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Interaction of Fast Waves with Ions

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Abstract. To fully utilize the available power sources in DIII-D (FW, NBI, ECH), understanding of the synergism between the heating mechanisms is important. In this paper the ion distribution, under simultaneous application of NBI and FW, is calculated from Fokker-Planck code CQL3D coupled to ray-tracing code CURRAY. It is found that interaction between energetic ions and FW can be minimized or maximized by adjusting various parameters such as magnetic field, density, beam energy, and FW frequency. Specifically, in DIII-D, we find negligible interactions above 1.8 T and above 80 MHz, while the interaction increases at lower fields and frequencies. The results are compared with experiments in DIII-D including the calculated neutron rate. Energetic ion orbit losses may play an important role in the ion distribution, and this effect is being investigated.

In order to reach steady-state reactor relevant tokamak parameters, DIII-D will have available various types of power sources for heating (H) and current drive (CD), including 6 MW of fast waves (FW) in the range of 30 to 120 MHz, 20 MW of neutral beam (NB), and electron cyclotron (EC) waves [1]. To optimize the use of available power, it is natural to consider simultaneous application of the sources. The question then arises as to whether there is any appreciable synergism between the power sources. In this paper, we consider the interactions between FW and the energetic ions produced by NB.

The ray tracing code CURRAY is used to determine the FW wave fields and polarizations. Typically, about 28 to 42 rays are calculated in the main spectra which are determined from a coupling code V1DARY (coupled to FELICE) [2]. The wave information from CURRAY are input into the bounce averaged Fokker-Planck code CQL3D [3] which calculates the ion distribution in the presence of the rf quasilinear diffusion and neutral beam sources on specified flux surfaces. We use typical DIII-D parameters: $B_0 = 1$ to 2 T, $n_e = 1$ to $3 \times 10^{13}/\text{cm}^3$, $T_e = T_i = 2$ to 3 keV, neutral beam power $P_b = 3$ MW, and $P_{\text{fw}} = 1.5$ MW. An ITER L-mode transport scaling is used to determine the relationship between density and temperature.

In the absence of rf, a slowing down distribution due to NBI is produced. When the rf is applied, a large energetic ion tail is developed in CQL3D. This results in enhancement of the absorption of FW by the tail species, and a decrease in electron absorption. The results of parameter scans are summarized in Figs. 1 and 2. It is seen that the ion absorption is decreased by increasing B , decreasing n_e . This behavior can be qualitatively understood as follows. The perpendicular phase velocity of FW can be approximated by the Alfvén velocity v_A , then, the relevant argument of the Bessel function expansion for harmonic absorbed power at the beam energy can be written as

$$\left(\frac{k_{\perp} u_{b\perp}}{2\Omega} \right)^2 \approx (0.073 n_{13} E_b / B^2) \sin^2 \theta, \quad (1)$$

where n_{13} is density in $10^{13}/\text{cm}^3$, E_b is beam energy in 10 keV, B is in tesla,

FIGURE 1. Absorption versus B-field; solid line is for $n_e = 3 \times 10^{13}/\text{cc}$, $T_{e0} = 2.3$ keV; dotted line is for $n_e = 2 \times 10^{13}/\text{cc}$, $T_{e0} = 3.3$ keV, NB power is 3 MW; rf power is 1.5 MW, E-beam is 75 keV, rf frequency = 60 MHz.

FIGURE 2. Absorption versus average density for two fields; solid line for $B = 1.4$ T; dotted line for $B = 1.8$ T; NB and rf parameters same as Fig. 3.

and θ is pitch angle. At the ℓ -th harmonic $\ell \approx \omega/\Omega$, the ratio of rf-diffusion to collisional diffusion is roughly given by

$$\frac{|D|}{\nu_s D_c} \propto \frac{P n_e^{\ell-2.5} E_b^\ell}{B^{2\ell-3} T_e} \quad (2)$$

Thus for DIII-D parameters where $\ell > 4$, the rf-diffusion decreases with decreasing density and beam energy, and with increasing B -field. The above were for an rf-power at 60 MHz. The interaction is shown to be much weaker at a higher frequency of 80 MHz (Fig. 3).

The neutron rate produced by the energetic ions has been measured with and without rf. We can calculate this rate in CQL3D. Comparison between theory and experiment, however, indicates that the quasilinear theory overestimates the energetic tail produced by rf. One likely origin of the discrepancy is that orbit loss is not taken into account in the calculations. Inclusion of this effect in CQL3D is being implemented and shall be discussed.

In conclusion, the interaction between NBI and FW is investigated. We find that the interaction can be minimized by reducing the density and beam energy, and by increasing the B -field and frequency. At 1.8 T and 80 MHz, the ion absorption becomes small compared with electron absorption. A comparison with experiment, however, indicates that the Fokker-Planck calculation considerably overestimates the energetic ion enhancement. We think a likely source is the neglect of orbit loss in the calculations.

FIGURE 3. Absorption versus B-field for two frequencies; solid line for $f = 60$ MHz, dotted line for $f = 80$ MHz, beam energy is 60 keV, $n_e = 2 \times 10^{13}/\text{cc}$, $T_{e0} = 3.3$ keV.

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