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the Second Alfvén Gap in TFTR

by

E. Fredrickson, J.W. Van Dam, R.V. Budny, D. Darrow, G.Y. Fu,
J. Hosea, C.K. Phillips, and J.R. Wilson

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Observation of Modes at Frequencies Near the Second Alfvén Gap in TFTR


Princeton Plasma Physics Laboratory, Princeton, New Jersey
*Institute for Fusion Studies, University of Texas at Austin, Austin, Texas

Abstract. Modes have been observed near the frequency of the second Alfvén gap during off-axis H-minority heating experiments in the circular cross-section Tokamak Fusion Test Reactor. The observation of these modes is surprising in that the second gap, which is generally opened with ellipticity, is expected to be small, of order \((r/R)^2\). A model is proposed in which the second gap is opened by the fast ion beta, which is shown to be able to introduce mode coupling, much as toroidal effects introduce mode coupling for Toroidal Alfvén Eigenmodes (TAE). The modes are seen with and without accompanying TAE mode activity.

INTRODUCTION

Ion Cyclotron Range of Frequency (ICRF) conditioning and H-minority heating experiments have been done under a wide range of conditions on the Tokamak Fusion Test Reactor (TFTR), predominantly with the resonance near the magnetic axis. Two types of modes at frequencies near the first Alfvén gap were commonly observed during these ICRF heating experiments. These modes were identified as Toroidal Alfvén Eigenmodes (TAE) and a fast ion mode [1], probably the Energetic Particle Mode (EPM) [2]. Damage to plasma facing components ascribed to fast ion losses driven by these modes [3,4] encouraged efforts to avoid them. Moving the ICRF resonance from near the magnetic axis to the high field side beyond the q=1 radius
was found to greatly reduce the frequency of occurrence and amplitude of the TAE modes.

The off-axis resonance on the high field side introduced a third mode. Approximately 90% of the plasmas with the resonance off-axis on the high field side and ≥ 4 MW of RF power had modes near the frequency of the second Alfvén gap. These off-axis RF experiments were performed primarily to condition the RF systems and were performed for a limited range of parameters. The “second gap” modes were observed under the following conditions: $I_p = 1.3$ MA, $B_{tor} = 2.5$ T, $f_{ICRF} = 43$ MHz, $R_0 = 2.62$ m, $R_{res}(H) = 2.35$ m, $a_p = 0.97$ m. The spectral peaks at the second gap frequency are not harmonics of TAE modes, nor non-linear mixing of the TAE activity. As desired, in these experiments the TAE activity is often weak or non-existent.

The observation of modes in the range of frequencies of the second gap is surprising in the circular TFTR plasmas since the second gap is generally thought of as the Alfvén gap that is induced primarily by ellipticity. Toroidicity effects on the second gap are negligibly small, of order $\varepsilon^2$, where $\varepsilon = r/R_0$ is the inverse aspect ratio. However, it has been predicted that the second Alfvén gap can be opened by the presence of an energetic trapped ion population [5,6]. Estimates suggest that there is sufficient fast ion beta in these TFTR plasmas to open the gap and drive the mode.

In a very few of the on-axis ICRF heating experiments, peaks in the magnetic fluctuation spectra were observed at twice the TAE frequency. These peaks were only observed in the presence of strong TAE or EPM activity and the frequency peaks were sums of the TAE frequencies, suggesting that these modes represent non-linear mode mixing of the TAE modes. The data will not be discussed in more detail in this paper.
EXPERIMENTAL OBSERVATIONS

The “second gap” modes are seen at frequencies near the second Alfvén frequency gap, approximately 407 kHz to 411 kHz. The modes appear as short bursts characteristically lasting less than or of order 50 msec. The modes are not harmonics of the TAE modes sometimes seen in these same plasmas. This can be seen in Figure 1 where the spectrogram of Mirnov coil data shows the TAE activity in the frequency range from 185 kHz to 200 kHz and the 2nd Gap Mode Activity in the range 398 kHz to 422 kHz. The time behavior of the 2nd Gap Modes is qualitatively different than that of the TAEs and the harmonics of the TAEs would be at frequencies lower than 400 kHz.

The mode behavior is correlated with the sawtooth period, illustrated in Fig. 1 by the inclusion of a central soft x-ray chord showing the sawtooth times. Most commonly the mode frequency chirps down, changing frequency in each burst by about 5 kHz in the 30–50 ms that it is present, and in this case by about 20 kHz over the period of the sawtooth. The bursts often consist of multiple modes, separated in frequency by a few kHz, e.g., the bursts between 4.3 and 4.4 s and between 4.4 and 4.5 s.

The TAE (and EPM) modes propagate in the ion diamagnetic drift direction [2] but the 2nd Gap Modes propagate in the opposite direction. This can be seen in Figures 2a and 2b. In Figure 2a is shown the phase information from the Fourier analysis of the 2nd Gap Mode burst between 4.28 and 4.3 sec in Fig. 1. The data is best fit assuming a dominant toroidal mode number, n = +4. The slope of the fitted line gives the toroidal propagation direction. For comparison, the phase information from the TAE mode in Fig. 1 in the same time period is shown in Fig. 2b. As can be seen, the direction of propagation is reversed. The phase data from this particular TAE mode is best fit with a toroidal mode number of n = -5. The lower frequency TAE mode is best fit with n = -4 and the higher frequency mode with n = -6.

In the course of the RF conditioning campaigns, variations in limiter conditions resulted in a range of central plasma densities from $3.2 \times 10^{19}$ m$^{-3}$ to $4.2 \times 10^{19}$ m$^{-3}$. From this data set it was possible to examine the scaling of the mode frequency with
density. The mode frequency is in good agreement with the predicted frequency for the second gap and scales as $n_e(0)^{1/2}$, consistent with Alfvén-like modes [Fig. 3]. The existing data set has no toroidal field variation from which to derive a magnetic field scaling.

With the off-axis resonance on the high field side, the sawteeth were not stabilized in these plasmas in the range of input power levels (< 5 MW). The sawteeth flatten the density profile in the core region, dropping the density within the $q=1$ surface. The chirping of the mode frequency is consistent with the time evolution of the central density. In Figure 4 the mode frequency evolution is overlaid with the second Gap frequency calculated using the central density, $f_{TAE} \equiv \frac{0.96 \, V_{Alfvén}(0)}{(2\pi q R_0)}$ including a 4% correction to qualitatively match the mode frequency evolution. The second gap mode frequency drop is consistent with the recovery of density peaking between sawtooth crashes, indicating that the modes are localized within the $q=1$ surface.

The linear growth rate of the mode, shown for two representative cases in Fig. 5, varies from about $2 \times 10^4$/s to $5 \times 10^2$/s. The rate can be nearly constant over three e-foldings of the mode amplitude. The mode amplitude then saturates and can remain nearly constant or have slow variations. Sawtooth crashes terminate the bursts. It is assumed that sawteeth flatten the fast ion beta within $q=1$. This observation is consistent with the mode’s drive being within $q=1$.

The saturation amplitude of the magnetic fluctuation (in Gauss) is roughly linearly proportional to the linear growth rate [Fig. 6]. This conclusion is contingent on the accuracy of the assumption that the mode structure does not change substantially during the period of the mode growth, so that the external magnetic fluctuation level is proportional to the internal magnetic fluctuation level. This linear correlation between the saturated magnetic fluctuation amplitude and the linear growth rate is not consistent with the normal saturation mechanism arising from wave trapping of fast ions where the saturated wave amplitude is proportional to the square root of the growth rate [7].
More complicated growth behavior is also seen as illustrated in Fig 7. The mode amplitude does not necessarily remain constant after “saturation”. This could reflect changes in equilibrium parameters such as the density profile or fast ion distribution. It could also result from the non-linear effects in a weakly driven system when the mode linear growth rate is slow compared to the fast ion replenishment rate [7].

**ANALYSIS**

The RF heat deposition and tail formation are modeled with the TRANSP code [8] and the SPRUCE RF package. Sawteeth are assumed to affect the fast ion tail population, and this is modeled in TRANSP with a mixing model that assumes a Kadomtsev-type flux reconnection at each sawtooth crash. The RF resonant position is off-axis on the high field side, presumably resulting in few deeply trapped fast ions. The fast ion beta is calculated to vary from flat to hollow during the sawtooth cycle. Prior to each sawtooth crash the fast ion profile becomes hollow, peaking off-axis at a minor radius of about 15 cm in a 96 cm minor radius plasma [Fig. 8]. The sawtooth model used in this TRANSP run mixes the fast ion population within the reconnection layer.

The calculated RF power deposition layer is broad. However, the width depends on assumptions about the fast ion redistribution at each sawtooth crash and the H minority fraction that is assumed. The peak fast ion beta varies from about 0.6% just prior to a sawtooth crash, to about 0.4% after the sawtooth crash mixes the fast ions, flattening the profile. This fast ion beta is estimated to broaden the 2nd Gap to about 10 kHz from 4 kHz.

The hollow fast ion pressure profile prior to sawtooth crashes is a possible explanation for the “backwards” mode propagation. An alternative explanation is that high field resonance ICRF heating can create an inverted energy population (at constant magnetic moment). This can also provide a source of free energy to drive Alfvénic modes and in some cases this mechanism can couple strongly to counter-propagating modes [9]. More detailed calculations of the damping and driving terms
still need to be done. The agreement between the frequency evolution and the central
density evolution suggest that the modes are most strongly affected by core (within
q=1) plasma parameters. This would be consistent with the supposition that the
“backward” wave propagation is due to the hollow fast ion profile.

The TRANSP and SPRUCE code modeling predict a fast ion tail temperature of order
30 - 50 keV at the center of plasma. The required tail energy for the precessional
resonance for the n=4 EAE at 400 kHz is $\geq 1\text{MeV}$, thus it is unlikely that the
precessional resonance is driving the mode. For the bounce resonance at 400 kHz, the
condition is $E_{\text{hot}} \geq 900 \text{ keV} / l^2$ where $l$ is an integer. Thus, the bounce resonance
condition can be approximately met for $l \approx 3$. This high harmonic bounce resonance
is peculiar to off-axis heating. The reason is that the tail temperature in off-axis heating
cases is much lower than the on-axis cases so that the precessional drift resonance is not
possible.

**SUMMARY**

Modes have been observed in the frequency range of the second Alfvénic gap in
H-minority ICRF heated plasmas in TFTR. This observation is surprising in that the
second gap is generally considered to be small in circular cross section plasmas. The
mode is inferred to be a “core” mode, i.e., localized in some sense within the q=1
surface. This follows from the observation that the time dependence of the mode
frequency is consistent with the changes in the central density, with the appearance of
the mode in the latter part of the sawtooth period when the central fast ion beta has
peaked up, and with the direction of propagation. The last of these being explained in
part by a hollow fast ion beta profile, which is only present in the core region. The
modes are generally not observed during on-axis H-minority heating, but commonly
observed during off-axis heating on the high field side (with the resonant layer
outside the q=1 surface). A model has been proposed that the beta of the fast ions
opens the second gap, allowing instability. For TFTR parameters, the model predicts
a gap width of approximately 10 kHz, which is 2.5% of the second gap frequency. If
the backwards mode propagation is due to a hollow fast ion profile (as indicated in
the TRANSP calculations), then instability due to wave-particle resonance at the magnetic curvature precessional frequency can occur only if the precessional frequency is reversed. This can indeed be the case for off-axis heating on the high field side. Thus, trapped fast ion pressure effects seem to explain several of the observed features of these second gap fluctuations.

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REFERENCES

FIGURE CAPTIONS

Figure 1. Spectrogram of the second gap modes and the TAE modes. A central soft X-ray chord is included to illustrate the sawtooth times. (87838).

Figure 2a. Best fit to phase analysis of EAE burst at 407 kHz, 4.28 sec using a toroidal mode number of +4, propagating in the counter-to-plasma current direction; nominally the ion diamagnetic drift direction.

Figure 2b. Best fit to phase analysis of TAE mode at 192 kHz, 4.28 sec using a toroidal mode number of −5, propagating in the nominal electron diamagnetic drift direction.

Figure 3. Frequency scaling of second gap modes with central density.

Figure 4. Spectrogram of the second gap modes. Dashed line shows the approximate central Alfvén frequency, \( f_{\text{Alfvén}} \approx \frac{V_{\text{Alfvén}}(0)}{(2\pi qR)} \).

Figure 5. The linear growth rate can vary by up to a factor of three, even within the same shot.

Figure 6. The saturated mode amplitude is proportional to the linear growth rate (here normalized to mode frequency).

Figure 7. The saturated amplitude of the bursts varies considerably in time during the bursts.

Figure 8. Fast ion beta before and after sawtooth crash. The q-profile is also shown.
Figure 1

Soft x-ray (a.u.)

Time (sec)

Frequency (kHz)

2.5

87838

Central Chord
Figure 2a

![Graph showing the relationship between Toroidal Angle (degrees) and Relative Phase (radians). The graph includes data points for Shot 87838.](image)
Figure 2b

Shot 87838

Toroidal Angle (degrees)

Relative Phase (radians)
Figure 3

\[ f_{\text{obs}} = f_{\text{Alfvén}} \]

\[ f_{\text{obs}} = 0.5 \times f_{\text{Alfvén}} \]

- 2nd Gap mode
- TAE

\[ f_{\text{Alfvén}} = \frac{V_{\text{Alfvén}}(0)}{2\pi R} \]
Figure 4
Figure 5

Amplitude (a.u.)

10 msec

$\tau_g = 0.6$ ms

$\tau_g = 1.8$ ms

10 msec
Figure 6
Figure 7

![Graph showing Amplitude vs Time](image-url)
Figure 8

![Graph showing major radius (m) vs. beta fast (%) before and after sawtooth.](image-url)