown in Figures 35, 36, 37, and 38. Neither gold nor silver lines were seen on any of the data. (Theoretical predictions had indicated that these tracer elements would not be exposed.) The relatively high excitation energies required for gold, the very small amount of gold in the fuel rods, and the intensities of the other lines indicate that the gold lines would have been very faint and possibly not detectable even if the gold had been exposed.

Other references used in estimating line intensities and interfering lines were NBS Monograph 32 (Part 1), Tables of Spectral Line Intensities - Arranged by Elements (United States Department of Commerce, and NBS Circular 467, Atomic Energy Levels (United States Department of Commerce).

## Data from Sandia Airborne Spectral Instruments

The TRSS's in AFSWC-461 and AFSWC-521 detected radiation throughout the reentry. These instruments were both in sharp focus and, because of the movement of the film and their relatively narrow field of view (about 3 degrees), they did not suffer much from overlapping spectra from different objects. These instruments were hand-tracked.

The film was to be pulled at $3 / 16-i n . / s e c$ past a $3 / 4$-inch slot so time resolution was to be about 4 seconds. BRT was to have been imprinted on the film, but it failed on the TRSS aboard AFSWC-461. Time on this film was estimated from filmtransport rate, associating events with other data. On the TRSS aboard AFSWC-461, the film-pull rate was actually about twice as fast as had been intended. Since variable-speed DC motors were used to drive the cameras, it was difficult to achieve a desired speed; when this TRSS was airborne, the friction between the film and the plate against which it was held (by a vacuum pump) varied from what it had been at ground level.

AFSWC-461 was off to the side of the trajectory, about even with the "flare region" where flare materials were expected to be exposed. Much good data were collected, but better data would have been recorded if the film rate had been $3 / 16$ in. $/ \mathrm{sec}$. AFSWC-521 was stationed farther along the trajectory and looking upstream, but even so, good data were recorded in the flare region and good exposures of the RV burnup were recorded.

The roll of the airplanes and effects of hand-tracking from an unstable platform are visible in the data. The hydraulic autopilot in AFSWC-521 was not as good as the electronic one in AFSWC-461. The results of the greater roll of AFSWC-521 are noticeable in the plate-camera data as well as in the hand-tracked TRSS records. A disadvantage of tracked instruments is that one has a record of what is being detected with respect to time (if time is operational) but not of the location of the events in space, so the data must be correlated with other data. A good timing record was obtained on the TRSS aboard AFSWC-521.

AFSWC-521 Time-Resolved Streak Spectrograph -- The data obtained from the TRSS aboard AFSWC-521 were first recorded at 339.5 seconds, and were made possible by a mistrack that resulted in light integration on the film. It is believed that data were recorded somewhat earlier by the TRSS aboard AFSWC-461, because of the relative positions of the two aircraft. The TRSS in AFSWC-521 tracked the rods, or objects in the "flare" region, until 357 seconds, and then had a period with no record. At about 374 seconds, the tracker apparently started tracking the RV, and this was tracked until 392 seconds, with the intensity much diminished after 385 seconds. This last-tracked object is believed to have been the re-entry vehicle, since no other data show any other object with intensity enough to record at this late time in the trajectory.

At 345 seconds an intensity integration due to mistrack allows spectra to be identified. These spectral lines follow: magnesium lines at 5183, 5172, and $5167 \AA$; lead line at $4058 \AA$ or potassium lines at 4044 and $4047 \AA$; aluminum lines at 3944 and $3961 \AA$, and more magnesium lines at 3829,3832 , and $3838 \AA$. Another intensity integration at 347.5 seconds has the same general characteristics as the one at 345 seconds, except that some radiation at about $4319 \AA$ is quite intense. This could be the strong iron lines at 4308 and $4326 \AA$, but since the two lines are not resolved, because of movement of the camera, the identification is
tentative. At 351 seconds occurs a flare with the band structure of what has been identified as strontium. There is also a line at the correct position to be the strong strontium line at $4607 \AA$. The flare at 351 seconds is believed to be the same one identified as strontiun at 350.72 seconds on the framing-camera color film. Another flare occurring at 357 seconds is believed to be a strontium flare (the one at 355.98 seconds from framing-camera data).

The several-second time resolution on the TRSS films allows only approximate determination of time. The prominent spectra on this spectrogram are sodium. Sodium spectral lines 5890 and $5896 \AA$ were visible from 339.5 until 353 seconds, and again from 374 until 385 seconds. Lithium is not present in the spectra from 339.5 until 353 seconds, but it is barely detectable from 374 until 382 seconds. The RV is the probable source for the latter spectra. The lithium does not appear as intense, in comparison with the sodium lines, as in the fourth-stage spectra. Furthermore, the intense aluminum-oxide bands associated with the fourth stage are completely absent here. Chromium and iron were undergoing re-entry burnup during this time. Lines detected follow: chromium lines at 4254, 4274, 4289, 5204, 5206, 5208 , and $5264 \AA$ including nearby iron lines, $5298 \AA$ including nearby iron lines, and $5345 \AA$; and a combination iron line composed of iron lines 5324 and $5328 \AA$.

AFSWC-461 Time-Resolved Streak Spectrograph -- As mentioned previously, there were no timing marks on the film from the AFSWC-461 TRSS. A calculation based on the total time anything was being tracked allows an estimate of $3 / 8 \mathrm{in} . / \mathrm{sec}$. All times given for these data will be based on this somewhat arbitrary estimate. Sodium lines at 5890 and $5896 \AA$ were visible from 341 to 350 seconds. Little or no sodium and no lithium are detectable on this record from 376.9 until $392 \mathrm{sec}-$ onds. There are two intensity increases or intensity integrations at $343.6 \mathrm{sec}-$ onds and at 344.3 seconds, and these have the same spectra. Spectral lines identified foll 18 : sodium at 5890 and $5896 \AA$; magnesium at $5183,5172,5167,3838,3823$, and $3829 \AA$; chromium at 4289,4274 , and $4254 \AA$; lead at $4058 \AA$ or potassium at 4044 and $4047 \AA$; and aluminum at 3944 and $3961 \AA$.

An intense flare at 350 seconds has the band structure associated with the strontium flares, and a line identified as the strontium line at $4607 \AA$. This flare was recorded better on this film than on any other. If it is the same flare identified as the most intense "strontium" flare on other data, its true time is probably 350.72 seconds. Another strontium-type flare occurs at 352.6 seconds, and yet another flare without such definite strontium characteristics occurs at 356.4 seconds. These may be the flares identified as strontium on other data at 352.00 and 355.98 seconds. If so, only the "strontium" flare at 353.88 was not detected. The zero order is multiple in the "flare" region, indicating that more than one object was being tracked; but it is single and sharp from 376.9 until 388.4 seconds, when the RV was the probable object being tracked. The spectra from the RV region are fuzzy and not well exposed. However, chromium lines and chromium-oxide bands appear to be present, and this agrees with other RV data. It must be emphasized that the timing was relatively uncertain on the TRSS on AFSWC-461. The last spectra from the RV probably were not after 384 seconds, or the end of the large reactor-RV flareup.

AFSWC-521 Photometer -- The photometer on AFSWC-521 was hand-tracked on the same mount as the TRSS. This photometer collected no usable data. Its field of view was only 24 minutes of angle, so the difficulties of hand-tracking a heavy assembly, compounded by the poor autopilot on this aircraft, caused the only detectable signal to appear as the photometer field of view passed over a re-entry object. Except for very brief intervals, there were only two periods when signals were recorded. These were 0.2 second at 380 seconds, and 0.05 second at 376.5 seconds. During the longer interval, the photometer was reading high on the silver 5209- $\AA$ filter, which was probably monitoring the chromium lines at 5204 , 5206 , and $5208 \AA$. This would be consistent with the other data for this time. The brief intervals of record indicate that the photometer functioned properly within its field-of-view limitations. The C.E.C. recorder provided excellent time resolution.

## Data from AFSWC Spectral Instruments

On AFSWC-521, spectral data were collected on $\mathrm{Z}-1, \mathrm{Z}-2$, and $\mathrm{Z}-10$, which were the K-24 spectral cameras discussed in Section III and Tables III and V. The Hulcher $70-\mathrm{mm}$ cinespectrograph on AFSWC-521 did not collect any usable data. The
timing did not record satisfactorily, and exposures believed to be of the RV reentry are underexposed and poorly resolved. The camera may have been out of focus. Timing was somewhat uncertain on the K-24 spectral cameras, which were supposed to cycle 2.5 seconds open and 1.5 seconds shut while the film was being pulled. In reality, the signals from the camera to the timing record show open times from 7 to 8 seconds down to zero. Since it was not possible to definitely relate one timing signal to a certain frame, the relation was arrived at deductively.

Z-1 Spectral Camera: The deduced times for the two exposed frames on the Z-1 camera were from 371.3 to 378 seconds, and from 378.6 to 380.8 seconds. This camera was aimed aft, and so should have been looking at the RV-burnup region in space; the spectrogram reproduced in Figure 40 agrees with the other data on RV burnup, and the timing agrees approximately with other data on RV intensity brightness. The intensity fluctuations in time seem to indicate that the timing may be off by several seconds. Since the spectrum lines and bands are identified in Figure 40, their identification will not be discussed further here.

The second spectrogram from the $\mathrm{Z}-1$ camera has the same features as the one reproduced. Camera vibration and motion are signified by the waviness of the lines. This camera was in sharp focus and probably provides the clearest identification of lines arising from burnup of the RV.

Z-2 Spectral Camera: Camera Z-2 was the center camera of the three spectral $\mathrm{K}-24$ 's. Three of its frames acquired data that appear to be of the same event recorded on the $Z-1$; therefore the data must be of the RV burnup. Since these data were not as good as those from Camera 2-1, and since they apparently had no features not contained in those from the $\mathrm{Z}-1$, the data were not reduced further.

Z-10 Spectral Camera: Camera Z-10, aimed forward, covered the "flare" region. The focus on this instrument was only fair and the film was underexposed. Data were collected on three frames. The underexposure is lamentable, because this camera was operated with an $f$ stop of $f / 5.6$ when it could have been opened up to $f / 2.5$ for a fivefold or sixfold increase in exposure. An event identified as a strontium flare on the AFSWC-521 TRSS record, at what is called 350.72 seconds from the motion-picture data, was identified on the Z-10 data. The line identified as the strontium line at $4607 \AA$ is visible. The film indicates that a wide-open camera would have acquired excellent data. Timing of the frames was not exactly determined, since the time record was bad. The spectral characteristics indicate that the exposed frames were of the "fuel-rod flare" region. The sodium lines at 5890 and $5896 \AA$ associated with the "bottom" object in the flare region were recorded, and chromium lines were recorded as well.

On AFSWC-461, none of the AFSWC-provided spectral instrumentation recorded data. Orientation of cameras on the aircraft may be the explanation.

## Data From NASA Airborne Spectral Instruments

The NASA aircraft was located well downrange. As a result, its optical instrumentation was viewing the trajectory somewhat end on. Most of the trajectory could be covered with single, fixed cameras; however, the aircraft position was far enough to the side to resolve events along the trajectory with the open plate trajectory and the spectral plate cameras. NASA instrumented the NASA aircraft completely. The data from the chopped trajectory plate camera AC-1 (see Figure 34) have been valuable in determining time for different events. The chops on this plate were coded and had a 0.25 -second open shutter and a 0.25 -second closed shutter. NASA had three spectral cameras (KG-24's) with gratings as indicated in Table IV. The NASA spectral cameras used Royal-X Pan film, as did the AFSWC cameras, and they suffered the same cutoff in the red region of the spectrum above 6300 A.


Figure 40. RFD-1 RV-reactor spectra
(from $-\mathbf{Z - 1}$ )

AC-37 Spectral Camera -- Spectral Camera $\mathrm{AC}-37$ used a 600 -lines/mm grating, and had a dispersion on the film of about $93 \AA / \mathrm{mm}$. Although this camera was not in sharp focus, the images were fairly good. The field of view covered the latter part of the "fuel-rod flare" region and the RV flareups in the zero order. The spectra from the fourth stage were on the plate, although the zero order was not. The aluminum-oxide molecular bands, and the sodium lines at 5890 and $5896 \AA$ associated with the four th stage, are very evident and compare well with the data recorded on the LA-24 TRSS. Because of the red-end cutoff of the film, the lithium line at $6707 \AA$ is not seen. Spectra arising from the RV flareups are not sharp and well resolved. These spectra are not as intense as would be expected, because they arise from the unblazed side of the grating and thus have much less intensity than those arising from the blazed side. In this instance these spectra were off the format. However, the spectra from the RV burnup are obviously the same as those in the data recorded by the other instruments in which the details were more distinct and have been or will be discussed. Therefore, AC-37 data will not be pursued further.

AC-85 Spectral Camera -- Spectral Camera AC-85 utilized a 400-1ines/mm grating, giving a dispersion on the film of about $136 \AA / \mathrm{mm}$. The data recorded by this instrument were the most comprehensive and distinct recorded by any single instrument. The focus was sharp, the regions of interest in space were well covered, and lines were, relatively speaking, well resolved (Figure 41). The NASA cameras all seemed to suffer much less from aircraft vibration than did the instruments in the AFSWC C-54's.

The regions in space covered by the zero order of $\mathrm{AC}-85$ were from the middle of the "fuel-rod flare" region through the RV flareups, or from about 351 seconds until after 386 seconds. Spectra from the entire "fuel-rod flare" region are on the film due to the angle of the grating lines. In addition to the first-order spectrum of the fuel-rod region, there are some second-order spectra from the RV flareup, some first-order spectra of the fourth stage, and some first-order spectra of the RV flareup from the unblazed side of the grating. AC-85 was the only camera or instrument to record indications of barium flares. Barium lines arising from the "fuel-rod region" occur from about 350.5 until 353.5 seconds, and again at 359.37 seconds. The flare observed at 356.62 seconds is possibly barium. The bariug lines seen in the flare at 350.5 to 353.5 seconds are the strong line at $5535 \AA$ and the weaker line at $4554 \AA$. A line that could be the barium line at $6142 \AA$ tends to confirm the identification. The barium line at $5535 \AA$ was the only barium line identified in the flare at 359.37 seconds. The strontium flares identified on the other spectral data are recognizable on this spectrogram and are labeled on the spectrogram.

In addition to the lines labeled on the spectrogram shown in Figure 41, other lines from the RV flareup were also identified. They will be listed, although they add little to the picture previously drawn by the other data. The lines and bands follow: chromium oxide band heads at 5554 and $5794 \AA$; a combination line made up of iron lines at 5265, 5266, 5269, 5281, and 5283 $\AA$, and chromium lines at 5264 and $5265 \AA$; a combination line made up of iron lines at 5324, $5328,5329,5332$, and $5339 \AA$, and chromium lines at 5345 and $5409 \AA$. At the very right edge of the spectrogram are some short magnesium lines at 5167, 5172, and $5183 \AA$. The bright line in the spectrum at the right of the picture that then trails to the left and down across the other spectra is the strong sodium doublet associated with the "bottom" object or objects. The sodium doublet is also visible in the RV spectra. In the trajectory cameras the bottom trace is seen to separate into two traces. The sodium spectrum is seen to fuzz out (in time), indicating that two sources of sodium radiation could be present.

AC-84 Spectral Camera -- Spectral Camera AC-84 was aimed somewhat more upstream than were $\mathrm{AC}-37$ and $\mathrm{AC}-85$ (see Figure 42). The zero order of the entire "fuel-rod region," the first RV flareup, and the fourth stage were recorded. This instrument used a $75-1$ ines $/ \mathrm{mm}$ grating, so the dispersion was only about $740 \mathrm{~A} / \mathrm{mm}$ on the film. Because of the small dispersion and somewhat fuzzy focus, relatively poor resolution was present in the first order. However, the second-order spectra of the fourth stage and some of the fuel-rod region are exposed well enough to be usable at its higher dispersion of $370 \AA / \mathrm{mm}$. There are even some third-order spectra arising from the fourth stage. The falling off in intensity of spectra with increasing order is one reason for using low-dispersion gratings. In cases where the first-order spectrum is overexposed, the second or third are very unlikely to be. In addition to the lines labeled on the spectrogram, a line identified as


Figure 41. RFD-1 re-entry spectra (from Camera AC-85)


Figure 42. RFD-1 re-entry spectra (from Camera AC-84)
either potassium lines 4044 and $4047 \AA$ or the lead line at $4058 \AA$ arises from the "bottom" object which has the strong sodium doublet associated with it. The strontium flares with their typical spectral bands are recognizable but could not be identified from the spectrogram alone. The aluminum-oxide bands and the sodium doublet are again pronounced spectral structures of the fourth-stage spectra.

## Motion-Picture Data

As explained in Table V, Section III, most of the framing-camera films were underexposed. Three films, however, did record usable information. These included:

1. The color films from Sandia ME-16-1 (35-mm Mitche 11 motion-picture camera with 18 -inch-f.1., f/3.0 Paxar lens).
2. The black-and-white film from NASA AC-101 (35-mm Flight Research motion-picture camera with 10 -inch-f.l., $f / 4.5$ Bell and Howell Eyeman lens).
3. The color film from NASA 113 ( 16 -mm Cine Kodak motion-picture camera with 2 -inch-f.1., f/2.0 Kodak Anastigmat lens).

Reduction of these three films constituted the motion-picture evaluation of the reentry. A discussion of their analysis follows:

Four of the re-entry objects are quite evident on the motion-picture films. Several others are just visible but too faint and blurred for identification or analysis. A complete tabulation of the observed events from the films versus time is given in Tables XXIII to XXVI. BRT, which is accurate to 0.01 second, was recorded on the border of the films. The small discrepancies in times from the three films are due to several factors. For example, since the frame rates of the two NASA films were 10 and 16 frames $/ \mathrm{sec}$, while the Sandia camera frame rate was 96 frames/sec, the NASA cameras could have missed some of the faster events. In addition, the Sandia camera was closer to the early portion of the re-entry flight path, while the NASA cameras were closer to the latter portion (see Figure 13). The Sandia camera could have missed some of the less intense flares due to its faster frame rate. Neither NASA Cameras AC-101 nor Sandia Camera ME-16-2 tracked the fuel rods to the completion of burnup, but NASA Camera 113 did. Even though the timing on the films was accurate to 0.01 second, determination of the precise times of flare beginning, peaks, and end was difficult. The times agree closely enough to permit correlation with other films and records. No determination of altitude or location in space is directly possible from these films. However, a comparison of the motion-picture times with the plate films and theoretical trajectories furnishes this information.

Figure 43 is a black-and-white reproduction of one frame from the $35-\mathrm{mm}$ color motion-picture film taken by Sandia Camera ME-16. Time of the events shown was 351.62 seconds. The lower object was identified as a strontium flare, and the upper object is believed to be a barium flare.

Figure 44 is an enlargement of one frame from the 35 -mm black-and-white motion-picture film taken by NASA Camera AC-101. Time of the events shown was 352.00 seconds. The objects included are, from left to right: the re-entry vehicle, strontium and barium flares, and the fuel-rod brackets.

## Re-entry Vehicle

Identification of the RV on the motion-picture films was easily made because of its size, brightness, and behavior during re-entry. Disassembly of the simulated SNAP-10A reactor was visible on the films (see Table XXVI and Figure 45). These events correlated well with the predicted, telemetered disassembly and spectral data. Reactor disassembly is discussed in SC-RR-64-515.

TABLE XXIII
Fuel-Rod Flare Times Recorded by NASA Camera AC-101

| Time from Launch (sec) |  |  | Intensity |  |  | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Flare Start | Flare Peak | Flare Out | Bright | Med. | Dim |  |
| 342.17 | 342.37 | 342.57 |  | x |  | Not fuel-rod flares |
| 342.27 | 342.37 | 342.47 |  | x |  | Bottom, strontium flare |
| 344.70 | 344.80 | 345.10 |  | x |  | Bottom, strontium flare |
| 345.20 | 345.20 | 345.20 |  | x |  | Bottom, strontium flare |
| 345.30 | 345.41 | 345.51 |  |  | x | Bottom, strontium flare |
| 345.51 | 345.51 | 345.51 |  |  | x | Not fuel-rod flares |
| 346.11 | 346.21 | 346.31 | x |  |  | Bottom, strontium flare |
| 346.21 | 346.31 | 346.41 | x |  |  | Bottom, strontium flare |
| 346.72 | 346.92 | 347.12 |  |  | x | Bottom, strontium flare |
| 348.13 | 348.13 | 348.13 |  |  | x | Bottom, strontium flare |
| 348.63 | 348.73 | 348.83 |  | x |  | Bottom, strontium flare |
| 348.64 | 348.94 | 349.45 |  | x |  | Bottom, strontium flare |
| 349.55 | 349.65 | 349.85 |  | x |  | Bottom, strontium flare |
| 349.65 | 349.85 | 349.95 |  | x |  | Bottom, strontium flare |
| 349.95 | 350.05 | 350.15 |  | x |  | Bottom, strontium flare |
| 350.15 | 350.25 | 350.46 |  | x |  | Bottom, strontium flare |
| 350.15 | 350.35 | 350.46 |  | x |  | Bottom, strontium flare |
| 350.15 | 350.35 | 350.55 | xx |  |  | Bottom, strontium flare |
| 350.15 | 350.35 | 350.55 | xx |  |  | Bottom, strontium flare |
| 350.55 | 350.75 | 350.85 | xx |  |  | Bottom, strontium flare |
| 351.46 | 351.66 | 352.16 | x x |  |  | Bottom, strontium flare |
| 350.25 | 350.75 | 352.26 | x x |  |  | Top, barium flare |
| 352.77 | 352.77 | 352.87 | x ${ }^{\text {r }}$ |  |  | Top, barium flare |
| 352.87 | 352.87 | 352.87 |  | x |  | Top, barium flare |
| 352.87 | 352.87 | 352.87 |  | x |  | Bottom, strontium flare |
| 352.97 | 352.97 | 352.97 |  |  | x | Top, barium flare |
| 352.97 | 352.97 | 352.97 |  |  | x | Bottom, strontium flare |
| 353.28 | 353.28 | 353.28 |  |  | x | Bottom, strontium flare |
| 353.58 | 353.88 | 354.39 | x |  |  | Bottom, strontium flare |
| 354.39 | 354.49 | 354.79 | x |  |  | Bottom, strontium flare |
| 354.79 | 354.99 | Dim, not out | x |  |  | Bottom, strontium flare |
|  | 355.29 | Dim, not out | x |  |  | Bottom, strontium flare |
|  | 355.90 | 356.00 | x |  |  | Bottom, strontium flare |
| 356.90 | 356.90 | 356.90 |  |  | x | Top, barium flare |
| 357.10 | 357.20 | 357.52 | x |  |  | Top, barium flare |
|  | 358.00 |  |  |  |  | Off edge of film |

TABLE XXIV


TABLE XXV
Fue1-Rod Flare Times Recorded by NASA Camera 113

| Time from Launch (sec) |  |  | Intensity |  |  | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Flare Start | Flare Peak | Flare Out | Bright | Med. | Dim |  |
| 345.18 | 345.24 | 345.43 |  | x |  | Strontium flare |
| 345.74 | 345.74 | 345.87 |  | x |  | Strontium flare |
| 346.10 | 346.12 | 346.21 |  | x |  | Not fuel-rod flares |
| 346.61 | 346.62 | 346.72 |  | x |  | Not fuel-rod flares |
| 348.75 | 348.94 | 349.00 |  | x |  | Strontium flare |
| 349.00 | 349.25 | 349.37 |  | x |  | Strontium flare |
| 349.44 | 349.44 | 349.56 |  | x |  | Strontium flare |
| 349.62 | 349.69 | 349.81 |  | x |  | Strontium flare |
| 350.00 | 350.13 | 350.25 |  |  | x | Strontium flare |
| 350.31 | 350.31 | 350.38 |  | x |  | Bottom, strontium flare |
| 350.31 | 350.38 | 350.44 |  | x |  | Top, barium flare |
| 350.44 | 350.44 | 350.56 | xx |  |  | Bottom, strontium flare |
| 350.50 | 351.31 | 351.94 | x |  |  | Top, barium flare |
| 350.63 | 350.69 | 350.75 |  |  | x | Bottom, strontium flare |
| 350.94 | 351.50 | 351.75 | x |  |  | Bottom, strontium flare |
| 352.32 | 352.38 | 352.38 |  |  | x | Bottom, strontium flare |
| 352.51 | 352.57 | 352.63 |  |  | x | Bottom, strontium flare |
| 352.56 | 352.56 | 352.62 |  |  | x | Top, barium flare |
| 352.76 | 352.82 | 352.88 |  | x |  | Bottom, strontium flare |
| 353.01 | 352.32 | 353.70 | x |  |  | Bottom, strontium flare |
| 353.76 | 353.76 | 353.82 |  | x |  | Bottom, strontium flare |
| 353.88 | 353.88 | 353.95 |  | x |  | Bottom, strontium flare |
| 354.01 | 354.01 | 354.07 |  | x |  | Bottom, strontium flare |
| 354.14 | 354.57 | 354.82 |  | x |  | Bottom, strontium flare |
| 354.89 | 355.08 | 355.20 | x |  |  | Bottom, strontium flare |
| 356.18 | 356.18 | 356.25 |  |  | x | Top, barium flare |
| 356.31 | 356.62 | 356.81 | x |  |  | Top, barium flare |
| 357.01 | 357.01 | 357.01 |  |  | xX | Bottom, strontium flare |
| 357.14 | 357.14 | 357.14 |  |  | xx | Bottom, strontium flare |
| 359.12 | 359.37 | 359.56 | x |  |  | Top, barium flare |
| 359.68 | 359.74 | 359.87 |  |  | x | Top, barium flare |
| 360.83 | 361.14 | 361.45 |  | x |  | Top, barium flare |
| 361.76 | 361.83 | 361.89 |  | x |  | Top, barium flare |

NOTE: Color of top (barium) flares was yellow-white. Color of bottom (strontium) flares was red.

TABLE XXVI
RV Event Times Recorded by NASA Camera AC-101

| ```Time from Launch (sec)``` | Event | Remarks |
| :---: | :---: | :---: |
| 340.38 | Brackets show | Tumbling |
| 342.64 | RV shows |  |
| 343.64 | RV and reflectors brighter |  |
| 344.64 | Object off RV |  |
| 346.04 | RV bright and object off |  |
| 348.00 | Object off |  |
| 352.43 | Object off |  |
| 352.62 | RV flat face on front of flare | Coning |
| 352.82 | Stable | Coning |
| 352.94 | RV front of flare slanted | Coning |
| 353.04 | RV front of flare slanted opposite | Coning |
| 353.15 | RV front of flare slanted | Coning |
| 353.25 | Stable | Coning |
| 353.35 | Very flat face - object off | Coning |
| 354.45 | Two objects from previous object off |  |
| 354.85 | Rear object breaks into 2 parts - RV stable |  |
| 356.06 | Object off |  |
| 357.67 | Severe coning starts | Coning |
| 359.07 | RV bright | Coning |
| 359.47 | RV appears to be canted - large - bright | Coning |
| 359.67 | RV stable | Coning |
| 362.50 | Object off |  |
| 363.71 | RV very bright | Overexposed frame |
| 364.41 | Object off |  |
| 365.00 | Many objects off | Reactor disassembly |
| 366.02 | Object off |  |
| 366.72 | Rear objects break up |  |
| 371.86 | Off edge of film |  |



Figure 43. Strontium and barium flares during re-entry ( 351.62 seconds)


Figure 44. RFD-1 re-entry at 352 seconds (blown up from 35-mm motion-picture film, Camera AC-101)


Figure 45. RFD-1 re-entry at 366.9 seconds
(blown up from $35-\mathrm{mm}$ motion-
picture film, Camera AC-101)

## External Fuel Rods

Identification of the strontium and barium tracers on the motion-picture films was determined from the flare colors and from the pattern of pulsations. This analysis is thought to be conclusive for the following reasons:

1. When strontium is excited in an air atmosphere, the predominance of radiation from it is in the red spectral region (6000 to $7000 \AA$ ). This is a consequence of heavy molecular bands plus some atomic lines in that region. Barium appears as a white light with a yellow-pink cast. The white light results from an even distribution of lines over the spectrum from 4000 to $7000 \AA$. Strong barium-oxide molecular bands exist in the yellow-orange region. These bands, plus atomic line radiation in the orange region, contribute to the yellow-pink cast. Comparison of colors in the color photoprints from which Figure 12 (preflight electric-arc tunnel tests) and Figure 43 (ME-16-2) were reproduced suggests that the two flares were strontium and barium. The top flare of Figure 43 was noted pulsating 10 times from 350.25 seconds ( 198,000 feet) to 357.52 seconds ( 178,000 feet) on framing-camera films. The bottom flare pulsed 28 times from 342.27 seconds (220, 125 feet) to 361.83 seconds ( 165,400 feet). Color film does not always depict true colors, because of overexposure, underexposure, atmospheric effects, film emulsion, and processing. However, the flare colors recorded on the films remained essentially the same through many cycles of varying intensity, pulsating from very bright to not visible. This lends some validity to the color comparisons. As previously mentioned the spectral data corroborates these conclusions.
2. The pattern of flare occurrence on the motion-picture films is similar to that noted in the electric-arc tunnel preflight ablation tests. During these tests, even though the specimens were rotating, a hot spot invariably caused excessive ablation at one or two locations. This resulted in a hole or holes through the rod wall, which exposed the tracer. As rotation and ablation continued, the flare from the tracer pulsated in an irregular pattern of intensity similar to the one mentioned above. The irregular flare pattern from the fuel-rod tracers on the flighttest motion-picture films was thought to be caused by intermittent exposure of the tracers. This was probably due to tumbling and to uneven ablation of the rods. It should be noted that the
re-entry trails from the RV and brackets were essentially steady and continuous streaks. These phenomena are also very evident on the plate-camera films.

## Fue1-Rod Brackets

Even though Bracket No. 1 was the first visible re-entry object, and among the brightest, identification of the fuel-rod brackets from the motion-picture films was not possible. Identification was initially determined from the spectral films and theoretical trajectories. It was first visible on the motion-picture films at 340.38 seconds ( 227,000 feet) and appeared to have a tumbling motion. Peak intensity was reached at 343.64 seconds ( 218,000 feet). However, the light from this object was essentially continuous and did not pulsate or flare appreciably. From peak intensity to 364.50 seconds ( 173,000 feet), it grew steadily dimmer, and it was not visible after that time. The brightness of Bracket No. 1 was believed caused by the sodium content of the fiberglass shield and the contents of the battery it supported. Bracket No. 2 did not include a battery and was therefore not as bright on the motion-picture films.

The rest of the re-entry items are faintly visible in Figure 46. They do not, however, present any positive characteristics to allow identification.


Figure 46. RFD-1 re-entry at 348.9 seconds (blown up from 35-mm motionpicture film, Camera AC-101)

Identification of the re-entry objects from the motion-picture films should not be construed as final proof. Confirmation of these observations by the spectral and plate-camera films and by the theoretical trajectories is discussed on pp. 106 to 114.

Reduction of the RFD-1 films and records to establish re-entry events and object identification was pursued according to individual type of film or record involved. The three major types included plate-camera films, spectral data, and motion-picture films. This method was employed because each type displayed the information in varying manners and required different procedures for analysis.

Generally, each record revealed information about the re-entry that by itself was not absolute. However, coincidental agreement of the three data categories on almost every conclusion appears to verify and confirm those conclusions. For clarity, discussion of the data correlations will be according to individual re-entry objects. Time is considered as the common dimension of comparison. Figures 47-50 summarize the finding discussed in this section.


Figure 47. Summary of RFD-1 re-entry events

A. From Plate Camera T-2

B. From Plate Camera T-3

Figure 48. Summary of RFD-1 re-entry events ( $0.1^{\prime \prime} \approx 2800^{\prime}$ )


Figure 49. Summary of RFD-1 re-entry events (from Camera AC-34)


Identification of the reactor and RV from the films was obvious, especially when compared with visual observation of the actual re-entry. They were the leading object in the re-entry sequence (as predicted by the preflight trajectory). They were also much larger and brighter than all other items except the fourthstage motor. The fourth-stage motor, however, was far enough removed from the re-entry area so as not to interfere with their identification.

The theoretical RV trajectory discussed previously (pp. 46-79) agreed to within 300 feet with the measured FPS-16 radar trajectory, from approximately 167 seconds ( 450,000 feet) to 360 seconds ( 168,800 feet). This theoretical trajectory also agreed to within 400 feet with the measured plate-camera trajectory, from 325.4 seconds ( 266,000 feet) to 392.1 seconds $(80,400$ feet). At this altitude, the velocity of the RV and reactor was slow enough to cause cooling below the temperature which permitted them to be self-luminous.

Spectral films from the TRSS on AFSWC 521 indicated the presence of aluminum at 345 seconds ( 212,000 feet) and 347.5 seconds ( 205,000 feet). The TRSS on AFSWC 461 showed aluminum at 343.6 seconds ( 218,200 feet) and 344.3 seconds (214,000 feet). These times and altitudes compare with the indicated melting time of 336 seconds ( 237,500 feet) to 342 seconds (220,400 feet) of the aluminum fins on the reactor, as telemetered from the thermocouples. However, these spectra cannot be definitely confirmed as having radiated from the zero-order streak known to be the RV-reactor. Aluminum brackets and reflectors were also melting in this vicinity at the same time, and could have contributed at least part of the aluminum spectra.

Chrome and iron lines were continuously evident (with intermittent variations in intensity) between 343.6 seconds ( 218,200 feet) and 388.4 seconds ( 87,300 feet) on most of the spectral films. Many of the lines were directly associated with the zero-order trail of the RV-reactor. Both chrome and iron are constituents of stainless steel, and all the reactor components, except the aluminum fins, were constructed of stainless steel. The strongest indication of stainless steel occurred from 365 seconds ( 153,900 feet) to 388.4 seconds ( 87,300 feet). This was the period of very intense reactor flaring. Additional discussion of the reactor disassembly and spectral verifications can be found in SC-RR-64-515.

Sodium spectra associated with the zero-order RV-reactor streak is also continuously in evidence between 343.6 seconds ( 218,200 feet) and 385 seconds ( 96,300 feet). This element was found in the fiberglass ablative material from the RV. Although the specification for that material specifies only trace amounts of sodium (which probably would not have been very prominent spectrally), laboratory spectral analysis of an actual RFD-1 specimen revealed greater than 1 percent sodium content. This amount can be readily seen under the re-entry condition of RFD-1. It was noted that all objects containing fiberglass (the RV, the fourth-stage motor, and the fuel-rod brackets) appeared very intense on all the plate, spectral, and framing-camera films.

Motion-picture data confirm the RV-reactor identification and the reactor disassembly. Figure 45 shows a typical reactor-object separation event. This photograph was enlarged from the film in NASA framing-camera AC-101. The time of this particular event was 366.9 seconds. Table XXVI tabulated all the RV-reactor re-entry events from that film. Approximately 12 objects can be seen separating from the reactor between 344.6 seconds ( 213,500 feet) and 371.9 seconds ( 133,600 feet). These events and times agree reasonably well with the thermocouple and spectral data.

## Strontium-Loaded Fue 1 Rods

From 342.60 seconds ( 218,600 feet) to 361.90 seconds ( 165,200 feet), the theoretical trajectory for the strontium-loaded fuel rods compared closely with the trajectory as measured from the plate-camera film. However, at $348.64 \mathrm{sec}-$ onds ( 202,000 feet), the measured trajectory separated into two streaks, with the lower one following a somewhat steeper trajectory to 355.98 seconds ( 181,500 feet), when it disappeared. Both of these trails are known to be the strontium-loaded rods because of the bright flares identified as strontium on the spectral films. The discrepancy in the latter portion of the trajectory is thought to be due to the final stages of ablation, breakup of the rods, and melting of the strontium tracer. As can be seen on the plate-camera films, the strontium-loaded rods
pulsated and flared from 342.27 seconds ( 220,100 feet) to 361.83 seconds ( 165,400 feet). This long period of intermittent strontium flares was probably caused by molten strontium spilling out of the rods through the vent holes and/or ablated areas. There were five major flares during this time, three in the upper strontium streak and two in the lower. The lower strontium streak flared at 353.88 seconds ( 187,300 feet) and 355.98 seconds ( 181,500 feet). The lower trail definitely ended at this flare, with no further evidence of either strontium or the UZrH fuel rod. The upper trail flared at 350.72 seconds (196,200 feet), 352.00 seconds (192,200 feet), and 361.83 seconds ( 165,400 feet). A faint streak assumed to be caused by final consumption of the UZrH fuel element continued to approximately 363.32 seconds ( 161,400 feet) on the NASA plate-camera films. This point is considered as the time and altitude of burnup for purposes of calculating heat input versus volume consumed (see Figure 50).

Four of the major flares in the re-entry trails mentioned above were identified as strontium on the spectral films. They were evident on spectral films from Cameras S2, S5, TRSS 521, TRSS 461, AC-84, and AC-85. The first-order spectra of the strontium bands were definitely associated with the flares shown in the zero order. It was also obvious that these flares were the same flares shown on the plate-camera films. The last flare in this group, at 361.83 seconds, was evident in the zero order of the plate-camera film but could not be identified as strontium in any of the first-order spectra. This was thought to be caused by lack of intensity. Even the four flares that were identified were extremely faint on the films.

The motion-picture films discussed on pp. 98 to 105 also identified the strontium flares. As shown in Table XXIII, the strontium flares were first visible at 342.27 seconds (220,100 feet). They were intermittently visible, pulsating from not visible to very bright, down to 361.83 seconds ( 165,400 feet). These pulsations can definitely be matched to the flares shown in the plate-camera and spectral films. Also, as explained on p. 104 and 105, the flare colors on the films compare with the color for strontium extremely well.

## Barium-Loaded Fue 1 Rods

Identification of the barium-loaded rods was not as positive as the identification of the strontium-loaded rods. However, enough corroborative evidence was extracted from the films to strongly imply their presence and ablation altitudes.

The theoretical and measured trajectories for the barium-loaded rods agreed almost exactly for the entire re-entry sequence. Their patterns of behavior were very similar. Also, no other theoretical trajectory for the remaining re-entry objects showed above the RV as did the barium rods (see Figure 31). Sandia platecamera film T-2 and P-2 indicate some of the barium flares above the RV from 348.00 seconds ( 205,200 feet) to 352.77 seconds ( 191,900 feet). After that time, the barium is superimposed on the RV trail until 356.62 seconds ( 181,200 feet), then slightly below the RV to 359.37 seconds (173,600 feet). Bright flares are noted at 356.62 seconds ( 181,200 feet) and 359.37 seconds ( 173,600 feet). The change in trajectory of the barium-loaded rods from above to below the RV is explainable. The barium-loaded rods were ejected up when released at $282.09 \mathrm{sec}-$ onds. However, their W/CDA was 44.4 as compared with 259.4 for the RV. This caused the steeper trajectory for the rods to eventually fall below the RV trajectory. Also, in the latter portion of the re-entry, the barium-loaded rods were ablating and breaking up, as were the strontium-loaded rods. Another reason could have been separation of the three rods into slightly different trajectory altitudes. The NASA plate films from AC-36 and AC-38, taken much closer to the trajectory and at a different angle, show these same streaks and flares very clearly, and they are not superimposed on the RV streak (see Figures 49 and 50). After the last barium flare on the NASA plate-camera films at 359.37 seconds ( 173,600 feet), a faint trail continued down to approximately 371.50 seconds ( 142,000 feet). This was assumed to be the altitude of final burnup of the barium-1oaded rods.

Only one spectral film (NASA AC-85) indicated the presence of barium. This was probably because the barium flares were not so intense as the strontium flares and because they occurred closer to the NASA aircraft Camera AC-85 than to the
other stations. However, the zero-order events producing the spectral lines were the same events identified as barium from the plate and framing cameras. Since the NASA AC-85 film was not time-chopped, all times were assigned to the barium flares from the NASA and Sandia framing cameras and the Sandia chopped-trajectory camera.

The framing-camera films showed the pulsating barium flares from $350.25 \mathrm{sec}-$ onds ( 198,900 feet) to 357.52 seconds ( 178,000 feet). Evidently the barium tracer, like the strontium tracer, was prematurely exposed. The times and durations of the flares (see Tables XXIII-XXV) agree with the plate and spectral films. The strongest indication from the motion-picture film that these were the barium-1oaded rods was their pulsating pattern, and the color of the flares which matched those of movies made of laboratory tests under simulated re-entry conditions.

## Gold- and Silver-Loaded Fuel Rods

The theoretical trajectories for the gold- and silver-loaded fuel rods correlated with two re-entry trails on the plate cameras (see Figures 31 and 32). Both groups of rods were superimposed on the RV trail until 355.5 seconds ( 184,000 feet), when their trajectories fell below the RV. They were dim compared with the barium and strontium fuel-rod trails. The extreme brightness of the strontiumand barium-loaded rods was due to the strontium and barium, not the UZrH. Since the gold and silver tracers were not exposed, only the radiation from the UZrH was visible. These trails can be traced down to 368.00 seconds ( 149,000 feet) where they apparently were no longer luminous. There was no evidence of either gold or silver on the spectral films. The framing-camera films also furnished no indication of the gold- or silver-ioaded fuel rods.

## Fuel-Rod Brackets

Identification of the remaining re-entry objects was completed only so that they could be eliminated from consideration and not be mistaken for re-entering fuel rods. However, since the fuel-rod brackets were so strongly in evidence on the plate, spectral, and framing-camera films, an explanation of their reduction is in order.

The complete fuel-rod mount consisted of two brackets (see Figure 10B). One of the brackets included an MC1192 battery for initiation of the explosive bolts. Both brackets were made of aluminum and magnesium, and included a fiberglass heat shield on the front surface. Theoretical trajectories indicated that the lower re-entry trail on the plate-camera films was Bracket No. 1. The trail of Bracket No. 2 was somewhat higher as a consequence of ejection perturbations. However, these ejection conditions could not be measured accurately and, therefore, the theoretical trajectories are not necessarily positive proof of exact actual altitudes. Their main function was to duplicate re-entry patterns. Spectral films S-2, S-5, TRSS 521, TRSS 461 and AC-84 all revealed the presence of aluminum, magnesium, potassium, lead, and sodium. These spectral lines radiated from the zero order of the bottom re-entry trail. All these elements are ingredients of some part of the bracket, battery, or fiberglass shield. The zero order on the spectral film shows two objects in this bottom trail. The plate-camera and framing-camera films confirm this observation. Consequently, it was concluded that the bottom re-entry streak was from the two fuel-rod brackets. These objects are also very bright on the motion-picture films, but cannot be individually identified.

## Fourth-Stage Motor

The fourth-stage motor was retro-rocketed away from the flight path after ejection from the re-entry vehicle at 282.09 seconds. This perturbation was included as a reduction in the velocity of the initial conditions of the theoretical trajectory. Several unsuccessful attempts were made to compute a theoretical trajectory which would match the measured fourth-stage trajectory from the plate films. It was concluded that unknowns such as pitch angle and residual thrust at ejection made comparison impossible.

Spectral reduction of the combination LA- 24 TRSS and photometer indicated that this instrument tracked the fourth-stage motor. This observation was based on the presence of sodium, lithium, and aluminum-oxide spectra. Cameras $\mathrm{S}-2, \mathrm{~S}-4$, AC-37, and AC-84 confirmed their presence and also showed the same band structure and intensity pattern. In addition, on those data having timing, the timing agreed. This was discussed previously on pp. 89 to 91.

The fourth-stage motor was not tracked by any of the motion-picture cameras because it was out of the fields of view of all cameras except the LA-24 framing camera, which did not function.

Fuel-Rod Ablation Analysis

The trajectory actually followed by the fuel rods was somewhat different from the preflight nominal prediction. The theoretical aerodynamic heating associated with the actual flight of the strontium-, barium-, silver-, and gold-loaded rods is given in Figures 51 to 54 . An analysis of the response of the fuel rods to this heating will be made in this section, together with predictions of the resulting burnup altitudes.

The temperatures of the materials within the insulation were measured in neither the qualification nor the flight tests. Since some heat will be conducted through the insulation to the flares, an iterative process must be used together with theoretically predicted temperatures to estimate and evaluate this effect. The reduction of data from the qualification tests is presented in detail in Sandia Corporation report SCDR 124-63. The calculated aerodynamic heat absorbed by the UZrH is given below for convenience. No calculation was made of possible oxidation heating effects, since the initial model conditions and the surface temperature response were not known.

## Qualification Test Results

| Tracer | $Q_{T}(B T U / 1 \mathrm{~b})$ |
| :--- | :---: |
| Strontium | 749 |
| Silver | 761 |
| Solid UZrH | 639 |

This analysis of the specimens also neglected any heat conducted through the insulation. This heat was neglected because the insulation presented a rather severe resistance to heat flow as compared to that absorbed by the UZrH. For purposes of thoroughness, it seems advisable to neglect this heat in an initial analysis of the rods during the re-entry, and to include it in a later analysis so as to provide a limiting comparison of the internal effects.

The models used in the qualification tests were similar in cross section to the re-entry rods, except that the cladding was not included on the tunnel-test models. For purposes of predicting for the re-entry, the required heat is that determined from the qualification tests to ablate the UZrH, plus that necessary to heat and melt the cladding on the outside of the rods. This may be written as

$$
\begin{equation*}
Q_{r}=(W C \Delta \Gamma)_{c}+(W H)_{c}+\mathrm{MQ}_{\mathrm{T}}, \tag{1}
\end{equation*}
$$



Figure 51. Actual trajectory and heating for strontium-loaded rods


Figure 52. Actual trajectory and heating for barium-loaded rods


Figure 53. Actual trajectory and heating for silver-loaded rods


Figure 54. Actual trajectory and heating for gold-loaded rods
where

$$
\begin{aligned}
\mathrm{Q}_{\mathbf{r}} & =\text { calculated aerodynamic heat required, } \mathrm{BTU} \\
(\mathrm{WC} \Delta \mathrm{~T})_{\mathrm{C}} & =\text { heat required to raise temperature of cladding, BTU } \\
(\mathrm{WH})_{c} & =\text { heat required to melt cladding, BTU } \\
\mathrm{MQ}_{T} & =\text { heat required for } \mathrm{UZrH} \text { from tests, } \mathrm{BTU}
\end{aligned}
$$

For the strontium-loaded rod, the heat required for complete ablation is

$$
\mathrm{Q}_{\mathrm{r}}=(0.336)(0.095)(2350)+(0.336)(130)+(0.995) \mathrm{Q}_{\mathrm{T}}=118.69+0.955 \mathrm{Q}_{\mathrm{T}}
$$

Using the values obtained in the tunnel tests,

$$
\begin{gather*}
Q_{r}=118.69+(0.955)(761)=845.44 \mathrm{BTU}  \tag{la}\\
Q_{r}^{\prime}=118.69+(0.955)(749)=854 \mathrm{BTU}  \tag{lb}\\
Q_{r}^{\prime \prime}=118.69+(0.955)(639)=729 \mathrm{BTU} \tag{lc}
\end{gather*}
$$

(In all cases, three values of heat and an altitude range will be calculated, since it is not obvious which one of the results from the tunnel tests is more nearly correct.)

The cold-wall stagnation-point heat flux taken from the theoretical trajectory is adjusted for radiation from the surface and for the effect of increasing surface temperature on the enthalpy gradient across the boundary layer. The net heat available is

$$
\begin{equation*}
\dot{q}_{n e t}=F A \dot{q}_{c w}\left(1-\frac{h_{w}}{h_{s}}\right)-A q_{r a d}, \tag{2}
\end{equation*}
$$

where

$$
\begin{aligned}
& \mathrm{q}_{\text {net }}=\text { available heat into the fuel rod for purposes of } \\
& \text { material ablation, BTU/sec } \\
& \dot{\mathrm{q}}_{\mathrm{cw}}=\text { cold-wall stagnation heat } \mathrm{flux}, B T U / \mathrm{ft}^{2}-\mathrm{sec} \\
& h_{w}=C_{p} T_{W}=\text { gas enthalpy evaluated at the wall temperature, } \\
& \text { BTU/lb } \\
& h_{s}=\text { free-stream stagnation enthalpy, } B T U / 1 \mathrm{~b} \\
& \dot{\mathrm{q}}_{\mathrm{rad}}=\epsilon \sigma \mathrm{Tw}^{4}=\text { radiation loss, } \mathrm{BTU} / \mathrm{ft}^{2}-\mathrm{sec} \\
& A=\text { fuel } 1-\text { rod surface area }=0.327 \mathrm{ft}^{2} \\
& F=0.179=0.167+\frac{0.242}{2(\mathrm{~L} / \mathrm{D})}=\text { factor to reflect stagnation } \\
& \text { heating on a hemisphere to }
\end{aligned}
$$

Then for the re-entry

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{a}}=\int_{\theta_{1}}^{\theta_{2}} \dot{\mathrm{q}}_{\text {net }} \mathrm{d} \theta \tag{3}
\end{equation*}
$$

where

$$
\begin{aligned}
\mathrm{Q}_{\mathrm{a}} & =\text { total heat available at time } \theta, \mathrm{BTU} \\
\theta & =\text { time, seconds. }
\end{aligned}
$$

It is necessary in this analysis to know the fuel-rod surface temperature as a function of time during the re-entry. For this purpose a theoretical study was performed on both a "thermalog" analog computer and a CDC 1604 digital computer. The results of this study, which will be discussed in detail in a later part of this section, are presented in Figures 55 to 58 . These surface-temperature responses are used to calculate radiation loss from the rods for the type of analysis being discussed in this section.


Figure 55. Predicted temperatures for strontium-loaded rods



Figure 58. Predicted temperatures for gold-loaded rods

The net total aerodynamic heat available to the rods during the re-entry for purposes of ablating the UZrH, as computed from Eq. (3), is shown in Figures 59 to 62. With values for the heat required as computed from Eq. (1), one can predict the range of altitude burnup of the strontium-loaded rod as given below. Following the same analysis for the rod containing the barium tracer, the predicted range of burnup altitude is also given below. The same type of analysis performed on the silver- and gold-loaded rods indicates that insufficient heat is available to ablate the UZrH and expose the tracer. Further analysis of the silver- and gold-loaded rods will not be pursued in this report.

| Tracer | Predicted Burnup <br> Altitude (ft) |
| :--- | ---: |
| Strontium | 152,000 to <br>  <br> Barium |
|  | 167,000 <br> $\quad 120,000$ to |
|  |  |



Figure 59. Net aerodynamic heating for strontium-loaded rods


Figure 60. Net aerodynamic heating for barium-loaded rods


Figure 61. Net aerodynamic heating for silver-loaded rods


Figure 62. Net aerodynamic heating for gold-loaded rods

The preceding calculation neglects any heat conducted into the tracer, and, as such, it presents an upper limit on the altitudes of tracer exposure predicted by means of this analysis. Computer analyses indicate that quantities of heat which may not be negligible are conducted through the insulation during the reentry period. Estimates of this quantity of heat are made utilizing computerpredicted temperatures for the qualification tests, as shown in Figures 63 and 64.


Figure 63
Predicted temperature response of strontium-loaded rod in qualification tests

Figure 64
Predicted temperature response of barium-loaded rod in qualification tests


It is convenient to compute heat conducted through the insulation on a unitlength basis. The computation should include the heat required to raise the temperature of the flare, the lead ballast, and the insulation to their predicted values. For the tunnel tests, the aerodynamic heat remaining in the UZrH to cause its ablation may be expressed as

$$
\begin{equation*}
Q_{U Z r H}=Q_{T}-(W C \Delta T)_{T}-(W C \Delta T)_{B}-(W C \Delta T)_{I} \tag{4}
\end{equation*}
$$

where

$$
\begin{aligned}
& (\mathrm{WC} \Delta \mathrm{~T})_{\mathrm{T}}=\text { heat absorbed by tracer, BTU/inch } \\
& (\mathrm{WC} \Delta \mathrm{~T})_{\mathrm{B}}=\text { heat absorbed by ballast, BTU/inch } \\
& (\mathrm{WC} \Delta \mathrm{~T})_{\mathrm{I}}=\text { heat absorbed by insulation, BTU/inch. }
\end{aligned}
$$

For the strontium-loaded rod in the qualification test, this equation results in

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{UZrH}}=59.62-1.532-3.98=54.11 \mathrm{BTU} / \text { inch } \tag{4a}
\end{equation*}
$$

or

$$
Q_{U Z r H}=\frac{54.11}{0.0796}=680 \mathrm{BTU} / 1 \mathrm{~b} .
$$

Likewise, for the silver-loaded rod this calculates as

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{UZrH}}=146.87-(3.36+1.33)=142.18 \mathrm{BTU} / \text { inch } \tag{4b}
\end{equation*}
$$

or

$$
\mathrm{Q}_{\mathrm{UZrH}}=\frac{142.18}{0.193}-736.7 \mathrm{BTU} / 1 \mathrm{~b} .
$$

Of course the solid rod had no internal flare, so the required heat for ablation remains

$$
\mathrm{Q}_{\mathrm{UZrH}}=639 \mathrm{BTU} / 1 \mathrm{~b}
$$

Using the estimates of altitudes and time of flare exposure made previously, and the computer surface-temperature predictions shown in Figures 55 and 56 , it is possible to calculate the heat absorbed by the tracer, ballast, and insulation during the flight test. The total heat necessary for ablation of the rod under these conditions is

$$
\begin{equation*}
Q_{F}=Q_{U Z r H}+(W C \Delta T)_{T}+(W C \Delta T+W H)_{B}+(W C \Delta T)_{I}+(W C \Delta T+W H)_{C} . \tag{5}
\end{equation*}
$$

For the strontium-loaded rod, using the calculated net values of heat required in the qualification tests for the UZrH, this results in

$$
\begin{align*}
Q_{F} & =894 \mathrm{BTU}  \tag{5a}\\
Q_{F}^{\prime} & =944 \mathrm{BTU} \\
Q_{F}^{\prime \prime} & =988 \mathrm{BTU} .
\end{align*}
$$

From Figure 59 (the net input into the strontium-loaded rod during re-entry), the predicted altitudes of burnup may be made as shown below. Likewise, calculations for the barium-loaded rod result in the predicted altitudes of flare exposure as shown below. Similar calculations on the remaining two types of rods indicated that those flares would not be exposed.

| Tracer | Predicted Exposure Altitude (ft) |
| :---: | :---: |
| Strontium | 144,000 to 129,000 |
| Barium | 110,000 |

As discussed in other portions of this report, optical instrumentation employed to record the events of the fuel-rod re-entry resulted in the following observed altitudes of complete burnup of the various tracers.

| $\frac{\text { Tracer }}{\text { Strontium }}$ | Observed Burnup <br> Altitude (ft) |
| :--- | :---: |
| $\frac{161,400}{\text { Barium }}$ | 142,000 |

Since these altitudes are somewhat higher than the predicted values for the trajectory, some investigative analysis is warranted. Figures 51 and 52 show the cold-wall heat flux and total heat of this trajectory, together with the times of tracer burnup. A reasonable estimate of the net total aerodynamic heat input to the rod may be made with Eq. (3), and these values are shown in Figures 59 and 60. From these curves, values of aerodynamic heat associated with the rods at the time of burnup may be determined as:

| Tracer | Aerodynamic Heat <br> Absorbed (BTU) |
| :--- | :---: |
| Strontium | 770 |
| Barium | 980 |

It is apparent that these values are somewhat smaller than the corresponding values predicted with data from the qualification tests. Again, analyses may be made by (1) calculating the heat absorbed by the UZrH and the cladding, while neglecting any heat conducted through the insulation, and (2) estimating the heat absorbed by the internal components and considering it as absorbing a portion of the input. The first of these operations may be performed as follows:

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{a}}=\mathrm{Q}_{\mathrm{UZrH}}+(\mathrm{WC} \Delta \mathrm{~T})_{\mathrm{C}}+(\mathrm{WH})_{\mathrm{C}} \tag{6}
\end{equation*}
$$

Applying this expression to the strontium-loaded rod, the amount of aerodynamic heating absorbed by the UZrH at the time of tracer exposure as recorded during re-entry is

$$
\begin{equation*}
Q_{U Z r H}=\frac{Q_{a}-(W C \Delta T+W H)_{C}}{0.955}=673 \mathrm{BTU} / 1 \mathrm{~b} . \tag{6a}
\end{equation*}
$$

Likewise, the calculation for the barium-loaded rod is

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{UZrH}}=\frac{\mathrm{Q}_{\mathrm{a}}-(\mathrm{WC} \Delta \mathrm{~T}+\mathrm{WH})_{\mathrm{C}}}{1.543}=554 \mathrm{BTU} / 1 \mathrm{~b} . \tag{6b}
\end{equation*}
$$

The second of the analyses may be performed with Eq. (5). Using theoretical predictions of the internal temperatures at the time of exposure yields

$$
\begin{equation*}
Q_{U Z r H}=\frac{580}{0.955}=608 \mathrm{BTU} / 1 \mathrm{~b} \tag{5b}
\end{equation*}
$$

A similar calculation for the barium-loaded rod results in

$$
\begin{equation*}
Q_{U Z r H}=\frac{831}{1.543}=538 \mathrm{BTU} / 1 \mathrm{~b} . \tag{5c}
\end{equation*}
$$

The table below gives the calculated and observed values for the aerodynamic heat absorbed by the rods to cause flare exposure that results from the different assumptions in the analyses. The calculated values are based on the results of the hyperthermal tunnel tests, while the observed values are based on the actual RFD-I data reduction. It is apparent from this table that the calculated and observed values of heat required are not very different, especially for the strontiumloaded rod.

Aerodynamic Heat Absorbed (BTU)
Calculated

| Tracer |  |  | Observed |
| :---: | :---: | :---: | :---: |
|  | Neglect Tracer | Consider Tracer |  |
| Strontium | $\begin{aligned} & 729 \\ & 854 \end{aligned} \text { to }$ | $\begin{aligned} & 894 \text { to } \\ & 988 \end{aligned}$ | 770 |
| Barium | $\begin{aligned} & 1107 \text { to } \\ & 1293 \end{aligned}$ | 1350 | 1150 |

In an attempt to better correlate all the results, an investigation of oxidation contributions was undertaken.

An examination of possible contributions to heating made by oxidation may be conducted on the RFD-1 fuel rods in the following manner. The total heating rate on the rod after the UZrH surface is exposed to the high-enthalpy flow may be expressed as

$$
\begin{equation*}
\dot{q}_{t}=\dot{q}_{c}+\dot{q}_{o x} \tag{7}
\end{equation*}
$$

where

$$
\begin{aligned}
\dot{\mathrm{q}}_{\mathrm{t}} & =\text { total surface heat rate, } \mathrm{BTU} / \mathrm{ft}^{2}-\mathrm{sec} \\
\dot{\mathrm{q}}_{\mathrm{c}} & =\text { convective heat rate, } \mathrm{BTU} / \mathrm{ft}^{2}-\mathrm{sec} \\
\dot{\mathrm{q}}_{\mathrm{ox}} & =\text { effective oxidation heat rate, } \mathrm{BTU} / \mathrm{ft}^{2}-\mathrm{sec} .
\end{aligned}
$$

The convection heat rate is available from the trajectory, and the oxidation heating rate may be evaluated as

$$
\begin{equation*}
\dot{\mathrm{q}}_{\mathrm{ox}}=\mathrm{W}_{\mathrm{o}} \mathrm{H}_{\mathrm{ox}} \tag{8}
\end{equation*}
$$

where

$$
\begin{aligned}
W_{o} & =\text { effective rate of surface oxygen } f 1 \mathrm{ux}, \mathrm{lb} / \mathrm{ft}^{2}-\mathrm{sec} \\
\mathrm{H}_{\mathrm{ox}} & =\text { heat of oxidation, BTU/lb-oxygen. }
\end{aligned}
$$

The rate of oxygen flux to the surface is

$$
\begin{equation*}
\mathrm{W}_{\mathrm{o}}=0.215 \mathrm{~W}_{\mathrm{A}} \tag{9}
\end{equation*}
$$

where

$$
\begin{aligned}
\mathrm{W}_{\mathrm{A}}= & \text { effective rate which air diffuses to the body surface, } \\
& \mathrm{lb} / \mathrm{ft}^{2}-\sec .
\end{aligned}
$$

The mass-flow rate of the air immediately ahead of the re-entering body and its associated shock is based on a unit area, $\dot{M}=\rho \mathrm{V}$. Also, the local cold-wall heattransfer rate to a body may be defined as

$$
\begin{equation*}
\dot{\mathrm{q}}_{\mathrm{cw}}=\mathrm{C}_{\mathrm{H}} \rho \mathrm{~V} \mathrm{~h}_{\mathrm{s}} \tag{10}
\end{equation*}
$$

where

$$
\begin{aligned}
\mathrm{C}_{\mathrm{H}} & =\text { Stanton number } \\
\rho & =\text { gas density }, \mathrm{lb} / \mathrm{ft}^{2} \\
\mathrm{~V} & =\text { gas velocity }, \mathrm{ft} / \mathrm{sec} \\
\mathrm{~h}_{\mathrm{s}} & =\text { gas entha1py }, \mathrm{BTU} / 1 \mathrm{~b} .
\end{aligned}
$$

The rate at which the air molecules actually strike the surface of the material and release their energy to the body depends on the particle transfer within the boundary layer, and is defined as

$$
\begin{equation*}
\mathrm{W}_{\mathrm{A}}=\mathrm{C}_{\mathrm{H}} \dot{\mathrm{M}}=\mathrm{C}_{\mathrm{H}} \rho \mathrm{~V} \tag{11}
\end{equation*}
$$

Equations (9), (10), and (11) may be combined to yield

$$
\begin{equation*}
\mathrm{W}_{\mathrm{o}}=0.215 \frac{\dot{\mathrm{q}}_{\mathrm{cw}}}{\mathrm{~h}_{\mathrm{s}}} \tag{12}
\end{equation*}
$$

The process assumed to account for the significant portion of any heat of oxidation released is

$$
\begin{equation*}
\mathrm{Zr}+\mathrm{O}_{2} \rightarrow \mathrm{ZrO}_{2}+\text { energy . } \tag{13}
\end{equation*}
$$

For each mole of Zr and $\mathrm{O}_{2}$, one mole of $\mathrm{ZrO}_{2}$ is formed, so that the heat of oxidation is

$$
\begin{equation*}
\mathrm{H}_{\mathrm{ox}}=2815 \frac{(\text { Mo1. wt })_{\mathrm{Zr}}}{(\text { Mo1. wt })_{\mathrm{O}}} \tag{14}
\end{equation*}
$$

or

$$
\mathrm{H}_{\mathrm{OX}}=8.0245 \frac{\mathrm{CAL}}{\mathrm{MG}-\mathrm{O}_{2}}=14,350 \mathrm{BTU} / 1 \mathrm{~b}-\mathrm{O}_{2} .
$$

Combining Eqs. (12), (13), and (14) produces an expression for the total surface heating due to convection and oxidation as follows:

$$
\begin{equation*}
\dot{\mathrm{q}}_{\mathrm{T}}=\dot{\mathrm{q}}_{\mathrm{C}}+0.215 \frac{\dot{\mathrm{q}}_{\mathrm{c}}}{\mathrm{~h}_{\mathrm{s}}}(14,350), \tag{15}
\end{equation*}
$$

or

$$
\dot{\mathrm{q}}_{\mathrm{T}}=\dot{\mathrm{q}}_{\mathrm{C}}\left(1+\frac{3085}{\mathrm{~h}_{\mathrm{S}}}\right) \mathrm{BTU} / \mathrm{ft}^{2}-\mathrm{sec} .
$$

This equation may be evaluated from the information available on the theoretical trajectories. It is shown in Figures 65 and 66 for the strontium-loaded and the barium-loaded rods.


Figure 65. Net re-entry aerodynamic and oxidation heating to strontium-loaded rods


Figure 66. Net re-entry aerodynamic and oxidation heating to barium-loaded rods

The computer study mentioned earlier included the predicted effects of oxidation heating after the cladding material was removed. From the initial ejection of the rods from the RV until the cladding reached its melting temperature, the heating was assumed to be aerodynamic only, as given by Figures 59 and 60. The temperatures shown on Figures 55 and 56 resulted from the assumption that a 0.003 inch air gap existed between the cladding and the UZrH material. There is reason to believe that this assumption is reasonable and, in fact, that real effects such as local hot spots, internal pressure, and shear may cause the UZrH to be exposed to the flow even earlier than this calculation indicates. Once the cladding is removed, the total heat input was assumed to be convection- and oxidationinduced. It was calculated for the re-entry using Eq. (15). The computer program includes the nonhomogeneous cross section and also the surface radiation. No effects of hydrogen combustion or injection in the boundary layer are included in the calculated heat input.

These calculations indicate that the rods containing strontium and those containing barium should reach their melting temperature, while those containing the silver and those containing the gold would not reach sufficient temperature to cause ablation. The predicted burnup altitude from this analysis is

| Tracer | Predicted Burnup <br> Altitude (ft) |
| :--- | :---: |
| Strontium | 167,000 |
| Barium | 138,000 |

The temperature data from these analyses are used in the oxidation analysis to determine the temperature of the UZrH at the time of the cladding removal.

The energy required to ablate the UZrH from this condition is as shown in Eq. (16), with the temperature $T_{0}$ equal to the UZrH temperature at the time of cladding melt, as given on Figures 55 and 56:

$$
\begin{equation*}
Q=W C\left(T_{m}-T_{o}\right)+W D-W M \tag{16}
\end{equation*}
$$

where

$$
\begin{aligned}
\mathrm{Q} & =\text { heat to complete ablation, } \mathrm{BTU} \\
\mathrm{~W} & =\mathrm{UZrH} \text { weight, } \mathrm{lb} \\
\mathrm{C} & =\mathrm{UZrH} \text { specific heat } \mathrm{BTU} / 1 \mathrm{~b}-{ }^{\circ} \mathrm{F} \\
\mathrm{D} & =\text { heat to dissociate the hydrogen, } \mathrm{BTU} / 1 \mathrm{~b} \\
\mathrm{M} & =\mathrm{UZrH} \text { heat of fusion, } \mathrm{BTU} / 1 \mathrm{~b} \\
\mathrm{~T}_{\mathrm{m}} & =\text { temperature of } \mathrm{UZrH} \text { melt, }{ }^{\circ} \mathrm{F} .
\end{aligned}
$$

For the strontium-loaded rod, this may be evaluated as

$$
\begin{aligned}
& Q=(0.9548)(0.1)(3400-1600)+(0.9548)(650)+(0.9548)(56.5) \\
& Q=846 \mathrm{BTU} .
\end{aligned}
$$

Likewise, for the barium-loaded rod, is obtained:

$$
\begin{align*}
& Q=(1.543)(0.1)(3400-1600)+(1.543)(650)+(1.543)(56.5) \\
& Q=278+1003+87 \\
& Q=1368 \mathrm{BTU} . \tag{16b}
\end{align*}
$$

Corrections (Eq. 2) were made as described previously for the surface-radiation loss and the effect of the hot wall on the boundary-layer enthalpy gradient on the convection heating, and this was added to the oxidation heating. The net total heat into the model is calculated as

$$
\begin{equation*}
\mathrm{Q}_{\mathrm{T}}=\int_{\theta_{1}}^{\theta_{2}}\left(\dot{\mathrm{q}}_{\mathrm{c}}+\dot{\mathrm{q}}_{\mathrm{ox}}\right) \mathrm{d} \theta \tag{17}
\end{equation*}
$$

Figures 65 and 66 present this calculation for the strontium- and barium-loaded rods.

With the values of heat required to ablate the rod as shown in Eq. (16a) and (16b), the predicted altitudes of burnup and the corresponding aerodynamic heat absorbed are:

| Tracer | Altitude (ft) | Aerodynamic Heat Absorbed (BTU) |
| :---: | :---: | :---: |
| Strontium | 158,000 | 795 |
| Barium | 135,000 | 1215 |

## Discussion

The analysis and correlation of the flight-test results, the qualificationtest results, and the oxidation presentation were performed in two ways: (1) a total aerodynamic heat calculation to the time at which a significant event occurred, and (2) a computer analysis to describe the transient surface-temperature response. Each of these two methods has certain advantages, but the results of both are shown in the analysis for completeness. First, the total-heat method allows correlation between the varying pulse of re-entry heating and the constant hyperthermal input of tunnel heating, and second, the various thermo-chemical surface effects of the ionized gases are not sufficiently known to allow them to be included in a satisfactory computer program. Evaluation of these effects and incorporation of them into a computer program is presently an object of effort, but the work is incomplete. As a result, calculations of the transient temperature of a material such as UZrH, showing melting and/or ablation, are not completely satisfactory.

For the fuel-rod heating analysis, several assumptions were made which somewhat affect the results. As a summary of these results, the following table is presented:
Conditions

| Based on RFD-1 fuel-rod qualification |
| :--- |
| tests; neglects heat conducted into |
| tracer |
| Based on RFD-1 fuel-rod qualification |
| tests; estimates heat conducted into |
| tracer |
| Computer program; includes oxidation |
| after Hastelloy is removed |
| Oxidation analysis based on total |
| heat available |
| Observed on RFD-1 re-entry |

## Conditions

Resulting Predicted or Observed Altitude for Flare Exposure (ft)

| $\frac{\text { Strontium }}{152,000 \text { to }}$ | $\frac{\text { Barium }}{120,000 \text { to }}$ |
| :--- | :--- |
| 129,000 to | 136,000 |
| 144,000 | 110,000 |
| 167,000 | 138,000 |
| 158,000 | 135,000 |
| 161,400 | 142,000 |

Comparing observed values with the predictions made under the different assumptions shows that, in general, good agreement exists. However, assumptions were necessary for each of these analyses (and in fact cause the different predicted values), and these should be examined for their validity.

The basic assumption necessary in using data from the hyperthermal tunnel tests in the re-entry analysis is that a material responds to a particular tunnel environment in the same manner that it would to the same environment during reentry. This is obviously a major assumption, but computer analyses and the observed data from the re-entry indicate that, at least in this instance, the assumption is adequate. The total heat which had to be applied in the constant environment of the hyperthermal tunnel to result in complete model ablation is assumed to be the total necessary to cause a similar model to completely ablate in the varying re-entry environment.

Only estimates of the heat conducted through the insulation into the tracer materials could be made in the qualification tests, flight test, and the oxidation analysis. The estimates were based on computer analyses and, as has been explained, these analyses are not exact. It is obvious from the preceding table immediately above that the analysis with estimations made for the heat conducted into the tracers resulted in predicted altitudes which are the farthest from agreement with the observed conditions. These are included for completeness, but in view of the manner in which the heat conduction was assumed, they should probably be regarded as the least meaningful data in the analyses.

The corrections applied to the cold-wall heating taken from the trajectory were functions of the surface temperature, which were in turn determined from a computer. To neglect this correction completely results in an error of only a few percent (from 5 to 10 percent), but this, of course is enough to account for some of the difference between predicted and actual values. Assumed properties of the material are also factors that could affect predictions of burnup; but until material properties are known that are typical of the materials in re-entry, little would be gained by merely changing the material properties so as to have a particular program give data that correlate with a known result.

In summary, the method used to predict burnup conditions and to reduce data from the tests is thought to be satisfactory. From the table on the previous page it can be seen that the several analyses employed produced burnup altitudes not significantly different from the observed conditions. The inclusion of all the analyses in this report serves to illustrate the gross data correlation between the actual re-entry, empirical data from the hyperthermal tunnel, and analytical analyses employing a computer.

One manner in which a tracer might be exposed in a manner different from the one assumed in this analysis would be either a structural failure of the rods or a re-entry in which the rods stabilized in either end-on or stationary crossflow attitudes. Either of these would cause tracer exposures at a higher altitude than the one at which they would be exposed as a result of tumbling re-entry. It is not thought likely that either of these occurred, however, since either would result in continued traces of the UZrH remaining after the tracer was exposed. The altitudes given for observed burnup are those at which a given identified trace was determined to have completely stopped, whereas traces of the remaining rods continued. As a matter of fact, the paths of the rods containing the silver and gold tracers were identified during the re-entry. These traces could be seen to grow more dim during the latter part of the re-entry, and this agrees with the predictions of insufficient heat to ablate the UZrH and expose the tracer.

In summary, it can be said that the data resulting from the re-entry tracer exposures can be correlated fairly satisfactorily with predictions both from the hyperthermal tunnel tests and from the computer analyses. The agreement is gross in nature, since many transient effects can be postulated which cannot at present be evaluated. Since factors which affect and contribute to sources of heat other than convective heating cannot be evaluated, no attempt is made in this report to extrapolate to an orbital-decay re-entry environment. It is not recommended at this time that net values of aerodynamic heat as observed in the RFD-1 re-entry be used as the effective aerodynamic heating necessary to cause ablation of UZrH in other environments. Partial agreements, as mentioned previously, indicate that the flight-test data may be accurate, but the factors which result in partial agreement and cause the variations between the flight-test values and those predicted both from hyperthermal tunnel tests and oxidation analysis render problematical the question of how an orbital-decay environment would actually affect the aerodynamic heating required. This suggests that further studies, both experimental and analytical, must be done before one can predict the fate of UZrH in a wide range of re-entry conditions with any degree of analytical certainty. Data on the topic are being accumulated from several sources. Additional studies are being made of oxidation, combustion, environmental effects, and other factors.

## SECTION V -- CONCLUSIONS

Accurate monitoring of the ablation of UZrH fuel elements is difficult even under laboratory conditions. An attempt to measure the rate and volume of UZrH ablation at a distance of from 100 to 200 miles with spectral tracers and optical instrumentation therefore had a marginal chance of success. This was recognized from the initial concept of the RFD-1 fuel-rod experiment. Also, there was little information on such an experimental method to provide guidance. It was realized that, at best, only gross information on the ablation of the fuel rods would be obtained. However, even gross information was considered valuable enough to justify including the external fuel-rod experiment on RFD-1. Monitoring the reactor disassembly was a more straightforward operation. Optical coverage was utilized to supplement the TM data from the RV.

Most of the optical data obtained on the RFD-1 flight test were somewhat circumstantial. No attempt was made in this report to rationalize or mitigate any of the uncertainties involved in reduction of the films. However, coincidental agreement among the data from the various films and records indicated that the results are positive enough to justify interpretation and acceptance of them. If the data and results presented in this report are accepted as valid, the following conclusions can be drawn:

1. The strontium-loaded group of RFD-1 experimental fuel rods was completely ablated away at 363.32 seconds ( 161,400 feet). The average weight and volume of UZrH in each rod was 0.955 pound and 4.30 cubic inches. The total aerodynamic heat input to each rod at $363.32 \mathrm{sec}-$ onds was 770 BTU. Therefore, a total of $806 \mathrm{BTU} / \mathrm{lb}$ (aerodynamic heating) was required to complete the ablation at the observed altitude. The predicted time and altitude of complete ablation, based on theoretical calculations and experimental data, and assuming aerodynamic plus oxidation heating, was approximately 364.3 seconds (158,000 feet) (see p. 133). The calculated average BTU/Ib (aerodynamic heating) required for ablation was 843.
2. The barium-loaded group of RFD-1 experimental fuel rods was completely ablated away at 371.50 seconds ( 142,000 feet). The average weight and volume of UZrH in each rod was 1.543 pounds and 7.10 cubic inches. The total aerodynamic heat input to each rod at 371.50 seconds was 980 BTU . Therefore, a total of $635 \mathrm{BTU} / 1 \mathrm{~b}$ (aerodynamic heating) was required to complete the ablation at the observed altitude. The predicted time and altitude of complete ablation, based on theoretical calculations and experimental data, and assuming aerodynamic plus oxidation heating, was approximately 374.6 seconds (135,000 feet) (see p. 133). The calculated average BTU/lb (aerodynamic heating) required for ablation was 787.
3. The silver-loaded rods did not ablate enough to expose the silver tracer. The UZrH glow was not visible after 368.00 seconds ( 149,100 feet). Calculations indicate that the silver-loaded rods would have reached a maximum surface temperature of $3500^{\circ} \mathrm{F}$ at 397.0 seconds ( 96,000 feet). This corresponds to 1375 BTU , or $595 \mathrm{BTU} / \mathrm{lb}$ aerodynamic heat input. This is not enough total heat to complete ablation.
4. The gold-loaded rods did not ablate enough to expose the gold tracer. The UZrH glow was not visible after 368.00 seconds ( 149,600 feet). Calculations indicate that the gold-loaded rods would have reached a maximum surface temperature of $3000^{\circ} \mathrm{F}$ at 408.0 seconds ( 85,100 feet). This corresponds to 1325 BTU , or $485 \mathrm{BTU} / 1 \mathrm{~b}$ aerodynamic heat input. This is not enough total heat to complete ablation.
5. The discrepancy between the observed and calculated ablation is considered to be within the accuracy of the overall experiment. However, the apparently higher observed ablation rate could be due to (a) oxidation heating effects greater than had been predicted, (b) possible exothermic effects of nitrogen, and/or (c) effects of hydrogen combustion.

The observed disappearance of the silver- and gold-loaded rods before disappearance of the barium-loaded rods is not explainable. However, several possible reasons are offered: (a) runaway exothermic reaction of the barium due to higher $\mathrm{BTU} / \mathrm{lb}$ heat input, and (b) increased heat input to the barium-loaded rods due to increased velocity caused by breakup and subsequent higher $W / C_{D A}$. It is conceivable that the two traces identified as the silverand gold-loaded rods were not actually made by these rods. The only identification of these objects was made from a comparison of the theoretical trajectories with the measured trajectories from the plate films. The strontium- and barium-loaded rods were identified from the spectral and framing-camera films as well as from the trajectory comparisons, which makes their identification much more positive.
6. The volume of ablated material was as stated. However, the rate of ablation is not constant enough to be defined.
7. The final sizes of the ablated particles from the RFD-1 external fuel rods is not known.
8. Disassembly events of the reactor were observed (see SC-RR-64-515 for times, altitudes, and final conclusions).


#### Abstract

Considering the factors involved, optical coverage and reduction of the films and records from RFD-1 were highly successful. The experience gained in this operation should not be underestimated. During the design of the experiment and of the optical instrumentation, as well as during data acquisition, computation of theoretical trajectories, and data evaluation, many possible modifications and refinements which would improve the quality of the data obtained in future similar efforts became apparent. The extreme importance of some instruments and operational procedures in acquisition and reduction of the data emerged clearly. In what follows, these features are discussed so that their value can be realized and so that they may be used even more efficiently in future flight tests. Where possible, these recommendations have already been incorporated into the plan for RFD-2.


## Radar

The FPS-16 radar was extremely important to the RFD-l operation and data reduction. The FPS-16 directed the ME-16 and LA-24 cameras to the re-entry, indicated the correct starting time for the cameras, and recorded the re-entry trajectory of the RV. This record was used to confirm the theoretical trajectory for identification of the remaining objects. The theoretical trajectory program has since been modified to compute the position of the RV in terms of spatial coordinates, which will eliminate the need to convert the plate-camera films to terms of azimuth and elevation. The FPS-16 radar is necessary to establish the basic trajectory used in computing these coordinates. It is considered crucial enough to the program to warrant radar redundancy to assure tracking of the RV. Raw (uncorrected) data from the radar track should be furnished for data reduction, since the present TTA program includes built-in correction factors.

The verlort radar is also very necessary. To reduce the films from the aircraft cameras, the position, heading, turn rate, and bank angle of the aircraft, versus time, must be known. Film from the NASA aircraft would have been much more useful if that aircraft had been tracked by radar. Conceivably, however, the verlort radar could be switched between the RV and the aircraft during re-entry, so that spot checks of the positions of the RV and aircraft would be available.

> Re-entry Objects

The number of extraneous objects re-entering the atmosphere along with objects from the primary experiments should be kept to a minimum. Items such as brackets, bolts, and bands interfere with the identification of the primary objects and events of the experiment. Of course, most of the items are necessary, but where possible, vertical separation of the re-entering parts should be attempted by means of ejection or retro-rockets. This would lessen the effects of superposition of the plate-camera film streaks.

## Ejection Measurements

Monitoring the orientation of the $R V$ when objects are ejected is essential, since the directions of ejection are applied as perturbations to the initial conditions for computing the theoretical trajectories of individual objects. The accuracy, or lag, of the roll-stable gyro system should be determined before the flight test.

Velocities of all ejected objects should be measured during development of the test unit in the laboratory. Framing cameras furnished an accurate history of the ejection velocities and directions of the RFD-1 external fuel rods and brackets during preflight tests. Calculation of the velocities and directions is not accurate enough, because objects do not release instantaneously.

Tracers

Tracer elements should be carefully selected on the basis of special criteria. The most important of these criteria are low excitation energies for intense lines (see pp. 84 to 86).

## Identification of Materials

During reduction of the RFD-1 spectral films, many elements were identified as present in very small amounts in the RV, fourth-stage motor, brackets, and fuel rods. This indicates the possibility of identifying many objects, in addition to the tracer materials, by spectral means. The specifications of all materials included in any of the re-entry items should be recorded for use in postflight data reduction. In addition, specimens of the actual materials should be retained for laboratory spectral analysis. Often, impurities or elements not listed in the specification but present in the materials can be identified in the laboratory.

## Aircraft


#### Abstract

Airborne optical instrumentation has both advantages and disadvantages for re-entry investigations. The disadvantages of vibration, power limitations, lack of availability when needed, window restrictions, and other problems warrant con-


 centrated efforts at solution.Among the advantages, the positions of the stations are flexible: they can be located anywhere along the trajectory, and, if desired, the trajectory and stations can even be moved to another part of the world. In addition, such stations can be above clouds or ground-level haze. However, by far the biggest advantage of aircraft stations is the lesser atmospheric attenuation of the radiation reaching the detectors, even under clear-sky conditions. An approximate calculation based on the tables for a tropical maritime atmosphere, taken from Edgerton, Germeshausen and Grier, Inc., Report No. B-2621, November 15, 1963, Spectral Transmission of Slant Paths Through the Atmosphere, Vol. II, will exemplify the problem existing for distance objects, particularly at the shorter wavelengths. These calculations neglect the familiar $R^{2}$ falloff of intensity, which adds to the problem when a distant, point-source object is involved. At $4000 \AA$ and a slant range of 50 nautical miles from a sea-level detector to a source object at 50,000 feet, about 0.35 percent of the radiation is transmitted. At $4000 \AA$ and a slant range of 50 nautical miles from a detector at 10,000 feet to a source object at 50,000 feet, about 9.3 percent of the radiation is transmitted. At $6000 \AA$ and a slant range of 50 nautical miles from a detector at sea level to a source object at 50,000 feet, about 3.0 percent of the radiation is transmitted. At $6000 \AA$ and a slant range of 50 nautical miles from a detector at 10,000 feet to a source object at 50,000 feet, about 43.5 percent of the radiation is transmitted.

If the atmospheric conditions were poorer than "very clear," the results would be even more exaggerated. Assumptions and calculations based on aerosol and Rayleigh scattering are discussed in the cited report. The examples were not worked out for slant ranges greater than 50 nautical miles, because it was believed that the additional range and higher altitude involved in RFD-1-type re-entry experiments would reduce the atmospheric transmission by about the same amount for a ground station as for a station at 10,000 feet altitude.

The value of different optical stations was apparent during the RFD-1 data reduction. Trajectories that lay in the line of sight from one station were seen as different trajectories from other stations. Downrange stations, such as the NASA aircraft, were able to cover the entire portion of the flight path of prime optical interest. Two or more airborne stations would be preferrable, if most of the optical instruments have fixed axes, because if one aircraft should be off course for any reason, all coverage by that station would be lost. Two or more aircraft are desirable, even if they do not cover the same region in space, because they can then be positioned closer to the flight path. The light intensity reaching the cameras would then be increased by both the distance-squared effect and the atmospheric-attenuation effect. The advantages of different view angles, atmospheric-attenuation factors, and general redundancy in data acquisition favor having two or more airborne.optical stations. Aircraft mobility, of advantage in data acquisition, becomes a distinct disadvantage in the reduction of the platecamera films. To accurately identify the spatial position of the re-entry objects, the exact position and orientation of the observer (cameras) must be known. Consequently, tracking of the aircraft by the verlort radar during the time of picture taking is essential in order to locate the latitude and longitude of the aircraft. In addition to the location of the aircraft, its heading, turn rate, bank angle and velocity, versus time, must be known. A flight recorder is excellent for the purpose.

Orientation (azimuth and elevation) of the cameras with reference to the aircraft also enters into the reduction of the data. This can be measured either before or after the flight, since orientation is not variable with time. However, one method proposed for accurate measurement is to align the aircraft on the ground in a known direction and then photograph a distant object of known azimuth and elevation (such as a mountain range or building) with all the cameras. Reduction of these films would furnish a complete picture of the relative orientation of all the cameras for later reference. Aircraft cameras must be mounted to minimize all effects of aircraft vibration. Vibration causes the plate-camera film details to be blurred and superimposed; several of the RFD-1 films were unusable because of this.

## Photometer and Filters

The quality of the filters used in the photometers on RFD-1 must be upgraded for future experiments. The effective band passes of the filters used on RFD-1 were too wide to yield conclusive information. Narrow band-pass filters with high transmission are a necessity for experiments similar to RFD-1 if the shorter wavelengths are of interest. The filters should be checked in the laboratory to determine the precise peak of their transmission maxima at the temperature at which they will be used, since the transmission peak shifts with temperature. The band pass of the filters should also be checked.

Atmospheric attenuation is such that the photometer is probably the only ground-based instrument with sensitivity enough to be able to detect radiation in the blue and ultraviolet region of the spectrum. When the ultraviolet is of interest, ultraviolet-transmitting optics must be used. The advantages of aircraft for the short wavelengths have been discussed, and must again be emphasized. Many elements have their most intense spectral lines in the ultraviolet, and use of this spectral region would increase the available information on spectral events during re-entry. To adequately investigate the ultraviolet would require obtaining, or designing and building, special instruments in addition to the photometers.

Color film was used for most motion-picture coverage of the RFD-1 re-entry. The color rendition seemed to be quite good, and it aided in flare identification. However, with time as the only reference, it was often difficult to determine just what was being tracked with any of the tracked instruments. The color-film exposures indicated that the emulsion was quite fast. In the past, people have often used black-and-white film to maximize exposure, but the color film used seemed to have good sensitivity. The underexposures which occurred are attributable to shutter speed and aperture problems. Color film for some of the open-plate trajectory cameras would provide another means of identifying objects which emit characteristic colors. The black-and-white trajectory plates on RFD-1 were good, so the use of color plates in some cameras should be considered as an addition to, rather than a substitution for, black-and-white emulsions. Some of the emulsions used for RFD-1 proved to have imitations, particularly for the spectral cameras. The Royal-X Pan (RXP) film used by NASA and AFSWC lacks the requisite sensitivity in the red region of the spectrum beyond about $6350 \AA$. This meant that none of the cameras using RXP detected the lithium line at $6707 \AA$ or any other radiation in the red region. RXP seemed very satisfactory as far as speed and graininess were concerned. In general, the 103-F or Aero-Recon Tri-X used by Sandia on RFD-1 would be preferable when the red region is of interest. The range of some of the limited-range films has supposedly been extended. If use of these new emulsions is contemplated, they should be laboratory-tested first for sensitivity and wavelength range. Faster, finer-grained emulsions with greater wavelengths would be desirable if they can be obtained.

After events on the framing cameras had been viewed and evaluated, it was often difficult and time-consuming to find, later, the particular frames of interest. The printing of sequencing numbers on the edge of these films would be a very valuable aid in reducing the data from them.

## Focus

The focus on several instruments was not quite as good as it should have been. If possible, all cameras should have complete focusing runs under operating conditions as close to the actual ones as possible. Camera bodies, lenses, and backs should be considered one unit after alignment. Gratings should all be checked for quality before they are used. Reports from other investigators indicate that gratings are sometimes not up to specifications and could create poor images.

## Frame Rates - Focal Length and Exposure

Shutter speeds were too high on most of the framing cameras to give adequate exposures for RFD-1. NASA cameras operated at lower frame rates than did Sandia cameras, and their exposures were better. On the other hand, in case intensities should be higher than expected, not all cameras should be operated at very low frame rates. More exposure can also be obtained by using optics of larger aperture. Higher frame rates allow more frames of an event to be recorded if the event has a duration of a goodly portion of a second; higher rates thus reduce the difficulty of distinguishing between film imperfections and actual exposure. Also, in the case of short-duration events, the length of camera "dead time" between any two frames is reduced. The preferred method would seem to be to run most cameras at slow framing rates, but to run those with large-aperture optics which are tracking objects expected to be bright at frame rates up to 100 frames/ sec. The cinespectrograph data were radically underexposed, so this instrument should have larger-aperture optics and a slower frame rate. If it were not for the better time resolution of the cinespectrograph, it would be better to replace it with a streak spectrograph. Two-thirds of the cinespectrograph's operating time is "dead time." The best solution is to use both instruments and obtain the advantages of both.

Reduction of data from the trajectory camera was difficult because of the absence of separation between the traces recorded by different re-entry objects. The separation of two traces on the photographic plate is proportional to the focal length of the recording camera, so it is advisable to have lenses with focal lengths as long as possible while still remaining compatible with the other restrictions imposed by longer focal lengths. Objects at a distance of from 100 to 200 miles can be regarded as point sources. Since recording rate along the plate is doubled when the focal length is doubled, maintaining the same exposure requires doubling the aperture. However, large lenses of good quality are expensive.

## Low-Rate Framing Cameras

Plate-trajectory cameras with short-duration timing chop are essential for accurate data reduction. However, such cameras do have limitations, for they show the various re-entry streaks superimposed on one another. Not only are the timing marks somewhat obliterated as a result, but also the individual objects are difficult to distinguish. A low-rate framing camera (such as a modified K-37) would be free of some of these problems. By exposing each frame for only a short time (1 to 4 seconds), the problem of overlap, except from immediately adjacent objects, would be lessened. Since the framing pulses could be recorded, a measure of time could be obtained for correlation with other films. If two alternating cameras were employed, "dead time" between frames would be eliminated. This system would work equally well for spectral cameras.

Camera Orientation

Camera orientation should be carefully considered for nontracked instruments both on the ground and in aircraft. The spectral and open-plate-trajectory camera should have overlapping fields. When gratings are in front of the lenses and light intensity is expected to be a problem, vignetting by the grating should be minimized. If light intensities are expected to vary over a larger range, a grating with reduced dispersion should be considered, since it would allow both first-order and second-order spectra to record. Such a grating would enable two intensity levels to be recorded and hence would increase the probability of proper exposure, but it would have as a disadvantage the fact that the lens behind the grating would be operating farther off axis. If lens aberrations were a problem, these distortions could become serious.

## Tracking Cameras

Power scopes should be included on all hand-tracked cameras as well as on the ME-16 and LA-24 tracking telescopes, since visual acquisition of items as small as the external fuel rods is difficult. The 12 external rods in the RFD-1 re-entry were individually visible through the 5.6 - and 32 -power scopes.

A tape recorder for the optical trackers on ME-16 and LA-24 tracking telescopes to permit verbal recording of the re-entry picture by the operators during the re-entry would be advisable. Recalling the entire re-entry sequence after the conclusion of the test is difficult, and accounts frequently conflict.

The reduction of data from a large number of optical instruments creates a heavy work load. An effort should be made to computer-program as much of the data reduction as possible. In any case, computer reduction of data is usually more thorough and accurate than manual reduction. The theoretical trajectory program is one example already in operation. Other aspects of the data reduction which are adaptable to computerization are the photometer records and the plate, trajectory, and spectral films (by densitometry). Some records, such as the framingcamera films, are not adaptable to computer programs. However, their reduction can be made more accurate and efficient through the use of analyzing projectors, film editors, and edge-numbered films.

A record of the ME-16 and LA-24 tracking mount azimuth and elevations versus time would be of benefit in the data reduction. Motion-picture films furnish no indication of the position in space of the objects being tracked. Correlation of their azimuths, elevations, and time with those of the other films would provide a more complete picture of the re-entry sequence for preliminary and final reduction. This would also aid in relating the various films to one another. If azimuths and elevations versus time had been known for the LA-24 during RFD-1, a positive identification of what it had tracked would be available.

Inclusion of a barometer and thermometer at the High Point site is suggested. These would provide a measure of pressure and temperature during the re-entry, permitting atmospheric-refraction corrections to be made in preliminary reduction of the data until the reduced meteorological data are received. High-altitude radiosonde weather balloon soundings would provide adequate atmospheric data during the re-entry. Atmospheric conditions above 100,000 feet are reasonably stable. The standard 1959 atmosphere was very similar to the measured RFD-1 atmosphere.

The members of the data-reduction group who will be charged with reducing the flight-test data should be consulted about the design of the experiment during development of the instrumentation and flight-test details. Preparations for data reduction should be initiated before the test and included in the design.

Handling and use of the original films from the test should be kept to a minimum. Whenever possible, analysis and reduction should be done with prints of the original. No cutting or modification to the films should be allowed, since some reduction techniques utilize the frame dimensions as pseudo-fiducials. In the event of damage or breakage of the framing-camera film, repairs should be made so that the original length is maintained. Interpolation between timing marks is often necessary. Recording of the timing, camera functions, and photometer pulses would be better done on magnetic tape than on oscillograph paper, since playback of a tape can be condensed or expanded to the optimum scale for any given data of interest, and since tapes can be digitized for computer reduction of the data.

## Timing

The lack of timing on some RFD-1 data proved to be a serious limitation. Redundant timing systems would be desirable. In addition, more accurate time resolution would make it easier to correlate events recorded on various instruments.

Chopped-trajectory plate-camera films should employ a chopping code which allows a long exposure time relative to closed-shutter time. The RFD-1 chopping code included 7 seconds exposure time versus 2 seconds closed-shutter time. This ratio was adequate, but the long exposure time made interpolation of event times inaccurate, and the long closed-shutter time made completion of the chopped-out portion extremely difficult.

The RFD-1 unchopped plate-camera films furnished no indication of camera orientation, because the position of the star trail background versus time was not known. Consequently, corrections for atmospheric refraction and lens distortion were difficult to make, and accurate correlation with other data was not possible. It has been suggested that in future similar tests, these cameras be uncapped for a short time before and/or after the re-entry. The exact times of these operations should be recorded on the tape or oscillograph in terms of WWV or BRT. This would allow the star trail background to be used in reduction of the films.

## Micro-Densitometer

The theoretical trajectory program has been modified to furnish positions of the re-entry objects in terms of film coordinates. Employment of a microdensitometer to read the plate films in terms of a two-dimensional coordinate system would expedite and improve the quality of the data reduction. The resolution of the densitometer is much finer in distinguishing superimposed objects than that of the Mann Comparator, and it also provides a measure of the intensity of the streaks. This instrument could be digitized to furnish the data in a tabulated form for direct comparison with the theoretical trajectory. Experiments using the micro-densitometer on RFD-1 plate films have indicated its usefulness; it may also have possibilities for the analysis of spectral plates. This will be investigated further before the RFD-2 flight test.

## Weather Versus Launch

The decision to launch RFD-1 was delayed several weeks because of clouds and haze in the re-entry area. To determine the amount of haze and clouds in the area of the re-entry trajectory just before launch is extremely difficult. It has been suggested that an optical instrument be developed to scan the re-entry envelope to determine the probable attenuation of the re-entry light due to the atmospheric conditions. The equitorial telescope located at High Point, Bermuda, could be employed for this purpose. With the telescope, the brightness of known stars in the re-entry envelope could be measured periodically. The decision to launch could then be based on a comparison of star brightness with a predetermined brightness required for adequate optical coverage of the re-entry.

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[^0]:    *Includes insulation and alumina end plugs.

[^1]:    Excellent film; shows RV, fuel rods, and brackets; good timing; used extensively in data analysis.

    Good film; fuel-rod flare colors confirmed identification of tracers.

[^2]:    As mentioned previously, the objective of the experiment was to monitor fuel-rod ablation. Rate and volume of fuel-rod ablation were to be monitored by detecting flare materials inside the fuel rods and determining when and where in the trajectory these flare materials were exposed. The flare materials chosen were ones not likely to be present as part of or as impurities in the re-entering objects. Furthermore, they were materials with physical properties expected to be compatible with the re-entry environment, including: (1) vapor pressures such that the fuel rods would not be ruptured before ablation down to the tracer material, (2) moderate refractoriness, so as not to resist vaporization altogether and so preclude spectral excitation, and (3) spectral characteristics such as to be relatively easily detected, even in small abundances. Since Item (3) was obviously very important to the success of the experiment, excitation energies required for specific spectral lines, and the transition probabilities for these lines (the best available transition probabilities, whether theoretical or empirical) were carefully considered. The elements finally chosen for tracers were barium, strontium, silver, and gold.

