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## VOLUME 2

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# The Art and Science of Magnet Design Selected Notes of Klaus Halbach 



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## Preface

This volume contains a compilation of 57 notes written by Dr. Klaus Halbach selected from his collection of over 1650 such documents. It provides an historic snapshot of the evolution of magnet technology and related fields as the notes range from as early as 1965 to the present, and is intended to show the breadth of Dr. Halbach's interest and ability that have long been an inspiration to his many friends and colleagues.

As Halbach is an experimental physicist whose scientific interests span many areas, and who does his most innovative work with pencil and paper rather than at the workbench or with a computer, the vast majority of the notes in this volume were handwritten and their content varies greatly-some reflect original work or work for a specific project, while others are mere clarifications of mathematical calculations or design specifications. As we converted the notes to electronic form, some were superficially edited and corrected, while others were extensively re-written to reflect current knowledge and notation.

The notes are organized under five categories which reflect their primary content: Beam Position Monitors (bpm), Current Sheet Electron Magnets (csem), Magnet Theory (thry), Undulators and Wigglers (u-w), and Miscellaneous (misc). Within the category, they are presented chronologically starting from the most recent note and working backwards in time. The note number, listed in the Table of Contents and at the bottom of each note's first page, comes from a database we have created which includes the titles of the entire collection of notes, and a recently added sixth category, Conformal Transformations (ctr). The appendixes contain a table of all the notes in the database and a list of Dr. Halbach's publications.

The extensive use of hand-written notes by Dr. Halbach leads us to believe that there may be many that were sent to colleagues which were not retained in Dr. Halbach's files, and thus are missing from the database. If you happen to have a note of scientific interest from Dr. Halbach and believe it to be an original, we would appreciate receiving a copy.

Brian M. Kincaid

November 1994
Simonetta Turek

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## Halbach Geometries

The Halbach Geometries, referred to in the notes as Gm, are a collection of simple geometric shapes, simple function representations, and 2-dimensional electromagnetic geometries for which conformal mapping calculations have been done to compute basic features such as capacitance, excess flux, etc. For examples of calculations of excess flux, see documents 0336csem (p. 5), 0332csem (p. 11), 0183csem (p. 23), and 0131u-w (p. 175).

The following two pages summarize Dr. Halbach's representations and shorthand notations of his "Geometries." The reader is encouraged to refer back to them when encountering such abbreviations as Gm 3 or Gm 21 while reading the notes. (Note: Not all the Halbach Geometries are referenced in this collection.)


## Exact, Complete Proofs of Reciprocity Theorems for Electrostatic and Magnetostatic Beam Monitors

The following is an excercise in Maxwell's equations in a region that is bounded by perfect metal walls and contains nothing but moving electric charges.

$$
\begin{equation*}
\nabla \times \dot{\mathbf{H}}=\mathbf{j}+\dot{\mathbf{D}} \quad \text { and } \quad \nabla \cdot \mathbf{H}=0 \tag{1.1}
\end{equation*}
$$

where j comes from the moving charges represented by charge density $\varrho(x, y, z, t)$ in the beam, and

$$
\begin{equation*}
\nabla \times \mathbf{E}=-\dot{\mathbf{B}} \quad \text { and } \quad \nabla \cdot \mathbf{D}=\varrho \tag{2.1}
\end{equation*}
$$

We integrate all equations over time, starting before the front-end wake fields begin, and ending after the end wake fields and RF are gone.

$$
\begin{align*}
& \int \dot{\mathbf{D}} d t=0 \quad \text { and } \quad \int \dot{\mathbf{B}} d t=0 \\
& \nabla \times \mathcal{H}=\mathbf{J} \quad \text { and } \quad \nabla \cdot \mathcal{H}=0  \tag{3,1}\\
& \nabla \times \mathcal{V}=0 \quad \text { and } \quad \nabla \cdot \mathcal{V}=R \tag{4,1}
\end{align*}
$$

The new symbols stand for integrals over time at every $x, y, z$.

## Electrostatic Pick-up.

The beam with $R$ produces $V_{1}(x, y, z)$, with $V_{I}=0$ on wall everywhere, including on the electrode. An electrode on $V_{20}$ produces $V_{2}(x, y, z)$, with $V_{2}=0$ on wall except on the electrode*.
Using (4.1) and (4.2) we notice that $V_{2}$ is the actual potential and $V_{1}$ is the integrated potential over time. With $\mathcal{H}=-\Delta V$,

$$
\begin{aligned}
U & =\int \nabla \cdot\left(V_{1} \nabla V_{2}-V_{2} \nabla V_{1}\right) d v \\
& =\int(V_{1} \underbrace{\nabla \cdot \nabla V_{2}}_{0}-V_{2} \underbrace{\nabla \cdot \nabla V_{1}}_{-R}) d v \\
& =\int V_{2} R d v .
\end{aligned}
$$

[^0]But it is also true that, for the charge $q_{e}$,

$$
U=\int\left(V_{1} \nabla V_{2}-V_{2} \nabla V_{1}\right) \cdot d \mathbf{a}=V_{20} \cdot Q \quad \text { with } \quad Q=\int q_{e} d t
$$

induced by electrons on electrode.
The total charge in the bunch, $Q_{B}$, is related to $\varrho(x, y, z, t)$ through

$$
\begin{gathered}
\int \varrho d t d a=Q_{B} / v \\
\int R v d a=\int \varrho d t v d a=\int Q_{B}^{\prime} d a \\
Q \cdot V_{20}=\frac{1}{v} \int V_{2} Q_{B}^{\prime} d v=\frac{1}{v} \int Q_{B} V_{2} d z
\end{gathered}
$$

where $Q_{B}^{\prime}$ is independent of $z$, and is the charge going through area da, divided by $d a$. The units for $\left[Q_{B}\right]=A \mathrm{sec}$, and $[Q]=A \mathrm{sec}^{2}$.

## Magnetostatic Pick-up.

We use (3.1) and (3.2). The beam with $\mathbf{J}$ produces $\mathrm{A}_{1}(x, y, z)$, and $\mathcal{H}_{1}(x, y, z)$. In the coil, the flux from $\mathbf{J}$ integrated over $t$ is

$$
\Phi_{2}=\int \mu_{0} \mathcal{H}_{1} \cdot d \mathbf{a}=\int \nabla \times \mathbf{A}_{1} \cdot d \mathbf{a}=\mu_{0} \oint \mathbf{A}_{1} \cdot d \mathbf{s}
$$

In addition, we use a coil with a current, $I_{2}$, that produces $\mathbf{A}_{2}(x, y, z), \mathcal{H}_{2}(x, y, z)$. We now use, equivalently to the electrostatic case,

$$
U=\int \nabla \cdot\left(\mathbf{A}_{1} \times \mathcal{H}_{2}-\mathbf{A}_{2} \times \mathcal{H}_{1}\right) d v
$$

where $\mathbf{A}_{1}$ is the integrated vector potential, and $\mathbf{A}_{2}$ is the actual potential associated with $I_{2}$.

$$
\nabla \cdot \mathbf{A}_{1} \times \mathcal{H}_{2}=\mathcal{H}_{2} \cdot \nabla \times \mathbf{A}_{1}-\mathbf{A}_{1} \cdot \nabla \times \mathcal{H}_{2}=\mathcal{H}_{2} \cdot \mathcal{H}_{1}-\mathbf{A}_{1} \cdot \mathbf{J}_{2}
$$

thus,

$$
U=\int\left(\mathbf{A}_{2} \cdot \mathbf{J}_{1}-\mathbf{A}_{1} \cdot \mathbf{J}_{2}\right) d v
$$

With

$$
\mathrm{J}_{1}=J_{1} \mathrm{e}_{\mathrm{z}} \quad \text { and } \quad \int J_{1} d a=\int j_{1} d a d t=Q_{B}
$$

we get

$$
\begin{gathered}
\int \mathbf{A}_{2} \cdot \mathbf{J}_{1} d v=Q_{B} \cdot \int A_{2 z} d z \\
\int \mathbf{A}_{1} \cdot \mathbf{J}_{2} d v=I_{2} \int \mathbf{A}_{1} \cdot d \mathbf{s}=I_{2} \Phi_{2} / \mu_{0}
\end{gathered}
$$

and get

$$
U=Q_{B} \cdot \int A_{2 z} d z-I_{2} \Phi_{2} / \mu_{0}=\int\left(\mathbf{A}_{1} \times \mathcal{H}_{2}-\mathbf{A}_{2} \times \mathcal{H}_{1}\right) \cdot d \mathbf{a}
$$

with the last integral taken over the "superconducting" wall.
In the vicinity of the wall we use $\mathcal{H}=-\nabla V$. Thus,

$$
U=\int\left(\mathbf{A}_{2} \times \nabla V_{1}-\mathbf{A}_{1} \times \nabla V_{2}\right) \cdot d \mathrm{a}
$$

In general, $\nabla \times\left(V_{1} \mathbf{A}_{2}\right)=V_{1} \nabla \times \mathbf{A}_{2}-\mathbf{A}_{2} \times \nabla V_{1}$, thus

$$
U=\int\left(V_{1} \nabla \times \mathbf{A}_{2}-V_{2} \nabla \times \mathbf{A}_{1}\right) \cdot d \mathbf{a}=\int\left(V_{1} \mathcal{H}_{2}-V_{2} \mathcal{H}_{1}\right) \cdot d \mathbf{a}=0
$$

The last integral vanishes because on the superconducting wall the component of $\mathcal{H}$ perpendicular to the wall (i.e. parallel to $d \mathrm{a}$ ) is zero. We therefore get

$$
\Phi_{2}=\mu_{0} Q_{B} \int A_{2 z} d z / I_{2}
$$

The units are $\left[\Phi_{2}\right]=\mu_{0} A \mathrm{~m} \mathrm{sec},\left[\mathbf{A}_{2}\right]=A,\left[\mu_{0} Q_{B} \int A_{2 z} d z / I_{2}\right]=\mu_{0} A \mathrm{~m}$ sec. It is important to notice that $\Phi_{2}$ is the integrated flux, and the flux is the integrated induced voltage.

## Integral for Excess Flux Calculation

$$
J=\int_{i_{1}}^{t_{2}} \underbrace{\frac{f(t)}{\left(t-t_{1}\right)^{n_{1}}\left(t_{3}-t\right)^{n_{3}}}}_{G} d t
$$

We have shown in an earlier note that for $n_{1}=n_{3}=1 / 2$,

$$
J=3 \int_{-1}^{1} \frac{f(t)}{\sqrt{4-x^{2}}} d x, \quad t=\frac{1}{4}\left(2\left(t_{2}+t_{1}\right)+\left(t_{2}-t_{1}\right) x\left(3-x^{2}\right)\right)
$$

For $n_{1}, n_{3} \neq 1 / 2$, the approach that gave the above equation becomes very complicated, especially if one wants to have generally valid and simple integration. For the general case, we use (arbitrarily, for simplicity)

$$
t_{2}=1 / 2\left(t_{3}+t_{1}\right) \quad \text { and } \quad J=J_{1}+J_{3}
$$

where

$$
J_{1}=\int_{i_{1}}^{t_{2}} G(t) d t \quad \text { and } \quad J_{3}=\int_{t_{2}}^{t_{3}} G(t) d t
$$

We solve for $J_{1}$ :

$$
A d x=\left(t-t_{1}\right)^{-n_{1}}, \quad A x=\frac{\left(t-t_{1}\right)^{m_{1}}}{m_{1}} \text { and } \quad A=\frac{\left(t_{2}-t_{1}\right)^{m_{1}}}{m_{1}} \text { when } x\left(t_{2}\right)=1
$$

with

$$
\begin{gathered}
m_{1}=1-n_{1} \quad p_{1}=\frac{1}{m_{1}} \\
t=t_{1}+\left(t_{2}-t_{1}\right) x^{p_{1}}, \quad t_{3}-t=\left(t_{2}-t_{1}\right)\left(2-x^{p_{1}}\right)
\end{gathered}
$$

Thus,

$$
\begin{aligned}
J_{1} & =\frac{\left(t_{2}-t_{1}\right)^{m_{1}}}{m 1\left(t_{2}-t_{1}\right)^{n_{3}}} \int_{0}^{1} \frac{f(t)}{\left(2-x^{p_{1}}\right)^{n_{3}}} d x \\
& =\frac{\left(t_{2}-t_{1}\right)^{1-n_{1}-n_{3}}}{1-n_{1}} \int_{0}^{1} \frac{f(t)}{\left(2-x^{p_{1}}\right)^{n_{3}}} d x .
\end{aligned}
$$

June, 1993. Note 0336csem.

Equivalently, solving for $J_{3}$ :

$$
\begin{gathered}
-B d x=\left(t_{3}-t\right)^{n_{3}} d t, \quad B x=\frac{\left(t_{3}-t\right)^{m_{3}}}{m_{3}} \quad \text { and } \quad B=\frac{\left(t_{2}-t_{1}\right)^{m_{3}}}{m_{3}} \\
t=t_{3}-\left(t_{2}-t_{1}\right) x^{p_{3}} \quad t-t_{1}=\left(t_{2}-t_{1}\right)\left(2-x^{p_{3}}\right)
\end{gathered}
$$

Thus,

$$
J_{3}=\frac{\left(t_{2}-t_{1}\right)^{1-n_{1}-n_{3}}}{1-n_{3}} \int_{0}^{1} \frac{f(t) d x}{\left(2-x^{p_{3}}\right)^{n_{1}}} .
$$

We may now now conclude that

$$
\begin{aligned}
J & =\int_{t_{1}}^{t_{3}} \frac{f(t)}{\left(t-t_{1}\right)^{n_{1}}\left(t_{3}-t\right)^{n_{3}}} d t \\
& =\left(t_{2}-t_{1}\right)^{1-n_{1}-n_{3}}\left\{\int_{0}^{1} \frac{f\left(t_{1}+\Delta t x^{\frac{1}{1-n_{1}}}\right) d x}{\left(2-x^{\frac{1}{1-n_{1}}}\right)^{n_{3}}\left(1-n_{1}\right)}+\int_{0}^{1} \frac{f\left(t_{3}-\Delta t x^{\frac{1}{1-n_{3}}}\right) d x}{\left(2-x^{\frac{1}{1-n_{3}}}\right)^{n_{1}}\left(1-n_{3}\right)}\right\} .
\end{aligned}
$$

We examine a specific case of excess flux in the pole in the geometry of Figure 1,


Figure 1.
where,

$$
\begin{aligned}
& \pi E_{12}=\frac{1}{n_{1}} \ln \left(\left(n_{1}+n_{2}\right) I_{1}\right) \\
& n_{1}=\frac{\alpha}{\pi} \quad \text { and } \quad n_{2}=\frac{\beta}{\pi}
\end{aligned}
$$

$$
\begin{aligned}
I_{1} & =\int_{0}^{1} \frac{d t}{t^{n_{1}}(1-t)^{1-\left(n_{1}+n_{2}\right)}} \\
& =\left(\frac{1}{2}\right)^{n_{2}}\left\{\frac{1}{1-n_{1}} \int_{0}^{1} \frac{d x}{\left(2-x^{\frac{1}{1-n_{1}}}\right)^{1-\left(n_{1}+n_{2}\right)}}+\frac{1}{n_{1}+n_{2}} \int_{0}^{1} \frac{d x}{\left(2-x^{\frac{1}{n_{1}+n_{2}}}\right)^{n_{1}}}\right\}
\end{aligned}
$$

We may conclude that $I_{2}=\left(I_{1}\right)_{n_{1} \Leftrightarrow n_{2}}$. Further, the expression $I_{1} \frac{\sin \alpha}{\alpha}=I_{2} \frac{\sin \beta}{\beta}$ should be true. This is a non-trivial assertion and comes from a derivation of the expression for $E_{12}$ in an earlier note.

## $H^{*}$ at End of CSEM Block



Figure 1.

$$
H^{*}(z)=\frac{I}{2 \pi i} \cdot \frac{1}{z-z_{0}} \longrightarrow-\frac{I^{\prime}}{2 \pi i} \ln \frac{z}{z+x_{3}} \cdot \frac{z-z_{2}}{z-z_{1}} .
$$

In the vicinity of $z=0$,

$$
H^{*}=-\frac{H_{c}}{2 \pi i} \ln \frac{z z_{2}}{z_{1} x_{3}}
$$

where

$$
\begin{aligned}
\frac{z_{1} x_{3}}{z_{2}} & =\frac{i y_{1} x_{3}}{i y_{1}-x_{3}}=\frac{y_{1} x_{3}}{y_{1}+i x_{3}}=\frac{x_{3}}{1+i x_{3} / y_{1}}=x_{4} e^{-i \alpha} \\
x_{4} & =\frac{x_{3}}{\sqrt{1+x_{3}^{2} / y_{1}^{2}}}=x_{3} \cos \alpha \text { and } z=r e^{i \varphi}
\end{aligned}
$$

Thus,

$$
H^{*}=-\frac{H_{c}}{2 \pi i} \ln \frac{r e^{i(\varphi+\alpha)}}{x_{4}}=-\frac{H_{c}}{2 \pi i}\left(\varphi+\alpha+i \ln \frac{x_{4}}{r}\right)
$$

Field "blows up" at $r=0$. Thus, for scaling purposes, at location where $\ln \frac{x_{4}}{r}=2 \pi$, $r=x_{4} e^{-2 \pi}=x_{4} \cdot 1.9 \times 10^{-3}$.
There is a strong local field perpendicular to the "current sheet side", which is not problematic when easy axis is parallel to the "current sheet side". It is easier to see with charge sheet, and it leads to the same answer.
Interesting damage results for block not magnetized in either a perpendicular or parallel direction to the sides.

[^1]

Figure 2.
No damage will result in corner $A$, but there is a potential of demagnetization at corner $B$, and at symmetrically located corners.

## Summary of Excess Flux Formulae for Gm3, Gm18 and Gm40

Unless otherwise noted, the following definitions hold for all geometries in this Note

$$
F=\pi \frac{\Delta A}{V_{0}}, \quad n=\frac{\alpha}{\pi}, \quad \text { and } \quad a=\frac{h_{2}}{h_{1}}
$$



Figure 1.

$$
\begin{align*}
F_{01} & =(1+b) \int_{0}^{1} \frac{1-x^{n}}{(1-x)(b+x)} d x  \tag{1}\\
F_{23} & =F_{01}+\ln b, \quad \text { with } \quad b=a^{1 / n} \tag{2}
\end{align*}
$$



Figure 2.
For Figures 2, 3, and $4(\mathrm{Gm} 18$ and Gm 40$) \alpha=\pi / 2$.

$$
\begin{gather*}
F_{01}=\ln \frac{1+1 / a^{2}}{4}+2 a \arctan \frac{1}{a}  \tag{3}\\
F_{23}=\ln \frac{1+a^{2}}{4}+2 a \arctan \frac{1}{a} \tag{4}
\end{gather*}
$$

April, 1989. Note 0332csem.

We summarize here the the sum of excess fluxes for (1) and (3). For (1), we get

$$
\begin{equation*}
\left(F_{01}(a)+F_{01}(1 / a)\right)=(1+b)^{2} \int_{0}^{1} \frac{\left(1-x^{n}\right)(1+x)}{(1-x)(b+x)(1+x b)} d x . \tag{1D}
\end{equation*}
$$

And for (3), we have

$$
\begin{align*}
\left(F_{01}(a)+F_{01}(1 / a)\right)= & 2 \ln \frac{a+1 / a}{4} \\
& +2 a \arctan (1 / a)+2(1 / a) \arctan a . \tag{3D}
\end{align*}
$$



Figure 3.

$$
\begin{gather*}
F_{34}=\ln \left(1+a^{2}\right)+2 a \arctan (1 / a),  \tag{5}\\
F_{12}=2 \ln \left(a+\sqrt{1+a^{2}}\right),  \tag{6}\\
F_{01}=F_{34}-F_{12} . \tag{7}
\end{gather*}
$$

Continued on following page.


Figure 4.

$$
\begin{gather*}
F_{-11}=\ln \left(1+a^{2}\right)+2 a \arctan (1 / a)  \tag{8}\\
F_{67}=\ln \frac{1+a^{2}}{2 a\left(a+\sqrt{1+a^{2}}\right)}+2 a \arctan \frac{1}{a}  \tag{9}\\
F_{56}=\ln \frac{\sqrt{1+a^{2}}\left(a+\sqrt{1+a^{2}}\right)}{2}+\frac{1}{a} \arctan a  \tag{10}\\
F_{567}=\ln \frac{\left(\sqrt{1+a^{2}}\right)^{3}}{4 a}+2 a \arctan \frac{1}{a}+\frac{1}{a} \arctan a  \tag{11}\\
F_{234}=\ln \frac{\sqrt{1+1 / a^{2}}}{4}+\frac{1}{a} \arctan a \tag{12}
\end{gather*}
$$

## Anti-Symmetric Undulator to Make Vertically Polarized or Circularly Polarized Light



Figure 1.

We have

$$
\begin{equation*}
-V=\frac{B_{0}}{k_{0}} \sin k_{0} z \cosh k_{0} y \tag{1}
\end{equation*}
$$

with fields anti-symmetric to the midplane $y=0$.

$$
\begin{equation*}
B_{z}=B_{0} \cosh k_{0} y \cos k_{0} z \quad \text { and } \quad B_{y}=B_{0} \sinh k_{0} y \sin k_{0} z \tag{2}
\end{equation*}
$$

in direction $\delta$ relative to the $z$-axis.

$$
\begin{gather*}
\lambda_{z}=\lambda_{0} / \cos \delta  \tag{3.1}\\
B_{\perp}=\underbrace{B_{0} \cosh k_{0} y \sin \delta}_{B_{2}} \cos k_{0} z  \tag{3.2}\\
B_{y}=\underbrace{B_{0} \sinh k_{0} y}_{B_{1}} \sin k_{0} z \tag{3.3}
\end{gather*}
$$

## Linearly Polarized Light.

Let $y=0, B_{y}=0$,

$$
\begin{equation*}
B_{2}=\sin \delta B_{0}=\sin \delta B_{0}\left(\frac{g}{\lambda_{u}} \cdot \frac{1}{\cos \delta}\right) \tag{4}
\end{equation*}
$$

By the above $B_{0}$ is indicated the achievable $B_{0}$ as a function of $g / \lambda_{0}$, where $\lambda_{0}$ is the period in the $z$-direction, and $g$ is the magnet gap.

October, 1984. Note 0208csem.

To have a better understanding, we look at the pure CSEM undulator:

$$
B_{0}=B_{3} e^{-\pi g / \lambda_{0}}
$$

with $B_{3}$ equal to the product of $2 B_{\tau}$, the segmentation factor and the finite height factor.

$$
\begin{equation*}
B_{2}=B_{3} \sin \delta e^{-a / \cos \delta}, \quad \text { with } \quad a=\pi g / \lambda_{u} \tag{5.1}
\end{equation*}
$$

We optimize with $\delta$ for given $g / \lambda_{u}$. With $B_{2}^{\prime}=0$,

$$
\cos \delta-a \tan ^{2} \delta=0
$$

Instead of solving for given $a$, we make a table of $a v s . \delta$.

$$
\begin{gather*}
a=\frac{\cos ^{3} \delta}{\sin ^{2} \delta}  \tag{5.2}\\
B_{2}=B_{3} \sin \delta e^{-\cot ^{2} \delta}=B_{3} y
\end{gather*}
$$

Compared to a "normal" undulator:

$$
\begin{gather*}
B_{2 n}=B_{3} e^{-a}=B_{3} e^{-\cos \delta / \tan ^{2} \delta}=B_{3} y  \tag{5.4}\\
\frac{B_{2}}{B_{2 n}}=\sin \delta e^{-\cot ^{2} \delta} e^{\cos \delta \cot ^{2} \delta}=\frac{y}{y_{o}} \tag{5.5}
\end{gather*}
$$

| $\delta$ | $g / \lambda_{u}$ | $y$ | $y / y_{0}$ |
| :---: | :---: | :---: | :---: |
| 30.00 | 0.83 | 0.02 | 0.33 |
| 35.00 | 0.53 | 0.07 | 0.40 |
| 40.00 | 0.35 | 0.16 | 0.46 |
| 45.00 | 0.23 | 0.26 | 0.53 |
| 50.00 | 0.14 | 0.38 | 0.60 |
| 55.00 | 0.09 | 0.50 | 0.66 |
| 60.00 | 0.05 | 0.62 | 0.73 |

Table 1.

## Circularly Polarized Light.

We set $y=y_{1}$, and $k_{0} y_{1}=\beta$. From (3.2) and (3.3), we see we need to satisfy $\cosh \beta \sin \delta=\sinh \beta$ for the helical undulator action, thus

$$
\begin{equation*}
\sin \delta=\tanh \beta, \tag{6.1}
\end{equation*}
$$

or

$$
\begin{equation*}
\cos \delta=\sqrt{1-\tanh ^{2} \beta}=1 / \cosh \beta \tag{6.2}
\end{equation*}
$$

or

$$
\begin{gather*}
\tan \delta=\sinh \beta  \tag{6.3}\\
B_{u}=B_{0} \sinh \beta=B_{0} \tan \delta=B_{0}\left(\frac{g / \lambda_{u}}{\cos \delta}\right) \tan \delta \tag{7}
\end{gather*}
$$

We assess the reasonableness and feasability of the above analysis.
Clearly,

$$
\begin{equation*}
\varepsilon=\frac{2 y_{1}}{g} \tag{8.1}
\end{equation*}
$$

is an important parameter.

$$
\beta=\frac{2 \pi y_{1}}{\lambda_{0}}=\frac{2 y_{1}}{g} \cdot \pi \cdot \frac{g}{\lambda_{u} \cos \delta},
$$

and for $p=g / \lambda_{u}$,

$$
\begin{gather*}
\beta=\varepsilon p \frac{\pi}{\cos \delta}=\varepsilon \frac{a}{\cos \delta}  \tag{8.2}\\
\varepsilon=\frac{\beta \cos \delta}{a} . \tag{8.3}
\end{gather*}
$$

The indicated procedure is as follows. Given $p=g / \lambda_{u}$ and $B_{0}\left(g / \lambda_{0}\right)=B_{0}(p / \cos \delta)$, we optimize $B_{u}$ with $\delta$ and get $\varepsilon$ from (8.3).
For a pure REC undulator,

$$
\begin{gathered}
B_{0}=B_{3} e^{-a / \cos \delta} \text { with } a=\pi g / \lambda_{u}=\pi p \\
B_{u}=B_{3} \tan \delta \cdot e^{-a / \cos \delta}
\end{gathered}
$$

With $B_{u}^{\prime}=0$,

$$
\frac{1}{\cos ^{2} \delta}-a \tan \delta \cdot \frac{\sin \delta}{\cos ^{2} \delta}=0
$$

$$
\begin{align*}
& a \frac{\sin ^{2} \delta}{\cos \delta}=a \frac{1-\cos ^{2} \delta}{\cos \delta}=1 \\
& \cos \delta=-\frac{1}{2 a}+\sqrt{\frac{1}{4 a^{2}}+1} \tag{9.1}
\end{align*}
$$

For $\cosh \beta=1 / \cos \delta$,

$$
\begin{gather*}
\beta=\ln \left(\sqrt{\cosh ^{2} \beta-1}+\cosh \beta\right) .  \tag{9.2}\\
\varepsilon=\beta \cos \frac{\delta}{a} \cdot  \tag{9.3}\\
\frac{B_{u}}{B_{3}}=\tan \delta \cdot e^{-a / \cos \delta} . \tag{9.4}
\end{gather*}
$$

For extreme "legal" $\varepsilon=1$ :

$$
\begin{gathered}
\beta \cos \frac{\delta}{a}=\sin ^{2} \delta \cdot \beta=\sin ^{2} \delta \cdot \ln \left(\sqrt{1+\tan ^{2} \delta}+\tan \delta\right)=\sin ^{2} \delta \cdot \ln \frac{1+\sin \delta}{\cos \delta} \\
\beta=60.27^{\circ}, \quad g / \lambda_{u}=.21, \quad \text { and } \quad B_{u} / B_{3}=.46
\end{gathered}
$$

## Hybrid Undulator with Superimposed Quadrupole Field

With the electron beam in the $z$-direction, and the midplane the $x z$-plane, the normal undulator fields can be described by

$$
B_{z}=i B_{y}=B_{1}^{*}=i F_{1}^{\prime}=i B_{1} \cos k(z+i y)
$$

For the complex potential $F_{1}=A_{1}+V_{1}$, with $A_{1}$ and $V_{1}$ the vector and scalar potentials respectively, it follows that

$$
F_{1}=\frac{B_{1}}{k} \sin k(z+i y) \quad \text { and } \quad V_{1}=\frac{B_{1}}{k} \cos k z \sinh k y
$$

The desired normal quadrupole fields are given by

$$
B_{0}^{*}=i F_{0}^{\prime}=i B_{0}^{\prime} z \quad F_{0}=\frac{1}{2} B_{0}^{\prime} z^{2}, \quad \text { and } \quad V_{0}=B_{0}^{\prime} x y
$$

For the scalar potential and the combined undulator and quadrupole fields, we therefore have

$$
V=V_{1}+V_{0}=\frac{B_{1}}{k} \cos k z \sinh k y+B_{0}^{\prime} x y
$$

Setting this equal to a constant gives the associated surface of a pole made with infinitely permeable material. With $y_{0}$ the half-gap of the pole at $z=x=0$,

$$
0=\cos k z \sinh k y-\sinh k y_{0}+\frac{B_{0}^{\prime}}{k B_{1}} k x k y
$$

With the following substitutions:

$$
a=\cos k z, \quad \mathcal{E}=\frac{B_{0}^{\prime}}{k B_{1}}, \quad u=k x, \quad \text { and } \quad v=k y
$$

we arrive at the following equation for the ideal 3D pole:

$$
a \sinh v-\sinh v_{0}+\mathcal{E} u v=G=0
$$

To understand what this means, we look at some derived quantities:

$$
y^{\prime}=\frac{d y}{d x}=-\frac{G_{x}^{\prime}}{G_{y}^{\prime}}=-\frac{G_{u}^{\prime}}{G_{v}^{\prime}}=-\frac{\mathcal{E} v}{a \cosh v+\mathcal{E} u}
$$

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is the slope of the pole in the $x y$-plane. For $x=z=0$ it is reduced to

$$
y_{0}^{\prime}=-\frac{B_{0}^{\prime} y_{0} / B_{1}}{\cosh v_{0}}=-\frac{\mathcal{E} v_{0}}{\cosh v_{0}}
$$

Looking at the slope just above the axis of the system, i.e. for $x=0$,

$$
a \sinh v_{1}=\sinh v_{0}
$$

and

$$
y^{\prime}=-\frac{\mathcal{E} v_{1}}{a \cosh v_{1}}
$$

where the subscript 1 refers to the case of $x=0$ and $z$ equal to anything. For $z=0$ this reduces to

$$
y^{\prime}=-\frac{B_{0}^{\prime} y_{0} / B_{1}}{\cosh v_{0}}=-\frac{\mathcal{E} v_{0}}{\cosh v_{0}} .
$$

Eliminating $a=\cos k z$ gives

$$
y^{\prime}=-\mathcal{E} v_{1} \frac{\tanh v_{1}}{\sinh v_{0}}
$$

For the curvature of the pole in the $x y$-plane we need

$$
\begin{aligned}
y^{\prime \prime} & =\frac{\partial y^{\prime}}{\partial x}+\frac{d y}{d x} \cdot \frac{\partial y^{\prime}}{\partial y} \\
& =k\left(\frac{\partial y^{\prime}}{\partial u}+y^{\prime} \frac{\partial y^{\prime}}{\partial v}\right) \\
& =k\left(\frac{\mathcal{E}^{2} v}{(a \cosh v+\mathcal{E} u)^{2}}+\frac{\mathcal{E}^{2} v}{(a \cosh v+\mathcal{E} u)} \cdot \frac{a \cosh v+\mathcal{E} u-a v \sinh v}{(a \cosh v+\mathcal{E} u)^{2}}\right) .
\end{aligned}
$$

For $u=0$, this reduces to

$$
\begin{aligned}
y^{\prime \prime} & =\frac{k \mathcal{E}^{2} v_{1} a}{a^{3} \cosh ^{3} v_{1}}\left(2 \cosh v_{1}-v_{1} \sinh v_{1}\right) \\
& =\mathcal{E}^{2} k \frac{v_{1} \tanh ^{2} v_{1}}{\sinh ^{2} v_{0}}\left(2-v_{1} \tanh v_{1}\right)
\end{aligned}
$$

We let $\left|y^{\prime}\right| \ll 1$, and therefore $\sqrt{1+\left(y^{\prime}\right)^{2}}{ }^{3} \approx 1$. With radius of curvature $R$, we get

$$
\frac{1}{k \mathcal{E}^{2}} \cdot \frac{\sinh ^{2} v_{0}}{v_{1} \tanh ^{2} v_{1}\left(2-v_{1} \tanh v_{1}\right)}=R=\frac{y_{0}}{\mathcal{E}^{2}} \cdot \frac{\sinh ^{2} v_{0}}{v_{0} v_{1} \tanh ^{2} v_{1}\left(2-v_{1} \tanh v_{1}\right)},
$$

For $v_{1}=v_{0}$, and $y_{0}^{\prime}=-\mathcal{E} v_{0} / \cosh v_{0}$ :

$$
R_{0}=\frac{y_{0}}{\mathcal{E}^{2}} \cdot \frac{\cosh ^{2} v_{0}}{v_{0}^{2}\left(2-v_{0} \tanh v_{0}\right)}=\frac{y_{0}}{\left(y_{0}^{\prime}\right)^{2}} \cdot \frac{1}{v_{0}^{2}\left(2-v_{0} \tanh v_{0}\right)}
$$

We make the following assignments:

$$
\frac{1}{k \mathcal{E}^{2}}=\frac{k B_{1}^{2}}{\left(B_{0}^{\prime}\right)^{2}}=y_{0} \frac{B_{1}^{2}}{y_{0}^{2}\left(B_{0}^{\prime}\right)^{2}} v_{0} \quad \text { and } \quad b=\frac{B_{1}^{2}}{y_{0}^{2}\left(B_{0}^{\prime}\right)^{2}}
$$

and re-write

$$
R=y_{0} b \cdot \frac{v_{0}}{v_{1}} \cdot \frac{\sinh ^{2} v_{0}}{\tanh ^{2} v_{1}\left(2-v_{1} \tanh v_{1}\right)}
$$

where for $2-v \tanh v=0, v=\frac{2 \pi y}{\lambda}=2.0653$.

## Excess Flux into Gm13



Figure 1.

For the above graph,

$$
\frac{\alpha}{\pi}=n_{1}, \quad \frac{\beta}{\pi}=n_{2}, \quad \text { and } \quad n_{1}+n_{2}=n_{3}
$$

The conformal map is described by

$$
\dot{z}=a \frac{t^{n_{3}}}{(t-1)^{1+n_{2}}}, \quad a \in \Re
$$

From

$$
\pi \dot{F}=\frac{1}{t-1} \quad \text { follows } \quad \pi F=\ln (t-1)
$$

We describe the flux into surface 2 of Figure 1 as

$$
\Phi_{2}=-\frac{1}{\pi} \ln \left(1-t_{1}\right)=\int_{s_{0}}^{s_{1}} \frac{d s}{\beta s}+\Delta A_{2}
$$

Thus

$$
\begin{aligned}
\beta \Delta A_{2} & =-n_{2} \ln \left(1-t_{1}\right)-\ln \left(1+\frac{s_{1}-s_{0}}{s_{0}}\right) \\
& =\ln \frac{\left(1-t_{1}\right)^{-n_{2}}}{\left(1+\frac{a}{s_{0}} \int_{0}^{t_{1}} \frac{t^{n_{3}} d t}{(1-t)^{1+n_{2}}}\right)}
\end{aligned}
$$

and by l'Hôpital's Rule,

$$
\begin{aligned}
& =\ln \frac{n_{2}\left(1-t_{1}\right)^{-n_{2}-1}}{\left(\frac{a}{s_{0}} t_{1}^{n_{3}}\left(1-t_{1}\right)^{-n_{2}-1}\right)} \\
& =\ln \frac{n_{2} s_{0}}{a},
\end{aligned}
$$

and therefore

$$
\begin{equation*}
\Delta A_{2}=\frac{1}{\beta} \ln \left(\frac{h}{\pi} \frac{\beta}{a \sin \beta}\right) . \tag{1}
\end{equation*}
$$

We need to calculate $h / a$, and we begin with

$$
\begin{equation*}
y_{1}=h+a \sin \beta \int_{0}^{t_{1}} \frac{t^{n_{3}} d t}{(1-t)^{n_{2}+1}} \tag{2}
\end{equation*}
$$

We now calculate $y_{1}$ by going around the singularity at $t=1$ on circle with $\varrho=\varrho_{1}=$ $1-t_{1}$, that is

$$
t=1+\varrho_{1} e^{i \varphi} \quad \text { and } \quad d t=i \varrho_{1} e^{i \varphi} d \varphi
$$

and thus

$$
\begin{aligned}
y_{1} & =\Im i \varrho_{1} a \int_{0}^{\pi} \frac{\left(1+\varrho_{1} e^{i \varphi}\right)^{n_{3}}}{\varrho_{1}^{1+n_{2}} e^{i \varphi\left(1+n_{2}\right)}} e^{i \varphi} d \varphi \\
& =\Im \frac{i a}{\varrho_{1}^{n_{2}}} \int_{0}^{\pi}\left(1+\varrho_{1} e^{i \varphi}\right)^{n_{3}} e^{-i n_{2} \varphi} d \varphi
\end{aligned}
$$

For $\varrho_{1}$ sufficiently small,

$$
y_{1}=\Im \frac{i a}{\varrho_{1}^{n_{2}}} \frac{e^{-i n_{2} \pi}-1}{-i n_{2}}=\frac{a \sin \beta}{n_{2} \varrho_{1}^{n_{2}}}
$$

Re-writing (2) with $t=1-\varrho$

$$
y_{1}=h+a \sin \beta \int_{\varrho_{1}}^{1} \frac{(1-\varrho)^{n_{3}}}{\varrho^{n_{2}+1}} d \varrho=\frac{a \sin \beta}{n_{2} \varrho_{1}^{n_{2}}} .
$$

Thus

$$
\begin{aligned}
\frac{h}{a \sin \beta} & =\frac{1}{n_{2} \varrho_{1}^{n_{2}}}-\int_{\varrho_{1}}^{1} \frac{(1-\varrho)^{n_{3}}}{\varrho^{n_{2}+1}} d \varrho \\
& =\frac{1}{n_{2} \varrho_{1}^{n_{2}}}+\left.\frac{(1-\varrho)^{n_{3}}}{n_{2} \varrho^{n_{2}}}\right|_{\varrho_{1}} ^{1}+\frac{n_{3}}{n_{2}} \int_{\varrho_{1}}^{1} \frac{(1-\varrho)^{n_{3}-1}}{\varrho^{n_{2}}} d \varrho
\end{aligned}
$$

and

$$
\frac{h n_{2}}{a \sin \beta}=\frac{h \beta}{\pi a \sin \beta}=n_{3} I_{2}
$$

where

$$
I_{2}=\int_{0}^{1} \frac{d t}{t^{n_{2}}(1-t)^{1-n_{3}}} \quad \text { and } \quad I_{1}=\int_{0}^{1} \frac{d t}{t^{n_{1}}(1-t)^{1-n_{3}}}
$$

Therefore we may now summarize from (1) that

$$
\Delta A_{2}=\frac{1}{\beta} \ln \left(n_{3} I_{2}\right)
$$

and equivalently

$$
\Delta A_{1}=\frac{1}{\alpha} \ln \left(n_{3} I_{1}\right) .
$$

Further, since

$$
\begin{gathered}
\frac{h}{\pi a}=n_{3} I_{2} \frac{\sin \beta}{\beta}=n_{3} I_{1} \frac{\sin \alpha}{\alpha} \\
n_{3} I_{1}=n_{3} I_{2} \frac{\sin \beta / \beta}{\sin \alpha / \alpha} .
\end{gathered}
$$

A different way to look at what was done earlier:


Figure 2.

Since, clearly, $h=\Im \int_{1+}^{0} \dot{z} d t$, one may take a path from any point on the real $t$-axis to the right of $t=1$ to $t=0$.
In this note the path followed a $\varrho_{1} \Rightarrow 0$ half-circle around $t=1$, and then on axis to $t=0$.

## Flux Distribution Symmetry Theorem

Even though this case is the same as the electrostatic case, it is formulated for magnetic fields because of the application to CSEM in hybrid systems.

## Theorem.

There are $N$ bodies with $\mu=\infty$. The matrix $M$, which describes the relationship between potentials $V_{n}$ on fluxes $F_{n}$ leaving body $n$, is symmetrical, i.e.: $F=M V$, $M^{t}=M$. In this notation, the subscript 0 indicates the reference body on potential $V=0$. Thus, the theorem states

$$
M_{n m}=M_{m n}
$$

## Analysis.

Stored energy in the field is unique: it depends only on the initial state (we assume $V_{n}=0$ ) and the end state. By going from the initial to the end state by bringing bodies in any sequence from $V_{n}=0$ to the final $V_{n}$, and doing so by moving magnetic charges from infinity, we get

$$
\mathcal{E}=\int V^{t} d F=\int V^{t} M d V
$$

' $\mathcal{E}$ must be independent of sequence in which this is done: the right hand side must be a complete differential leading to $M_{n m}=M_{m n}$. We show this explicitly for $V_{1}, V_{2}$ and all other $V_{n}=0$ :

$$
\begin{aligned}
\mathcal{E} & =\int\left(\begin{array}{ll}
V_{1} & V_{2}
\end{array}\right)\binom{M_{11} d V_{1}-M_{12} d V_{2}}{-M_{21} d V_{1}+M_{22} d V_{2}} \\
& =\int M_{11} V_{1} d V_{1}+\int M_{22} V_{2} d V_{2}-\underbrace{\int\left(M_{12} V_{1} d V_{2}+M_{21} V_{2} d V_{1}\right)}_{G}
\end{aligned}
$$

We simplify $G$ by making the following substitutions:

$$
\begin{gather*}
M_{12}=S+D, \quad M_{21}=S-D, \quad \text { and } \quad S=\frac{1}{2}\left(M_{12}+M_{21}\right), \quad D=\frac{1}{2}\left(M_{12}-M_{21}\right) \\
G=S \int \underbrace{\left(V_{1} d V_{2}+V_{2} d V_{1}\right)}_{(a)}+D \underbrace{\int\left(V_{1} d V_{2}-V_{2} d V_{1}\right)}_{(b)} \tag{b}
\end{gather*}
$$

where (a) is $d\left(\begin{array}{ll}V_{1} & V_{2}\end{array}\right)$ and is therefore independent of the sequence, and (b) would
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be dependent on our sequence. Since $G$ must be independent of the sequence, it follows that $D=0$.
In a CSEM circuit, $\mathbf{F}$ equals the vector of charges deposited by the CSEM on the surfaces (with all $V=0$ ). Therefore, a hybrid system can be represented by magnetic capacitors and sources that deposit DC charges. If one finds this more convenient, one may also do this analysis with capacitors and AC current sources, or with resistors and DC currents.
The theorems known for these circuits apply, such as Kirchhoff's reciprocity theorem, i.e. the nodal equations, etc. One can also use $2 \times 2$ matrix methods for systems like ladder networks, and apply them directly to hybrid wigglers. One can use concepts like characteristic impedance of networks and quadrupole thecry, i.e. all the tools that have been developed for analog network analysis.

## Stored Energy in CSEM

We define the energy density as

$$
\mathcal{E}=\int \mathbf{H} \cdot d \mathbf{B}
$$

We can look at the easy-axis direction and the direction perpendicular to the easy-axis separately. The lower integral limit is arbitrary, but must be fixed.
With $B_{\|}=B_{r}+\mu_{0} \mu_{\|} H_{\|}$and $B_{\perp}=\mu_{0} \mu_{\perp} H_{\perp}$ :

$$
\mathcal{E}_{\|}=\int_{H_{\|}=0}^{H_{\|}} H_{\|} d B_{\|}=\int_{H_{\|}=0}^{H_{\|}} H_{\|} \frac{d B_{\|}}{d H_{\|}} d H_{\|}=\mu_{0} \mu_{\|} \int_{H_{\|}=0}^{H_{\|}} H_{\|} d H_{\| \mid}
$$

Thus,

$$
\mathcal{E}_{\|}=\frac{1}{2} \mu_{0} \mu_{\|} H_{\|}^{2}
$$

and similarly,

$$
\varepsilon_{\perp}=\frac{1}{2} \mu_{0} \mu_{\perp} H_{\perp}^{2}
$$

with

$$
\mathcal{E}_{\text {tot }}=\varepsilon_{\|}+\mathcal{E}_{\perp \cdot}
$$

This obviously gives $H_{\|}$a unique role.

## Earnshaw's Theorem for Non-Permeable Material

Problem: There is an assembly of "frozen" magnetic charges $\varrho\left(r^{\prime}\right)^{\dagger}$ in an external magnetic field (produced by solenoid or other REC assembly) without any permeable material in the system.
Question: What is the restoring force for small displacements?
Analysis: The force components in the ( $x_{1}, x_{2}, x_{3}$ ) coordinate system are

$$
F_{1}=-\int \varrho V_{1}^{\prime} d v, \quad F_{2}=-\int \varrho V_{2}^{\prime} d v, \quad \text { and } \quad F_{3}=-\int \varrho V_{3}^{\prime} d v .
$$

We displace the system by $\Delta x_{1}, \Delta x_{2}, \Delta x_{3}$, which is the same as displacing the external fields by $-\Delta x_{1},-\Delta x_{2},-\Delta x_{3}$ without changing $\varrho$, and get

$$
\begin{aligned}
\Delta F_{n} & =\int \sum \Delta x_{m} \frac{\partial}{\partial x_{m}} \varrho \frac{\partial V}{\partial x_{n}} d v \\
& =\sum M_{n m} \Delta x_{m} \quad \text { with } \quad M_{n m}=\int \varrho \frac{\partial}{\partial x_{n}} \frac{\partial}{\partial x_{m}} V d v,
\end{aligned}
$$

$$
\Delta \mathbf{F}=M \cdot \Delta \mathrm{x} .
$$

In general, $M$ will not be a diagonal matrix. We assume that it can be made diagonal (by going to a new coordinate system) with matrix $C$, where

$$
\begin{gathered}
C \Delta \mathbf{F}=\Delta \mathrm{F}_{d}=C M C^{-1} \cdot C \Delta \mathrm{x}=C M C^{-1} \cdot \Delta \mathrm{x}_{d}, \\
C M C^{-1}=M_{d}=\int \varrho\left(\begin{array}{ccc}
V_{x x}^{\prime \prime} & 0 & 0 \\
0 & V_{y y}^{\prime \prime} & 0 \\
0 & 0 & V_{z z}^{\prime \prime}
\end{array}\right) d v,
\end{gathered}
$$

where $x, y, z$ are the new coordinates. From this, it follows that

$$
\frac{\Delta F_{x}}{\Delta x}+\frac{\Delta F_{y}}{\Delta y}+\frac{\Delta F_{z}}{\Delta z}=\int \varrho \underbrace{\left(V_{x x}^{\prime \prime}+V_{y y}^{\prime \prime}+V_{z z}^{\prime \prime}\right)}_{=\nabla^{2} V=0} d v=0 .
$$

Since a stable system requires a negative restoring force in each of the three coordinate directions, any such system will be unstable.

[^2]In applications, it will often be clear a priori in which coordinate system the matrix $M$ will be diagonal, and the above equation can then be used directly in its final form. This means that for systems with cylindrical symmetry about the $z$-axis, that because

$$
\frac{\Delta F_{y}}{\Delta y}=\frac{\Delta F_{x}}{\Delta x} \Rightarrow \frac{\Delta F_{r}}{\Delta r} \quad \text { and } \quad 2 \frac{\Delta F_{r}}{\Delta r}+\frac{\Delta F_{z}}{\Delta z}=0
$$

only one of the stiffnesses needs to be calculated from basics.
It is interesting to note that Earnshaw's Theorem is often stated in an overly broad fashion. For instance, stable support is possible if one allows forces not derived from a potential satisfying $\nabla^{2} V=0$. The forces between contacting solid materials, for example, such as mechanical bearings, are due to the quantum nature of solids, and hence do not obey Earnshaw's Theorem. It is also clear without any mathematics that a permanent magnet is stably supported in a superconducting bowl. This is similarly true for an extreme diamagnetic bowl.

## Harmonics Produced by Rectangular REC Block .



Figure 1.

In this document we refer to the above diagram and (4b) and (16b) of the Nuclear Instruments and Methods article ${ }^{\dagger}$

$$
\begin{gathered}
\underline{B}^{*}\left(z_{0}\right)=\sum_{n=1} b_{n} z_{0}^{n-1} \\
b_{n}=\frac{\underline{B}_{r}}{2 \pi i} \oint \frac{d x}{z^{n}}
\end{gathered}
$$

For $n=1$,

$$
\begin{aligned}
b_{1} & =\frac{\underline{B}_{r}}{2 \pi i} \ln \frac{z_{2}}{z_{1}} \cdot \frac{z_{4}}{z_{3}}=\frac{\underline{B}_{r}}{2 \pi i} \ln \frac{z_{2} / z_{1}}{z_{3} / z_{4}} \\
& =\frac{\underline{B}_{r}}{\pi} \Im \ln \frac{z_{2}}{z_{1}} .
\end{aligned}
$$

For $n \geq 2$,

$$
\begin{aligned}
b_{n} & =-\frac{\underline{B}_{r}}{2 \pi i} \cdot \frac{1}{n-1}\left(\frac{1}{z_{2}^{n-1}}-\frac{1}{z_{1}^{n-1}}+\frac{1}{z_{4}^{n-1}}-\frac{1}{z_{3}^{n-1}}\right) \\
& =-\frac{\underline{B}_{r}}{\pi} \cdot \frac{1}{n-1} \Im\left(\frac{1}{z_{2}^{n-1}}-\frac{1}{z_{1}^{n-1}}\right) \cdot
\end{aligned}
$$

$\dagger$ Halbach, H., Nuclear Instruments and Methods 169, 1 (1980).

## A Possible REC Undulator for SSRL

## I. Reason for REC.

It may be possibe to use some very specific ferrite. All other materials, like the Alnicos, are not only significantly inferior in performance, but would probably also have to be magnetized in final assembly which may be difficult to do.

A potential future advantage is that the permanent magnet undulator can be scaled down in physical size without difficulty. One can therefore envision the following scheme: design a REC undulator for very small gap and $\lambda$ and have it inside vacuum. Move the two arrays apart to have the large gap necessary during beam formation. When beam is established, move the 2 arrays together to form the design gap that the beam allows. Clearly, this would be nearly impossible with either a conventional, or even a superconducting, undulator.

## II. Use of Nomogram and Notation.



Figure 1.

$$
\begin{gathered}
B_{y}+i B_{z}=B_{0} \cos \frac{2 \pi(z+i y)}{\lambda}, \\
K=B_{0} \lambda \\
k=\frac{B_{r}}{K} \underbrace{\frac{\sin \pi / M^{\prime}}{\pi / M^{\prime}}}_{E_{1}} \underbrace{\left(1-e^{-2 \pi L / \lambda}\right)}_{E_{2}}
\end{gathered}
$$

where $B_{r}$ is the remanent field of REC, $M^{\prime}$ is the number of pieces with fixed easy-axis

April, 1979. Note 0038 csem .
per period.

$$
2 k \lambda e^{-\pi g / \lambda}=1 .
$$



Figure 2.

Because $E_{1}, E_{2}$ are close to 1 , and one usually chooses $K \approx 1 \mathrm{~T}(\mathrm{~cm}), k\left(\mathrm{~cm}^{-1}\right)$ is numerically close to $B_{r}$.
In this document, all lengths will be in cm , and $B$ 's in Tesla.

## III. Desigñ Considerations.

- End Section. Normalize center to $z=0$ and that piece has easy-axis parallel to the $y$ axis. The last pieces at both ends must have the same easy-axis as piece at $z=0$, but should have only half of normal length in the $z$-direction. One may want to use coils to fine-tune the end sections, but it would not be surprising if this were unnecessary.

In order to reduce the effects from finite length in $x$-direction (or to get away with shorter length in that direction) and to avoid 3D fringe effects at the ends in $z$ direction by cutting end fields off rapidly, one should back-up REC with a soft steel plate with reasonable overhang in $z$ and $x$ directions. This will not affect the amplitude of the $\cos (2 \pi(z+i y) / \lambda)$ term, but will introduce a very weak, unnoticeable in the midplane, third harmonic (for $M^{\prime}=4$ ).

- Length of $R E C$ in $x$-direction. The present estimate is that it should be approximately the sum of the width of the beam and $2 g$. The 3D effects discussed in the previous section are easily analyzed computationally and should be done before ordering materials!
- Choice of $M^{\prime}$ and $L$. It is recommended, at least for the first undulator, to use $M^{\prime}=4\left(\right.$ giving $\left.E_{1}=.9\right)$ and $L=\lambda / 2$ (giving $E_{2}=.96$ ) or $L=\lambda / 4$ (giving $E_{2}=.79$ ). With these choices, the undulator can be assembled from identical REC pieces with square cross-section and the easy-axis parallel to a side. The exception would be with the end pieces which could be obtained by cutting or grinding the normal pieces.


## IV. Specific Calculations.

For a realistic undulator with $g=2.8 \mathrm{~cm}, B_{r}=.95 \mathrm{~T}, K=1 \mathrm{Tcm}$ and $M^{\prime}=4$ :

| $L(\mathrm{~cm})$ | $k\left(\mathrm{~cm}^{-1}\right)$ | $\lambda(\mathrm{cm})$ |
| :---: | :---: | :---: |
| $\lambda / 4=1.18$ | .68 | 4.73 |
| $\lambda / 2=2.22$ | .82 | 4.44 |

Table 1.

Since $\lambda / 4$ uses only half the REC of the $\lambda / 2$ case and $\lambda$ is only less than $10 \%$ larger, this is the preferable design. The volume for $\lambda / 4$ is $V=3540 \mathrm{~cm}^{3}$, and for $\lambda / 2$ is $V=6660 \mathrm{~cm}^{3}$.
The REC price would probably be approximately $\$ 1-2 / \mathrm{cm}^{-3}$.

## A Simple Derivation of the Lorentz Transformation Without Talking About Light

Postulate: Physics is independent of location, time and uniform motion of the system in which the experiment is performed.

We look at two systems that move with velocity, $v$, relative to each other. We establish clocks and space ( $x$ ) markers in each system.

$\qquad$


Figure 1.

We locate the origins and synchronize the clocks so that $x_{1}=0, t_{1}=0, x_{1.5}=0$, and $t_{1.5}=0$. Also notice that the " 1.5 " system has $x$ increasing in the opposite direction from the " 1 " system.

We want to know ( $x_{1.5}, t_{1.5}$ ) as a function of $\left(x_{1}, t_{1}\right)$.
We know that $\Delta x_{1.5} / \Delta x_{1}, \Delta x_{1.5} / \Delta t_{1}, \Delta t_{1.5} / \Delta x_{1}$, and $\Delta t_{1.5} / \Delta t_{1}$ can not depend on $x_{1}, t_{1}$ because of our postulate. This means that the relationship between the two systems is linear, and can be expressed as a 2 by 2 matrix.

$$
\left.\begin{array}{l}
x_{1.5}=a_{11} x_{1}+a_{12} t_{1} \\
t_{1.5}=a_{21} x_{1}+a_{22} t_{1}
\end{array}\right\} \Rightarrow r_{1.5}=\binom{x_{1.5}}{t_{1.5}}=A \cdot r_{1}
$$

The velocity of a particle in system "1.5" (e.g. at $x_{1.5}=0$ ) as seen from system " 1 " is $v=-a_{12} / a_{11}$. Thus

$$
a_{12}=-a_{11} v
$$

with $a_{11} \neq 0$ always true.
The choice of the relative sign of $x$ in the two systems means that the observer in each system sees the other system move in the positive $x$-direction with velocity $v$.

July 30, 1992. Note 0287misc.

Therefore,

$$
r_{1}=A \cdot r_{1.5}=A \cdot A \cdot r_{1},
$$

and also,

$$
A^{2}=\left(\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right)
$$

must be satisfied. By multiplication and substitution,

$$
A^{2}=\left(\begin{array}{cc}
a_{11}\left(a_{11}-a_{21}\right) v & -a_{11} v\left(a_{11}+a_{22}\right) \\
a_{21}\left(a_{11}+a_{22}\right) & a_{22}^{2}-a_{21} a_{11} v
\end{array}\right) .
$$

Therefore,

$$
a_{11}=-a_{22}
$$

and

$$
a_{21}=\left(1-a_{22}^{2}\right) /\left(a_{22} v\right) .
$$

By further substitution,

$$
A=\left(\begin{array}{cc}
-a_{22} & a_{22} v \\
\left(1 / a_{22}-a_{22}\right) / v & a_{22}
\end{array}\right) .
$$

We introduce $x_{2}=-x_{1.5}, t_{2}=t_{1.5}$ and

$$
r_{2}=B \cdot r_{1}, \quad B=\left(\begin{array}{cc}
a_{22} & -a_{22} v \\
\left(1 / a_{22}-a_{22}\right) / v & a_{22}
\end{array}\right) .
$$

We further define

$$
\begin{gathered}
\gamma=a_{22} . \\
g=1 / a_{22}^{2}-1=1 / \gamma^{2}-1,
\end{gathered}
$$

and therefore

$$
B=\gamma\left(\begin{array}{cc}
1 & -v \\
g / v & 1
\end{array}\right) .
$$

It is important to notice that the diagonal elements are identical.


Figure 2.

We define

$$
\begin{aligned}
r_{2} & =B_{1 \rightarrow 2} \cdot r_{1} \\
r_{3} & =B_{2 \rightarrow 3} \cdot r_{2} \\
& =B_{2 \rightarrow 3} \cdot B_{1 \rightarrow 2} \cdot r_{1}
\end{aligned}
$$

and thus

$$
B_{1 \rightarrow 3}=\gamma_{1 \rightarrow 2} \gamma_{2 \rightarrow 3}\left(\begin{array}{cc}
1 & -v_{2 \rightarrow 3} \\
g_{2 \rightarrow 3} / v_{2 \rightarrow 3} & 1
\end{array}\right)\left(\begin{array}{cc}
1 & -v_{1 \rightarrow 2} \\
g_{1 \rightarrow 2} / v_{1 \rightarrow 2} & 1
\end{array}\right)
$$

and further,

$$
B_{1 \rightarrow 3}=\gamma_{1 \rightarrow 2} \gamma_{2 \rightarrow 3}\left(\begin{array}{cc}
1-v_{2 \rightarrow 3} g_{1 \rightarrow 2} / v_{1 \rightarrow 2} & -\left(v_{1 \rightarrow 2}+v_{2 \rightarrow 3}\right) \\
g_{2 \rightarrow 3} / v_{2 \rightarrow 3}+g_{1 \rightarrow 2} / v_{1 \rightarrow 2} & 1-v_{1 \rightarrow 2} g_{2 \rightarrow 3} / v_{2 \rightarrow 3}
\end{array}\right)
$$

By the identical diagonal elements we have:

$$
\frac{v_{2 \rightarrow 3} g_{1 \rightarrow 2}}{v_{1 \rightarrow 2}}=\frac{v_{1 \rightarrow 2} g_{2 \rightarrow 3}}{v_{2 \rightarrow 3}} \Longrightarrow \frac{g_{1 \rightarrow 2}}{v_{1 \rightarrow 2}^{2}}=\frac{g_{2 \rightarrow 3}}{v_{2 \rightarrow 3}^{2}}
$$

Thus, we may generalize our equation and we have

$$
g / v^{2}=\frac{\left(1 / \gamma^{2}-1\right)}{v^{2}}=k=\text { constant of nature }
$$

Here

$$
\gamma=\frac{1}{\sqrt{1+k v^{2}}}
$$

We have to verify that other relationships are also satisfied (e.g. relation between elements [11] and [12], etc.). We have shown that if the postulate is true, the relationship between $x$ and $t$ of the systems moving with velocity $v$ relative to each other must be

$$
\binom{x_{2}}{t_{2}}=\gamma\left(\begin{array}{cc}
1 & -v \\
k v & 1
\end{array}\right)\binom{x_{1}}{t_{1}}, \quad \gamma=\frac{1}{\sqrt{1+k v^{2}}}
$$

We have not shown that $k \neq 0$, but the value of $k$ can be obtained from "any" experiment, e.g. lifetime of meson, etc., and experiments do not have to use light.

Figure 3.

$$
\begin{gathered}
\binom{x_{2}}{t_{2}}=\gamma\left(\begin{array}{cc}
1 & -v \\
-v / c^{2} & 1
\end{array}\right)\binom{x_{1}}{t_{1}},\binom{x_{1}}{t_{1}}=\gamma\left(\begin{array}{cc}
1 & -v \\
-v / c^{2} & 1
\end{array}\right)\binom{x_{2}}{t_{2}} \\
\gamma=\frac{1}{\sqrt{1-v^{2} / c^{2}}}
\end{gathered}
$$

1) The lifetime of particle at rest at $x_{1}=0$ in system "1" is $\tau_{1}$. What is it in system " 2 "?

$$
x_{1}=0, \quad t_{1}=\tau_{1} \Longrightarrow t_{2}=\tau_{2}=\gamma \tau_{1}
$$

2) The distance $x_{1}$ covered by observer at rest at $x_{2}=0$ in time $t_{2}$ :

$$
x_{2}=0, \quad x_{1}=\gamma v t_{2}, \quad x_{1} / t_{2}=\gamma v
$$

Note that

$$
v \gamma=c \text { for } \mathrm{v} / \mathrm{c}=1 / \sqrt{2}
$$

## Dimensional Analysis of the Trajectory of Non-Relativistic

 Charged Particles in Stationary Electric and Magnetic Fields (MKS units, with and without space charge)
## Motivation.

To explain the structure of trajectory equations to engineers working on cyclotronmass spectrometer.
We use linear scaling length $D_{0}$, and represent $\mathbf{B}$ and $\mathbf{E}$ fields by the scaling quantities $B_{0}$ and $E_{0}=V_{0} D_{0}$ times the appropriate dimensionless functions of $x / D_{0}, y / D_{0}$, and $z / D_{0}$. We must be able to represent the trajectory $x(t)(t=0$ at start of trajectory) as the product of $D_{0}$ and a function of dimensionless products $P_{n}$. The list of parameters to form $P$ 's has, in addition to $D_{0}, B_{0}, V_{0}\left(E_{0}\right), t$, the quantity $m / e$ due to the equation of motion. Thus, the complete list consists of $m / e, D_{0}, B_{0}, V_{0}\left(E_{0}\right), t$.
We construct $P$ 's by first finding the appropriate physics relationship, then re-writing them in product form with parameters from the above list, and finally by solving for $P$, i.e.:

$$
m \ddot{\mathbf{x}}=e \mathbf{E} \quad \rightarrow \quad m \frac{D_{0}}{t^{2}}=e \frac{V_{0}}{D_{0}} \quad \rightarrow \quad P_{1}=t^{2} \frac{e}{m} \frac{V_{0}}{D_{0}^{2}} .
$$

For construction of next P , consider the parameter list without $e / m$ :

$$
\mathbf{E}=\mathrm{v} \times \mathbf{B} \quad \rightarrow \quad E_{0}=B_{0} \frac{D_{0}}{t} \quad \rightarrow \quad P_{2}=\frac{B_{0}}{E_{0}} \frac{D_{0}}{t}=\frac{B_{0} D_{0}^{2}}{V_{0} t} .
$$

We use $P_{1}$ to form $P_{3}$ without $t$, and use $P_{3}$ instead of $P_{2}$ :

$$
P_{3}=P_{2}^{2} P_{1}=\frac{B_{0}^{2}}{E_{0}^{2}} \frac{e}{m} V_{0} \quad \rightarrow \quad P_{3}=\frac{e}{m} \frac{B_{0}^{2} D_{0}^{2}}{V_{0}}
$$

We now remove $B_{0}$ from the parameter list, leaving only $D_{0}, V_{0}\left(E_{0}\right), t$, and we see that no additional $P$ 's are possible. Thus, we have

$$
x(t)=D_{0} F_{x}\left(P_{1}, P_{3}\right)=D_{0} F_{x}\left(t^{2} \frac{e}{m} \frac{V_{0}}{D_{0}^{2}}, \frac{e}{m} \frac{B_{0}^{2} D_{0}^{2}}{V_{0}}\right),
$$

and this is equivalently true for $y(t)$ and $z(t)$.

April, 1992. Note 0278 misc.

These expressions show the available options for changing the values of parameters if one of these has to be changed in a particular way and if one does not want to change the trajectory. If one does not care how long it takes for the particle to traverse its trajectory, then $P_{3}$ is the only $P$ that is to be kept constant. $P_{1}$ can be considered to be an expression for the time to traverse the system.

If one wants to include space charge effects, one must include $I, e$, and $\varepsilon_{0}$ to the remaining list of parameters $D_{0}, V_{0}, t$. If the magnetic fields produced by the moving charges are important, one must add $\mu_{0}$ as well.
When writing the $P$ 's, we shall use the fact that space charge and magnetic fields for charged particles go to 0 as $1 / \varepsilon_{0}$ and $\mu_{0}$ go to 0 . The space charge effects come from

$$
\nabla \cdot \varepsilon_{0} \mathbf{E}=\varrho \quad \rightarrow \quad \varepsilon_{0} \frac{E_{0}}{D_{0}}=I \frac{t}{D_{0}^{3}} \quad \rightarrow \quad P_{4}=\frac{I}{\varepsilon_{0}} \frac{t}{V_{0} D_{0}}
$$

We remove t with $P_{2}$ to get

$$
P_{5}=P_{4} P_{2}=\frac{I}{\varepsilon_{0}} \frac{1}{V_{0}} \frac{B_{0}}{E_{0}} \quad \rightarrow \quad P_{5}=\frac{I}{\varepsilon_{0}} \frac{B_{0} D_{0}}{V_{0}^{2}}
$$

We remove $B_{0}$ with $P_{3}$ to get

$$
P_{6}=\frac{P_{5}^{2}}{P_{3}}=\left(\frac{I}{\varepsilon_{0}}\right)^{2} \frac{D_{0}^{2}}{V_{0}^{4}} \frac{m}{e} \frac{V_{0}}{D_{0}^{2}} \rightarrow P_{6}=\left(\frac{I}{\varepsilon_{0}}\right)^{2} \frac{m / e}{V_{0}^{3}}
$$

with $P_{4}, P_{5}$ discarded.
We remove $I$ from our parameter list and are left with $e, \varepsilon_{0}, D_{0}, t, V_{0}$ for

$$
\nabla \cdot \varepsilon_{0} \mathbf{E}=\varrho, \quad \varepsilon_{0} E_{0}=\frac{e}{D_{0}^{2}} \quad \rightarrow \quad P_{7}=\frac{e}{\varepsilon_{0} D_{0} V_{0}}
$$

By removing $\varepsilon_{0}$ we see that no more $P$ 's are possible with just $e, D_{0}, t, V_{0}$. Thus, we have

$$
\begin{aligned}
x(t) & =D_{0} F_{x}\left(P_{1}, P_{3}, P_{6}, P_{7}\right) \\
& =D_{0} F_{x}\left(\frac{e}{m} \frac{V_{0} t^{2}}{D_{0}}, \frac{e}{m} \frac{B_{0}^{2} D_{0}^{2}}{V_{0}}, \frac{I^{2}}{\varepsilon_{0}^{2}} \frac{m / e}{V_{0}^{3}}, \frac{e}{\varepsilon_{0} D_{0} V_{0}}\right)
\end{aligned}
$$

and this is equivalently true for $y(t)$ and $z(t)$.
We expect that some $P$ 's are not significant if $\varepsilon_{0}$ is large enough so that $P_{6}$ and/or $P_{7}$ are small enough. For instance, for $D_{0}=10^{-2} \mathrm{~m}, V_{0}=10^{3} \mathrm{~V}$, and $e=$ charge of electron, we have $P_{7}=1.8 \times 10^{-9} \ll 1$, thus $P_{7}$ is probably not important.

We now add $\mu_{0}$ to the parameter list, and for $e, D_{0}, t, V_{0}, \mu_{0}$, we have

$$
\begin{gathered}
\nabla \times \mathbf{E}=-\dot{\mathbf{B}} \rightarrow \frac{E_{0}}{D_{0}}=\frac{V_{0}}{D_{0}^{2}}=\frac{\mu_{0}(e / t) / D_{0}}{t}=\frac{\mu_{0} e}{t^{2} D_{0}}, \\
P_{8}=\frac{\mu_{0} e D_{0}}{t^{2} V_{0}} .
\end{gathered}
$$

Using other $P$ 's, we get

$$
P_{9}=\frac{\mu_{0} I}{D_{0} B_{0}} .
$$

We remove $\mu_{0}$ from the parameter list, and see that no more $P$ 's are possible, and we have

$$
x(t)=D_{0} F_{x}\left(P_{1}, P_{3}, P_{6}, P_{7}, P_{9}\right),
$$

and this is equivalently true for $y(t)$ and $z(t)$.

## Application to System with Fixed $D_{0}, B_{0}$ (Cyclotrino).

We ignore $P_{1}$ since it determines traversal time. Without space charge and currentfield effects, we must keep $V_{0} m / e$ constant to get same behavior when the particle is changed, i.e. $V_{0} \sim e / m$ is necessary. To see how space charge limitation affects "permissible" current, one must look at $P_{6}$ :

$$
\frac{I^{2} m / e}{V_{0}^{3}}=\frac{I^{2}(m / e)^{4}}{\left(V_{0} m / e\right)^{3}}=\text { constant }
$$

and this implies that $(m / e)^{2} \cdot I$ should be a constant or small enough. As stated earlier, $P_{7}$ will be small enough to cause no problems, and the same will be true for $P_{9}$.

## Further Comments.

While this theory was formulated with scale factors in mind, the $P$ 's also have local meaning. That is, if the "local" $V^{\dagger}$ is interpreted as potential energy (divided by e), it becomes clear that $P_{6}$ and $P_{7}$ (with the local $V$ and $D$ ) cannot be sufficiently small to be ignored everywhere since the particles start somewhere with $e V=0$. But if the ion source is considered as a separate entity the ignorability argument will hold. It is also clear that looking at the $P$ 's with subscripted $V, V_{0}$ applies not only to applied potentials within the structure, but also to the energy of the incoming beam. ${ }^{\frac{+}{F}}$
$\dagger$ Without the subscript 0 that identifies the "global" scale.
$\ddagger$ This study was motivated by Tony Young's question about how $V_{0}$ has to be changed when $B_{0}$ differs from its original design value. Using $P_{3}$ we must have $B_{0}^{2} / V_{0}=$ constant.

## Some Practical Numbers.

Use $\gamma$ sufficiently smaller than 1 to make a $P$ ignorable. Then $P_{7}, \sqrt{P_{6}}$ can be ignored if

$$
\begin{gathered}
P_{7}=\frac{e}{\varepsilon_{0} D_{0} V_{0}}<\gamma \rightarrow \quad D_{0} V_{0}=\frac{e}{\varepsilon_{0}} \frac{1}{\gamma}>\frac{1.8 \times 10^{-8}}{\gamma}, \\
\sqrt{P_{6}}=\frac{I}{\varepsilon_{0}} \frac{\sqrt{m / e}}{V_{0}^{3 / 2}}<\gamma \rightarrow I<V_{0}^{3 / 2} \gamma \varepsilon_{0} \sqrt{e / m} .
\end{gathered}
$$

With $e_{c}=$ electron charge, and $m_{p}=$ proton mass,

$$
\begin{gathered}
I<V_{0}^{3 / 2} \gamma \sqrt{\frac{e / e_{c}}{m / m_{p}}} \varepsilon_{0} \sqrt{\frac{e_{c}}{m_{p}}}=V_{0}^{3 / 2} \gamma \sqrt{\frac{e / e_{c}}{m / m_{p}}} 8.7 \times 10^{-8} \\
I<V_{0}^{3 / 2} \gamma \sqrt{\frac{e / e_{c}}{m / m_{p}}} 8.7 \times 10^{-8} .
\end{gathered}
$$

## Analog Integrator Dynamics

Contrary to conventional analysis, which expresses the output signal in terms of the input signal, the quantity one wants (time integral over input integral) is expressed in terms of output signal (in digital form or as a scope trace), with all dynamic effects taken into account. In addition to dynamic terms being caused by the frequency response of the operational amplifier, the first order sensitivity is also affected by its dynamic behavior.


Figure 1.

For $p$ the Laplace transform variable, and $R C=\tau_{1}$,

$$
\frac{V_{0}+\varepsilon V_{2}}{R}=-V_{2}(1+\varepsilon) p C
$$

where $\mu=1 / \varepsilon \gg 1$ is the open loop gain of the operational amplifier, and

$$
V_{0}=-V_{2}\left(p \tau_{1}(1+\varepsilon)+\varepsilon\right)=-V_{2} \cdot F .
$$

We use the following rough numbers:

$$
\varepsilon=0, \quad \int V_{0} d t=10^{-6} \mathrm{Vsec}, \quad \tau_{1}=10^{-3} \mathrm{sec}, \quad \text { and } \quad-V_{2}=\frac{\int V_{0} d t}{\tau_{1}}=10^{-3} \mathrm{~V}
$$

The frequency response of the operational amplifier is

$$
\mu_{1}(p)=\frac{\mu_{0}}{1+p \tau_{0}}
$$

It actually behaves in this fashion until the open loop gain is much less than 1 . If the operational amplifier were not to behave this way, it would be useless for many

January, 1989. Note 0267misc.
applications. We characterize the frequency response either by the time constant $\tau_{0}$, or the frequency (times $2 \pi$ ) where the amplitude gain is reduced to 1 :

$$
1=\frac{\mu_{0}}{\omega_{2} \tau_{0}}=\frac{1}{\omega_{2} \varepsilon_{0} \tau_{0}}=\frac{1}{\omega_{2} \tau_{2}}
$$

with

$$
\tau_{2}=\frac{1}{\omega_{2}} \approx 10^{-7} \sec , \quad \mu_{0}=1 / \varepsilon_{0} \approx 10^{6}, \quad \text { and } \quad \varepsilon=\varepsilon_{0}\left(1+p \tau_{0}\right)=\varepsilon_{0}+p \tau_{2}
$$

and for $\tau_{1}^{*}=\left(\tau_{1}+\tau_{2}\right)$ :

$$
\begin{aligned}
F & =p \tau_{1}\left(1+p \tau_{2}\right)+\varepsilon_{0}+p \tau_{2} \\
& =p \tau_{1}^{*}+\varepsilon_{0}+p^{2} \tau_{1} \tau_{2} \\
& =p \tau_{1}^{*}+\varepsilon_{0}+p^{2} \tau_{2}\left(\tau_{1}^{*}-\tau_{2}\right) \\
-\frac{V_{0}}{p \tau_{1}^{*}} & =V_{2}\left(1+\frac{1}{p \mu_{0} \tau_{1}^{*}}+\frac{p \tau_{2} \tau_{1}}{\tau_{1}^{*}}\right) .
\end{aligned}
$$

In the time domain, the quantity of interest, $\int V_{0}(\tau) d \tau$, is expressed in terms of the measured quantity $V_{2}(t)$ by

$$
\int V_{0}(\tau) d \tau=V_{2}(t) \cdot \tau_{1}^{*}+\frac{\int V_{2}(\tau) d \tau}{\mu_{0}}+\dot{V}_{2}(t) \tau_{2} \tau_{1}
$$

One has to choose the time constants and open loop gain such that the second and third terms are small compared to the first term so that they can effectively be ignored or corrections can easily be made. It should be noted that the frequency response of the operational amplifier can make a small, but noticeable, correction to the effective time constant $\tau_{1}^{*}$ through $\tau_{2}$.

## Local Interpolation with Continuous Function and its First N Derivatives



Figure 1.

1. Real function $y_{0}(x)$ must have known values at $x=x_{0}, x_{1}, \ldots, x_{n}$.
2. Establish interpolation functions $P_{1, \ldots, n-1}(x)$, that have properties appropriate to model $y_{0}(x)$ in small regions. This necessitates continuous functions, and continuous and meaningful first $N$ derivatives. $P_{j}(x)$ must reproduce $y_{0}(x)$ exactly for $x=x_{j-1}$, $x=x_{j}$ and $x=x_{j+1}$, for $1 \leq j \leq n-1$.
3. Calculate the approximate function $y(x)$ from $y(x)=P_{1}(x) W_{1}(x)+P_{2}(x) W_{2}(x)$ in interval $x_{1} \leq x \leq x_{2}$, and similarly in other intervals. Make the choices, to some degree arbitrary, for the weight functions $W_{1, \ldots, n-1}(x)$ so that the desired goal is obtained in a reasonable fashion.
4. If, $P_{1}$ and $P_{2}$ are the same as $y_{0}(x)$, we do not want the interpolation scheme to destroy the relationship $y(x)=y_{0}(x)$. Therefore, we must have that

$$
\text { Condition 1: } \quad W_{1}(x)+W_{2}(x)=1
$$

And if the above is satisfied, it is also true that

$$
\text { Condition 2: } \quad W_{1}^{(n)}(x)+W_{2}^{(n)}(x)=0
$$

5. We examine $y^{(n)}(x)$ :

$$
\begin{aligned}
y^{(n)}(x) & =P_{1}^{(n)} W_{1}+n P_{1}^{(n-1)} W_{1}^{(1)}+\ldots+n P_{1}^{(1)} W_{1}^{(n-1)}+P_{1} W_{1}^{(n)} \\
& +P_{2}^{(n)} W_{2}+n P_{2}^{(n-1)} W_{2}^{(1)}+\ldots+n P_{2}^{(1)} W_{2}^{(n-1)}+P_{2} W_{2}^{(n)}
\end{aligned}
$$

May, 1981. Note 0177misc.
for $n=1,2, \ldots, N$. We choose $W^{(1)}$ so that all needed derivatives exist. At $x=x_{1}$ or $x=x_{2}$,

$$
P_{1} W_{1}^{(n)}+P_{2} W_{2}^{(n)}=P_{1}\left(W_{1}^{(n)}+W_{2}^{(n)}\right)=0
$$

because $P_{1}, P_{2}$ fit $y_{0}(x)$ exactly at $x=x_{1}, x=x_{2}$, and due to Condition 2 . We now choose the weight functions such that at $x=x_{1}, y^{(n)}=P_{1}^{(n)}$, and at $x=x_{2}$, $y^{(n)}=P_{2}^{(n)}$. We do this by requiring that weight functions fulfill

Condition 3: $\quad W_{1}\left(x_{1}\right)=1, \quad W_{1}\left(x_{2}\right)=0, \quad$ and

$$
W_{2}\left(x_{1}\right)=0, \quad W_{2}\left(x_{2}\right)=1
$$

and fulfill
Condition 4: $\quad W_{1}^{n}\left(x_{1}\right)=W_{1}^{n}\left(x_{2}\right)=0 \quad$ and

$$
W_{2}^{n}\left(x_{1}\right)=W_{2}^{n}\left(x_{2}\right)=0 \quad \text { for } \quad n=1,2, \ldots, N-1
$$

With the above choices, $y$ and its first $N$ derivatives at $x=x_{n}$ depend only on $P_{n}$, independently of whether we get to $x_{n}$ from an upper or lower interval, i.e., $y$ and its derivatives are continuous everywhere.
6. The construction of the weight functions that satisfy Conditions 1 (and therefore Condition 2), 3, and 4, is not unique. We introduce

$$
u(x)=\frac{x-\left(x_{1}+x_{2}\right) / 2}{\left(x_{2}-x_{1}\right) / 2}=\frac{2 x-\left(x_{1}+x_{2}\right)}{\left(x_{2}-x_{1}\right)}
$$

This gives us

$$
\begin{gathered}
u\left(x_{1}\right)=-1 \\
u\left(\left(x_{1}+x_{2}\right) / 2\right)=0 \\
u\left(x_{2}\right)=1
\end{gathered}
$$

We now have that

$$
W_{2}(x)=\frac{1}{2}\left(1+g_{N}(x)\right),
$$

$$
W_{1}(x)=\frac{1}{2}\left(1-g_{N}(x)\right)
$$

$$
\begin{aligned}
y(x) & =\frac{1}{2}\left(P_{2}(x)+P_{1}(x)+\left(P_{2}(x)-P_{1}(x)\right) \cdot g_{N}(x)\right) \\
& =P_{1}(x) W_{1}(x)+P_{2}(x) W_{2}(x)
\end{aligned}
$$

where,

$$
g_{N}(x)=a_{N} \int_{0}^{u(x)}\left(1-v^{2}\right)^{N-1} d v \text { and } \frac{1}{a_{N}}=\int_{0}^{1}\left(1-v^{2}\right)^{N-1} d v
$$

We may now conclude that, clearly, Conditions 1 and 3 are satisfied, and from

$$
\begin{aligned}
W_{2}^{(1)}(x) & =\frac{g_{N}(x)}{x_{2}-x_{1}} \\
& =\frac{a_{N}}{x_{2}-x_{1}}\left(1-u^{2}\right)^{N-1} \\
& =\frac{a_{N}}{x_{2}-x_{1}}(1-u)^{N-1}(1+u)^{N-1}
\end{aligned}
$$

it follows that Condition 4 is satisfied as well.

We introduce here some further details. Given

$$
\frac{1}{a_{N}}=\int_{0}^{1}\left(1-v^{2}\right)^{N-1} d v=b_{N}
$$

we have that

$$
\begin{aligned}
b_{N} & =\int_{0}^{1}\left(1-v^{2}\right)^{N-2}\left(1-v^{2}\right) d v \\
& =b_{N-1}-\int_{0}^{1}\left(1-v^{2}\right)^{N-2} v^{2} d v
\end{aligned}
$$

For

$$
\begin{gathered}
d u=-1\left(1-v^{2}\right)^{N-2} v d v \quad \text { and } \quad u=\frac{\left(1-v^{2}\right)^{N-1}}{2(N-1)}, \\
r=v \text { and } d r=d v
\end{gathered}
$$

we have that

$$
\begin{gathered}
b_{N}=\frac{b_{N-1}}{1+\frac{1}{2(N-1)}} \\
a_{N}=a_{N-1}\left(1+\frac{1}{2(N-1)}\right)
\end{gathered}
$$

Thus, for $a_{1}=1$

$$
\dot{a}_{N}=a_{N-1}\left(1+\frac{1}{2(N-1)}\right)=\prod_{n=1}^{N-1}\left(1+\frac{1}{2(N-1)}\right) .
$$

And further,

$$
a_{N}=\prod_{1}^{N-1} \frac{2 n+1}{2 n}=\frac{3 \cdot 5 \cdots(2 N-1)}{2^{N-1}(N-1)!}=\frac{(2 N-1)!}{2^{N-1} 2^{N-1}(N-1)!^{2}}=\frac{(2 N-1)!}{4^{N-1}(N-1)!^{2}}
$$

## Summary.

$P_{1}$ fits $y_{0}$ exactly at $x=x_{0}, x_{1}, x_{2}$.
$P_{2}$ fits $y_{0}$ exactly at $x=x_{1}, x_{2}, x_{3}$.

$$
\begin{gathered}
y(x)=P_{1}(x) W_{1}(x)+P_{2}(x) W_{2}(x), \\
W_{2}(x)=\frac{1}{2}\left(1+g_{N}(x)\right), \\
W_{1}(x)=\frac{1}{2}\left(1-g_{N}(x)\right), \\
u(x)=\frac{2 x-\left(x_{1}+x_{2}\right)}{\left(x_{2}-x_{1}\right)}, \\
\frac{1}{a_{N}}=\int_{0}^{1}\left(1-v^{2}\right)^{N-1} d v, \\
g_{N}(x)=a_{N} \int_{0}^{u(x)}\left(1-v^{2}\right)^{N-1} d v .
\end{gathered}
$$

## Special Cases.

$$
\begin{array}{ll}
N=1: & g_{1}=u . \\
N=2: & g_{2}=a_{2} \int_{0}^{u}\left(1-v^{2}\right) d v=a_{2} u\left(1-u^{2} / 3\right), \quad 1 / a_{2}=2 / 3 \\
& g_{2}=\frac{1}{2} u\left(3-u^{2}\right) . \\
N=3: & g_{3}=a_{3} \int_{0}^{u}\left(1-2 v^{2}+v^{4}\right)=a_{3} u\left(1-\frac{2}{3} u^{2}+\frac{u^{4}}{5}\right), \quad \frac{1}{a^{3}}=\frac{1}{3}+\frac{1}{5}=\frac{8}{15}, \\
& g_{3}=\frac{1}{8} u\left(15-10 u^{2}+3 u^{4}\right) .
\end{array}
$$

## Linear Least Squares with Erroneous Matrix

When one is dealing with a system in which the relationships between parameter changes, $\Delta p$, and the system performance changes, $\Delta s$, are in good approximation represented by the linear relationship

$$
\Delta s=M \Delta p
$$

achieving a desired performance change is simply accomplished by parameter changes

$$
\Delta p=M^{-1} \Delta s
$$

as long as one has as many parameters as system performance characteristics.
When the desired change, $\Delta s$, has more components than $\Delta p$, it is often adequate to minimize the weighted sum of the deviations from the desired performance, i.e. one minimizes

$$
S=\Delta s^{t} W \Delta s
$$

where $W$ is a diagonal square matrix with appropriate weights on the diagonal. $S$ is minimized in the first iteration if the parameter vector is changed by

$$
\Delta p_{1}=A \Delta s_{1} \quad \text { where } \quad A=\left(M^{t} W M\right)^{-1} M^{t} W
$$

If the matrix $M$ used for this operation deviates by $\Delta M=M_{\Re}-M$ from the real matrix $M_{\Re}$, the desired change $\Delta s_{2}$ with the new parameters is given by

$$
\Delta s_{2}=\Delta s_{1}-(M+\Delta M) \Delta p_{1}=(I-M A-\Delta M A) \Delta s_{1}
$$

If the effort to determine $M_{\Re}$ (often by elaborate measurements) is too large one can iterate the procedure, and it would be of interest to estimate the asymptotic $\Delta s_{\infty}$. To obtain this, we introduce

$$
I-M A=B \text { and }-\triangle M A=D
$$

Thus,

$$
\Delta s_{n}=(B+D)^{n-1} \Delta s_{1} \text { and } \Delta p_{n}=A(B+D)^{n-1} \Delta s_{1}
$$

Notice that $A M=I, A B=0, D B=0$ and $B^{2}=I-2 M A+M A M A=B$.

November, 1971. Note 0038 misc.

Further,

$$
\begin{gathered}
(B+D)^{2}=B(I+D)+D^{2} \\
(B+D)^{3}=B\left(I+D+D^{2}\right)+D^{3}
\end{gathered}
$$

and so forth such that

$$
(B+D)^{n}=B(I-D)^{-1}\left(1-D^{n}\right)+D^{n}
$$

Therefore,

$$
\frac{\Delta s_{n}=\left(B(I-D)^{-1}\left(1-D^{n-1}\right)+D^{n-1}\right) \Delta s_{1}}{\Delta p_{n}=A D^{n-1} \Delta s_{1}}
$$

as it must be, because for $\Delta M=0$ and $n \geq 2, \Delta p_{n}=0$ and $\Delta s_{n}=\Delta s_{2}$.
If $\Delta M$ is small enough, the absolute values of the eigenvalues of $D$ will be less than 1 , resulting in the following for large enough $n$ :

$$
\Delta s_{\infty}=B(I-D)^{-1} \Delta s_{1}=(I-M A)(1+\Delta M A)^{-1} \Delta s_{1}
$$

Judging whether one is close to this value is possible by observing the decrease in $\Delta p_{n}$ with increasing $n$.

## Matrix Describing Second Order Effects to Second Order in One Dimension

## No Momentum Errors.

The normalized equation of motion is

$$
y^{\prime \prime}=y+b y^{2}
$$

Expand $y$ in terms of initial conditions $y_{0}, y_{0}^{\prime}$ up to $2^{\text {nd }}$ order:

$$
y=a_{11} y_{0}+a_{12} y_{0}^{\prime}+a_{13} y_{0}^{2}+a_{14} y_{0} y_{0}^{\prime}+a_{15} y_{0}^{\prime 2}
$$

Initial conditions for $a(x)$

$$
a_{11}(0)=a_{12}^{\prime}(0)=1
$$

all others are 0 . The equation for $a(x)$ is:

$$
\begin{gathered}
a_{11}^{\prime \prime}=a_{11}, \quad \text { and } \quad a_{12}^{\prime \prime}=a_{12} \Longrightarrow \begin{cases}a_{11}=\cosh x, & a_{12}=\sinh x \\
a_{21}=\sinh x, & a_{22}=\cosh x\end{cases} \\
a_{13}^{\prime \prime}=a_{13}+b a_{11}^{2}, \quad a_{14}^{\prime \prime}=a_{14}+2 b a_{11} a_{12}, \quad \text { and } \quad a_{15}^{\prime \prime}=a_{15}+b a_{12}^{2}
\end{gathered}
$$

Because in all three cases $a(0)=a^{\prime}(0)=0: \mathcal{L}\left(a^{\prime \prime}\right)=p^{2} \mathcal{L}(a)$.
For $a_{13}$,

$$
a_{11}^{2}=\frac{1}{4}\left(e^{2 x}+e^{-2 x}+2\right) \rightleftharpoons \frac{1}{4}\left(\frac{1}{p-2}+\frac{1}{p+2}+\frac{2}{p}\right) .
$$

In general,

$$
\frac{1}{(\dot{p}-1)(p+1)(p+c)} \rightleftharpoons \frac{e^{x}}{2(1+c)}+\frac{e^{-x}}{2(1-c)}+\frac{e^{-c x}}{c^{2}-1}
$$

thus,

$$
\begin{aligned}
\frac{4 a_{11}^{2}}{p^{2}-1} & \rightleftharpoons \frac{4}{3} \cdot \frac{e^{x}}{2}+\frac{4}{3} \cdot \frac{e^{-x}}{2}+\frac{e^{2 x}+e^{-2 x}}{3}-2 \\
& =\frac{4}{3} \cosh x+\frac{2}{3} \cosh 2 x-2
\end{aligned}
$$

Therefore,

$$
a_{13}=\frac{b}{6}(2 \cosh x+\cosh 2 x-3) \text { and } a_{13}^{\prime}=a_{23}=\frac{b}{3}(\sinh x+\sinh 2 x) .
$$

February, 1966. Note 0006misc.

For $a_{14}$,

$$
a_{11} a_{12}=\frac{1}{4}\left(e^{2 x}-e^{-2 x}\right) \rightleftharpoons \frac{1}{4}\left(\frac{1}{p-2}-\frac{1}{p+2}\right),
$$

and thus,

$$
\begin{aligned}
\frac{4 a_{11} a_{12}}{p^{2}-1} & \rightleftharpoons \frac{-4}{3} \cdot \frac{e^{x}}{2}+\frac{4}{3} \cdot \frac{e^{-x}}{2}+\frac{e^{2 x}-e^{-2 x}}{3} \\
& =-\frac{4}{3} \sinh x+\frac{2}{3} \sinh 2 x
\end{aligned}
$$

Therefore,

$$
a_{14}=\frac{b}{3}(\sinh 2 x-2 \sinh x) \text { and } a_{14}^{\prime}=a_{24}=\frac{2 b}{3}(\cosh 2 x-\cosh x) .
$$

Similarly, for $a_{15}$,

$$
a_{12}^{2}=\frac{1}{4}\left(e^{2 x}+e^{-2 x}-2\right) \rightleftharpoons \frac{1}{4}\left(\frac{1}{p-2}+\frac{1}{p+2}-\frac{2}{p}\right)
$$

thus,

$$
\begin{aligned}
\frac{4 a_{12}^{2}}{p^{2}-1} & \rightleftharpoons \frac{-8}{3} \cdot \frac{e^{x}}{2}+\frac{-8}{3} \cdot \frac{e^{-x}}{2}+\frac{e^{2 x}+e^{-2 x}}{3}+2 \\
& =-\frac{8}{3} \cosh x+\frac{2}{3} \cosh 2 x+2
\end{aligned}
$$

Therefore,

$$
a_{15}=\frac{b}{6}(\cosh 2 x-4 \cosh x+3) \quad \text { and } \quad a_{15}^{\prime}=a_{25}=\frac{b}{3}(\sinh 2 x-2 \sinh x) .
$$

## Inclusion of Momentum Error $\alpha$.

The normalized equation of motion is

$$
y^{\prime \prime}=y+\alpha+b y^{2}
$$

First, add the term linear in $\alpha$ to the expansion in $y_{0}, y_{0}^{\prime}$ : add $a_{16} \alpha$. The initial conditions are

$$
a_{16}(0)=a_{16}^{\prime}(0)=0, \quad a_{16}^{\prime \prime}=a_{16}+1 \quad \Longrightarrow \quad a_{16}=\cosh x-1
$$

Second, take the terms $\alpha^{2}, \alpha y_{0}, \alpha y_{0}^{\prime}$ into account, where the procedure is the same as in the calculation of $a_{13}, a_{14}$, and $a_{15}$.
Third, do not add any terms, but introduce $z=y+\beta$ ( $\beta$ is a constant) in the differential equation. Thus,

$$
\begin{gathered}
z^{\prime \prime}=z-\beta+\alpha+b\left(z^{2}-2 z \beta+\beta^{2}\right) \\
\beta^{2}-\frac{\beta}{b}+\frac{\alpha}{b}=0 \Longrightarrow \beta=\frac{1}{2 b}(1-\sqrt{1-4 \alpha \beta}) \\
z(1-2 b \beta)+b z^{2}=z^{\prime \prime}=z \sqrt{1-4 \alpha \beta}+b z^{2}
\end{gathered}
$$

This procedure requires the calculation of a new" matrix for every $\alpha$ of interest, but this will give more insight in return.

## General Procedure.

Description of higher order effects with power expansion and the consequences for stability.
We describe deviations from the closed orbit by the column vector

$$
\mathrm{v}=\left(\begin{array}{c}
y_{1} \\
y_{2} \\
\vdots \\
y_{k}
\end{array}\right)
$$

where the components of $y_{1}$ are $y$ and $y^{\prime}$, and components of $y_{2}, y_{3}, \ldots, y_{k}$ are, respectively, the second, third, $\ldots, k^{\text {th }}$ order contributions of $y$ and $y^{\prime}$. Then,

$$
\mathbf{v}_{2}=M \mathrm{v}_{1}, \quad \text { with } \quad M=\left(\begin{array}{ccccc}
A_{11} & A_{12} & A_{13} & \ldots & A_{1 k} \\
0 & A_{22} & A_{23} & \ldots & A_{2 k} \\
0 & 0 & A_{33} & \ldots & A_{3 k} \\
\vdots & \vdots & \vdots & \ddots & \vdots \\
0 & 0 & 0 & \ldots & A_{k k}
\end{array}\right) .
$$

In $M, A_{11}$ describes the first order effects, $A_{12}$ the second order effects, etc. The
other matrices reproduce the higher than first order components of $v$. The diagonal elements $A_{22}, A_{33}, \ldots, A_{k k}$ depend only on the matrix element of $A_{11}$. The eigenvalues of $M$ do not depend on $A_{12}, \ldots, A_{1 k}$. Thus, the stability of the system does not depend, in this approximation, on the non-linear effects described by these elements. Since stability obviously can depend on non-linear effects, this implies that the power expansion for many passes through the system has a progressively shrinking radius of convergence. One can thus conclude that although this method is worthless to evaluate the effect of non-linearities on stability, it might still yield valuable information provided the system does not become unstable because of the non-linearities.
We show rough numbers for

$$
\begin{gathered}
F_{3}=a_{13} / b, \quad F_{4}=a_{14} / b, \quad F_{5}=a_{15} / b \\
F_{3}^{\prime}=a_{13}^{\prime} / b=a_{23} / b, \quad F_{4}^{\prime}=a_{14}^{\prime} / b=a_{24} / b, \quad F_{5}^{\prime}=a_{15}^{\prime} / b=a_{25} / b
\end{gathered}
$$

| $x$ | $\pi / 4$ | $\pi / 2$ | $\pi$ |
| :---: | :---: | :---: | :---: |
| $e^{x}$ | 2.2 | 4.8 | 23 |
| $e^{-x}$ | 0.46 | 0.21 | 0 |
| $\cosh x$ | 1.33 | 2.5 | 11.5 |
| $\sinh x$ | 0.87 | 2.3 | 11.5 |


| $x$ | $\pi / 4$ | $\pi / 2$ |
| :---: | :---: | :---: |
| $F_{3}(x)$ | 0.77 | 2.25 |
| $F_{4}(x)$ | 0.18 | 2.3 |
| $F_{5}(x)$ | 0.03 | 0.75 |


| $x$ | $\pi / 4$ | $\pi / 2$ |
| :---: | :---: | :---: |
| $F_{3}^{\prime}(x)$ | 1.06 | 4.6 |
| $F_{4}^{\prime}(x)$ | 0.78 | 6 |
| $F_{5}^{\prime}(x)$ | 0.18 | 2.3 |

## Curvature of 2D Magnetic Field Lines and Scalar Potential Lines

## I. Preparation and Background.

Magnetic fields. 2D magnetic fields can be derived from a scalar potential $V$ or a vector potential $A$, or the complex potential $F(z)=A+i V$, an analytic function of the complex variable $z=x+i y$, according to

$$
\begin{equation*}
B_{x}-i B_{y}=B^{*}=i \frac{d F}{d z}=i F^{\prime} \tag{1}
\end{equation*}
$$

Field lines and scalar potential lines in the $z$-plane are the $z(F)$ maps of straight lines parallel to either the imaginary or real axis of the $F$-plane.
Modification of the curvature by a conformal map. If a curve in the $z$-plane has a local tangent in the direction $e^{i \alpha_{z}}$, the conformal map $w(z)$ of that region has a local tangent in the direction

$$
\begin{equation*}
e^{i \alpha_{z}}=e^{i \alpha_{z}} \frac{w^{\prime}}{\left|w^{\prime}\right|} \tag{2}
\end{equation*}
$$

This equation shows that the angle of intersection of any two curves in the $z$-plane is preserved under the transformation $w(z)$, hence the name conformal transformation. If the curve at that location in the $z$-plane is $k_{z}$, then the curvature of the map of that point can be shown to be.

$$
\begin{equation*}
k_{w}=\frac{\left(k_{z}+\Im\left(\frac{w^{\prime \prime}}{w^{\prime}} e^{i \alpha_{z}}\right)\right)}{\left|w^{\prime}\right|} \tag{3}
\end{equation*}
$$

The sign convention used for this formula is such that a positive curvature means that if one proceeds in the direction of the tangent, the curve turns to the left, i.e. the conventional mathematically positive direction.

## II. Application of (2).

Fundamental relationships. There are several ways to apply (2) to this problem. The most natural way to do so seems to be, at least at first, to assign quantities $w$ and $z$ in (2) to the the variables $F$ and $z$ of our problem, since we are looking at the map of a region of the $F$-plane to the $z$-plane. For most applications, this is not very practical since one then gets the curvature of the maps of constant potential lines as a function of $A$ and $V$, when in fact one wants the curvature as a function of $x$ and

[^3]$y$. We therefore proceed in the following manner: we assign $z$ and $w$ to $z$ and $F$, and look in
\[

$$
\begin{equation*}
k_{F}=\frac{\left(k_{z}+\Im\left(\frac{F^{\prime \prime}}{F^{\prime}} e^{i \alpha_{z}}\right)\right)}{\left|F^{\prime}\right|} \tag{4}
\end{equation*}
$$

\]

for $k_{F}=0$, i.e. the curvature of maps of straight lines in the $F$-plane is given by

$$
\begin{equation*}
k_{z}=-\Im\left(\frac{F^{\prime \prime}}{F^{\prime}} e^{i \alpha_{z}}\right) \tag{5}
\end{equation*}
$$

To get a more practical formula, we express $e^{i \alpha_{z}}$ with the help of (2) through

$$
\begin{equation*}
e^{i \alpha_{z}}=e^{i \alpha_{F}} \frac{\left|F^{\prime}\right|}{F^{u}} \tag{6}
\end{equation*}
$$

and the derivatives of $F$ through the fields as given in (1), yielding

$$
\begin{equation*}
k_{z}=-\Re\left(\frac{B^{*^{\prime}}}{B^{* 2}}|B| e^{i \alpha_{F}}\right) \tag{7}
\end{equation*}
$$

For some expressions of the fields, it is more convenient to write this as

$$
\begin{equation*}
k_{z}=+\Re\left(\left(\frac{1}{B^{*}}\right)^{\prime}|B| e^{i \alpha_{F}}\right) \tag{8}
\end{equation*}
$$

In both (7) and (8), $e^{i \alpha_{F}}$ has the absolute value 1 and is real if one is looking at a scalar potential line, and is purely imaginary if one looks at a field line.

Comments. It is worth noting that in order to calculate the curvatures of interest, one needs only the expressions for the complex field, not the complex potential. Under most circumstances, the expression for the complex potential is not more complicated than the expression for the complex field. There are, however, exceptions. For instance, the field of a modified sextupole is given by

$$
\begin{equation*}
B^{*}=i z^{2} e^{a z^{2}} \tag{9}
\end{equation*}
$$

Integrating this to get the complex potential, (1), leads to the error function in the complex plane.

## III. Applications.



Figures $1(a, b)$.
(i) The regular multipole. For a multipole of order $n$ with the field perpendicular to the midplane, the field is given by

$$
\begin{equation*}
B^{*}=i z^{n-1} \tag{10}
\end{equation*}
$$

Substituting in (8) gives directly

$$
\begin{equation*}
k_{z}=(n-1)\left|z^{n-1}\right| \Re\left(i z^{-n} e^{i \alpha_{w}}\right) . \tag{11}
\end{equation*}
$$

Using, for $e^{i \alpha_{F}}$, the phases corresponding to the arrows in Figures 1(a) and 1(b), and using $z=r e^{i \varphi}$, gives, for the curvature of the field line and the scalar equipotential:

$$
\begin{align*}
& k_{z}=(n-1) \cos \frac{n \varphi}{r}  \tag{12A}\\
& k_{z}=(n-1) \sin \frac{n \varphi}{r} \tag{12V}
\end{align*}
$$

(ii) The modified sextupole. This particular implementation of a modified sextupole has the field in the midplane perpendicular to the midplane, and behaves like a good sextupole close to the origin, but has a stronger modified field, proportional to $x^{2} e^{a x^{2}}$, $a \in \Re$, as one moves away from the origin of the coordinate system. The complex field is therefore given by

$$
B^{*}=i z e^{a x^{2}}
$$

- $\quad \therefore \quad$ • $\because \square$


## Fringe Field Model Function for Dipoles

For a number of beam optics tasks, it is important to have an analytical function that describes the field in the fringe field region of a dipole ${ }^{i}$. We restrict ourselves to the simple case of a dipole that has a straght effective field boundary, making this a very simple problem of describing two dimensional fields. Putting the $x$-axis into the midplane of a dipole whose half gap is normalized to be equal to 1 , with large $x>0$ describing the outside of the magnet, and the far negative end of the $x$-axis the deep inside region of the magnet, the field in the region of interest can be described by

$$
\begin{align*}
\frac{B_{y}(x, y)+i B_{x}(x, y)}{B_{0}}= & G(z)=D_{1}(z)+D_{2}(z)+D_{3}(z)  \tag{0.0}\\
& z=x+i y \tag{0.1}
\end{align*}
$$

and the functions $D_{1}, D_{2}, D_{3}$ chosen such that the asymptotic behavior of $G(z)$ reflects the properties of the fields in the regions deep inside and far outside the magnet. In addition, $G(z)$ should not have any singularities for the space within $-1<y<1$. The following functions satisfy these conditions:

$$
\begin{gather*}
D_{1}(z)=\frac{1+n A e^{\pi z / 2}}{\left(1+A e^{\pi z / 2}\right)^{n}}  \tag{1}\\
D_{2}(z)=C_{1} e^{-C_{2}\left(z-x_{2}\right)^{2}}  \tag{2}\\
D_{3}(z)=K_{1} \cdot \frac{\left(1+K_{3} e^{-\pi z / 2}\right)^{m}}{\left(1+K_{2}\left(z-x_{3}\right)^{2}\right)^{3 / 2}}, \tag{3}
\end{gather*}
$$

with all coefficients real, $n \geq 2, K_{2}>1$, and $A, C_{2}, K_{3}, m>0$. The fields deep inside the magnet are dominated by the "longest surviving" term $e^{\pi z}$ from $D_{1}(z)$, while far outside the magnet the field is dominated, as desired, by the "longest surviving" term proportional to $1 / z^{3}$ from $D_{3}$, with clearly no singularities for $-1<y<1$. $D_{2}(z)$ has been added (and one could add more such terms) to allow a good fit of $G(z)$ to measured or computed data in the transition region. While this suggested model function $G(z)$ has enough free parameters to fit data, the quality of such a fit has not been tested on a real problem, but the $G(z)$ given here should contain a sufficient number of suggestions that this approach to the Enge function promises to be successful.
$\dagger$ See document 0609thry, Comments about RAYTRACE.

## Comments about RAYTRACE

## Introduction.

Several years ago I was asked at a workshop to comment on the representation of magnetic fields in the RAYTRACE code, a computer program that was developed by H. A. Enge and his students in the 1960 's ${ }^{\dagger}$. Since my comments contained not only some academically interesting points, but also suggestions for improvement of this enormously successful code, several people asked me to put my thoughts on this subject on paper. After describing the specific aspects of the code that I want to discuss, I will elaborate on what I would characterize as shortcomings, together with suggestions for eliminating them, and a description of some mathematical detail at the end.

## Fields in RAYTRACE.

Even though it is not a major effort to generalize my comments, I restrict the discussion to the case of the fringe field region of a dipole magnet that has a straight effective field boundary in the region of interest. This means that we are dealing with two dimensional fields, with all the associated simplifications that make it possible to address the core of the problem without unnecessary distractions.
Using the midplane of the magnet as the $x$-axis of the $x y$ coordinate system, with large positive $x$ representing the region far outside the dipole, and the other extreme the region deep inside the magnet, the field is characterized by the following function, commonly called the Enge function :

$$
\begin{equation*}
B_{y}(x, 0)=\frac{B_{0}}{1+e^{P(x)}} \tag{1a}
\end{equation*}
$$

where

$$
\begin{equation*}
P(x)=C_{0}+C_{1} x+C_{2} x^{2}+\ldots+C_{n} x^{n}, \tag{1b}
\end{equation*}
$$

with $n$ an odd integer and $C_{n}>0$.
The coefficients are obtained by fitting measured or computed field values in the midplane to (1), and fields off the midplane are obtained by using a Taylor series expansion, with the derivatives obtained from (1).

## Comments and Suggestions.

I have' problems with three tightly linked aspects of this procedure:

[^4](A) It is true, in general, that if one fits parameters of a function so that the fields on the surface of a volume are well represented by that function, the quality of the fields computed with that function inside the volume is at least as good as (but usually better than) the original data on the surface. It is, of course, assumed that the function and field calculation algorithm satisfy all the relevant vacuum field equations. Conversely, calculating fields in the volume from a function whose free parameters were determined on a line inside the volume gives fields that are not nearly as accurate as the original data. These facts are qualitatively clear if one thinks of the fringe fields in the midplane of the dipole: significantly different pole contours produce very similar fields in the midplane. That means that if one calculates fields off the midplane accurately from the fields in the midplane, small differences in the function there will give significantly different fields far away from the midplane.
(B) Calculating fields off the midplane with a Taylor series expansion makes no sense in this case for the following reasons: since $B_{x}-i B_{y}$ or, more conveniently in this case, $B_{y}+i B_{x}$, is an analytical function of the complex variable $z=x+i y$, the field at location $(x, y)$ can be obtained directly, without any approximation, by evaluating (1) for complex argument:
\[

$$
\begin{equation*}
B_{y}(x, y)+i B_{x}(x, y)=B_{y}(x+i y, 0)=\frac{B_{0}}{1+e^{P(x+i y)}} \tag{2}
\end{equation*}
$$

\]

This very simple evaluation of fields from a midplane model function makes it obviously easy to fit the parameters of the model, no matter the nature of that function, to fields off the midplane, thus eliminating the objection raised in (A).
(C) It seems to me that the form of the Enge function is not well suited to this problem for two reasons: 1) the function does not have ther appropriate asymptotic behavior far away from either end of the magnet; and 2) unless one makes a careful study of the Enge function, it may have one or more singularities in the "business" region. Avoiding that kind of disaster by evaluating the field only approximately is clearly not a satisfactory answer to this problem. While it is fairly easy with the help of (2) to make the singularity check (see Appendix), it might be simpler to "design" a function that can not have that kind of singularity, in addition to having the proper asymptotic behavior. I have some very promising candidates but have not made the effort to test them on some real problems.

## Appendix.

For the Enge function to have no damaging singularity it is necessary and sufficient that the equation

$$
\begin{equation*}
P(z)=i m \pi \text { with } m=\text { odd integer } \neq 0 \tag{3}
\end{equation*}
$$

has no solution for $z$ between the midplane and a line parallel to the midplane one half gap, $h$, away from the midplane. This test is most easily carried out with the argument principle that states, in this case, that the number of zeroes of $w(\approx)$ within a region of the $z$-plane equals the number of times $w(z)$ goes around $w=0$ when $z$ traces the boundary of the region. Since, in this case,

$$
\begin{equation*}
w(z)=P(z)+i m \pi, \tag{4}
\end{equation*}
$$

with all $C_{n}$ in $P(z)$ real, it is only necessary to find the locations where the map of the straight line parallel to the midplane at distance $h$ intersects the imaginary axis of $\mathrm{P}(\mathrm{z})$, i.e. one has to find $\Im P(z)$ at the locations where $\Re P(x+i h)=0$. Since $\Re P(x+i h)=0$ means nothing more than finding the real roots (in $x$ ) of a polynomial of order $n$, this a very simple exercise for a computer. Having these points, it is trivial to see whether $w(z)=0$ is possible for any odd $m$. I have carried out that test for the example given by Spencer and Enge, and for four cases given to me by S. Kowalski. I am happy to report that while none of these cases had singularities within one half gap of the midplane, there were some singularities just outside the end of the dipole only a little more than a half gap away from the midplane.

## Stored Energy in H-Magnet for $\mu=\infty$



Figure 1.

$$
2 \mathcal{E}=\int \mathbf{B} \cdot \mathbf{H} d v=\int \mathbf{H} \cdot \nabla \times \mathbf{A} d \dot{v}
$$

From

$$
\nabla \cdot(\mathbf{A} \times \mathbf{H})=\mathbf{H} \cdot(\nabla \times \mathbf{A})-\mathbf{A} \cdot(\nabla \times \mathbf{H})
$$

we have that, with $\mathbf{j}=j \mathbf{e}_{z}$,

$$
\begin{aligned}
2 \mathcal{E} & =\int \mathbf{A} \cdot \mathbf{j} d v+\int(\mathbf{A} \times \mathbf{H}) \cdot d \mathbf{a} \\
& =\int \mathbf{A} \cdot \mathbf{j} d v+\int \mathbf{A} \cdot(\mathbf{H} \times d \mathbf{a})
\end{aligned}
$$

where $\mathrm{H} \times d \mathrm{a}=0$ on $\mu=\infty$ surface.
In case of a long magnet, $\int j d a \equiv 0$ which means that we can add any constant to $A$ without changing anything. We make $A=0$ along the $y$-axis. We now use $\mathrm{A}=\mu_{0} A \mathrm{e}_{z}$, so that the total energy per unit length is

$$
\mathcal{E}^{\prime}=\frac{1}{2} \mu_{0} j \int A d x d y
$$

where the integral is evaluated over the coil in the first quadrant.

To get

$$
J_{1}=\int A(x, y) d x d y
$$

we look at

$$
\begin{aligned}
J_{2}(y) & =\int_{0}^{y<y_{2}} \int_{0}^{h_{2}} H_{x} d x d y=\iint \frac{\partial A}{\partial y} d y d x \\
& =\int\left(A(x, y)-A_{t}\right) d x=\int A(x, y) d x-A_{t} h_{2}
\end{aligned}
$$

where $A_{t}$ is $A$ at top of the coil slot.
We integrate the original expression for $J_{2}$ over $x$ first, and by Ampère's Law,

$$
J_{2}(y)=\int_{0}^{y} \frac{I}{y_{2}} y d y=\frac{I y^{2}}{2 y_{2}} .
$$

Therefore,

$$
\frac{\int A(x, y) d x=\frac{I y^{2}}{2 y_{2}}+A_{t} h_{2}}{J_{1}=\int_{0}^{y_{2}} \int A(x, y) d x d y=\frac{I y_{2}^{2}}{6}+A_{t} h_{2} y_{2} .}
$$

From H-Magnet With Minimal Yoke Flux Density ${ }^{\dot{\dagger}}$ we know that

$$
A_{t}=I\left(\frac{W_{1}}{h_{1}}+\frac{D+y_{2} / 2}{h_{2}}+E_{1}\right)
$$

and thus we have that

$$
\begin{gathered}
J_{1}=I h_{2} y_{2}\left(\frac{W_{1}}{h_{1}}+\frac{D+y_{2} / 2}{h_{2}}+E_{1}+\frac{y_{2}}{6 h_{2}}\right) . \\
j J_{1}=I^{2}\left(\frac{W_{1}}{h_{1}}+\frac{2 y_{2}}{3 h_{2}}+\frac{D}{h_{2}}+E_{1}\left(\frac{h_{1}}{h_{2}}\right)\right) . \\
\quad . \\
\mathcal{E}^{\prime}=\frac{1}{2} \mu_{0}\left(j J_{1}\right),
\end{gathered}
$$

$\dagger$ Document 0606thry.
where

$$
\begin{gathered}
I=H_{1} h_{1}=j h_{2} y_{2}, \\
E_{1}(a)=a+\frac{2}{\pi}\left(\ln \frac{a+1 / a}{4}+\left(\frac{1}{a}-a\right) \arctan a\right) \quad \text { with } a=h_{2} / h_{1} .
\end{gathered}
$$

## H-Magnet With Minimal Yoke Flux Density



Figure 1.
$W_{1}, H_{1}, W_{3}, D$, and $h_{1}$ are given. We want to minimize $\bar{B}_{\text {yoke,max }}$ for $\mu=\infty$ by chosing the proper $h_{2}$.

## Procedure.

Calculate flux for 0-thickness coil at top of coil slot using excess flux coefficient $E_{1}$ for corner. Subtract "window frame flux" from combination of real coil and 0 -thickness coil.
For $V=H_{1} h_{1}=j h_{2} y_{2}$ we chose $h_{2}$ and $j$. Thus,

$$
\begin{aligned}
A= & V \cdot\left(\frac{W_{1}}{h_{1}}+\frac{D+y_{2}}{h_{2}}+E_{1}\left(\frac{h_{1}}{h_{2}}\right)-\frac{y_{2}}{2 h_{2}}\right) \mu_{0} \\
= & \mu_{0} V\left(\frac{W_{1}}{h_{1}}+\frac{D+y_{2} / 2}{h_{2}}+E_{1}\left(\frac{h_{1}}{h_{2}}\right)\right) \\
& \bar{B}_{\text {yoke,max }}=\frac{A}{W_{3}-W_{1}-h_{2}} .
\end{aligned}
$$

We determine the minimum value of $\bar{B}_{\text {yoke,max }}$ by varying $h_{2}$, and we define

$$
E_{1}(a)=a+\frac{2}{\pi}\left(\ln \frac{a+1 / a}{4}+\left(\frac{1}{a}-a\right) \arctan a\right) \quad \text { with } \quad a=h_{2} / h_{1}
$$

[^5]
## Dipole with Small Gap Bypass



Figure 1.

$$
\begin{equation*}
B_{0} h_{0}+h_{1} \mu_{0} H\left(B_{0}\right)=B_{1}\left(h_{1}+h_{0}\right) \tag{1}
\end{equation*}
$$

For $W_{0} / W_{2}=\varepsilon_{0}$,

$$
\begin{gathered}
B_{2} W_{2}=B_{0} W_{0}+B_{1} W_{1} \text { and thus } B_{2}=B_{0} \varepsilon_{0}+B_{1}\left(1-\varepsilon_{0}\right) \\
V_{00}=\frac{B_{1}}{\mu_{0}}\left(h_{0}+h_{1}\right)+h_{2} H\left(B_{0} \varepsilon_{0}+B_{1}\left(1-\varepsilon_{0}\right)\right)
\end{gathered}
$$

The exact equation

$$
\begin{equation*}
\mu_{0} V_{00}=B_{1}\left(h_{0}+h_{1}\right)+\mu_{0} h_{2} H\left(B_{0} \varepsilon_{0}+B_{1}\left(1-\varepsilon_{0}\right)\right) \tag{2}
\end{equation*}
$$

has the following implementations

$$
B_{1} \rightarrow B_{0} \text { and then } B_{1}, B_{0} \rightarrow V_{00}
$$

We will now examine three special cases.

June, 1993. Note 0591thry.

In Case 1, $V_{00}$ is so small that $\mu_{0} H(B)=\gamma B$,

$$
\begin{gathered}
B_{0}\left(h_{0}+\gamma h_{1}\right)=B_{1}\left(h_{0}+h_{1}\right), \\
\mu_{0} V_{00}=B_{1} \underbrace{\left(h_{0}+h_{1}+\gamma h_{2}\left(\varepsilon_{0} \frac{h_{0}+h_{1}}{h_{0}+\gamma h_{1}}+1-\varepsilon_{0}\right)\right)}_{K},
\end{gathered}
$$

where

$$
\begin{gathered}
\frac{h_{0}+h_{1}}{h_{0}+\gamma h_{1}}=1+\frac{(1-\gamma) h_{1}}{h_{0}+\gamma h 1} \text { and } K_{h_{0}}^{\prime}=1-\frac{\gamma h_{1} \varepsilon_{0} h_{2}(1-\gamma)}{\left(h_{0}+\gamma h_{1}\right)^{2}}, \\
B_{1}=\frac{\mu_{0} V_{00}}{K}, \quad B_{1}^{\prime}=\frac{-\mu_{0} V_{00}}{K^{2}} K^{\prime}=\frac{\mu_{0} V_{00}}{K^{2}}\left(\frac{\gamma h_{1} \varepsilon_{0} h_{2}(1-\gamma)}{\left(h_{0}+\gamma h_{1}\right)^{2}}-1\right) .
\end{gathered}
$$

For $h_{0} \ll \gamma h_{1}$ :

$$
B_{1}^{\prime}=\frac{\mu_{0} V_{00}}{K^{2}}\left(\frac{\varepsilon_{0} h_{2}(1-\gamma)}{\gamma h_{1}}-1\right)>0
$$

$K^{\prime}=0$ for $h_{0}+\gamma h_{1}=\sqrt{\varepsilon_{0} \gamma h_{1} h_{2}(1-\gamma)}$ such that

$$
h_{0}=\gamma h_{1}(\underbrace{\sqrt{\frac{\varepsilon_{0} h_{2}(1-\gamma)}{\gamma h_{1}}}-1}_{\gg 1}) \approx \sqrt{\varepsilon_{0} h_{1} h_{2}(1-\gamma) \gamma} .
$$

For $h_{1}=1, h_{2}=5, \varepsilon_{0}=1 / 2, \gamma=10^{-3}$, we have

$$
h_{0}=\sqrt{1 / 2 \cdot 1 \cdot 5 \cdot 10^{-3}}=\frac{1}{20} \mathrm{~cm} .
$$

In Case 2, we need $V_{00}$ large enough so that $B_{0} \approx B_{s}$, but small enough so that for (2) $\mu_{0} H(B)=\gamma B$, thus

$$
\mu_{0} V_{00}=B_{1}\left(h_{0}+h_{1}\right)+B_{1} h_{2} \gamma\left(1-\varepsilon_{0}\right)+B_{s} h_{2} \gamma \varepsilon_{0}
$$

where

$$
B_{1}=\frac{\mu_{0} V_{00}-B_{s} \gamma h_{2} \Xi_{0}}{h_{0}+h_{1}+h_{2} \gamma\left(1-\varepsilon_{0}\right)} .
$$

In a third, simple case, Case 3 , for a still higher $V_{00}$,

$$
B_{s} \varepsilon_{0}+B_{1}\left(1-\varepsilon_{0}\right) \rightarrow B_{s}
$$

i.e. it is independent of $V_{00}$, and thus

$$
B_{1}=B_{s} .
$$

## Boundary Condition at Iron-Air Interface for AC and Application to 2-Dimensional Cylinder

Interface at $y=0$,

$$
\begin{gathered}
\frac{\partial}{\partial x}=\frac{\partial}{\partial z}=0 \quad \text { and } \quad \frac{\partial}{\partial y}={ }^{\prime} \\
\mathbf{H}=H \mathrm{e}_{x}, \quad \mathrm{j}=\mathrm{e}_{z} \sigma E \text { and } \mathbf{E}=\mathrm{e}_{z} E . \\
(\nabla \times \mathbf{H})_{z}=-H^{\prime}=\sigma E \text { and } E=-\varrho H^{\prime} \\
(\nabla \times \mathbf{E})_{x}=E^{\prime}=-\mu_{0} \mu p H=-\varrho H^{\prime \prime} \quad \text { and } H^{\prime \prime}-k^{2} H=0
\end{gathered}
$$

where, depending on the application, $p$ is either the Laplace transform variable or, for sinusoidal excitation, $i \omega$.

$$
\begin{gathered}
k=\sqrt{\sigma \mu_{0} \mu p}=\frac{1}{D_{1}} \\
H=H_{0} e^{-k y} \text { and } \Phi=\mu_{0} \mu \int_{0}^{\infty} H d y=\frac{H_{0} \mu_{0} \mu}{k}=\mu_{0} H_{0} \mu D_{1} .
\end{gathered}
$$

Therefore, with

$$
\mu D_{1}=D_{2}
$$

$$
\begin{gathered}
\frac{d \Phi}{d x} \Delta x=\mu_{0} D_{2} \frac{\partial H_{0}}{\partial x} \Delta x=\mu_{0} H_{y} \Delta x \\
H_{y}=D_{2} \frac{\partial H_{0}}{\partial x} \Longrightarrow H_{\perp}=D_{2} \frac{\partial H_{\|}}{\partial s_{\|}} .
\end{gathered}
$$

Given an iron cylinder with $D_{2}$, of radius 1 , in far-field $\mathbf{H}=H_{\infty} \mathbf{e}_{x}$, we try to solve for the complex potential $F$. Ansatz: the superposition of the macro and micro dipoles, with normalized units.

$$
F=-i H_{1}(z+1 / z)-i H_{2}(z-1 / z)
$$

with $z=x+i y$, normalized with radius $r_{0}$ of the cylinder.

On $|z|=1$

$$
\begin{gathered}
F=A+i V=2 H_{2} \sin \varphi-i 2 H_{1} \cos \varphi \\
H^{*}=i F^{\prime}=H_{x}-i H_{y} \text { and } \mathcal{H}=H_{r}+i H_{\varphi}=H e^{-i \varphi} \\
\mathcal{H}^{*}=H^{*} e^{i \varphi}
\end{gathered}
$$

and the boundary condition is, with

$$
\begin{gathered}
H_{\|}=H_{\varphi} \quad \text { and } \quad H_{\perp}=-H_{r} \\
H_{\tau}=-D_{2} \frac{\partial H_{\varphi}}{\partial \varphi} \\
H^{*}=H_{1}\left(1-1 / z^{2}\right)+H_{2}\left(1+1 / z^{2}\right)=e^{-i \varphi} \mathcal{H}^{*}
\end{gathered}
$$

On the surface,

$$
\begin{aligned}
& \mathcal{H}^{*}=H_{r}-i H_{\varphi}=2 i H_{1} \sin \varphi+2 H_{2} \cos \varphi, \\
& H_{r}=2 H_{2} \cos \varphi \text { and } H_{\varphi}=2 H_{1} \sin \varphi
\end{aligned}
$$

and the boundary condition is, with

$$
2 H_{2} \cos \varphi=D_{2} 2 H_{1} \sin \varphi+2 H_{2} \cos \varphi \quad \text { and } \quad H_{2}=D_{2} H_{1}
$$

$$
H_{\infty}=H_{1}+H_{2}=H_{1}\left(1+D_{2}\right)
$$

$$
H_{1}=\frac{H_{\infty}}{1+D_{2}}, \quad H_{2}=\frac{H_{\infty} D_{2}}{1+D_{2}}
$$

$$
H_{\tau}=2 H_{\infty} \frac{D_{2}}{1+D_{2}} \cos \varphi \text { and } H_{\varphi}=-2 H_{\infty} \frac{\sin \varphi}{1+D_{2}}
$$

normalized with radius $r_{0}$ of the cylinder.

Using SI units, we choose $\sigma \mu_{0}=10, \mu=10^{4}$ and $\omega=2 \pi \cdot 60 \mathrm{~Hz}$, and therefore we have

$$
\left|D_{2}\right|=\sqrt{\frac{\mu}{\sigma \mu_{0} \omega}}=\frac{10^{2}}{\sqrt{10^{4}(.12 \pi)}},
$$

$$
\left|D_{2}\right|=1.6 \mathrm{~m}
$$

$$
\left|D_{1}\right|=.16 \mathrm{~mm} \quad \text { and } \quad\left|D_{1}\right| \sqrt{2}=.23 \mathrm{~mm}
$$

For sinusoidal excitation,

$$
\begin{gathered}
D_{2}=\left|D_{2}\right| \frac{(1-i)}{\sqrt{2}} \\
\left|1+D_{2}\right|=\sqrt{\left(1+\frac{\left|D_{2}\right|}{\sqrt{2}}\right)^{2}+\frac{\left|D_{2}\right|^{2}}{2}}=\sqrt{1+\left|D_{2}\right|^{2}+\left|D_{2}\right| \sqrt{2}}
\end{gathered}
$$

Normalized, where $r_{0}$ is the radius of the cylinder:

$$
D_{2}=\frac{D_{2}(\mathrm{~m})}{\mathrm{r}_{0}(\mathrm{~m})}
$$

That is, for same material and frequency, $\left|D_{2}\right|$ is large for a small cylinder and $\left|D_{2}\right|$ is small for a large cylinder.
Unfortunately, if $H_{\perp}=D_{2} \frac{\partial H_{\|}}{\partial s_{\|}}$is valid in z-geometry, it is not satisfied in conformally mapped $w$-geometry, i.e. dealing with this problem in mapped geometry is not practical.

## Flux Into A Rectangular Box



Figure 1.

$$
\dot{z}=c \frac{\sqrt{t^{2}-\varepsilon^{2}}}{\sqrt{t^{2}-1}}
$$

For

$$
t=\varepsilon \sin \varphi, \quad d t=\varepsilon \cos \varphi d \varphi \quad \sqrt{1-\varepsilon^{2}}=\varepsilon_{1}
$$

we have

$$
\begin{gathered}
\frac{a}{c}=\int_{0}^{\varepsilon} \frac{\sqrt{\varepsilon^{2}-t^{2}}}{\sqrt{1-t^{2}}}=\varepsilon^{2} \int_{0}^{\pi / 2} \frac{\cos ^{2} \varphi d \varphi}{\sqrt{1-\varepsilon^{2} \sin ^{2} \varphi}}=\int_{0}^{\pi / 2} \frac{\left(1-\varepsilon^{2} \sin ^{2} \varphi\right)+\left(\varepsilon^{2}-1\right)}{\sqrt{1-\varepsilon^{2} \sin ^{2} \varphi}} d \varphi \\
\frac{a}{c}=\int_{0}^{\pi / 2} \sqrt{1-\varepsilon^{2} \sin ^{2} \varphi} d \varphi-\varepsilon_{1}^{2} \int_{0}^{\pi / 2} \frac{d \varphi}{\sqrt{1-\varepsilon^{2} \sin ^{2} \varphi}} \\
\frac{a}{c}=E\left(\varepsilon^{2}\right)-\varepsilon_{1}^{2} K\left(\varepsilon^{2}\right) .
\end{gathered}
$$

For

$$
t=\cos \varphi, d t=-\sin \varphi d \varphi, \quad \varepsilon=\cos \alpha, \quad \varepsilon_{1}=\sin \alpha
$$

May, 1988. Note 0491thry.

$$
\sin \varphi=\varepsilon_{1} \sin \psi, d \varphi=\frac{\varepsilon_{1} \cos \psi d \psi}{\sqrt{1-\varepsilon_{1}^{2} \sin ^{2} \psi}}
$$

we have

$$
\begin{aligned}
\frac{b}{c}=\int_{\varepsilon}^{1} \frac{\sqrt{t^{2}-\varepsilon^{2}}}{\sqrt{1-t^{2}}} d t= & \int_{0}^{\alpha} \sqrt{\varepsilon_{1}^{2}-\sin ^{2} \varphi} d \varphi=\varepsilon_{1}^{2} \int_{0}^{\pi / 2} \frac{\cos ^{2} \psi}{\sqrt{1-\varepsilon_{1}^{2} \sin ^{2} \psi}} d \psi \\
& \frac{b}{c}=E\left(\varepsilon_{1}^{2}\right)-\varepsilon^{2} K\left(\varepsilon_{1}^{2}\right) .
\end{aligned}
$$

Thus

$$
\frac{a}{b}=\frac{E\left(\varepsilon^{2}\right)-\varepsilon_{1}^{2} K\left(\varepsilon^{2}\right)}{E\left(\varepsilon_{1}^{2}\right)-\varepsilon^{2} K\left(\varepsilon_{1}^{2}\right)} .
$$

For

$$
F(t)=c B_{\infty} t, \quad F^{\prime}=B_{\infty} \frac{\sqrt{t^{2}-1}}{\sqrt{t^{2}-\varepsilon^{2}}}, \quad B_{0}=\frac{B_{\infty}}{\varepsilon} .
$$

and therefore

$$
F(\varepsilon)=\frac{a B_{\infty} \varepsilon}{E\left(\varepsilon^{2}\right)-\varepsilon_{1}^{2} K\left(\varepsilon^{2}\right)},
$$

$$
F(1)=\frac{a B_{\infty}}{E\left(\varepsilon^{2}\right)-\varepsilon_{1}^{2} K\left(\varepsilon^{2}\right)}
$$

Given a square box, with dimensions $\varepsilon^{2}=1 / 2, E(1 / 2)=1.3506, K(1 / 2)=1.8541$,

$$
F(\varepsilon)=F(\sqrt{1 / 2})=1.67 a B_{\infty}, \quad F(1)=2.361 a B_{\infty}, \quad B_{0}=1.41 B_{\infty}
$$

## Propagation of Fast Perturbation in Dipole

We describe the boundary condition as

$$
H_{y}(h)=D_{2} \frac{\partial H_{x}}{\partial x} \cdot \text { with } \quad D_{2}=\mu D_{1}=\frac{\mu}{\sqrt{i \omega \sigma \mu_{0} \mu}}
$$

Ansatz:

$$
H_{y}(x, y)=\sum a_{n} \cos k_{n} y e^{-k_{n} x}
$$

where we look to satisfy the $H_{y}(-y)=\dot{H}_{y}(y)$ symmetry only. $\nabla^{2} H_{y}=0$ is obviously satisfied.

$$
\begin{gathered}
\frac{\partial H_{x}}{\partial y}=\frac{\partial H_{y}}{\partial x}=-\sum a_{n} k_{n} \cos k_{n} y e^{-k_{n} x} \\
H_{x}=-\sum a_{n} \sin k_{n} y e^{-k_{n} x}
\end{gathered}
$$

At the boundary we have

$$
\sum a_{n} \cos k_{n} h e^{-k_{n} x}=\sum a_{n} D_{2} k_{n} \sin k_{n} h e^{-k_{n} x}
$$

For

$$
D_{2} k_{n} \tan k_{n} \dot{h}=1 \quad \text { and } \quad \alpha_{n}=k_{n} h
$$

we therefore have

$$
\alpha_{n} \tan \alpha_{n}=\frac{h}{D_{2}}
$$

where

$$
\frac{1}{D_{2}}=\sqrt{\frac{i \omega \sigma \mu_{0}}{\mu}}
$$

November, 1987. Note 0489thry.

Case 1: "normal" case,

$$
\mu \rightarrow \infty \quad \alpha_{n}=n \pi
$$

Case 2: superconducting case, $\quad \sigma \rightarrow \infty \quad \alpha_{n}=(n+1 / 2) \pi$.
Case 3: using iron with $\omega=2 \pi 60 \mathrm{~Hz}$, and given that $D_{1}=\frac{1}{\sqrt{i \omega \sigma \mu_{0} \mu}}$,

$$
\left|D_{1}\right|=\frac{1}{\sqrt{2 \pi \cdot 60 \cdot 10^{1+3}}}=5.2 \cdot 10^{-4} \mathrm{~m}=.52 \mathrm{~mm}
$$

$$
\left|D_{2}\right|=\mu\left|D_{1}\right|=52 \mathrm{~cm}
$$

We introduce

$$
\alpha_{0} \tan \alpha_{0}=\frac{h}{D_{2}}=\varepsilon=h \sqrt{\frac{i \omega \sigma \mu_{0}}{\mu}}
$$

and for $|\varepsilon| \ll 1, \alpha_{0} \approx \sqrt{\varepsilon}$. Thus, for $\alpha_{n}=n \pi+\delta_{n}$,

$$
\left(n \pi+\delta_{n}\right) \tan \delta_{n}=\varepsilon \Longrightarrow \delta_{n} \approx \frac{\varepsilon}{n \pi}
$$

For a better notation of $\alpha_{0}$ we have that, for

$$
\alpha_{0}^{2}+\alpha_{0}^{4} / 3=\varepsilon \quad \Longrightarrow \quad \alpha_{0}^{2}=-3 / 2+\sqrt{9 / 4+3 \varepsilon}
$$

and it follows that

$$
\alpha_{0}^{2}=\frac{3 \varepsilon}{\frac{3}{2}+\sqrt{\frac{9}{4}+3 \varepsilon}}=\frac{2 \varepsilon}{1+\sqrt{1+\frac{4 \varepsilon}{3}}}
$$

$$
\alpha_{0}^{2}=\frac{2 \varepsilon}{2+\frac{2 \varepsilon}{3}}=\frac{\varepsilon}{1+\frac{\varepsilon}{3}} .
$$

To determine $a_{n}$ from $H_{y}(y)$ at $x=0$ we try

$$
\int_{0}^{h} H_{y}(y) \cos k_{m} y d y=\sum a_{n} \int_{0}^{h} \cos k_{n} y \cos k_{m} y d y
$$

Since $2 \cos k_{n} y \cos k_{m} y=\cos \left(k_{n}+k_{m}\right) y+\cos \left(k_{n}-k_{m}\right) y$,

$$
\begin{aligned}
& \frac{2}{h} \int_{0}^{h} \cos k_{n} y \cos k_{m} y d y=\frac{\sin \left(\alpha_{n}+\alpha_{m}\right)}{\alpha_{n}+\alpha_{m}}+\frac{\sin \left(\alpha_{n}-\alpha_{m}\right)}{\alpha_{n}-\alpha_{m}} \\
&=\frac{\alpha_{n}\left(\sin \left(\alpha_{n}+\alpha_{m}\right)+\sin \left(\alpha_{n}-\alpha_{m}\right)\right)}{\alpha_{n}^{2}-\alpha_{m}^{2}} \\
&-\frac{\alpha_{m}\left(\sin \left(\alpha_{n}+\alpha_{m}\right)-\sin \left(\alpha_{n}-\alpha_{m}\right)\right)}{\alpha_{n}^{2}-\alpha_{m}^{2}} \\
&=\frac{2\left(\alpha_{n} \sin \alpha_{n} \cos \alpha_{m}-\alpha_{m} \cos \alpha_{n} \sin \alpha_{m}\right)}{\alpha_{n}^{2}-\alpha_{m}^{2}} \\
&=\frac{2 \cos \alpha_{n} \cos \alpha_{m}\left(\alpha_{n} \tan \alpha_{n}-\alpha_{m} \tan \alpha_{m}\right)}{\alpha_{n}^{2}-\alpha_{m}^{2}} \\
&=0 \text { for } n \neq m .
\end{aligned}
$$

Note: this orthogonality condition is not satisfied for $\int_{0}^{h} \sin k_{n} y \sin k_{m} y d y$. So, for instance, $V(0, y)$ would not work "directly". One would have to first calculate $H_{y}(0, y)$.

## Description of the Properties of an Ellipse

For many problems, one needs integrals over the circumference of an ellipse, whose equation is

$$
\frac{x^{2}}{a^{2}}+\frac{y^{2}}{b^{2}}=1
$$

One may describe the ellipse by the parametric representation

$$
z=a \cos \varphi+i b \sin \varphi
$$

and use $\varphi$ as the integration variable.
However, in many cases, it is mathematically more convenient to use $z$ on the circumference as the integrations variable. If one can represent all quantities of interest on the circumference as analytic functions of $z$, one can then use the Cauchy Theorem to execute the integration.
In using the parametric representation, one usually does something similar by introducing $e^{i \varphi}$ as the new integration variable. While this oftens works very well, it can lead to difficulties: for example, when $e^{k z}$ appears in the function to be integrated.

In general, problems are much simpler for circles, where $a=b$. When $b \neq a$ it often becomes so difficult to execute the integral that it is most convenient to expand in a quantity that is equivalent to $a-b$ and thus the formulas will be easily written and therefore the expansion will be similarly easy.
Thus,

$$
\begin{align*}
& z=a \cos \varphi+i b \sin \varphi=e^{i \varphi} \frac{a+b}{2}+e^{-i \varphi} \frac{a-b}{2}  \tag{1}\\
& \quad 0=e^{i \varphi}-2 \frac{z}{a+b}+e^{-i \varphi} \frac{a-b}{a+b}=0 \\
& \quad=e^{2 i \varphi}(a+b)^{2}-2 z e^{i \varphi}(a+b)+\varepsilon \\
& \quad \varepsilon=a^{2}-b^{2} \text { and } W_{1}=\sqrt{1-\varepsilon / z^{2}} \tag{2}
\end{align*}
$$

$$
\begin{equation*}
e^{i \varphi}=\frac{z+\sqrt{z^{2}-\varepsilon}}{a+b}=z \frac{1+W_{1}}{a+b} \tag{3.1}
\end{equation*}
$$

January, 1988. Note 0476thry.

$$
\begin{align*}
e^{-i \varphi} & =2 \frac{z}{a-b}-e^{i \varphi} \frac{a+b}{a-b} \\
& =\frac{z-\sqrt{z^{2}-\varepsilon}}{a-b}=z \frac{1-W_{1}}{a-b}  \tag{3.2}\\
& =\frac{(a+b) / z}{1+W_{1}} \tag{3.3}
\end{align*}
$$

For $\cos \varphi, \sin \varphi$ :

$$
\begin{gather*}
\frac{1}{a+b}+\frac{1}{a-b}=\frac{2 a}{\varepsilon} \text { and } \frac{1}{a+b}-\frac{1}{a-b}=\frac{-2 b}{\varepsilon}, \\
\cos \varphi=\frac{a z-b \sqrt{z^{2}-\varepsilon}}{\varepsilon}=\frac{z^{2}+b^{2}}{a z+b \sqrt{z^{2}-\varepsilon}}=\frac{z^{2}+b^{2}}{z\left(a+b W_{1}\right)},  \tag{4.1}\\
i \sin \varphi=\frac{-b z+a \sqrt{z^{2}-\varepsilon}}{\varepsilon}=\frac{z^{2}-a^{2}}{b z+a \sqrt{z^{2}-\varepsilon}}=\frac{z^{2}-a^{2}}{z\left(b+a W_{1}\right)} .  \tag{4.2}\\
d s=\sqrt{a^{2} \sin ^{2} \varphi+b^{2} \cos ^{2} \varphi d \varphi .} \tag{5}
\end{gather*}
$$

From (3.1) we have

$$
\begin{gather*}
i e^{i \varphi} d \varphi=\frac{\left(1+\left(z / \sqrt{z^{2}-\varepsilon}\right)\right) d z}{a+b}=\frac{e^{i \varphi} d z}{z W_{1}} \\
d \varphi=\frac{1}{i} \frac{d z}{z W_{1}} \tag{6.0}
\end{gather*}
$$

Thus,

$$
\begin{aligned}
\varepsilon^{2} G= & \varepsilon^{2}\left(a^{2} \sin ^{2} \varphi+b^{2} \cos ^{2} \varphi\right) \\
= & z^{2}\left(b^{2}\left(a-b W_{1}\right)^{2}-a^{2}\left(b-a W_{1}\right)^{2}\right) \\
= & z^{2}\left(W_{1}^{2}\left(b^{4}-a^{4}\right)+2 a b W_{1}\left(a^{2}-b^{2}\right)\right) \\
= & z^{2}\left(\frac{\varepsilon\left(a^{2}+b^{2}\right)}{z^{2}}-(a-b)^{2}-2 a b \frac{1-W_{1}^{2}}{1+W_{1}}\right) \\
& G=a^{2}+b^{2}-z^{2} \frac{\varepsilon}{(a+b)^{2}}-\frac{2 a b}{1+W_{1}}
\end{aligned}
$$

where

$$
\frac{2}{1+W_{1}}=1+\frac{1-W_{1}}{1+W_{1}}
$$

and thus

$$
\begin{align*}
G & =a^{2}+b^{2}-a b-z^{2} \frac{\varepsilon}{(a+b)^{2}}-a b \frac{1-W_{1}}{1+W_{1}} \\
& =(a-b)^{2}+a b-z^{2} \frac{\varepsilon}{(a+b)^{2}}-a b \frac{1-W_{1}}{1+W_{1}} \\
& =\frac{\varepsilon^{2}}{(a+b)^{2}}+a b-z^{2} \frac{\varepsilon}{(a+b)^{2}}-a b \frac{1-W_{1}}{1+W_{1}} \\
& =a b-\varepsilon \frac{z^{2}-\varepsilon}{(a+b)^{2}}-a b \frac{1-W_{1}}{1+W_{1}} . \tag{7.1}
\end{align*}
$$

To expand an expression like

$$
\frac{1-W_{1}}{1+W_{1}} \quad \text { with } \quad W_{1}=\sqrt{1-\varepsilon / z^{2}}
$$

in $\varepsilon$, it is often convenient to break it up into an even and odd part in $\varepsilon$ :

$$
\begin{gathered}
2 F(\varepsilon)=F(\varepsilon)+F(-\varepsilon)+F(\varepsilon)-F(-\varepsilon) \text { with } W_{2}=\sqrt{1+\varepsilon / z^{2}}, \\
2 \frac{1-W_{1}}{1+W_{1}}=2 H=\frac{1-W_{1}}{1+W_{1}}+\frac{1-W_{2}}{1+W_{2}}+\frac{1-W_{1}}{1+W_{1}}-\frac{1-W_{2}}{1+W_{2}}, \\
2 H\left(1+W_{1} W_{2}+W_{1}+W_{2}\right)=\left(1-W_{1}\right)\left(1+W_{2}\right)+\left(1+W_{1}\right)\left(1-W_{2}\right) \\
\quad+\left(1-W_{1}\right)\left(1+W_{2}\right)-\left(1+W_{1}\right)\left(1-W_{2}\right) \\
=2\left(1-W_{1} W_{2}+W_{2}-W_{1}\right), \\
\cdot \\
\cdot \begin{array}{c}
\frac{1-W_{1}}{1+W_{1}}=\frac{1-\sqrt{1-\varepsilon^{2} / z^{4}}+\sqrt{1+\varepsilon / z^{2}}-\sqrt{1-\varepsilon / z^{2}}}{1+\sqrt{1-\varepsilon^{2} / z^{4}}+\sqrt{1+\varepsilon / z^{2}}+\sqrt{1-\varepsilon / z^{2}}}
\end{array}
\end{gathered}
$$

To second order in $\varepsilon$ :

$$
\begin{equation*}
\frac{1-W_{1}}{1+W_{1}}=\frac{\varepsilon}{z^{2}} \frac{1+\varepsilon / 2 z^{2}}{4} \tag{7.2}
\end{equation*}
$$

A comment about the expansion in $\varepsilon$ and subsequent integration: the expansion has to be valid and good for $z$ on the ellipse. If, to carry out the integration, one later modifies the integration path (in particular, to a very small circle around $z=0$ ), this will not invalidate the original expansion.

Addendum. A different way to derive G.
For

$$
\begin{gathered}
W_{0}=\sqrt{z^{2}-\varepsilon} \\
\varepsilon=a^{2}-b^{2} \quad \text { and } s=a^{2}+b^{2} \\
s^{2}-\varepsilon^{2}=4 a^{2} b^{2} \quad \text { and } \quad 2 a b=\sqrt{s^{2}-\varepsilon^{2}}
\end{gathered}
$$

we have

$$
\begin{aligned}
\varepsilon^{2} G & =(b \varepsilon \cos \varphi+i a \varepsilon \sin \varphi)(b \varepsilon \cos \varphi-i a \varepsilon \sin \varphi) \\
& =W_{0}\left(2 a b z-W_{0}\left(a^{2}+b^{2}\right)\right) \\
& =s W_{0}\left(z \sqrt{1-\varepsilon^{2} / s^{2}}-W_{0}\right) \\
& =s W_{0} \frac{z^{2}\left(1-\varepsilon^{2} / s^{2}\right)-z^{2}+\varepsilon}{z \sqrt{1-\varepsilon^{2} / s^{2}}+W_{0}} \\
G=s W_{0} & \frac{1-\varepsilon z^{2} / s^{2}}{z \sqrt{1-\varepsilon^{2} / s^{2}}+W_{0}}=s W_{1} \frac{1-\varepsilon z^{2} / s^{2}}{\sqrt{1-\varepsilon^{2} / s^{2}}+W_{1}} .
\end{aligned}
$$

To first order in $\varepsilon$ :

$$
G=s \frac{1}{2}\left(1-\frac{\varepsilon}{2 z^{2}}\right)\left(1-\frac{\varepsilon z^{2}}{s^{2}}\right)\left(1+\frac{\varepsilon}{4 z^{2}}\right)=\frac{s}{2}\left(1-\varepsilon\left(\frac{z^{2}}{s^{2}}+\frac{1}{4 z^{2}}\right)\right)
$$

for $s=2 a^{2}$ and $s^{2}=4 a^{4}$,

$$
\sqrt{G}=a\left(1-\frac{\varepsilon}{2}\left(\frac{z^{2}}{s^{2}}+\frac{1}{4 z^{2}}\right)\right)
$$

$$
\sqrt{G}=a\left(1-\frac{\varepsilon}{8 a^{2}}\left(\frac{z^{2}}{a^{2}}+\frac{a^{2}}{z^{2}}\right)\right)
$$

## Characterization of Dipole Fringe Fields with Field Integrals



Figure 1.

## Background and Introduction.

The quantity $\int B_{y}(x, y, z) d z$ was measured as a function of $y$ for a fixed $x$, with integration region beginning in the homogenous field region inside the dipole magnet and reaching into the essentially field-free region outside. This resulted in the approximate plotted curve of Figure 1 below.


Figure 2.

The conclusion reached pointed to the coil being too close to or too far from the midplane. For didactic purposes this is a very interesting problem for two reasons.
(1) The coil position is only indirectly responsible. The fact that $\int B_{y} d z$ depends on $y$ indicates that this is a 3 D problem: namely, $\int B_{y}(x, 0, z) d z$ will have a curvature of opposite polarity (i.e. effective field boundary is curved). This is due either to a curvature of the pole ends (when projected into the $x z$-plane) or to the finite width in the $x$-direction. If the latter is the cause, the problem is magnified by the absence (or incorrect design) of the field clamp and by a coil that is too far from the midplane.
(2) The characterization of the fringe field by measuring $\int B_{y}(x, 0, z) d z$ gives, in case of midplane symmetry, more information than $\int B_{y}(0, y, z) d z$ alone.

May, 1986. Note 0438thry.

## Analysis.

We assume midplane symmetry. Violation of midplane symmetry should be detected and/or measured, preferably with null method.
In vacuum, full 3D, the following hold:

$$
\begin{gather*}
\frac{\partial B_{y}}{\partial x}-\frac{\partial B_{x}}{\partial y}=0  \tag{1.1}\\
\frac{\partial B_{x}}{\partial x}+\frac{\partial B_{y}}{\partial y}+\frac{\partial B_{z}}{\partial z}=0 \tag{1.2}
\end{gather*}
$$

We now investigate the properties of

$$
\begin{equation*}
\int_{z_{1}}^{z_{2}} B_{x}(x, y, z) d z / L=b_{x}(x, y) \text { and } \int_{z_{1}}^{z_{2}} B_{y}(x, y, z) d z / L=b_{y}(x, y) \tag{2a,b}
\end{equation*}
$$

where $z_{1}, z_{2}$ are constants, i.e. they are not considered variables; $L$ is a convenient length that is used only for normalization purposes. Integration is performed over (1.1). Integration and differentiation can be interchanged and thus

$$
\begin{equation*}
\frac{\partial b_{y}}{\partial x}-\frac{\partial b_{x}}{\partial y}=0 \tag{3.1}
\end{equation*}
$$

Integration is performed over (1.2) and

$$
\begin{equation*}
\frac{\partial b_{x}}{\partial x}+\frac{\partial b_{y}}{\partial y}=\left(B_{z}\left(x, y, z_{1}\right)-B_{z}\left(x, y, z_{2}\right)\right) / L \tag{3.2}
\end{equation*}
$$

If, independently of $x$ and $y, B_{z}$ at the two end-points is the same ${ }^{\dagger}$, we have

$$
\begin{equation*}
\frac{\partial b_{x}}{\partial x}+\frac{\partial b_{y}}{\partial y}=0 \tag{3.3}
\end{equation*}
$$

(3.1) and (3.2) mean that $b_{x}$ and $-b_{y}$ satisfy the Cauchy-Riemann conditions of real and imaginary parts of the analytic function of $Z=x+i y$ :

$$
\begin{equation*}
b_{x}-i b_{y}=b^{*}(Z) \tag{4}
\end{equation*}
$$

We use the Taylor series to represent $b^{*}(Z)$ :

$$
b^{*}(Z)=-i \sum_{0}^{n} a_{n} Z^{n}
$$

where $n=$ multipole order -1 , i.e. $n=0 \Longrightarrow$ dipole, $n=1 \Longrightarrow$ quadrupole, etc. Because of midplane symmetry all $a_{n}$ must be real. Notice that for $y=0$,

[^6]$b_{y}(x, 0)=\sum a_{n} x^{n}$, i.e. all harmonics contribute; while for $x=0$, only $a_{n}$ with even $n$ contributes to $b_{y}(0, y)$. If one measures for $x=0$ both $b_{y}$ and $b_{x}$, then one gets information about all harmonics. There is an advantage to measuring both $b_{y}(0, y)$ and $b_{x}(0, y)$ since one gives only the odd harmonics and the other only the even, while they are all mixed when calculating $b_{y}(x, 0)$.
Looking at (5) it is obvious that if $b_{y}(0, y)$ is not constant, but depends on $y$, then $b_{y}(x, 0)$ must depend on $x$. That is, one is in fact dealing with 3D fields which must be due to curved (in projection on $x y$-plane) poles or poles of insufficient width in the $x$-direction, and failure to use a field clamp. It is also qualitatively clear that these 3D effects get more pronounced with increasing distance of the coils from the midplane.
Specifically, for a known value of $b_{y}(0, y)$, what is $b_{y}(x, 0)$ ?
\[

$$
\begin{gather*}
b_{y}(0, y)=\sum_{0}^{m} a_{2 m} y^{2 m}(-1)^{m}  \tag{6.1}\\
b_{y}(x, 0)=\sum_{0}^{m} a_{2 m} x^{2 m}+\sum_{0}^{m} a_{2 m+1} \tag{6.2}
\end{gather*}
$$
\]

where in (6.2) $a_{2 m+1}$ is not obtainable from $b_{y}(0, y)$, but can be obtained from $b_{x}(0, y)$. If one measures $b_{x}(0, y)$, one gets

$$
\begin{equation*}
b_{x}(0, y)=\sum_{0}^{m} a_{2 m+1} y^{2 m+1}(-1)^{m} \tag{6.3}
\end{equation*}
$$

For simple analysis, one should plot $b_{y}(0, y)$ vs $y^{2}$, and $b_{x}(0, y) / y$ vs $y^{2}$. One has to be careful to make the measurements in such a way that they really mean something. The flux loop and integrator method is perhaps best because it can practically always be done in such a way that one makes a null measurment.

## Penetration of Solenoidal Field through Conducting Shell

Preliminaries: Solenoid, shield infinitely long. Thin shell, circular cross-section: treat eddy currents in it in plane geometry with proper boundary values. Only one shell: matrix formulation not needed.


Figure 1.

At $z=0$,

$$
H_{y}=H_{0}(p)=\text { given solenoid current. }
$$

At $z=D$,

$$
E_{x}=\frac{\mu_{0} p H_{1} \pi r_{1}^{2}}{2 \pi r_{1}}=\frac{r_{1}}{2} \mu_{0} p H_{1}
$$

In shell,

$$
\begin{gathered}
\mathbf{H}=\mathbf{e}_{y} H, \quad \mathbf{E}=\mathbf{e}_{x} E, \quad \text { and } \quad \sigma=\frac{\partial}{\partial z} \\
\nabla \times \mathbf{H}=\sigma \mathbf{E} \quad \Longrightarrow \quad-H^{\prime}=\sigma E \\
\nabla \times \mathbf{E}=-\mu_{0} \mu p \mathbf{H} \quad \Longrightarrow \quad E^{\prime}=-\mu_{0} \mu p H
\end{gathered}
$$

For $\mu_{0} \mu \sigma p=k^{2}$,

$$
\begin{gathered}
H^{\prime \prime}=-\sigma E^{\prime}=\mu_{0} \mu \sigma p H=k^{2} H \\
H=H_{0} \cosh k z+b \sinh k z
\end{gathered}
$$

and for $\gamma=k D$,

$$
H_{1}=H_{0} \cosh \gamma+b \sinh \gamma, \quad \text { and thus } \quad b=\frac{H_{1}-H_{0} \cosh \gamma}{\sinh \gamma}
$$

May, 1986. Note 0437thry.

$$
\begin{gathered}
E_{1}=\frac{r_{1}}{2} \mu_{0} p H_{1}=-\frac{H_{1}^{\prime}}{\sigma}=-\frac{k}{\sigma}\left(H_{0} \sinh \gamma+b \cosh \gamma\right), \\
H_{1}=\frac{\mu_{0} \sigma p r_{1}}{2 k}=-\left(H_{0} \sinh \gamma+\frac{\cosh \gamma\left(H_{1}-H_{0} \cosh \gamma\right)}{\sinh \gamma}\right),
\end{gathered}
$$

where, for

$$
\frac{\mu_{0} \sigma p r_{1}}{2 k}=\frac{k^{2} r_{1} / \mu}{2 k}=\frac{r_{1}}{2} \frac{k}{\mu}=\gamma \varepsilon, \quad \text { and thus } \quad \varepsilon=\frac{r_{1}}{2 \mu D},
$$

where diffusion occurs for $\mu \gg 1$. Thus, for

$$
\begin{gathered}
H_{1}(\varepsilon \gamma \sinh \gamma+\cosh \gamma)=H_{0}, \\
H_{1}(p)=\frac{H_{0}}{\cosh \gamma+\varepsilon \gamma \sinh \gamma} .
\end{gathered}
$$

The zeroes of $\cosh \gamma+\varepsilon \gamma \sinh \gamma$ are given, with

$$
\gamma_{n}=D \sqrt{\mu_{0} \mu \sigma p_{n}}=i \alpha_{n}
$$

And

$$
\begin{gathered}
\cosh \gamma+\varepsilon \gamma \sinh \gamma=\cos \alpha_{n}-\varepsilon \alpha_{n} \sin \alpha_{n}=0 \\
\frac{1}{\varepsilon}=\alpha_{n} \tan \alpha_{n} \\
p_{n}=\frac{-\alpha_{n}^{2}}{\mu_{0} \mu \sigma D^{2}}
\end{gathered}
$$

One may expect difficulties for some calculations because $\gamma=D \sqrt{\mu_{0} \mu \sigma p}$, but this is not so, for

$$
\begin{aligned}
\cosh \gamma+\varepsilon \gamma \sinh \gamma & =1+\gamma^{2}\left(\frac{1}{2}+\varepsilon\right)+\gamma^{4}\left(\frac{1}{4!}+\frac{\varepsilon}{3!}\right)+\gamma^{6}\left(\frac{1}{6!}+\frac{\varepsilon}{5!}\right)+\cdots \\
& =1+\sum_{n=1}^{\infty} \gamma^{2 n}\left(\frac{1}{(2 n)!}+\frac{\varepsilon}{(2 n-1)!}\right)
\end{aligned}
$$

and $\cosh \gamma+\varepsilon \gamma \sinh \gamma$ is a simple function of $\gamma^{2}$, not a complicated function of $\gamma$. For

$$
\varepsilon \gg 1 \Rightarrow \alpha_{0}^{2} \approx \frac{1}{\varepsilon}=\mu_{0} \mu \sigma p_{0} D^{2}=\frac{2 \mu D}{r_{1}} \Longrightarrow \mu_{0} \sigma D r_{1} p_{0}=2
$$

with $\sigma D$ implying non-diffusion.

The roots, in quadrants 1 and 3 , are given by,

$$
\tan \alpha_{n}=\frac{1}{\alpha_{n} \varepsilon}
$$

Where, for $\alpha_{n} \varepsilon \gg 1, \alpha_{n}=n \pi+\sigma_{n}$, and thus

$$
\sigma=\frac{1}{\varepsilon\left(n \pi+\sigma_{n}\right)} \approx \frac{1}{\varepsilon n \pi}
$$

The residue contribution from $\cosh \gamma+\varepsilon \gamma \sinh \gamma$ is given by

$$
\frac{1}{\cosh \gamma+\varepsilon \gamma \sinh \gamma} \quad \Longrightarrow \quad R=\frac{1}{\frac{d \gamma}{d p}(\sinh \gamma+\varepsilon \sinh \gamma+\varepsilon \gamma \cosh \gamma)}
$$

where $d \gamma / d p=\gamma / 2 p$ and $\varepsilon \gamma=-\cosh \gamma / \sinh \gamma$. Thus,

$$
R=\frac{2 p}{\gamma} \frac{\sinh \gamma}{\varepsilon \sinh ^{2} \gamma-1}=\frac{2 p}{\gamma^{2}} \frac{\gamma \sinh \gamma}{\varepsilon \sinh ^{2} \gamma-1}=\frac{2 p}{\gamma^{2}} \frac{\alpha_{n} \sin \alpha_{n}}{\varepsilon \sin ^{2} \alpha_{n}+1}
$$

For $\cot \alpha_{n}=\alpha_{n} \varepsilon$, and $\sin \alpha_{n}=1 / \sqrt{1+\cot ^{2} \alpha^{2}}=1 / \sqrt{1+\alpha_{n}^{2} \varepsilon^{2}}$,

$$
R_{n}=\frac{2}{\mu_{0} \mu D^{2} \sigma} \frac{\alpha_{n}}{\sqrt{1+\alpha_{n}^{2} \varepsilon^{2}}\left(1+\frac{\varepsilon}{1+\alpha_{n}^{2} \varepsilon^{2}}\right)}=\frac{2}{\mu_{0} \mu D^{2} \sigma} \frac{\alpha_{n} \sqrt{1+\alpha_{n}^{2} \varepsilon^{2}}}{1+\alpha_{n}^{2} \varepsilon^{2}+\varepsilon}
$$

For $\alpha_{0}^{2}=1 / \varepsilon$,

$$
R_{0}=\frac{2}{\mu_{0} \mu \sigma D^{2}} \frac{1}{\sqrt{\varepsilon}} \frac{\sqrt{1+\varepsilon}}{1+2 \varepsilon}=\frac{1}{\mu_{0} \mu \sigma D^{2} \varepsilon} \frac{\sqrt{1+1 / \varepsilon}}{1+1 / 2 \varepsilon}
$$

## Rogowski Dipole



Figure 1.

$$
\begin{gathered}
B=B_{0} \cos \alpha \cdot \frac{e^{i \alpha}}{i} \text { and } B^{*}=i B_{0} \cos \alpha \cdot e^{-i \alpha}, \\
\frac{i B_{0}}{B^{*}}=G=\frac{e^{i \alpha}}{\cos \alpha}=1+i \tan \alpha \cdot \text { with } G=\frac{1}{F^{\prime}}=\frac{\dot{z}}{\dot{F}} .
\end{gathered}
$$

On pole: $\Re G=0$. On $0,1, \infty: \Im G=0$.



Figures $2(\mathrm{a}, \mathrm{b})$.

$$
\begin{gathered}
\dot{G}=\frac{a / 2}{\sqrt{t}} \text { and } G=1+a \sqrt{t} \\
\dot{F}=\frac{b / 2}{\sqrt{t} \sqrt{t-1}} \text { and } F=b \ln (\sqrt{t}+\sqrt{t-1}) .
\end{gathered}
$$

Thus

$$
\dot{z}=\dot{F} G=\frac{b}{2}\left(\frac{1}{\sqrt{t} \sqrt{t-1}}+\frac{a}{\sqrt{t-1}}\right) \text { and } z=b(\ln (\sqrt{t}+\sqrt{t-1})+a \sqrt{t-1}) .
$$

February, 1981. Note 0397thry.

Since $i h=b(\ln i+i a)=i b(\pi / 2+a)$, we let

$$
\pi / 2+a=C \text { and } h=b C .
$$

On pole: $t=-\tau<0$,

$$
x+i y=b(i \pi / 2+\ln (\sqrt{\tau}+\sqrt{\tau+1})+i a \sqrt{\tau+1})
$$

with

$$
\tau=\sinh ^{2} \alpha, \quad x=b \alpha
$$

$$
b\left(\frac{\pi}{2}+a \cosh \alpha\right)=\frac{h}{C}\left(C+a\left(\cosh \frac{x}{b}-1\right)\right)=y=h\left(1+\frac{a}{c}\left(\cosh C \frac{x}{h}-1\right)\right)
$$

and

$$
G(1)=1+a=\frac{B(z=i \hbar)}{B(z=0)}
$$

## Rogowski Quadrupole: Formulation of Problem



$$
\mu=\infty, \quad \text { and } \quad B^{*}=+i B_{0}
$$

$$
B=B_{0} \cos \alpha \cdot \frac{e^{i \alpha}}{i}, \quad \text { and } \quad B^{*}=i B_{0} \cos \alpha \cdot e^{-i \alpha}
$$

Thus,

$$
\frac{i B_{0}}{B^{*}}=G=\frac{e^{i \alpha}}{\cos \alpha}=1+i \tan \alpha
$$

on the pole, with $\Re G=$ constant $=1$, and

$$
|G|=\frac{B_{0}}{|B|} .
$$

On the $45^{\circ}$ line,

$$
B \approx e^{i \pi / 4} / i, \quad B^{*} \approx i e^{i \pi / 4}, \quad \text { and } \quad G \approx \epsilon^{i \pi / 4}
$$

February, 1981. Note 0326thry.

Observe Figures 2(a,b,c):


Figures 2(a,b,c).

$$
\frac{d G}{d t}=b \frac{t-a}{t^{5 / 4} \sqrt{t+1}} e^{i \pi / 4}
$$

$$
B^{*}=i B_{0} \frac{d F}{d z}, \quad \frac{d z}{d F}=\frac{\dot{z}}{\dot{F}}=G \quad \text { and } \quad \dot{z}=G(t) \cdot \dot{F} .
$$

In the $F$-plane, for $-1 \leq t \leq 0$ :

$$
\pi \dot{F}=\frac{c:}{\sqrt{t} \sqrt{t+1}}
$$

The mathematical difficulty arises in the integration of $\dot{G}$. For

$$
t=w^{4}, \quad d t=4 w^{3} d w, \quad \frac{d t}{t^{5 / 4}}=\frac{4 w^{3} d w}{w^{5}}=\frac{4 d w}{w^{2}}
$$

we solve the elliptic integral

$$
G=b \int \frac{w^{4}-a}{w^{2} \sqrt{1+w^{4}}} d w
$$

The integration of $\dot{z}=(\dot{G} F)-\dot{G} F$ leads to

$$
z=G F-b e^{i \pi / 4} \int \frac{t-1}{t^{5 / 4} \sqrt{t+1}} F d t
$$

and therefore

$$
F=2 C \ln (\sqrt{t}+\sqrt{t+1})
$$

## Eddy Currents for Fast Permanent Magnet Magnetization

Sometimes permanent magnets are magnetized by "hitting" them for a short time with high $H$. It is of interest to know how the magnetization front propagates through the material. Since this is a highly non-linear problem, a strongly simplified model is used in this document, but one which has all the essential features of the real process.

At the left edge of the material, assume a step function excitation starting at $t=0$, with amplitude $H_{00}$. Our model is a 1-dimensional block of material with the left edge at $x=0$, and

$$
\frac{\partial}{\partial y}=\frac{\partial}{\partial z}=0, \quad \mathbf{H}=\mathrm{e}_{y} H, \quad \mathbf{E}=\mathrm{e}_{z} E, \quad \text { and } \quad \mathbf{j}=\sigma \mathbf{E}
$$

For the times of interest, $H \geq 0$ everywhere. We use a strongly simplified $B(H)$ curve.


Figure 1.

Figure 1 shows that in the beginning, the material "sees" no $H$ and $B=0$. As soon as it "sees" $H>0$, it becomes magnetized according to the above curve.

To reach an initial understanding of the problem, only the propagation in a medium that is, at least at first, unlimited to the right is treated.

$$
\begin{gathered}
\nabla \times \mathbf{H}=\mathbf{j} \quad \Longrightarrow \quad H^{\prime}=\sigma E, \quad \text { with } H^{\prime}=\frac{\partial}{\partial x} \\
\nabla \times \mathbf{E}=-\dot{\mathbf{B}} \quad \Longrightarrow \quad+E^{\prime}=+\dot{B}
\end{gathered}
$$

August, 1977. Note 0264thry.


Figure 2.
The "location" of the front is designated by $x_{0}(t)$. We integrate $E^{\prime}=\dot{B}$ over $x$ across $x_{0}(t)$ :

$$
E\left(x_{0}+\varepsilon\right)-E\left(x_{0}-\varepsilon\right)=-E\left(x_{0}\right)=\int \dot{B} d x=B_{0} \dot{x}_{0}
$$

giving the equation of the front for both Case I and Case II below.

Cāse I: $B(H)=B_{0}$.
For $x<x_{0}$ :

$$
\begin{gathered}
E^{\prime}=0, \quad \text { and } \quad E=-B_{0} \dot{x}_{0}=\varrho H^{\prime} . \\
H(x)=H_{00}-\sigma B_{0} x \dot{x}_{0},
\end{gathered}
$$

meaning that for this $B(H)$ model, the $H(x)$ curves of Figure 2 are straight lines.

$$
H\left(x_{0}\right)=0=H_{00}-\sigma B_{0} x_{0} \dot{x}_{0}
$$

Integrating over $t$ gives $H_{00} t-\sigma B_{0} x_{0}^{2} / 2$, giving the following result for the propagation of the front:

$$
\frac{2 t H_{00}}{B_{0} \sigma}=\frac{B_{00}}{B_{0}} \cdot \frac{2 t}{\mu_{0} \sigma}=x_{0}^{2}=r \frac{2 t}{\mu_{0} \sigma}, \quad \text { with } \quad r=\frac{\mu_{0} H_{00}}{B_{0}}=\frac{B_{00}}{B_{0}} .
$$

Case II: $B=B_{0}+\mu_{0} H$.
For $x<x_{0}: \quad H^{\prime}=\sigma E, \quad$ and $\quad E^{\prime}=\mu_{0} \dot{H}^{\dagger}$.
For $x=x_{0_{+}}: \quad H=0$.
$\dagger$ The right side of this last equation is now non-zero in contrast to Case I where $\dot{B}=0$ in the magnetized part of the material; i.e. $\dot{B} \neq 0$ only at the propagating front.

For $x=x_{0_{-}}: \quad B=B_{0}, \quad E=-B_{0} \dot{x}_{0}=\varrho H^{\prime}$.
For $x=0: \quad H=H_{00}$.
The differential equation is $H^{\prime \prime}=\mu_{0} \sigma \dot{H}$.
We introduce

$$
\sqrt{\frac{t}{\mu_{0} \sigma}}=\tau \text { and thus } t=\tau^{2} \mu_{0} \sigma
$$

and get

$$
\dot{H}=\frac{\partial H}{\partial \tau} \cdot \frac{d \tau}{d t}=\frac{\partial H}{\partial \tau} \cdot \frac{1}{2 \tau \mu_{0} \sigma}, \quad \text { and } \frac{\partial^{2} H}{\partial x^{2}}=\frac{\partial H}{\partial \tau} \cdot \frac{1}{2 \tau} .
$$

We use the dimensional analysis argument that $x$ and $\tau$ are the only dimensional quantities entering the problem. This means that $H$ must be a function of $x / \tau$. We let $u=x / 2 \tau$. For $H=F(u)$ :

$$
\begin{gathered}
\frac{\partial^{2} H}{\partial x^{2}}=\frac{F^{\prime \prime}}{4 \tau^{2}}=\frac{1}{2 \tau} \cdot F^{\prime} \cdot \frac{-u}{\tau}=\frac{-F^{\prime} u}{2 \tau^{2}}, \quad \text { and thus } F^{\prime \prime}+2 u F^{\prime}=0, \\
\ln \frac{F^{\prime}}{F_{0}^{\prime}}+u^{2}=0, \quad \text { and } \quad F^{\prime}=-a e^{-u^{2}},
\end{gathered}
$$

with $F_{0}^{\prime}=-a$ because $F^{\prime}$ has to be less than 0 .
The boundary conditions, with fixed $\tau>0$, are

$$
F=H(u)=H_{00}-a \int_{0}^{u} e^{-u^{2}} d u
$$

$u_{0}=x_{0} / 2 \tau$ and $x_{0}$ refer to the location of front.

$$
H\left(u_{0}\right)=H_{00}-a \int_{0}^{u_{0}} e^{-u^{2}} d u=0, \quad \text { with } \quad a=\frac{H_{00}}{\int_{0}^{u_{0}} e^{-u^{2}} d u}
$$

Just as in Case I, the location $x_{0}$ of the front is proportional to $\sqrt{t}$. Thus,

$$
\begin{aligned}
& \frac{\partial H\left(u_{0}\right)}{\partial x}=\frac{-a e^{-u_{0}^{2}}}{2 \tau}=\sigma E=-B_{0} \sigma \dot{x}_{0}=-B_{0} \sigma \cdot \frac{d x_{0}}{d \tau} \cdot \frac{1}{2 \tau \mu_{0} \sigma} \\
& a=e^{u_{0}^{2}} \cdot \frac{B_{0}}{\mu_{0}} \cdot \frac{d x_{0}}{d \tau} \text { and } \frac{B_{00}}{B_{0}}=r=\frac{d x_{0}}{d \tau} \cdot e^{-u_{0}^{2}} \cdot \int_{0}^{u_{0}} e^{-u^{2}} d u .
\end{aligned}
$$

Since $u_{0}=x_{0} / 2 \tau$, this is a first order differential equation that looks difficult at first.

However, it is obvious that from dimensional considerations the solution must be

$$
x_{0}=2 \tau g(r)
$$

where the factor of 2 is for neatness. That is, $u_{0}=g(r)$, and then $g(r)$ is determined by

$$
r=2 g e^{g^{2}} \int_{0}^{g} e^{-u^{2}} d u
$$

For small $g$ :

$$
r=2 g^{2} \quad \Longrightarrow \quad g=\sqrt{r / 2}, \quad \text { and } \quad x_{0}=\tau \sqrt{2 r}=\sqrt{r \cdot \frac{2 t}{\mu_{0} \sigma}}
$$

where $x_{0}$ has the same solution as the case of $B=B_{0}$, as it has to be.
Evaluation with a TI59 gives the following results:

| $g$ | 0.01 | 0.1 | 0.2 | 0.3 | 0.4 | 0.6 | 0.8 | 1.0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $r$ | $2.0 \times 10^{-4}$ | $2.01 \times 10^{-2}$ | $8.22 \times 10^{-2}$ | .191 | .356 | .920 | 2.00 | 4.06 |


| $g$ | 1.1 | 1.2 | 1.3 | 1.4 | 1.5 | 1.6 | 1.8 | 2.0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $r$ | 5.76 | 8.17 | 11.7 | 16.8 | 24.4 | 35.8 | 80.6 | 193 |

If $r$ is of order 4 , then

$$
\ddot{g} \approx 1 \Longrightarrow x_{0}=2 \tau=2 \sqrt{\frac{t}{\mu_{0} \sigma}} .
$$

Case III: $H$ after complete penetration of the slab.
When the front has reached $x_{0}=x_{1}$, where $2 x_{1}$ is the slab thickness, the boundary conditions change. In contrast to our earlier analysis, a given length $x_{1}$ enters the problem implicitly, and thus the dimensional analysis argument that $H$ must be equal to $F(x / 2 \tau)$ is no longer valid. If time is counted anew, with $t=0$ when $x_{0}=x_{1}$, we are dealing with a linear system with boundary condition

$$
H\left(x_{1}+\Delta x\right)=H\left(x_{1}-\Delta x\right) \quad \text { for } \quad t \geq 0
$$

with known and given $H(x)$ for $t=0$ and $0 \leq x \leq 2 x$.

If $H(x)-H_{00}$ is defined to be an odd function of $x$ with respect to $x=0$, and an even function with respect to $x= \pm x_{1}$, and this function is expanded into a Fourier series, then the period is $4 x_{1}$, and

$$
H(x)-H_{00}=\sum a_{n}(t) \sin \left(n \frac{2 \pi}{4 x_{1}} x\right)
$$

To satisfy these symmetry conditions $n$ must be odd, and we get

$$
H(x)=H_{00}+\sum a_{2 m+1}(t) \sin \left((2 m+1) \frac{\pi x}{2 x_{1}}\right)
$$

and $a_{2 m+1}(0)$ from known $H(x)$ at $t=0$, the time of complete penetration. Recalling the differential equation, $H^{\prime \prime}=\mu_{0} \sigma \dot{H}$, with $n=2 m+1$ we have

$$
\begin{gathered}
-a_{n}\left(\frac{n \pi}{2 x_{1}}\right)^{2}=\mu_{0} \sigma \dot{a}_{n} \\
a_{n}(t)=a_{n}(0) \cdot e^{-\left(\frac{n \pi}{2 x_{1}}\right)^{2} \frac{t}{\mu_{0} \sigma}}, \\
a_{2 m+1}(t)=a_{2 m+1}(0) \cdot e^{-(2 m+1)^{2} \frac{\pi^{2} t}{4 \pi_{1}^{2} \mu_{0} \sigma}}
\end{gathered}
$$

At $t=0:$

$$
\begin{gathered}
H(x)=H_{00}-H_{00} \frac{\int_{0}^{u} e^{-u^{2}} d u}{\int_{0}^{u_{0}} e^{-u^{2}} d u}=H_{00}\left(1-\frac{\int_{0}^{g \frac{x}{x_{1}}} e^{-u^{2}} d u}{\int_{0}^{g} e^{-u^{2}} d u}\right), \\
-H_{00} \frac{\int_{0}^{g \frac{x}{x_{1}}} e^{-u^{2}} d u}{\int_{0}^{g} e^{-u^{2}} d u}=\sum a_{2 m+1}(0) \sin \left((2 m+1) \frac{\pi x}{2 x_{1}}\right) .
\end{gathered}
$$

To determine $a_{2 m+1}(0)$ we let

$$
a_{n}(0) \cdot \int_{0}^{1} \sin ^{2}\left(n \frac{\pi}{2} v\right) d v=-\frac{H_{00}}{\int_{0}^{g} e^{-u^{2}} d u} \cdot \int_{0}^{1} \underbrace{\left(\int_{0}^{g v} e^{-u^{2}} d u\right)}_{\zeta} \underbrace{\sin \left(n \frac{\pi}{2} v\right) d v}_{d \eta}
$$

with

$$
\int_{0}^{1} \sin ^{2}\left(n \frac{\pi}{2} v\right) d v=\frac{1}{2} \int_{0}^{1}(1-\cos (n \pi v)) d v=\frac{1}{2}
$$

$$
\begin{gathered}
\zeta=\int_{0}^{g v} e^{-u^{2}} d u, \quad \text { and } \quad d \zeta=g e^{-g^{2} v^{2}}, \\
d \eta=\sin \left(n \frac{\pi}{2} v\right) d v, \text { and } \eta=-\frac{\cos \left(n \frac{\pi}{2} v\right)}{n \frac{\pi}{2} v}, \\
\left(\int_{0}^{1} \zeta d \eta\right)_{n=\mathrm{odd}}=\frac{2 g}{n \pi} \int_{0}^{1} e^{-g^{2} v^{2}} \cos \left(n \frac{\pi}{2} v\right) d v .
\end{gathered}
$$

Thus,

$$
a_{2 m+1}=-H_{00} \cdot \frac{4}{(2 m+1) \pi} \cdot \frac{\int_{0}^{1} e^{-g^{2} v^{2}} \cos \left((2 m+1) \frac{\pi}{2} v\right) d v}{\frac{1}{g} \int_{0}^{g} e^{-u^{2}} d u} .
$$

One would let

$$
\frac{1}{g} \int_{0}^{g} e^{-u^{2}} d u=\int_{0}^{1} e^{-g^{2} v^{2}} d v
$$

if one wants to evaluate it with the integral of the numerator. Otherwise, one may return to the definition of $g$ and use

$$
\frac{1}{g} \int_{0}^{g} e^{-u^{2}} d u=\frac{r e^{-g^{2}}}{2 g}
$$

The total time required to get good magnetization is determined as follows:
(1) Time to reach $x=x_{1}: \quad t_{1}=\mu_{0} \sigma\left(x_{1}^{2} / 4\right)$.
(2) Time for $a_{1}$ to decay by a factor of $e$ : $\quad t_{2}=\mu_{0} \sigma x_{1}^{2}\left(4 / \pi^{2}\right)$. Thus,

$$
\mu_{0} \sigma x_{1}^{2}\left(\frac{1}{4}+\frac{4}{\pi^{2}}\right)=t_{\text {total }} \approx .65 \mu_{0} \sigma x_{1}^{2}
$$

(3) If time for $a_{1}$ to decay by $e^{\pi}$ is used, $t_{2}=\mu_{0} \sigma x_{1}^{2}(4 / \pi)$, and

$$
\mu_{0} \sigma x_{1}^{2}(0.25+1.27)=t_{\text {total }} \approx 1.5 \mu_{0} \sigma x_{1}^{2} .
$$

## Change of Determinant for Small Changes of One Element of the

 Matrix that Describes a System that Is Least Squares Optimized with Restraints and Has Least Squares Limitations on ParametersWe let

$$
A=\left(\begin{array}{cc}
M^{t} W M+V & N^{t} \\
N & 0
\end{array}\right)
$$

and consider only those terms linear in $\Delta M_{n m}$ or $\Delta N_{n m}$.
1)

$$
\Delta M_{i k}=\Delta M_{n m} \delta(i-n) \delta(k-m) \quad \text { with } \quad \Delta M_{n m}=a
$$

Here and below we sum over indices appearing more than once.

$$
\begin{aligned}
\left(M^{t} W M\right)_{i k}= & M_{l i} W_{l l} M_{l k} \\
& \rightarrow\left(M_{l i}+a \delta(l-n) \delta(i-m)\right) W_{l l}\left(M_{l k}+a \delta(l-n) \delta(k-m)\right) \\
= & M_{l i} W_{l l} M_{l k}+a W_{n n}\left(M_{n i} \delta(k-m)+M_{n k} \delta(i-m) .\right)+a^{2}
\end{aligned}
$$

To get $\|A+\Delta A\|$ to first order in $a$, one must differentiate $\|A+\Delta A\|$ with respect to $a$ and then evaluate, knowing that $A^{-1}$ is Hermitian.

$$
\|A+\Delta A\|=\|A\|+a W_{n n} \cdot 2 M_{n k} K_{m k}
$$

where $K$ is the co-factor, and summation over $k$ is done only over values consistent with the number of rows in $M$.

$$
\frac{\|A+\Delta A\|}{\|A\|}=1+\frac{\Delta M_{n m}}{M_{n m}} \cdot 2 W_{n n} M_{n m} \cdot \sum_{k} M_{n k} A_{k m}^{-1}
$$

where $W_{n n} \sum_{k} M_{n k} A_{k m}^{-1}=\left(A^{-1} M^{t} W\right)_{m n}$.
2)

$$
\begin{gathered}
\Delta N_{i k}=\Delta N_{n m} \delta(i-n) \delta(k-m) \\
\|A+\Delta A\|=\|A\|+2 \Delta N_{n m} K_{n m} \\
\frac{\|A+\Delta A\|}{\|A\|}=1+\frac{\Delta N_{n m}}{N_{n m}} \cdot 2 V_{n m} A_{m n}^{-1}
\end{gathered}
$$

November, 1968. Note 0072thry.

## Sensitivity of Solution of Linear Equations to Change of an Individual Matrix Element

We let

$$
\begin{gathered}
M P=S, \quad M^{-1}=N \quad \text { and } P=N S \\
(M+\Delta M)(P+\Delta P)=M(I+A)(P+\Delta P)=S \quad \text { where } A=N \Delta M, \\
P+\Delta P=(I+A)^{-1} P \quad \text { and } \Delta P=\left((I+A)^{-1}-I\right) P .
\end{gathered}
$$

Further

$$
\begin{gathered}
\Delta M_{n m}=a \delta\left(n-n_{0}\right) \delta\left(m-m_{0}\right) \quad \text { where } \quad a=\Delta M_{n_{0} m_{0}}, \\
A_{k m}=N_{k n} \Delta M_{n m}=a N_{k n} \delta\left(n-n_{0}\right) \delta\left(m-m_{0}\right)=a N_{k n_{0}} \delta\left(m-m_{0}\right), \\
A_{k m}^{2}=A_{k i} A_{i m}=a^{2} N_{k n_{0}} \delta\left(i-m_{0}\right) N_{i n_{0}} \delta\left(m-m_{0}\right),
\end{gathered}
$$

where $A^{2}=a N_{m_{0} n_{0}} A$, and we let $\alpha=a N_{m_{0} n_{0}}$.

$$
(I+A)^{-1}=I+\gamma A \quad \text { and } \quad(I+A)(I+\gamma A)=I+A(I+\gamma+\gamma \alpha)=I,
$$

thus,

$$
\gamma=-\frac{1}{1+\Delta M_{n_{0} m_{0}} N_{m_{0} n_{0}}} \text { with } \Delta M_{n_{0} m_{0}} N_{m_{0} n_{0}} \neq-1
$$

$$
\begin{gathered}
\Delta P=\gamma A P \\
\Delta P_{k}=\gamma A_{k m} P_{m}=\gamma a N_{k n_{0}} \delta\left(m-m_{0}\right) P_{m}, \\
\Delta P_{k}=-\frac{\Delta M_{n_{0} m_{0}} P_{m_{0}}}{1+\Delta M_{n_{0} m_{0}} N_{m_{0} n_{0}}} \cdot N_{k n_{0}}
\end{gathered}
$$

That the matrix $M$ becomes exactly singular for $\Delta M_{n_{0} m_{0}}=-1 / N_{m_{0} n_{0}}$ is easily shown with Cramer's Rule. Let $K_{n m}$ be the co-factor to the $n m$ element, and $\|M+\Delta M\|=$ $\|M\|+\Delta M_{n_{0} m_{0}} K_{m_{0} n_{0}}:$

$$
\Delta M_{n_{0} m_{0}} \cdot \frac{K_{n_{0} m_{0}}}{\|M\|}=\Delta M_{n_{0} m_{0}} N_{m_{0} n_{0}}=-1
$$

which is the necessary condition for a singular matrix.

November, 1968. Note 0071thry:

This condition can easily be used to judge whether a matrix is "close" to being singular. One would test

$$
\frac{M_{n_{0} m_{0}}}{\Delta M_{n_{0} m_{0}}}=-M_{n_{0} m_{0}} N_{m_{0} n_{0}}
$$

and when the result is large compared to the inverse of the relative error of $M_{n_{0} m_{0}}$, one is likely to be in trouble. This is of particular importance when the matrix elements are experimentally determined.

## Fourier Analysis of Numerical Data

We assume that the spacing between data points is uniform, $2 \pi / N$. Representing $F(\varphi)$ by a Fourier series with unknown coefficients and making the coefficients such that

$$
\sum_{\varphi_{n}}\left(F\left(\varphi_{n}\right)-\sum_{m} a_{m} e^{i m \varphi_{n}}\right)^{2}=\min
$$

gives the same coefficients that one obtains by evaluating the integral

$$
\int F(\varphi) e^{-i m \varphi} d \varphi
$$

with trapezoidal rule applied to the whole integrand:

$$
a_{m}^{*}=\frac{1}{2 \pi} \int F(\varphi) e^{i m \varphi} d \varphi \Longrightarrow \frac{\Delta \varphi}{2 \pi} \sum F\left(\varphi_{n}\right) e^{i m \varphi_{n}}=\frac{1}{N} \sum F\left(\varphi_{n}\right) e^{i m \varphi_{n}}
$$

A better way to integrate would be to assume that not the whole integrand changes linearly over an individual interval, but that only $F(\varphi)$ changes linearly over the interval.
For one interval,

$$
\int F(\varphi) e^{i(m \varphi+\alpha)} d \varphi=\int(a+b \varphi) e^{i(m \varphi+\alpha)} d \varphi=(a+b \varphi) \frac{e^{i(m \varphi+\alpha)}}{i m}+\frac{b e^{i(m \varphi+\alpha)}}{m^{2}}
$$

When summing over the whole range of $\varphi$, the first term contributions cancel. With $b=\left(F_{2}-F_{1}\right) / \Delta \varphi$, we get

$$
\begin{aligned}
I & =\int_{\text {interval }} F(\varphi) e^{i(m \varphi+\alpha)} d \varphi \\
& =\frac{F_{2}-F_{1}}{m^{2} \Delta \varphi}\left(e^{i\left(m \varphi_{2}+\alpha\right)}-e^{i\left(m \varphi_{1}+\alpha\right)}\right) \\
& =\frac{F_{2} e^{i\left(m \varphi_{2}+\alpha\right)}\left(1-e^{-i m \Delta \varphi}\right)-F_{1} e^{i\left(m \varphi_{1}+\alpha\right)}\left(e^{i m \Delta \varphi}-1\right)}{m^{2} \Delta \varphi} .
\end{aligned}
$$

When summing over the whole circle, we get:

$$
\begin{aligned}
\int_{0}^{2 \pi} F(\varphi) e^{i(m \varphi+\alpha)} d \varphi & =\sum \frac{F\left(\varphi_{n}\right)\left(1-e^{-i m \Delta \varphi}-e^{i m \Delta \varphi}+1\right)}{m^{2} \Delta \varphi} e^{i\left(m \varphi_{n}+\alpha\right)} \\
& =\sum \frac{F\left(\varphi_{n}\right) 4 \sin ^{2} \varepsilon}{m^{2} \Delta \varphi} e^{i\left(m \varphi_{n}+\alpha\right)} \quad \text { with } \quad \varepsilon=\frac{m \Delta \varphi}{2} \\
& =\left(\frac{\sin \varepsilon}{\varepsilon}\right)^{2} \Delta \varphi \sum F\left(\varphi_{n}\right) e^{i\left(m \varphi_{n}+\alpha\right)}
\end{aligned}
$$

with

$$
\frac{m \Delta \varphi}{2}=\varepsilon=m \frac{\pi}{N}
$$

A parabolic approximation for $F$ over two intervals, $-\Delta \varphi \leq \varphi \leq \Delta \varphi$, without $e^{i \alpha}$, gives after some calculation:

$$
I_{2}=\left(\frac{\sin ^{4} \varepsilon}{\varepsilon^{2}}+\cos \varepsilon\left(\frac{\sin \varepsilon}{\varepsilon}\right)^{3}\right) \cdot \Delta \varphi \cdot \sum_{n} F\left(\varphi_{n}\right) e^{i\left(m \varphi_{n}+\alpha\right)}
$$

with $\varepsilon=m \pi / N$, and

$$
\left(\frac{\sin ^{4} \varepsilon}{\varepsilon^{2}}+\cos \varepsilon\left(\frac{\sin \varepsilon}{\varepsilon}\right)^{3}\right)=t=\left(\frac{\sin \varepsilon}{\varepsilon}\right)^{2} \cdot\left(\sin ^{2} \varepsilon+\cos \varepsilon \frac{\sin \varepsilon}{\varepsilon}\right)
$$

## Program to Calculate $t$, and Results.

5 CLS
10 FOR $N=1$ TO 18
$15 \mathrm{E}=\mathrm{F} * 3.14159265 / 36$
$20 \operatorname{PRINT}((\operatorname{SIN}(E) / E) \wedge 2 *((\operatorname{SIN}(E)) \wedge 2+\operatorname{COS}(E) * \operatorname{SIN}(E) / E)$
30 NEXT N
For $\varepsilon=N^{*} 5$ :

| $\varepsilon$ | $5^{\circ}$ | $10^{\circ}$ | $15^{\circ}$ | $20^{\circ}$ | $25^{\circ}$ | $30^{\circ}$ | $35^{\circ}$ | $40^{\circ}$ | $45^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $t$ | .9999 | .9998 | .9988 | .9962 | .9911 | .9821 | .9682 | .9482 | .9213 |


| $\varepsilon$ | $50^{\circ}$ | $55^{\circ}$ | $60^{\circ}$ | $65^{\circ}$ | $70^{\circ}$ | $75^{\circ}$ | $80^{\circ}$ | $85^{\circ}$ | $90^{\circ}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $t$ | .8870 | .8450 | .7957 | .7397 | .6780 | .6120 | .5434 | .4739 | .4053 |

Program to Calculate $K_{1} / \varepsilon^{2}$ and $K_{2} / \varepsilon^{2}$, and Results.

5 CLS
10 FOR N=1 TO 18
$20 \mathrm{E}=\mathrm{N} * 3.14159265 / 36$
$30 \operatorname{PRINT} N * 5,(3+\operatorname{COS}(4 * E)-\operatorname{SIN}(4 * E) / E) /\left(4 * E^{\sim} 2\right),(\operatorname{SIN}(2 * E) /(2 * E)-\operatorname{COS}(2 * E)) /\left(E^{-} 2\right)$ 40 NEXT N

| $\varepsilon$ | $K_{1} / \varepsilon^{2}$ | $K_{2} / \varepsilon^{2}$ |
| :---: | :---: | :---: |
| $5^{\circ}$ | . 67069299 | 1.3292762 |
| $10^{\circ}$ | . 68235368 | 1.3171576 |
| $15^{\circ}$ | . 70042915 | 1.2971353 |
| $20^{\circ}$ | . 72300072 | 1.2694693 |
| $25^{\circ}$ | . 74761149 | 1.2345172 |
| $30^{\circ}$ | . 77147163 | 1.1927287 |
| $35^{\circ}$ | . 79169091 | 1.1446375 |
| $40^{\circ}$ | . 80551841 | 1.0908528 |
| $45^{\circ}$ | . 81056947 | 1.0320491 |
| $50^{\circ}$ | . 8050208 | . 96895504 |
| $55^{\circ}$ | . 78775796 | . 90234147 |
| $60^{\circ}$ | . 75866353 | . 83300908 |
| $65^{\circ}$ | . 71763969 | . 76177546 |
| $70^{\circ}$ | . 6665645 | . 68946226 |
| $75^{\circ}$ | . 60718709 . | . 61688241 |
| $80^{\circ}$ | . 54197173 | . 54482766 |
| $85^{\circ}$ | . 47370526 | . 47405682 |
| $90^{\circ}$ | . 40528474 | . 40528474 |

## Curvature of Field Lines in a Quadrupole



Figure 1.

$$
F(z)=(z+r)^{2} \quad \text { with } \quad r=1 / K
$$

The field line is described by $\Re(z+r)^{2}=(x+r)^{2}-y^{2}=$ constant, thus

$$
(x+r)^{2}-y^{2}=\left(x_{0}+r\right)^{2} \quad \rightarrow \quad x=-r+\sqrt{y^{2}+\left(x_{0}+r\right)^{2}}
$$

Further, $1 / R=x^{\prime \prime} /\left(1+\left(x^{\prime}\right)^{2}\right)^{3 / 2}$, with

$$
\begin{gathered}
x^{\prime}=\frac{y}{\sqrt{y^{2}+\left(x_{0}+r\right)^{2}}} \text { and } x^{\prime \prime}=\frac{\left(x_{0}+r\right)^{2}}{\left(y^{2}+\left(x_{0}+r\right)^{2}\right)^{3 / 2}} \\
1+\left(x^{\prime}\right)^{2}=\frac{2 y^{2}+\left(x_{0}+r\right)^{2}}{\sqrt{y^{2}+\left(x_{0}+r\right)^{2}}}, \quad \text { and } \quad\left(1+\left(x^{\prime}\right)^{2}\right)^{3 / 2}=\frac{\left(2 y^{2}+\left(x_{0}+r\right)^{2}\right)^{3 / 2}}{\left(y^{2}+\left(x_{0}+r\right)^{2}\right)^{3 / 2}}
\end{gathered}
$$

Thus,

$$
\frac{1}{R}=\frac{\left(x_{0}+r\right)^{2}}{\left(2 y^{2}+\left(x_{0}+r\right)^{2}\right)^{3 / 2}} \text { and } R=\left(x_{0}+r\right)\left(1+\frac{2 y^{2}}{\left(x_{0}+r\right)^{2}}\right)^{3 / 2}
$$

for field line starting at $x_{0}$.

May, 1986. Note 0009thry.

The field line at $x, y$ is described by

$$
\frac{\left((x+r)^{2}+y^{2}\right)^{3 / 2}}{(x+r)^{2}-y^{2}}=R=(x+r) \frac{\left(1+\left(\frac{y}{x+r}\right)^{2}\right)^{3 / 2}}{1-\left(\frac{y}{x+r}\right)^{2}}
$$

We make the following substitutions:

$$
\frac{y}{x+r}=\tan \alpha, \quad 1+\tan ^{2} \alpha=\frac{1}{\cos ^{2} \alpha}, \quad \text { and } \quad 1-\tan ^{2} \alpha=\frac{\cos 2 \alpha}{\cos ^{2} \alpha}
$$

and thus,

$$
R=\frac{(x+r)}{\cos \alpha \cos 2 \alpha}
$$

Also,

$$
\sqrt{(x+r)^{2}+y^{2}}\left(\frac{1+\tan ^{2} \alpha}{1-\tan ^{2} \alpha}\right)=R=\frac{\sqrt{(x+r)^{2}+\dot{y}^{2}}}{\cos 2 \alpha}
$$

## Skin Effect in Fe



Figure 1.

We introduce initial conditions and definitions:

$$
\begin{gathered}
\mathbf{E}=\mathbf{e}_{x} E, \quad \text { and } \mathbf{B}=\mathrm{e}_{z} B \text { with } B=\mu_{0} \mu_{\mathrm{rel}} H=\mu H \\
\frac{\partial}{\partial z}=\frac{\partial}{\partial x}=0, \quad \text { and } \frac{\partial}{\partial y} \neq 0 \\
\nabla \times \mathbf{H}=H^{\prime}=\sigma E=j, \quad \text { and } \nabla \times \mathbf{E}=-E^{\prime}=-i \omega \mu H \\
\sigma E^{\prime}=j^{\prime}=i \omega \mu \sigma H
\end{gathered}
$$

We let $i \omega \mu \sigma=k^{2}$, and

$$
H^{\prime \prime}-k^{2} H=0, \quad H=H_{0} \cosh k y, \quad \text { and } \quad j=H_{0} k \sinh k y
$$

The average field in the sheet, $\bar{H}$, compared to the field outside, $H_{1}$, is given by

$$
\bar{H}=\frac{1}{2 y_{1}} \int_{-y_{1}}^{y_{1}} H d y=H_{0} \frac{\sin k y_{1}}{k y_{1}}, \quad H_{1}=H_{0} \cosh k y_{1}, \quad \text { and } \quad H_{0}=\frac{H_{1}}{\cosh k y_{1}}
$$

$$
\begin{equation*}
\frac{\bar{H}}{H_{1}}=\frac{\tanh k y_{1}}{k y_{1}} \tag{1}
\end{equation*}
$$

In (1), we let $x=k y_{1}$, and solve

February, 1966. Note 0007thry.

$$
\begin{align*}
\frac{\bar{H}}{H_{1}}=\frac{\tanh x}{x} & \approx \frac{1+x^{2} / 6+x^{4} / 120}{1+x^{2} / 2+x^{4} / 24} \\
& \approx\left(1+\frac{x^{2}}{6}+\frac{x^{4}}{120}\right)\left(1-\frac{x^{2}}{2}+\frac{5 x^{4}}{24}+\cdots\right) \\
& \approx 1-\frac{1 x^{2}}{3}+\frac{2 x^{4}}{15} . \tag{2}
\end{align*}
$$

The power dissipation per cubic meter is given by

$$
P=\frac{\varrho}{2}|j|^{2}=\frac{1}{2} \varrho H_{0}^{2}|k|^{2}\left|\sinh ^{2} k y\right|
$$

We let $k y=\alpha+i \alpha$ where $\alpha=|k| / \sqrt{2}$ thus

$$
\begin{aligned}
\sinh (\alpha & +i \alpha)=\sinh \alpha \cos \alpha+i \cosh \alpha \sin \alpha \\
|\sinh (\alpha+i \alpha)|^{2} & =\sinh ^{2} \alpha \cos ^{2} \alpha+\cosh ^{2} \alpha \sin ^{2} \alpha \\
& =\sinh ^{2} \alpha \cdot\left(1-\sin ^{2} \alpha\right)+\left(1-\sinh ^{2} \alpha\right) \cdot \sin ^{2} \alpha \\
& =\sinh ^{2} \alpha+\sin ^{2} \alpha \\
& =\frac{1}{2}(1-\cos 2 \alpha+\cosh 2 \alpha-1) \\
& =\frac{1}{2}(\cosh 2 \alpha-\cos 2 \alpha)
\end{aligned}
$$

and

$$
\begin{aligned}
\overline{|\sinh (\alpha+i \alpha)|^{2}} & =\frac{1}{2 \mathrm{y}_{1}} \int_{0}^{y_{1}}(\cosh 2 \alpha-\cos 2 \alpha) d y \\
& =\frac{1}{2 \alpha_{1}} \int_{0}^{\alpha_{1}}(\cosh 2 \alpha-\cos 2 \alpha) d \alpha \\
& =\frac{1}{4 \alpha_{1}}\left(\sinh 2 \alpha_{1}-\sin 2 \alpha_{1}\right)
\end{aligned}
$$

Thus,

$$
\bar{P}=\frac{H_{0}^{2} \mu}{2} \cdot \omega \cdot \frac{\sinh 2 \alpha_{1}-\sin 2 \alpha_{1}}{4 \alpha_{1}} \quad \text { with } \quad \alpha_{1}=y_{1} \sqrt{\frac{\omega \mu \sigma}{2}}
$$

For

$$
\frac{\sinh x-\sin x}{x} \approx \frac{2 \cdot\left(\frac{x^{3}}{3!}+\frac{x^{7}}{7!}\right)}{x}=\frac{x^{2}}{3}\left(1+\frac{x^{4}}{7!/ 3!}\right)=\frac{x^{2}}{3}\left(1+\frac{x^{4}}{840}\right) \approx \frac{x^{2}}{3}
$$

and thus,

$$
\frac{\sinh 2 \alpha_{1}-\sin 2 \alpha_{1}}{4 \alpha_{1}}=\frac{\left(2 \alpha_{1}\right)^{2}}{6}=\frac{\left(\sqrt{2} \gamma_{1}\right)^{2}}{6}=\frac{\gamma_{1}^{2}}{3}, \quad \text { with } \quad \gamma_{1}=y_{1} \sqrt{\omega \mu \sigma}=\sqrt{2} \alpha_{1}
$$

Therefore,

$$
\bar{P}=\frac{H_{0}^{2} \mu}{2} \cdot \omega \cdot \frac{y_{1}^{2} \omega \mu \sigma}{3}
$$

For $H_{0} \mu=B_{0}$,

$$
\bar{P}=\frac{B_{0}^{2}}{2} \cdot \frac{y_{1}^{2} \omega^{2} \sigma}{3}=\frac{B_{0}^{2}}{2} \cdot \frac{\left(2 y_{1}\right)^{2} \omega^{2} \sigma}{12}
$$

Resulting, thermally, in a trivial geometry:


Figure 2.

For heat conductivity, $S=\lambda T^{\prime}$ in power $/ \mathrm{m}^{2}$,

$$
\Delta S(x)=\bar{P} \Delta x, \quad \text { and thus } \quad \bar{P}=S^{t}=\lambda T^{\prime \prime}
$$

and thus,

$$
T_{\max }-T_{0}=\frac{x^{2} \bar{P}}{2 \lambda}
$$

## Typical Numbers for Dynamo Steel.

We let

$$
\begin{gathered}
\rho=46 \mu \Omega \mathrm{~cm}=4.6 \times 10^{-7} \Omega \mathrm{~m}, \\
\mu(14 \mathrm{kG})=2100, \quad \text { and } \mu(18 \mathrm{kG})=125, \\
\lambda=27-36 \frac{\mathrm{BTU} / \mathrm{h}}{\mathrm{ft}^{\circ} \mathrm{F}} \quad \text { with } \quad 1 \frac{\mathrm{BTU} / \mathrm{h}}{\mathrm{ft}^{\circ} \mathrm{F}}=\frac{0.293}{1.84} \frac{\mathrm{Watts}}{\mathrm{~m}^{\circ} \mathrm{C}} \\
=47.5-63.5 \mathrm{Watts} / \mathrm{m}^{\circ} \mathrm{C} .
\end{gathered}
$$

(1) For $\left|x^{2}\right| / 3=1 / 10$ we have

$$
\begin{gathered}
\frac{\left|x^{2}\right|}{3}=\frac{\omega \mu}{3 \varrho} \cdot y_{1}^{2}=\frac{1}{10}, \quad \text { and } \quad y_{1}=\sqrt{\frac{0.3 \varrho}{\omega \mu}}=10^{-3} \sqrt{\frac{2.3}{16.8}}=0.37 \times 10^{-3} \mathrm{~m}=0.37 \mathrm{~mm} \\
2 y_{1}=0.74 \mathrm{~mm} .
\end{gathered}
$$

(2) For $B_{0}=14 \mathrm{kG}=1.4 \mathrm{~T}$ and the above $2 y_{1}$,

$$
\bar{P}=\frac{60}{4.6} \times 10^{3}=13 \times 10^{3} \mathrm{Watts} / \mathrm{m}^{3}=13 \times 10^{-3} \mathrm{Watts} / \mathrm{cm}^{3}
$$

(3) For $x=0.4 \mathrm{~m}$ and the above $\bar{P}$,

$$
\Delta T=\frac{(.16)\left(13 \times 10^{3}\right)}{(2)(50)}=(13)(.16)^{\circ} \mathrm{C} \approx 21^{\circ} \mathrm{C}
$$

If the field is a sinusoidal function between $B=0 \mathrm{~T}$ and 14 kG , one has to use $B_{0}=$ 7 kG .

A More Detailed Expression for $\bar{H} / H_{1}$.
With $2 y_{1}=D$, we let

$$
x=\frac{k D}{2}=\frac{D}{2} \sqrt{i \omega \mu \sigma}=\frac{D}{2} \sqrt{2} \sqrt{i \frac{\omega \mu \sigma}{2}} .
$$

With $\lambda=\sqrt{2 / \omega \mu \sigma}$,

$$
x=\sqrt{i} \frac{D}{\sqrt{2} \lambda}=\sqrt{i} \varepsilon, \quad \text { where } \quad \varepsilon=\frac{D}{\sqrt{2} \lambda} .
$$

Therefore,

$$
\begin{gathered}
\frac{\bar{H}}{H_{1}}=1-i \frac{\varepsilon^{2}}{3}-\frac{2 \varepsilon^{4}}{15} \text { and } \tan \varphi \approx \frac{\varepsilon^{2}}{3}=\frac{(D / \lambda)^{2}}{6}, \\
\left|\frac{\bar{H}}{H_{1}}\right|^{2}=1-\frac{7 \varepsilon^{4}}{45}, \text { and }\left|\frac{\bar{H}}{H_{1}}\right|=1-\frac{7 \varepsilon^{4}}{90},
\end{gathered}
$$

and thus,

$$
\left|\frac{\bar{H}}{H_{1}}\right|=1-\frac{7}{(4)(90)} \frac{D^{4}}{\lambda}
$$

Results for $\mathrm{Al}, \mathrm{Cu}$ and Fe at $\mathbf{6 0 H z}$.
For Al and Cu , we let

$$
\begin{gathered}
\lambda=\sqrt{\frac{2 \varrho}{\sigma \mu_{0}}}=\sqrt{\frac{10^{4} \varrho}{2.4}} . \\
\varrho_{\mathrm{Al}}=2.8 \times 10^{-8} \quad \text { and } \quad \lambda_{\mathrm{Al}}=1.08 \mathrm{~cm}, \\
\varrho_{\mathrm{Cu}}=1.7 \times 10^{-8} \quad \text { and } \quad \lambda_{\mathrm{Cu}}=0.84 \mathrm{~cm},
\end{gathered}
$$

and $D=(1 / 4)$ in $=0.635 \mathrm{~cm}$ :

|  | $D / \lambda$ | $(D / \lambda)^{2}$ | $(D / \lambda)^{2} / 6$ | $0.7\left(D^{2} / 6 \lambda^{2}\right)^{2}$ |
| :---: | :---: | :---: | :---: | :---: |
| Cu | 0.755 | 0.570 | 0.095 | 0.00635 |
| Al | 0.587 | 0.345 | 0.0575 | .0023 |

For Fe with $\mu \approx 2000$,

$$
\varrho_{\mathrm{Fe}}=4.6 \times 10^{-7} \quad \text { and } \quad \lambda_{\mathrm{Fe}} \approx 1 \mathrm{~mm},
$$

and $D=14 \mathrm{~mm}$ :

|  | $D / \lambda$ | $(D / \lambda)^{2}$ | $(D / \lambda)^{2} / 6$ | $0.7\left(D^{2} / 6 \lambda^{2}\right)^{2}$ |
| :---: | :---: | :---: | :---: | :---: |
| Al | 0.350 | 0.1225 | 0.0205 | 0.00029 |

## Magnetic Field Energy Calculations

$$
E=\frac{1}{2} \int(\mathbf{B} \cdot \mathbf{H}) d \tau=\frac{1}{2} \int(\mathbf{B} \cdot \nabla V) d \tau=\frac{1}{2} \int \mathbf{H} \cdot(\nabla \times \mathbf{A}) d \tau
$$

with

$$
\begin{aligned}
& \mathbf{B} \cdot \nabla V=\nabla \cdot(V \mathbf{B})-V \nabla \cdot \mathbf{B}=\nabla \cdot(V \mathbf{B}) \\
& \mathbf{H} \cdot(\nabla \times \mathbf{A})=\mathbf{A} \cdot(\nabla \times \mathbf{H})=\mathbf{A} \cdot \mathbf{j}
\end{aligned}
$$

Field Energy in the Airspace of a Long, Symmetrical Bending Magnet.
The airspace is bounded by the midplane, an equipotential and two field lines (lines starting at two locations on the midplane).


Figure 1.

Derive $B$ from the potential: $\mathbf{B}=\nabla V$,

$$
2 \mu_{0} E=\int(\mathbf{B} \cdot \nabla V) d \tau=\int \nabla \cdot(V \mathbf{B}) d \tau=\int V \mathbf{B} \cdot d \sigma
$$

Normalize $V=0$ on equipotential, then contribution on equipotential is 0 , as well as being 0 along the field lines:

$$
2 \mu_{0} E=L V_{0} \int B_{y} d x
$$

where $L$ is the length of the magnet.

February, 1966. Note 0006thry.

For $B_{y}=B_{0}(1+K x)$ :

$$
\int_{-a}^{a} B_{y} d x=2 a B_{0}, \quad \text { and }\left.\quad V_{y}^{\prime}\right|_{x=0}=B_{y}=B_{0} \quad \Longrightarrow \quad V_{0}=y B_{0}
$$

Thus,

$$
2 \mu_{0} E=L B_{0} y_{0} 2 a B_{0} \quad \Longrightarrow \quad \frac{E}{L}=\frac{B_{0}^{2}}{2 \mu_{0}} 2 a y_{0} .
$$

$y_{0}$ equipotential is the hyperbola tangent to an ellipse with half-axis $a$ :

$$
y_{0}=\frac{c}{r_{0}}=\frac{b}{r_{0} a} \sqrt{\left(x+r_{0}\right)^{3} x}=b \frac{r_{0}}{a} \sqrt{\frac{x}{r_{0}}\left(\frac{x}{r_{0}}+1\right)^{3}} .
$$

For $a / r_{0}=\varepsilon$,

$$
s=\frac{x}{r_{0}}=\frac{2 \varepsilon^{2}}{1+\sqrt{1+8 \varepsilon^{2}}}=\frac{1}{4}\left(\sqrt{1+8 \varepsilon^{2}}-1\right)
$$

we redefine $y_{0}=b F(\varepsilon)$, where

$$
F(\varepsilon)=\sqrt{\frac{s(s+1)^{3}}{\varepsilon^{2}}}=\sqrt{\frac{2}{1+\sqrt{1+8 \varepsilon^{2}}}\left(\frac{3+\sqrt{1+8 \varepsilon^{2}}}{4}\right)^{3}} .
$$

For $8 \varepsilon^{2} \ll 1$ :

$$
\begin{gathered}
\sqrt{1+8 \varepsilon^{2}}=1+4 \varepsilon^{2} \\
F^{2}(\varepsilon)=\frac{2}{2+4 \varepsilon^{2}}\left(\frac{4+4 \varepsilon^{2}}{4}\right)^{3}=\left(1-2 \varepsilon^{2}\right)\left(1+3 \varepsilon^{2}\right) \\
F(\varepsilon)=1+\varepsilon^{2} / 2
\end{gathered}
$$

For $\varepsilon=1 / 2, F(1 / 2)=1.2$, while if $\varepsilon=1, F(1)=1.3$.

## Magnetic Energy of 2D Vacuum Field Inside Arbitrary Boundary.

Represent $\mathbf{B}$ by scalar potential: $\mathbf{B}=\nabla V$,

$$
2 \mu_{0} E=\int(\mathbf{B} \cdot \nabla V) d \tau=\int \nabla \cdot(V \mathbf{B}) d \tau=\int V(\mathbf{B} \cdot d \boldsymbol{\sigma})
$$

The expression for scalar product of two vectors in 2-dimensional space, when vectors are expressed by the complex numbers $a=a_{x}+i a_{y}$ and $b=b_{x}+i b_{y}$, is

$$
\mathbf{a} \cdot \mathbf{b}=a_{x} b_{x}+a_{y} b_{y}=\Re a b^{*}=\Re a^{*} b .
$$

Thus, for $d \sigma=i L d z$ :

$$
2 \mu_{0} E / L=\Re \int V i B^{*} d z
$$

For $V=v$, and $i B^{*}=F^{\prime}$ :

$$
\begin{aligned}
2 \mu_{0} E / L & =\Re \int v F^{\prime} d z=\Re \int v d F=\Re \int v(d u+i d v) \\
& =\int v d u=\int v\left(u_{x}^{\prime}+u_{y}^{\prime} y^{\prime}\right) d x .
\end{aligned}
$$

## Special Case.

The energy of field derived from $F=\left(B_{0} K / 2\right)\left(z+r_{0}\right)^{2}$, with $r_{0}=1 / K$, inside the ellipse described by $(x / a)^{2}+(y / b)^{2}=1$, is given by

$$
F=\frac{1}{2} B_{0} K\left(\left(x+r_{0}\right)^{2}-y^{2}+2 i y\left(x+r_{0}\right)\right) .
$$

With

$$
u_{x}^{\prime}=B_{0} K\left(x+r_{0}\right), \quad u_{y}^{\prime}=-B_{0} K y, \quad v=B_{0} K y\left(x+\dot{r}_{0}\right) ; \quad y^{\prime}=-\frac{b^{2}}{a^{2}} \frac{x}{y}
$$

and

$$
x=a \sin \varphi, \quad d x=a \cos \varphi d \varphi, \quad y=b \sqrt{1-(x / a)^{2}}=b \cos \varphi
$$

and $a / r_{0}=\varepsilon$,

$$
\begin{aligned}
2 \mu_{0} E / L & =B_{0}^{2} K^{2} \int y\left(x+r_{0}\right)\left(x+r_{0}+y \frac{b^{2} x}{a^{2} y}\right) d x \\
& =B_{0}^{2} K^{2} \int\left(x+r_{0}\right)\left(r_{0}+x\left(1+\frac{b^{2}}{a^{2}}\right)\right) y d x \\
& =B_{0}^{2} a b \int_{0}^{2 \pi}(1+\varepsilon \sin \varphi)\left(1+\varepsilon \sin \varphi\left(1+\frac{b^{2}}{a^{2}}\right)\right) \cos ^{2} \varphi d \varphi \\
& =B_{0}^{2} a b \int_{0}^{2 \pi}\left(1+\varepsilon \sin \varphi\left(2+\frac{b^{2}}{a^{2}}\right)+\varepsilon^{2} \sin ^{2} \varphi\left(1+\frac{b^{2}}{a^{2}}\right)\right) \cos ^{2} \varphi d \varphi \\
& =B_{0}^{2} a b\left(\pi+0+\varepsilon^{2} \frac{\pi}{4}\left(1+\frac{b^{2}}{a^{2}}\right)\right) \\
& =B_{0}^{2} a b \pi\left(1+\frac{a^{2}+b^{2}}{4 r_{0}^{2}}\right) .
\end{aligned}
$$

## Scalar Potential for 3D Fields in "Business Region" of Insertion Device with Finite Width Poles

Task.
In the absence of random errors, we are interested in the formulation of 3-dimensional $V(x, y, z)$ (with $\nabla^{2} V=0$ ) for $\mathbf{B}$ in the "business region" of an insertion device (hybrid or electro-magnetic) with finite width poles, and containing only a small number of free, easily measured constants.

## Notation and Coordinate System.

The beam will be in the direction of the $z$-axis. The midplane will be in the $x z$-plane. The field will be in the $y$-direction in the midplane.

Field Symmetries.
$B_{y}$ will be the even function of $x, y$ and $B_{x}$ will be the odd function of $x, y$.


Figure 1.
Representation of $V(x, y, z)$.
We represent $V(x, y, z)$ by a Fourier series in $z$ :

$$
\begin{equation*}
V(x, y, z)=\sum_{n=\mathrm{odd}} \cos n k_{3} z \cdot G_{n}(x, y) \quad \text { with } \quad k_{3}=2 \pi / \lambda \tag{1}
\end{equation*}
$$

where

$$
\begin{equation*}
\nabla^{2}=0 \quad \Longrightarrow \quad \nabla^{2} G_{n}=n^{2} k_{3}^{2} G_{n} \tag{2}
\end{equation*}
$$

February, 1992. Note 0144u-w.

For an infinitely wide pole:

$$
\partial G_{n} / \partial x=0 \quad \Longrightarrow \quad G_{n}(x, y)=a_{n} \sinh n k_{3} y
$$

which is a standard 2D solution. The effect of the finite width pole is described by

$$
G_{n}=a_{n} \sinh n k_{3} y+g_{n}(x, y),
$$

where $g_{n}$ is the effect of the finite width pole. Thus

$$
\begin{equation*}
\nabla^{2} g_{n}=n^{2} k_{3}^{2} g_{n} . \tag{3}
\end{equation*}
$$

Case 1:

$$
\begin{equation*}
\mu_{\mathrm{Fe}}=\infty \quad \Longrightarrow \quad B_{x}(x, \pm h, u \lambda / 2)=0 \tag{4}
\end{equation*}
$$

where $u$ is an integer.
1.1) We initially assume that

$$
G_{n}=g_{n}=0 \quad \text { for } \quad n>1
$$

and we see that the symmetries and (4) give $g_{1 x}^{\prime}(x, 0)=g_{1 x}^{\prime}(x, \pm h)=0$. We expand $g_{1 x}^{\prime}(x, y)$ in a Fourier series in $y$ with a $2 h$ period, and see that

$$
\begin{equation*}
g_{1 x}^{\prime}(x, y)=b_{m}^{\prime}(x) \sin m k_{2} y \quad \text { with } \quad k_{2}=\pi / h \tag{5}
\end{equation*}
$$

We now substitute (5) into (3). We have $b_{m}^{\prime \prime \prime}=b_{m}^{\prime}\left(m^{2} k_{2}^{2}+n^{2} k_{3}^{2}\right)$, with $n=1$. A solution which satisfies this equation and the symmetries is given by

$$
\begin{equation*}
b_{m}(x)=c_{m} \cosh k_{n m} x, \quad \text { where } \quad k_{n m}=\left(m^{2} k_{2}^{2}+n^{2} k_{3}^{2}\right)^{1 / 2} \tag{6}
\end{equation*}
$$

A complete solution is given by

$$
\begin{equation*}
V(x, y, z)=\cos k_{3} z \cdot a_{0}\left(\sinh k_{3} y+\sum_{m=1} \sin m k_{2} y \cosh k_{1 m} x \cdot \frac{a_{m}}{\cosh k_{1 m} w}\right) \tag{7}
\end{equation*}
$$

$c_{m}$ is chosen such that $a_{m}$ can be expected to be only weakly dependent on $w$.
1.2) We now assume that

$$
G_{n}=g_{n} \neq 0 \quad \text { for } \quad n \geq 1
$$

(7) can be generalized, but one must realize that the resulting formula does not have the same force of logic that was inherent in its original derivation. This generalized
formula allows the possibility that contributions from many $n$ could combine to make $B_{x}(x, \pm h, u \lambda / 2)=0$.

$$
\begin{align*}
V(x, y, z) & =\sum_{n=\mathrm{odd}} \cos n k_{3} z \cdot G_{n}(x, y) \\
G_{n}(x, y) & =A_{n 0}\left(\sinh n k_{3} y+\sum_{m=1} \sin m k_{2} y \cosh k_{n m} x \cdot \frac{a_{n m}}{\cosh k_{n m} w}\right)  \tag{8}\\
k_{n m} & =\left(n^{2} k_{3}^{2}+m^{2} k_{2}^{2}\right)^{1 / 2} \quad \text { with } \quad k_{3}=2 \pi / \lambda \quad \text { and } \quad k_{2}=\pi / h
\end{align*}
$$

Notice that in the above set of equations $A_{n 0}$ has units of Tesla-meters and $a_{n m}$ is dimensionless.
We expect that $a_{n m}$ is of the first order, the finite width effects decrease with increasing $n$ and $m$, and further, that only a few $a_{n m}$ are needed.

## Case 2:

$$
\begin{equation*}
\mu_{\mathrm{Fe}}<\infty \quad \Longrightarrow \quad B_{x}(x, \pm h, u \lambda / 2)=-B_{x}(-x, \pm h, u \lambda / 2) \neq 0 \tag{9}
\end{equation*}
$$

The contributions from Fe alone are given by the addition of $Q_{n}(x, y)$ :

$$
V=\sum \cos n k_{3} z \cdot Q_{n}(x, y) \quad \text { where } \quad \nabla^{2} Q_{n}=n^{2} k_{3}^{2} Q_{n}
$$

For the sake of simplification, we shall look at one $Q_{n}$, normalize lengths so that $n k_{3}=1$, and denormalize at the end

$$
\begin{equation*}
\nabla^{2} Q=Q \tag{10}
\end{equation*}
$$

We follow the logic of Case 1 as well as also satisfying $Q_{y}^{\prime}(x, 0) \approx x^{2}$ for sufficiently small $x$. Thus, we start with

$$
\begin{equation*}
Q(x, y)=(\cosh \eta x-1) P_{1}(y)+P_{2}(y) \tag{11}
\end{equation*}
$$

where $\eta$ is real and arbitrary. Later we will let $\eta \rightarrow 0$. We substitute (11) into (10) and get

$$
\begin{aligned}
\nabla^{2} Q & =\eta^{2} c P_{1}+(c-1) P_{1}^{\prime \prime}+P_{2}^{\prime \prime} \\
& =Q \\
& =(\dot{c}-1) P_{1}+P_{2}
\end{aligned}
$$

where $c=\cosh \eta x$ and $P_{1}, P_{2}$ are unknown. Separating into terms with and without
$c$, we have

$$
\begin{gather*}
\eta^{2} P_{1}+P_{1}^{\prime \prime}=P_{1}, \quad P_{2}^{\prime \prime}-P_{1}^{\prime \prime}=P_{2}-P_{1}  \tag{12}\\
p^{2}=1-\eta^{2}, \quad P_{1}^{\prime \prime}-p^{2} P_{1}=0 \Rightarrow P_{1}=\sinh p y  \tag{13}\\
P_{2}^{\prime \prime}-P_{2}=P_{1}^{\prime \prime}-P_{1}=-\eta^{2} P_{1}=-\eta^{2} \sinh p y \\
P_{2}=\sinh p y-p \sinh y \tag{14}
\end{gather*}
$$

in which the second term serves to satisfy the condition $P_{2 y}^{\prime}(0)=0$.

$$
\begin{equation*}
Q(x, y)=(\cosh \eta x-1) \sinh p y+\sinh p y-p \sinh y \tag{15}
\end{equation*}
$$

with $\eta \rightarrow 0$ and $p \rightarrow\left(1-\eta^{2} / 2\right)$.

$$
\begin{equation*}
Q(x, y)=\frac{\eta^{2}}{2}\left(x^{2} \sinh y-y \cosh y+\sinh y\right) \tag{16}
\end{equation*}
$$

We denormalize (16) and drop $\eta^{2} / 2$ and get

$$
\begin{gather*}
Q_{n}(x, y)=\left(n k_{3} x\right)^{2} \sinh n k_{3} y-n k_{3} y \cosh n k_{3} y+\sinh n k_{3} y  \tag{17}\\
V(x, y, z)=\sum_{n=\text { odd }} \cos n k_{3} z\left(G_{n}(x, y)+Q_{n}(x, y)\right) \tag{18}
\end{gather*}
$$

## Magnetic Measurements.

First, the simplest implementation consists of measuring the Fourier coefficients of the expansion of $B_{z}, B_{y}, B_{x}$ in $\sin n k z$ and $\cos n k z$ and determining the value of the free coefficients in $G_{n}$ and $Q_{n}$ that best fit the data. Use a "filter" to remove the random errors from the data sets.
Second, choose $x, y$ very carefully for each of these sets of measurements in order to take advantage of the properties of $G_{n}(x, y)$ and $Q_{n}(x, y)$ and its derivatives with respect to $x, y$. This is particularly important for the contributions originating from $\sin m k_{2} y$ in $G_{n}(x, y)$.
Third, investigate suitability of less conventional magnetic measurements, like a Hall probe or flux loop that vibrates in the $x$-direction, with phase sensitive de-modulation.

## Use of Model.

After verification of the validity region of the model is completed, it can be used for trajectory calculations. Furthermore, one can use this model to determine the maximal narrowness of the pole before detrimental effects become intolerable.
In application to existing hardware, one can break up the total field into the ideal 3D field and the random errors.

## Magnetic Measurement and Data Reduction to Identify Some Specific Error Field Consequences

## Measurement of Steering in Wigglers and Undulators.

Prefer null measurement method, if it can be done.
In "body" of wiggler or undulator: use coil with length equal to the product of period and integer.

In the end-region, from the field-free region to the periodic part: measure using long coil reaching from the outside to the periodic part, together with an attached compensation coil in the periodic part. This gives a signal that depends only on the steering integral, and is independent of position in the periodic part. It is an important tool for correcting the ends.
Normalized sensitivity of system, for $\varphi=k z=2 \pi z / \lambda$ is given by

$$
S(\varphi)=S_{0}(\varphi)+S_{1}(\varphi)
$$

where $S_{0}$ refers to the main coil, and $S_{1}$ to the compensation coil. With $\varphi=0$ referring to the end of the main coil, and $\varphi=-\alpha$ to the center of the correction coil (of length $2 \varphi_{1}$ ) relative to the end of the main coil, we have, in the coil coordinate system,

$$
\begin{aligned}
& S_{0}(\varphi)=1 \text { at }-\infty \leq \varphi \leq 0 \\
& S_{0}(\varphi)=0 \text { at } 0 \leq \varphi, \\
& S_{1}(\varphi)=\varepsilon \text { at }-\alpha-\varphi_{1} \leq \varphi \leq-\alpha+\varphi_{1} \\
& S_{1}(\varphi)=0 \text { at } \varphi \text { outside the above region. }
\end{aligned}
$$

For the periodic region, $\varphi>0$ and

$$
B=\sum_{n=\mathrm{odd}} n a_{n} \cos n \varphi=\Re \sum n a_{n} e^{i n \varphi}
$$

with the end of main coil at $\varphi_{0}>\pi$ in the field coordinate system, the signal from the main coil is

$$
F_{0}=\int_{-\infty}^{0} B d \varphi+\int_{0}^{\varphi_{0}} \Re \sum n a_{n} e^{i n \varphi} d \varphi=\text { Steering } \int+\Re \sum a_{n}\left(e^{i n \varphi_{0}}-1\right) / i
$$

and similarly, the signal from the compensating coil is

$$
F_{1}=\varepsilon \int_{\varphi_{0}-\alpha-\varphi_{1}}^{\varphi_{0}-\alpha+\varphi_{1}} \sum n a_{n} e^{i n \varphi} d \varphi=\Re \sum a_{n} e^{i n\left(\varphi_{0}-\alpha\right)} 2 \varepsilon \sin \varphi_{1}
$$

$$
F_{0}-F_{1}=\text { Steering } \int+\Re \sum a_{n} e^{i n \varphi_{0}}\left(\frac{1}{i}-2 \varepsilon e^{-i n \alpha} \sin n \varphi_{1}\right) .
$$

We want $F_{0}-F_{1}$ independent of $\varphi_{0}$, thus

$$
2 \varepsilon \sin n \alpha \sin n \varphi_{1}=1, \quad \text { and } \quad 2 \varepsilon \cos n \alpha \cos n \varphi_{1}=0
$$

When harmonics are weak (undulator), we need to satisfy these conditions only for $n=1$, but when strong harmonics are present (wiggler), we need to satisfy them for all odd $n$ : to get

$$
\begin{array}{ll|}
\cos n \alpha=0 & \text { choose } \alpha=\pi / 2 \\
\sin n \alpha=(-1)^{(n-1) / 2} & \text { choose }
\end{array}
$$

$\varphi_{1}=\pi / 2$ needs to be done by hardware, $\alpha=\pi / 2, \varepsilon=1 / 2$, can be done by "tuning" if one provides for it. Other solutions should be obvious.
This scheme can also be implemented with simple coil (or Hall probes) and software. However, software implementation is not a null method and therefore suffers much more from equipment imperfections.

## Phase Shifts of Emitted Light Due to Error Fields.

This is one of a number of ways to develop more insight into why or how synchrotron light properties deteriorate because of error fields.
We make the following definitions:

$$
\begin{gathered}
x^{\prime \prime}=\frac{g}{\gamma} B, \quad \text { with } \quad g=\frac{e}{m_{0}^{\prime} c} \\
\prime=\frac{\partial}{\partial z}, \quad \varphi=k z, \quad \text { and } \quad k=\frac{2 \pi}{\lambda} .
\end{gathered}
$$

For the reference trajectory:

$$
B(\varphi)=B_{0} \cos \varphi
$$

$$
\begin{gathered}
x_{0}^{\prime}=\frac{g B_{0}}{\gamma k} \sin \varphi=\frac{K}{\gamma} \sin \varphi, \quad \text { with } \frac{g B_{0}}{k}=K=.934 \cdot B_{0}(\mathrm{~T}) \cdot \lambda(\mathrm{cm}) \\
x_{0}=-\frac{K}{k \gamma} \cos \varphi \\
\cdot x_{W}=\frac{K}{k \gamma} .
\end{gathered}
$$

We define the trajectory length error as

$$
\begin{aligned}
\Delta s & =\int_{-\infty}\left(\sqrt{1+\left(x_{0}^{\prime}+\Delta x^{\prime}\right)^{2}}-\sqrt{1+{x_{0}^{\prime}}^{2}}\right) d z \\
& =\int\left(x_{0}^{\prime} \Delta x^{\prime}+\frac{1}{2} \Delta{x^{\prime}}^{2}\right) d z \\
& =x_{0} \Delta x^{\prime}-\int x_{0} \Delta x^{\prime \prime} d z+\frac{1}{2} \int \Delta{x^{\prime}}^{2} d z
\end{aligned}
$$

For $D=\Delta B / B_{0}$ as a function of $\varphi$ :

$$
\begin{gathered}
\Delta x^{\prime \prime}=\frac{g B_{0}}{\gamma} \frac{\Delta B}{B_{0}}=\frac{g B_{0}}{\gamma} D=\frac{K k}{\gamma} D \\
\Delta x^{\prime}=\frac{K}{\gamma} \int D d \varphi
\end{gathered}
$$

Thus,

$$
\Delta s=\left(\frac{K}{\gamma}\right)^{2} \frac{1}{k}\left(-\cos \varphi \int D d \varphi+\int \cos \varphi D d \varphi+\frac{1}{2} \int\left(\int \dot{D d \varphi}\right)^{2} d \varphi\right)
$$

With $\Delta t=\Delta s / c, \Delta \Phi=\omega_{L} \Delta t=\Delta s \omega_{L} / c=\Delta s k_{L}$, and

$$
\lambda_{L}=\frac{\lambda}{2 \gamma^{2}}\left(1+\frac{K^{2}}{2}\right)=\frac{\lambda K^{2}}{4 \gamma^{2}}\left(1+\frac{2}{K^{2}}\right)
$$

$$
\Delta \Phi=P(\underbrace{-\cos \varphi \int D d \varphi}_{\left(G_{1}\right)}+\underbrace{\int \cos \varphi D d \varphi}_{\left(G_{2}\right)}+\underbrace{\frac{1}{2} \int\left(\int D d \varphi\right)^{2} d \varphi}_{\left(G_{3}\right)})
$$

where

$$
P=\frac{4}{1+2 / K^{2}}
$$

Notice that, as a function of $K, \Delta \Phi \approx K^{2}$ for $K^{2} \ll 2$, and $\Delta \Phi$ is independent of $K^{2}$ for $K^{2} \gg 2 . G_{1}$ produces harmonics and reduces the intensity of the fundamental, but only if steering is not 0 . The $G_{2} \neq 0$ contribution depends on symmetry of $D$, not on the presence or absence of steering. In $G_{3}$ only steering errors contribute and it always gives $\Delta \Phi$ of the same sign. $G_{3}$ being of second order, we may expect a significant contribution only for a long undulator. We will see below that $G_{3}$ can be surprisingly large even for a short undulator.

We present order of magnitude estimates for $\overline{G_{2}^{2}}, \overline{G_{3}^{2}}$, for the ensemble.
For $G_{2}$ from one primary source:

$$
G_{2}=\varepsilon_{2} D_{2} \frac{\lambda}{2} \frac{2 \pi}{\lambda}=\varepsilon_{2} D_{2} \pi
$$

For $n_{2}$ sources per period: after $N_{1}=z / \lambda$ periods

$$
\overline{G_{2}^{2}}=\overline{\left(\varepsilon_{2} D_{2} \pi\right)^{2}} n_{2} \frac{z}{\lambda}=\overline{D_{2}^{2}} \pi^{2} \varepsilon_{2}^{2} n_{2} \frac{z}{\lambda}
$$

Without $\cos \varphi$ in the integrand, we may expect from steering:

$$
\begin{gathered}
\overline{G_{0}^{2}}=\overline{D_{0}^{2}} \pi^{2} \varepsilon_{0}^{2} n_{0} \frac{z}{\lambda}=\overline{D_{0}^{2}} \pi^{2} \varepsilon_{0}^{2} n_{0} N_{1} \\
\overline{G_{3}}=\frac{1}{2} \overline{D_{0}^{2}} \pi^{2} \varepsilon_{0}^{2} n_{0} \int_{0}^{L=N_{1} \lambda} \frac{z}{\lambda} d z \frac{2 \pi}{\lambda}=\frac{1}{2} \overline{D_{0}^{2}} \pi^{3} \varepsilon_{0}^{2} n_{0} N_{1}^{2} . \\
\varepsilon_{4}=\frac{\overline{G_{3}}}{\sqrt{G_{2}^{2}}}=\frac{\overline{D_{0}^{2}} \pi^{3} \varepsilon_{0}^{2} n_{0} N_{1}^{2} / 2}{\sqrt{\overline{D_{2}^{2}}} \pi \varepsilon_{2} \sqrt{n_{2}} \sqrt{N_{1}}}=\frac{\varepsilon_{0}^{2} n_{0} \overline{D_{0}^{2}}}{\sqrt{\varepsilon_{2} n_{2} \overline{D_{2}^{2}}}} \frac{\pi^{2}}{2} N_{1}^{3 / 2} .
\end{gathered}
$$

To get a feeling for the order of magnitude, we assume that the first order term, $\sqrt{\overline{G_{2}^{2}}}$, contributes twice the contribution of the second order term, $\overline{G_{3}}$, i.e. $\varepsilon_{4}=1 / 2$, and further, $\varepsilon_{0}=\varepsilon_{2}=1 n_{0}=n_{2}=4, N_{1}=10^{2} \overline{D_{0}^{2}}=\overline{D_{2}^{2}}$, then

$$
\sqrt{\overline{\overline{D_{0}^{2}}}}=\frac{1}{20 \times 10^{3}}=5 \times 10^{-5}
$$

has to be satisfied. And for $\sqrt{\overline{D_{0}^{2}}}=10^{-3}, \varepsilon_{0}=1, n_{0}=4, N_{1}=10^{2}$,

$$
\Delta \Phi=.6 \text { radians }
$$

$\Delta \Phi$ over-estimates the damage done to the emitted light because $\Delta \Phi=a+b z$ causes no real damage. We subtract the straight line from the original $\Delta \Phi(z)=f(z)$, and
then normalize the length of the undulator to 1 :

$$
H=\int_{0}^{1}(a+b z-f(z))^{2} d z
$$

and minimize $H$ with $a, b$,

$$
\int f(z) d z=F_{0}, \quad \int z f(z) d z=F_{1}, \quad \text { and } \quad \int f^{2}(z) d z=F_{2}
$$

The solution gives

$$
a=2\left(2 F_{0}-3 F_{1}\right), \quad \text { and } \quad b=6\left(2 F_{1}-F_{0}\right),
$$

and this gives

$$
H=F_{2}-4\left(F_{0}^{2}+3 F_{1}\left(F_{1}-F_{0}\right)\right)
$$

For a specific function $f(z)=u \sqrt{z}+v z^{2}$, corresponding to the ensemble model used above, and after optimization (see Appendix A for details),

$$
H=\left(\frac{1}{180}\right) \frac{2 u^{2}}{5}-\frac{8 u v}{7}+v^{2}
$$

With $\alpha=v / u$, we get the following improvement factor:

$$
\frac{H}{F_{2}}=\left(\frac{1}{36}\right) \frac{\alpha^{2}-8 \alpha / 7+2 / 5}{\alpha^{2}+20 \alpha / 7+5 / 2}
$$



Figure 1.

For Figure 1,

$$
\left(H / F_{2}\right)_{\min }=4.5 \times 10^{-4} \quad \text { at } \alpha=.605
$$

$$
\left(H / F_{2}\right)_{\max }=.274 \quad \text { at } \alpha=-1.65,
$$

Even if one does not consider the model of $f(z)=u \sqrt{z}+v z^{2}$ a realistic one, it is quite clear that (a) one should optimize not $\Delta \Phi(z)$, but $\Delta \Phi(z)$ minus "best" straight line, and that (b) gains can be remarkable. In other words, the quality of the light generated may be much better than one would think if one were to only look at $\Delta \Phi(z)$ or $D(z)$. We make a trivial, but interesting observation: since $H$.(before or after subtraction of straight line) is a quadratic function of $u, v$, an increase in the $G_{2}$ or $G_{3}$ contribution may lead to a decrease of $H$.

For the measurement of $G_{2}=\int \cos \varphi D d \varphi$, consider the following 2-coil configuration:


Figure 2.

The above design will measure the integral over $\cos \varphi \cdot B(\varphi)$. But, $\cos \varphi \cdot B_{0} \cos \varphi$ gives a large signal, therefore the null-coil system is needed. The proposed system cancels $B_{0} \cos \varphi$ but also "sees" steering; thus, it is fine only if steering is small enough or known. Therefore, we give below only the basic design and performance equations for system components, and one system.
Since $\Delta \Phi$ is only relevant for undulator, we ignore the harmonics.
The compensation coil is the same for measurement of steering at ends.
The main coil sensitivity is $S_{0}=\cos \left(\alpha_{0}+\varphi\right)$ at $-\varphi_{2} \leq \varphi \leq \varphi_{2}$, and $S_{0}=0$ outside this region. At the center of the coil, sensitivity is $\cos \alpha_{0}$, and $B=\cos \varphi_{0}=\Re e^{i \varphi_{0}}=B$.

$$
\begin{aligned}
F_{0} & =\Re \int_{-\varphi_{2}}^{\varphi_{2}} e^{i\left(\varphi_{0}+\varphi\right)} \cos \left(\alpha_{0}+\varphi\right) d \varphi \\
& =\frac{1}{2} \Re e^{i \varphi_{0}} \int_{-\varphi_{2}}^{\varphi_{2}}\left(e^{i \alpha_{0}} e^{2 i \varphi}+e^{-i \alpha_{0}}\right) d \varphi \\
& =\frac{1}{2} \Re e^{i \varphi_{0}}\left(e^{i \alpha_{0}} \sin 2 \varphi_{2}+e^{-i \alpha_{0}} 2 \varphi_{2}\right) .
\end{aligned}
$$

The compensation coil sensitivity is $S_{1}(\varphi)=\varepsilon_{1}$ at $-\varphi_{1} \leq \varphi \leq \varphi_{1}$, and $S_{1}(\varphi)=0$ outside this region. At the center of the coil, $B=\cos \left(\varphi_{0}+\beta\right)=\Re e^{i\left(\varphi_{0}+\beta\right)}$.

$$
\begin{aligned}
F_{1} & =\dot{\Re \varepsilon} \int_{-\varphi_{1}}^{\varphi_{1}} e^{i\left(\varphi_{0}+\beta+\alpha\right)} d \varphi \\
& =2 \varepsilon_{1} \Re e^{i\left(\varphi_{0}+\beta\right)}-\sin \varphi_{1} .
\end{aligned}
$$

With $\varepsilon=4 \varepsilon_{1} \sin \varphi_{1}$ and $2 \varphi_{2}=\gamma$,

$$
F_{0}+F_{1}=\frac{1}{2} \Re e^{i \varphi_{0}}\left(e^{i \alpha_{0}} \sin \gamma+e^{-i \alpha_{0}} \gamma+\varepsilon e^{i \beta}\right) .
$$

To get no signal for all $\varphi_{0}$, we must satisfy

$$
\cos \alpha_{0}(\sin \gamma+\gamma)=-\varepsilon \cos \beta, \quad \text { and } \quad \sin \alpha_{0}(\sin \gamma-\gamma)=-\varepsilon \sin \beta,
$$

and since there are four parameters to satisfy two equations there are many possible solutions. We pick one with $\beta=0$ and $\alpha_{0}=0$ and have

$$
\varepsilon=4 \varepsilon_{1} \sin \varphi_{1}=-(\gamma+\sin \gamma), \quad \varepsilon_{1}=-\frac{2 \varphi_{2}+\sin 2 \varphi_{2}}{4 \sin \varphi_{1}} .
$$

If $\left|\varepsilon_{1}\right|>1$, we can use a combined coil system as follows, with $\varphi_{2}=\varphi_{1}=3 \pi / 2$ and therefore $\varepsilon_{1}=3 \pi / 4$.


Figure 2.

This is not the ultimate answer, but only a first step, and it may start similar thinking on other issues.

## Appendix A.

For the execution of the optimization of $H$, we let

$$
\begin{gathered}
a+b / 2=F_{0}, \quad 2 F_{1}-F_{0}=b / 3, \quad a / 2+b / 3=F_{1}, \\
b=6\left(2 F_{1}-F_{0}\right) \quad a=2\left(2 F_{0}-3 F_{1}\right) .
\end{gathered}
$$

With the above, we have

$$
\begin{aligned}
H & =a^{2}+b^{2} / 3+F_{2}-2 a F_{0}-2 b F_{1}+a b \\
& =a \underbrace{(a+b / 2)}_{F_{0}}+b \underbrace{(b / 3+a / 2)}_{F_{1}}+F_{2}-2 a F_{0}-2 b F_{1} \\
& =F_{2}-a F_{0}-b F_{1} \\
& =F_{2}-2 F_{0}\left(2 F_{0}-3 F_{1}\right)-6 F_{1}\left(2 F_{1}-F_{0}\right) \\
& =F_{2}-4 F_{0}^{2}-12 F_{1}^{2}+12 F_{0} F_{1} .
\end{aligned}
$$

For the special case of $f(z)=\sqrt{z}+\alpha z^{2}$ :

$$
F_{0}=2 / 3+\alpha / 3=(\alpha+2) / 3, \quad \text { and } \quad F_{1}=2 / 5+\alpha / 4
$$

and for $f^{2}(z)=z+\alpha^{2} z^{4}+2 \alpha z^{2.5}, F_{2}=\frac{{ }_{2}}{}+\frac{\alpha^{2}}{5}+\frac{4 \alpha}{7}$. Therefore,

$$
\begin{aligned}
& F_{2}-H=\frac{4}{9}(\alpha+2)^{2}+12\left(\frac{\alpha}{4}+\frac{2}{5}\right)^{2}-4(\alpha+2)\left(\frac{\alpha}{4}+\frac{2}{5}\right) \\
&=\frac{7 \alpha^{2}}{36}+\frac{26 \alpha}{45}+\frac{112}{225}, \\
& H=\alpha^{2}\left(\frac{1}{5}-\frac{7}{36}\right)+\alpha\left(\frac{4}{7}-\frac{26}{45}\right)+\frac{1}{2}-\frac{112}{225} . \\
&=\frac{1}{180}\left(\alpha^{2}-\frac{8 \alpha}{7}+\frac{2}{5}\right)=\frac{1}{180}\left(v^{2}-\frac{8 u v}{7}+\frac{2 u^{2}}{5}\right) . \\
& \frac{H}{F_{2}}=\frac{1}{36} \frac{\alpha^{2}-8 \alpha / 7+2 / 5}{\alpha^{2}+20 \alpha / 7+5 / 2} .
\end{aligned}
$$

Least Square Fit of $f(z)$ with $a+b z$ in $0 \leq z \leq 1$

Origin and Purpose of Study. If $f(z)=$ phase shift, the difference between $f(z)$ and $a+b z$ is the only damaging property of $f(z)$ since $a$ would be an irrelevant shift of phase reference, and $b$ represents a shift of center of line without any broadening. We define

$$
S=\int(a+b z-f(z))^{2} d z
$$

For $S_{a}^{\prime}=0: \quad a+b / 2=F_{0}=\int f(z) d z$,
For $S_{b}^{\prime}=0: \quad a / 2+b / 3=F_{1}=\int z f(z) d z$.
Therefore,

$$
b=12 F_{1}-6 F_{0} \text { and } a=4 F_{0}-6 F_{1} .
$$

For $\int f(z)^{2} d z=F_{2}$,

$$
\begin{aligned}
S & =a^{2}+b^{2} / 3+F_{2}+a b-2 a F_{0}-2 b F_{1} \\
& =a(a+b / 2)+b(a / 2+b / 3)-2 a F_{0}-2 b F_{1}+F_{2} \\
& =F_{2}-a F_{0}-b F_{1} \\
& =F_{2}-F_{0}\left(4 F_{0}-6 F_{1}\right)-F_{1}\left(12 F_{1}-6 F_{0}\right) \\
& =F_{2}-4\left(F_{0}^{2}+3 F_{1}^{2}-3 F_{0} F_{1}\right) .
\end{aligned}
$$

For a specific function, $f(z)=\sqrt{z}+\alpha z^{2}$,

$$
\begin{gathered}
f(z)^{2}=z+2 \alpha z^{5 / 2}+\alpha^{2} z^{4} \\
F_{0}=\frac{2}{3}+\frac{\alpha}{3}, \quad F_{1}=\frac{2}{5}+\frac{\alpha}{4}, \quad \text { and } \quad F_{2}=\frac{1}{2}+\frac{4 \alpha}{7}+\frac{\alpha^{2}}{5}
\end{gathered}
$$

thus,

$$
\begin{aligned}
S & =\frac{\alpha^{2}}{5}+\frac{4 \alpha}{7}+\frac{1}{2}-4\left(\frac{1}{9}\left(\alpha^{2}+4 \alpha+4\right)+3\left(\frac{\alpha^{2}}{16}+\frac{\alpha}{5}+\frac{4}{25}\right)-(\alpha+2)\left(\frac{\alpha}{4}+\frac{2}{5}\right)\right) \\
& =\alpha^{2}\left(\frac{1}{180}\right)-\alpha\left(\frac{2}{315}\right)+\frac{1}{450} \\
& =\alpha^{2} a_{2}+\alpha a_{1}+a_{0}
\end{aligned}
$$

with

$$
a_{2}=\frac{1}{180}, \quad a_{1}=-\frac{2}{315}, \quad \text { and } \quad a_{0}=\frac{1}{450},
$$

therefore,

$$
\begin{aligned}
& S=\frac{1}{180}\left(\alpha^{2}-\frac{8 \alpha}{7}+\frac{2}{5}\right), \\
& \frac{S}{F_{2}}=\frac{1}{36} \cdot \frac{\alpha^{2}-\frac{8 \alpha}{7}+\frac{2}{5}}{\alpha^{2}+\frac{20 \alpha}{7}+\frac{5}{2}}
\end{aligned}
$$

For $\alpha \gg 1, \sqrt{S / F_{2}}$ is improved by a factor of 6 , and for $\alpha \ll 1, \sqrt{S / F_{2}}$ is improved by a factor of 15 .
$S / F_{2}$ is a strongly peaked function:
$\left(S / F_{2}\right)_{\max } \approx .274$ at $\alpha \approx-1.65$,
$S / F_{2} \approx .125$ at $\alpha \approx-3$,
$S / F_{2} \approx .004$ at $\alpha \approx 0$.
$\left(S / F_{2}\right)_{\text {min }} \approx 4.5 \times 10^{-4}$ at $\alpha \approx .605$,
Since $a+b z$ represents the error-free condition, looking at the deviation of phase shift from the straight line may represent the best way to characterize the consequences of the error fields.

$$
S_{\min }=\frac{1}{180}\left(\frac{2}{5}-\frac{16}{49}\right)=\frac{1}{50 \cdot 49} \quad \text { with } \quad \alpha=\frac{4}{7}
$$

and the values for $a, b$ are

$$
\begin{gathered}
a=4 F_{0}-6 F_{1}=\frac{4}{3}(\alpha+2)-6\left(\frac{\alpha}{4}+\frac{2}{5}\right)=-\frac{\alpha}{6}+\frac{4}{15} \\
b=12 F_{1}-6 F_{0}=3\left(\alpha+\frac{8}{5}\right)-2(\alpha-2)=\alpha+\frac{4}{5}
\end{gathered}
$$

At the center

$$
\begin{aligned}
\Delta B & =b(0) \\
& =\frac{V_{1}}{h} \frac{1}{k_{0}\left(x_{2}-x_{1}\right)} \ln \frac{\cosh k_{0} x_{2}+1}{\cosh k_{0} x_{1}+1} \\
& =\frac{V_{1}}{h} \cdot \frac{\left(\ln \frac{\sinh k_{0} x_{2} / 2}{\sinh k_{0} x_{1} / 2}\right)}{\left(\frac{k_{0}\left(x_{2}-x_{1}\right)}{2}\right)},
\end{aligned}
$$

with $k_{0}=\pi / h$ and $k_{1}=2 \pi / \lambda$. Further,

$$
\begin{aligned}
\int b(z) \cos \left(k_{1} z\right) k_{1} d z & =\frac{V_{1}}{h} \cdot \frac{2}{\pi} \cdot k_{1} \cdot \frac{2 \pi}{k_{0}} \cdot \int_{x_{1}}^{x_{2}} \frac{\sin k_{1} x}{\sinh \left(\pi k_{1} / k_{0}\right)} \cdot \frac{d x}{x_{2}-x_{1}} \\
& =\frac{V_{1}}{h} \cdot \underbrace{\frac{4 k_{1}}{k_{0}} \cdot \frac{\sin \left(k_{1} \Delta x / 2\right) /\left(k_{1} \Delta x / 2\right)}{\sinh \left(\pi k_{1} / k_{0}\right)}}_{g_{1}}
\end{aligned}
$$

$$
\frac{V_{1}}{h}=\frac{\Delta B}{g_{0}}
$$

$$
\int b(z) \cos \left(k_{1} z\right) k_{1} d z=\frac{\Delta B}{g_{0}} g_{1}
$$

$$
\int b d z=\frac{V_{1}}{h} \cdot \frac{\lambda}{2}=\frac{\Delta B \cdot \lambda / 2}{g_{0}}
$$

Thus,

$$
\int \frac{\Delta B}{B_{0}} k_{1} d z=\varepsilon_{2} \pi \frac{\Delta B}{B_{0}}=\frac{\Delta B}{B_{0}} \frac{\lambda / 2}{g_{0}} \frac{2 \pi}{\lambda},
$$

$$
\varepsilon_{2}=\frac{1}{g_{0}}
$$

Similarly,

$$
\begin{gathered}
\frac{1}{B_{0}} \int b \cos \left(k_{1} z\right) k_{1} d z=\frac{\Delta B}{B_{0}} \frac{g_{1}}{g_{0}}=\varepsilon_{1} \pi \frac{\Delta B}{B_{0}} \\
\varepsilon_{1}=\frac{g_{1}}{\pi g_{0}} .
\end{gathered}
$$

Thus,

$$
\frac{\varepsilon_{1}}{\varepsilon_{2}^{2}}=\frac{g_{0} g_{1}}{\pi},
$$

and

$$
\frac{g_{0} g_{1}}{\pi}=\frac{4}{\pi} \cdot \ln \left(\frac{\sinh k_{0} x_{2} / 2}{\sinh k_{0} x_{1} / 2}\right) \cdot \frac{\sin \left(k_{1} \Delta x / 2\right)}{\sinh \left(\pi k_{1} / k_{0}\right)} \cdot \frac{\left(k_{1} / k_{0}\right)^{2}}{\left(k_{1} \Delta x / 2\right)^{2}}
$$



Figure 1.

For the above figure,

$$
\begin{gathered}
x_{1}+\frac{1}{2}\left(x_{2}-x_{1}\right)=\frac{\lambda}{4}=\frac{x_{1}+x_{2}}{2}, \\
\frac{\lambda / 4}{x_{1}}=3, \quad x_{1}=\frac{\lambda}{12}, \\
x_{2}=\frac{\lambda}{2}-x_{1}=\lambda \frac{5}{12} .
\end{gathered}
$$

Further,

$$
\begin{gathered}
k_{1} \frac{\Delta x}{2}=\frac{2 \pi}{\lambda} \cdot \frac{1}{2} \cdot \frac{4 \lambda}{12}=\frac{\pi}{3}, \\
k_{0} \frac{x_{2}}{2}=\frac{\pi}{h} \cdot \frac{5 \lambda}{24}=\frac{\lambda}{h} \cdot \frac{5 \pi}{24}, \\
k_{0} \frac{x_{1}}{2}=\frac{\pi}{h} \cdot \frac{\lambda}{24}=\frac{\lambda}{h} \cdot \frac{\pi}{24}, \\
\frac{k_{1}}{k_{0}}=\frac{2 h}{\lambda}, \quad \frac{\lambda}{h}=a .
\end{gathered}
$$

Thus,

$$
\frac{g_{0} g_{1}}{\pi}=\frac{36 \cdot 4}{\pi^{3}} \cdot \frac{\sqrt{3}}{2} \cdot \ln \left(\frac{e^{a 5 \pi / 12}-1}{e^{a \pi / 12}-1}\right) \cdot \frac{1}{a^{2}} \cdot \frac{1}{\sinh (2 \pi / a)}
$$

where for $\frac{2 h}{\lambda}=b=\frac{2}{a}$, and $a=\frac{2}{b}$,

$$
\frac{g_{0} g_{1}}{\pi}=\frac{36 \sqrt{3}}{\pi^{3}} \cdot b^{2} \cdot \ln \left(\frac{e^{5 \pi / 6 b}-1}{e^{\pi / 6 b}-1}\right) \cdot \frac{1}{2 \sinh (\pi b)}
$$

## Comparison of First and Second Order Contributions of Error Fields to Phase Shift

We introduce, for $k z=\varphi, k d z=d \varphi$, and $\frac{\Delta B}{B_{0}}=D$,

$$
\Delta \Phi=P(\underbrace{-\cos \varphi \int D d \varphi}_{\left(A_{1}\right)}+\underbrace{\int \cos \varphi D d \varphi}_{\left(B_{1}\right)}+\underbrace{\frac{1}{2} \int\left(\int D d \varphi\right)^{2} d \varphi}_{\left(B_{2}\right)})
$$

where

$$
P=\frac{4}{1+2 / K^{2}}
$$

Notice that $\Delta \Phi \approx K^{2}$ for $K^{2} \ll 2$, and $\Delta \Phi$ is independent of $K^{2}$ for $K^{2} \gg 2$.
We denote the "typical" case of $B_{1}$ as

$$
B_{1}=\varepsilon_{1} D \frac{\lambda}{2} \frac{2 \pi}{\lambda}=\varepsilon_{1} \pi D
$$

At every error source, $B_{1}$ changes by the above "typical" value of $B_{1}{ }^{\ddagger}$. We assume $n$ contributions per period. After $N=z / \lambda$ periods, the total expectation value is

$$
\overline{B_{1}^{2}}=n \frac{z}{\lambda} \overline{\left(\varepsilon_{1} \pi D\right)^{2}}
$$

When

$$
B_{1}=\int D d \varphi
$$

we expect

$$
\overline{B_{1}^{2}}=n \frac{z}{\lambda} \overline{\left(\varepsilon_{2} \pi D\right)^{2}}
$$

At the end of insertion device with $N$ periods we expect

$$
\left\langle B_{1}^{2}\right\rangle=\varepsilon_{1}^{2} \pi^{2} \overline{D^{2}} n N
$$

[^7]\[

$$
\begin{aligned}
\left\langle B_{2}\right\rangle & =\frac{1}{2} \varepsilon_{2}^{2} \pi^{2} \overline{D^{2}} n \frac{(N \lambda)^{2}}{2 \lambda} \frac{2 \pi}{\lambda} \\
& =\frac{1}{2} \varepsilon_{2}^{2} \pi^{3} \overline{D^{2}} n N^{2} \\
& =\varepsilon_{3} \sqrt{\left\langle B_{1}^{2}\right\rangle} \quad \text { with } \quad \varepsilon_{3}<1
\end{aligned}
$$
\]

This means that

$$
\begin{gathered}
\frac{1}{2} \varepsilon_{2}^{2} \pi^{3} \overline{D^{2}} n N^{2}=\varepsilon_{3} \varepsilon_{1} \pi \sqrt{\overline{D^{2}}} \sqrt{n} \sqrt{N} \\
\sqrt{\overline{D^{2}}}=\frac{2 \varepsilon_{3} \varepsilon_{1} / \varepsilon_{2}^{2}}{\pi^{2} \sqrt{n}(\sqrt{N})^{3}}
\end{gathered}
$$

We make the following definitions:

$$
\varepsilon_{1}=\varepsilon_{2}=1, \quad 2 \varepsilon_{3}=1, \quad n=4 \quad \text { and } \quad N=81
$$

Thus,

$$
\sqrt{\overline{D^{2}}} \approx \frac{1}{20 \times 10^{3}} \approx 5 \times 10^{5} .
$$

This means that the second order contributions will dominate. Or, similarly

$$
\varepsilon_{3}=\frac{\pi^{2}}{2} \frac{\varepsilon_{2}^{2}}{\varepsilon_{1}} \sqrt{\overline{D^{2}}} \sqrt{n}(\sqrt{N})^{3},
$$

and for $\sqrt{\overline{D^{2}}}=10^{-3}$,

$$
\varepsilon_{3} \approx 5 \times 2 \times 10^{-3+} \times 10^{3} \approx 10
$$

still demonstrating that second order contributions will dominate.
The magnitude of the $\Delta \Phi$ shift along the length of the insertion device with second order contributions is

$$
\begin{aligned}
\Delta \Phi & =4 \frac{1}{2} \varepsilon_{2}^{2} \pi^{3} \overline{D^{2}} n N^{2} \\
& \approx 2 \times 30 \times 10^{-6} \times 4 \times 6.5 \times 10^{3} \\
& \approx 30 \times 50 \times 10^{-3} \\
& \approx 1.5 \text { radians. }
\end{aligned}
$$

Thus,

$$
\frac{\varepsilon_{1}}{\varepsilon_{2}^{2}} \approx .5 \rightarrow D
$$

for equal contributions (i.e.: $\varepsilon_{3}=1$ ), and for

$$
D \approx 5 \times 10^{-5} \rightarrow D \approx 5 \times 10^{-3} \rightarrow \alpha=100
$$

for representation of phase shift by straight line.

## Connection Between Undulator Field Errors and Optical Phase

We begin with the following definitions:

$$
x^{\prime \prime}=\frac{g}{\gamma} B \quad \text { and } \quad g=\frac{e}{m_{0} c}
$$

We introduce the following references:

$$
\begin{gathered}
B(z)=B_{0} \cos k z \\
x_{0}^{\prime}=\frac{g B_{0}}{\gamma k} \sin k z=\frac{K}{\gamma} \sin k z \quad \text { with } \quad \frac{g B_{0}}{k}=K \\
x_{0}=-\frac{K}{k \gamma} \cos k z \\
x_{W}=\frac{K}{k \gamma}
\end{gathered}
$$

We now proceed with the analysis.

$$
\Delta s=\int\left(\sqrt{1+\left(x_{0}^{\prime}+\Delta x^{\prime}\right)^{2}}-\sqrt{1+\left(x_{0}^{\prime}\right)^{2}}\right) d z=\int\left(x_{0}^{\prime} \Delta x^{\prime}+\frac{1}{2} \Delta\left(x^{\prime}\right)^{2}\right) d z
$$

By integration by parts, with $d u=x_{0}^{\prime} d z, u=x_{0}, v=\Delta x^{\prime}$, and $d v=\Delta x^{\prime \prime} d z$,

$$
\Delta s=x_{0} \Delta x^{\prime}-\int x_{0} \Delta x^{\prime \prime} d z+\frac{1}{2} \int \Delta\left(x^{\prime}\right)^{2} d z
$$

with

$$
\Delta x^{\prime}=\frac{g B_{0}}{\gamma k} \int \frac{\Delta B}{B_{0}} k d z \quad \text { and } \quad \Delta x^{\prime \prime}=\frac{g B_{0}}{\gamma} \frac{\Delta B}{B_{0}}
$$

where $g B_{0} / \gamma k=K / \gamma$, and thus

$$
\Delta s=\frac{K^{2}}{\gamma^{2} k}\left(-\cos k z \int \frac{\Delta B}{B_{0}} k d z+\int \cos (k z) \frac{\Delta B}{B_{0}} k d z+\frac{1}{2} \int\left(\int \frac{\Delta B}{B_{0}} k d z\right)^{2} k d z\right)
$$

Furthermore,

$$
\Delta t=\frac{\Delta s}{c}
$$

$$
\Delta \varphi=\omega \Delta t=\Delta s \frac{\omega}{c}=\Delta s \cdot k_{L}
$$

where,

$$
\begin{aligned}
& \lambda_{L}=\frac{\lambda}{2 \gamma^{2}}\left(1+\frac{K^{2}}{2}\right) \\
& k_{L}=k \frac{4 \gamma^{2}}{K^{2}\left(1+\frac{2}{K^{2}}\right)}
\end{aligned}
$$

and therefore,

$$
\Delta \varphi=P(\underbrace{-\cos k z \int \frac{\Delta B}{B_{0}} k d z}_{(a)}+\underbrace{\int \cos (k z) \frac{\Delta B}{B_{0}} k d z}_{(b)}+\underbrace{\frac{1}{2} \int\left(\int \frac{\Delta B}{B_{0}} k d z\right)^{2} k d z}_{(c)}),
$$

with

$$
P=\frac{4}{1+2 / K^{2}}
$$

Term (c) is of second order and is important only for a long insertion device. Term (a) gives harmonics and reduces line intensity for steering errors, but produces no effect if there is no steering. Term (b) produces phase shift and line broadening. Whether or not it is equal 0 depends on such elements as symmetry, but not on presence of net steering.

## $\varrho, A_{0} / B_{1}$ for Hybrid Insertion Device

This note is a result of ANL lecture notes, and from Simple Analytical Model For Fields From One Pole Of Hybrid Insertion Device, $\dagger$ with $k_{1}=2 \pi / \lambda$ and $k_{n}=n k_{1}$, and

$$
B(x)=\sum_{n=\mathrm{odd}} B_{n} \cos k_{n} x
$$

from poles with $\pm$ " 1 ", and

$$
B_{n}=\frac{2}{\pi} \int_{-\infty}^{\infty} b(x) \cos k_{n} x d k_{1} x
$$

where $b(x)$ is the field from one pole with excitation +1 .
For $V$ from 0 to $V_{1}$ at the edge of pole going from $-x_{1}$ to $x_{1}$,

$$
B_{n}=4 \frac{k_{1}}{k_{0}} \frac{\sin k_{n} x_{1}}{\sinh \left(\pi k_{n} / k_{0}\right)} \frac{V_{1}}{h} \quad \text { with } \quad k_{0}=\pi / h
$$

For $V$ going linearly from 0 to $\dot{V}_{1}$ over thickness of CSEM $=\left(x_{2}-x_{1}\right) / 2$, we have

$$
A_{0}=\int_{-\infty}^{\infty} b(x) d x=\frac{V_{1}}{h} \frac{\lambda}{2}
$$

$$
B_{n}=4 \frac{k_{1}}{k_{0}} \frac{V_{1}}{h} \frac{\sin \left(k_{n}\left(x_{2}-x_{1}\right) / 2\right)}{\left(k_{n}\left(x_{2}-x_{1}\right) / 2\right)} \frac{1}{\sinh \left(\pi k_{n} / k_{0}\right)} .
$$

For $k_{1}\left(x_{2}-x_{1}\right)^{2} / 6 \ll 1$,

$$
B_{1}=4 \frac{k_{1}}{k_{0}} \frac{V_{1}}{h} \frac{1}{\sinh \left(\pi k_{1} / k_{0}\right)}=\frac{4}{\pi} \frac{V_{1}}{h} \frac{\alpha}{\sinh \alpha} \quad \text { with } \quad \alpha=\pi \frac{k_{1}}{k_{0}}=2 \pi \frac{h}{\lambda}=\pi \frac{g}{\lambda} .
$$

We may now conclude that

$$
\frac{A_{0}}{B_{1}}=\lambda \frac{\pi}{8} \frac{\sinh \alpha}{\alpha}
$$

$\dagger$ Document 0136u-w

For

$$
\varrho=\frac{A_{0}}{\int_{-\lambda / 4}^{\lambda / 4} B_{1} \cos k_{1} x d x}=\frac{B_{1} \lambda \frac{\pi}{8} \frac{\sinh \alpha}{\alpha}}{B_{1}\left(2 / k_{1}\right)}=\frac{k_{1} \lambda \pi}{16} \frac{\sinh \alpha}{\alpha}=\frac{\pi^{2}}{8} \frac{\sinh \alpha}{\alpha}
$$

we have

| $\varrho$ | 1.5 | 2.0 | 2.5 | 3.0 | 4.0 | 5.0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $g / \lambda$ | .35 | .57 | .70 | .79 | .94 | 1.05 |

## Simple Analytical Model for Fields from One Pole of Hybrid Insertion Device



Figùre 1.

Model: midplane on $V=0$, and pole from $-\infty$ to $+\infty$ on $V=0$, except on $V=V_{1}$ for $-x_{1} \leq x \leq x_{1}$.

The above geometry is described by the following conformal map

$$
\pi \dot{z}=\frac{h}{t}
$$

and the following elements

$$
k_{0} z=\ln t, \quad t=e^{k_{0} z} \quad \text { and } \quad k_{0}=\frac{\pi}{h},
$$

and where $x_{1}$ is the half-width of the pole. Putting $\pm$ current filaments at $x= \pm x_{1}$,

$$
\begin{gathered}
\pi F=V_{1} \ln \frac{\left(t-t_{1}\right)}{\left(t-t_{2}\right)} \\
F^{\prime}=\frac{V_{1}}{h}\left(\frac{t}{t-t_{1}}-\frac{t}{t-t_{2}}\right) \\
=\frac{V_{1}}{h}\left(\frac{1}{t / t_{1}-1}-\frac{1}{t / t_{2}-1}\right) .
\end{gathered}
$$

where $t=e^{k_{0} z}, t_{1}=e^{k_{0}\left(x_{1}+i h\right)}=-e^{k_{0} x_{1}}$. Thus,

$$
\begin{align*}
F^{\prime} \frac{h}{V_{1}} & =\frac{1}{e^{k_{0}\left(z+x_{1}\right)}+1}-\frac{1}{e^{k_{0}\left(z-x_{1}\right)}+1} \\
& =-\frac{e^{k_{0} z}\left(e^{k_{0} x_{1}}-e^{-k_{0} x_{1}}\right)}{e^{2 k_{0} z}+1+e^{k_{0} z}\left(e^{k_{0} x_{1}}+e^{-k_{0} x_{1}}\right)} \\
& =-\frac{\sinh k_{0} x_{1}}{\cosh k_{0} z+\cosh k_{0} x_{1}}=-b(z) . \tag{1}
\end{align*}
$$

The odd harmonics of the field are described by

$$
\begin{equation*}
B_{N}=a \int_{-\infty}^{\infty} b(x) \cos \left(N k_{1} x\right) d x \tag{2}
\end{equation*}
$$

where $N=2 n+1, k_{N}=k_{1} N=k_{1}(2 n+1), k_{1}=2 \pi / \lambda$, and $a$ is a constant, thus

$$
\begin{align*}
\frac{B_{N}}{a} & =\sinh k_{0} x_{1} \int_{-\infty}^{\infty} \frac{\cos \left(k_{N} x\right) d x}{\cosh k_{0} x+\cosh k_{0} x_{1}} \\
& =\sinh k_{0} x_{1} \cdot G_{N}  \tag{3}\\
G_{N} & =\Re \int_{-\infty}^{\infty} \frac{e^{i k_{N} x}}{\cosh k_{0} x+\cosh k_{0} x_{1}} d x \tag{4}
\end{align*}
$$

That is, to evaluate this integral, one can integrate a line integral along the real axis of the complex $z$-plane, and close it at $\infty$ in the upper half-plane without changing its value.

There are singularities at $\cosh k_{0} z=-\cosh k_{0} x_{1}$ in the upper half-plane, with $k_{0} z=$ $\pm k_{0} x_{1}+i \pi(2 m+1)$, for $m=0,1,2, \ldots$. We take the first singularity at $k_{0} z=$ $+k_{0} x_{1}+i \pi M$ and do others later by replacing $x_{1}$ by $-x_{1}$. We integrate over the upper half-plane:

$$
\begin{aligned}
G_{N+} & =\Re \int \frac{e^{i k_{N} z} d z}{\cosh k_{0} z+\cosh k_{0} x_{1}} \\
& =\Re \sum_{m=0} \frac{e^{i k_{N}\left(x_{1}+i \pi M / k_{0}\right)}}{k_{0} \sinh k_{0} z_{m}} \cdot 2 \pi i \\
& =+\frac{2 \pi}{k_{0}} \cdot \frac{\sin k_{N} x_{1}}{\sinh k_{0} x_{1}} \cdot \sum e^{-\pi\left(k_{N} / k_{0}\right)(2 m+1)} \\
& =+\frac{2 \pi}{k_{0}} \cdot \frac{\sin k_{N} x_{1}}{\sinh k_{0} x_{1}} \cdot \frac{e^{-\pi k_{N} / k_{0}}}{1-e^{-2 \pi k_{N} / k_{0}}} \\
& =+\frac{2 \pi}{k_{0}} \cdot \frac{\sin k_{N} x_{1}}{\sinh k_{0} x_{1}} \cdot \frac{1}{2 \sinh \left(\pi k_{N} / k_{0}\right)} \\
& =\frac{\pi}{k_{0}} \cdot \frac{\sin k_{N} x_{1}}{\sinh k_{0} x_{1}} \cdot \frac{1}{\sinh \left(\pi k_{N} / k_{0}\right)}
\end{aligned}
$$

One solves for $G_{N-}$ similarly. Thus, we may re-write (3),

$$
\begin{equation*}
\frac{B_{N}}{a}=\frac{2 \pi}{k_{0}} \cdot \frac{\sin k_{N} x_{1}}{\sinh \left(\pi k_{N} / k_{0}\right)} \tag{5}
\end{equation*}
$$

and further,

$$
\begin{equation*}
\frac{B_{(2 n+1)}}{B_{1}}=\frac{\sin \left(k_{1}(2 n+1) x_{1}\right)}{\sinh \left(\left(\pi k_{1} / k_{0}\right)(2 n+1)\right)} \cdot \frac{\sinh \left(\pi k_{1} / k_{0}\right)}{\sin \left(k_{1} x_{1}\right)}, \tag{6}
\end{equation*}
$$

with $\quad k_{0}=\frac{\pi}{h}, \quad k_{1}=\frac{2 \pi}{\lambda}, \quad$ and thus $\pi \frac{k_{1}}{k_{0}}=2 \pi \frac{h}{\lambda}$.
This model of $b(z)$, and the resultant $B_{N}$, assume that the potential increases like a step function at the edge of the pole. As a next approximation, to improve this model, one would assume that the potential increases linearly over the size of the CSEM and represent this by the operation

$$
\frac{1}{x_{2}-x_{1}} \int_{x_{1}}^{x_{2}} d x_{1}
$$

which is easily executed on both $b(z)$ and $B_{N}$. For $b(z)$ we have

$$
\begin{equation*}
\int \frac{\sinh k_{0} x_{1}}{\cosh k_{0} z+\cosh k_{0} x_{1}} \cdot \frac{d x_{1}}{\left(x_{2}-x_{1}\right)}=\frac{1}{k_{0}\left(x_{2}-x_{1}\right)} \ln \frac{\cosh k_{0} z+\cosh k_{0} x_{2}}{\cosh k_{0} z+\cosh k_{0} x_{1}} \tag{7}
\end{equation*}
$$

and for $B_{N}$ we have

$$
\begin{align*}
\int \sin \left(N k_{1} x_{1}\right) \frac{d x_{1}}{\left(x_{2}-x_{1}\right)} & =\frac{\cos N k_{1} x_{1}-\cos N k_{1} x_{2}}{N k_{1}\left(x_{2}-x_{1}\right)} \\
& =-\sin \left((2 n+1) k_{1} \frac{x_{2}+x_{1}}{2}\right) \frac{\sin \left((2 n+1) k_{1} \frac{x_{2}-x_{1}}{2}\right)}{\frac{(2 n+1)}{k_{1}} \cdot \frac{x_{2}-x_{1}}{2}} \tag{8}
\end{align*}
$$

where

$$
\begin{gathered}
-\sin \left(k_{1}(2 n+1)\left(\frac{x_{2}+x_{1}}{2}\right)\right)=(-1)^{n+1}, \\
\frac{x_{2}+x_{1}}{2}=\frac{\lambda}{4} \text { and }(2 n+1) k_{1} \frac{x_{2}-x_{1}}{2}=\frac{\pi}{2}+n \pi .
\end{gathered}
$$

The argument of the log function can, and should be, operated on in the same manner, such that for

$$
\begin{gathered}
\cosh k_{0} z=C_{0}, \quad \cosh k_{0} x_{1}=C_{1}, \quad \text { and } \quad \cosh k_{0} x_{2}=C_{2}, \\
\frac{C_{0}+C_{2}}{C_{0}+C_{1}}=\frac{C_{0}+a+b}{C_{0}+a-b}=\frac{1+\frac{b}{C_{0}+a}}{1-\frac{b}{C_{0}+a}}
\end{gathered}
$$

where

$$
a=\frac{C_{2}+C_{1}}{2}, \quad b=\frac{C_{2}-C_{1}}{2}, \quad \text { and } \quad \ln \frac{H \varepsilon}{1-\varepsilon}=2\left(\varepsilon+\frac{\varepsilon^{3}}{3}+10\right)
$$

## Wiggler Parameter K Definitions

For $v=c$ we have

$$
\begin{gathered}
m_{0} \gamma v^{2} x^{\prime \prime}=e v B=e v A^{\prime}, \\
x^{\prime}=\frac{1}{\gamma} \frac{e}{m_{0} v} A \quad \text { and } \quad x_{\max }^{\prime}=\frac{K_{1}}{\gamma} .
\end{gathered}
$$

## Definition 1:

$$
K_{1}=\frac{e}{m_{0} c} A_{\max }=\frac{e}{2 \pi m_{0} c} 2 \pi A_{\max }
$$

For a pure sinusoidal field we have

$$
B=B_{0} \sin k z \quad \text { and } \quad A=\frac{B_{0}}{k} \cos k z
$$

Thus

$$
A_{\max }=\frac{B_{0}}{k} \Longrightarrow K_{1}=\frac{e}{2 \pi m_{0} c} B_{0} \lambda_{u}
$$

## Definition 2:

The "path length" slippage in $\lambda_{u}$ equals $\lambda_{\text {light }}$. (We shall refer to $\lambda_{\text {light }}$ as $\lambda_{\mathrm{L}}$ for the remainder of this document.)

$$
\Delta t=\frac{s}{c \beta}-\frac{\lambda_{u}}{c}=\frac{\lambda_{\mathrm{L}}}{c} \quad \text { and thus } \quad \lambda_{\mathrm{L}}=\lambda_{u}\left(\frac{s}{\lambda_{u} \beta}-1\right)
$$

where $s=$ path length over one period, and $s^{\prime}=\sqrt{1+x^{\prime 2}}$.
Proceeding from above, we now have that

$$
\begin{aligned}
\frac{\lambda_{\mathrm{L}}}{\lambda_{u}} & =\frac{1}{\beta \lambda_{u}} \int_{0}^{\lambda_{u}}\left(1+\frac{x^{\prime 2}}{2}\right) d z-1 \\
& =\frac{1}{\beta}-1+\frac{1}{2 \lambda_{u}} \int_{0}^{\lambda_{u}} x^{\prime 2} d z
\end{aligned}
$$

By introducing

$$
\beta^{2}+\frac{1}{\gamma^{2}}=1 \text { and thus } \beta^{-1}=\left(1-\frac{1}{\gamma^{2}}\right)^{-1 / 2}=1+\frac{1}{2 \gamma^{2}}
$$

we further simplify

$$
\begin{aligned}
\frac{\lambda_{\mathrm{L}}}{\lambda_{u}} & =\frac{1}{2 \gamma^{2}}+\frac{1}{2 \lambda_{u}} \int_{0}^{\lambda_{u}}{x^{\prime}}^{2} d z \\
& =\frac{1}{2 \gamma^{2}}(1+\underbrace{\left(\frac{e}{m_{0} c}\right)^{2} \frac{\int_{0}^{\lambda_{u}} A^{2} d z}{\lambda_{u}}}_{K_{2}^{2} / 2})
\end{aligned}
$$

and we now arrive at our definition

$$
K_{2}^{2}=\left(\frac{e}{m_{0} c}\right)^{2} \frac{2}{\lambda_{u}} \int_{0}^{\lambda_{u}} A^{2} d z
$$

For $A=\frac{B_{0}}{k} \cos k z$ we have

$$
K_{2}^{2}=\left(\frac{e}{m_{0} c}\right)^{2} \frac{2}{\lambda_{u}} \frac{B_{0}^{2}}{k^{2}} \frac{\lambda_{u}}{2}=\left(\frac{e}{2 \pi m_{0} c} B_{0} \lambda_{u}\right)^{2}
$$

where $\left(e / 2 \pi m_{0} c\right)=.934 \cdot 10^{2}$ in SI units.

We define $\left(2 \pi m_{0} c / e\right)=A_{e}$ and thus $1 / A_{e}=.934 \cdot 10^{2}$ MKS.
We now reformulate our Definitions 1 and 2 such that

$$
K_{1}=\frac{2 \pi}{A_{e}} A_{\max }
$$

and

$$
K_{2}=\frac{2 \pi}{A_{\dot{e}}} \sqrt{\frac{2 \int_{0}^{\lambda_{u} / 4} A^{2} d z}{\lambda_{u} / 4}}
$$

## Definition 3:

$$
K_{3}=\frac{B_{0} \lambda_{u}}{A_{e}}=\frac{V_{0}}{A_{e}} 4 \frac{\lambda_{u} / 4}{D_{4}}
$$

Where $D_{4}$ refers to the NPOLEI.BAS program variable which describes the distance factor in the transformation from scalar potential to the field.

## NPOLE

A recreation, with "Korea modification," of a program (for HP71B) to design and analyze $\lambda / 4$ of hybrid insertion device.

We will begin by establishing some background information for the conformal map and the limits for $t_{1}$ and $t_{2}$.


Figure 1.

For the map of Figure 1,

$$
\pi \dot{z}=-i \frac{1-t_{1}}{\sqrt{t}\left(t-t_{1}\right)(t-1)}
$$

$$
a=\frac{1}{\sqrt{t_{1}}} \text { and thus } t_{1}=\frac{1}{a^{2}} .
$$



Figure 2.

We proved in Korea in 1987 that

$$
0<t_{1}<1 / a^{2}<t_{2}<1 \text { and } \frac{1}{a^{2}}=\left(\frac{W_{1}}{W_{0}}\right)^{2}
$$

for the geometry of Figure 2.
Therefore, in program NPOLE1.BAS (included at the end of this document),

$$
t_{1}=\frac{1}{a^{2}} \operatorname{RANG}(\mathrm{C} 1), \quad t_{2}=\frac{1}{a^{2}}+\left(1-\frac{1}{a^{2}}\right) \operatorname{RANG}(\mathrm{C} 2),
$$

with $0<\operatorname{RANG}(x)<1$ as used in the first version, and $\operatorname{RANG}(x)=1 /\left(1+e^{x}\right)$ as specifically used now.
The map for geometry with corners at $t_{1}, t_{2}$ is described by

$$
\dot{z}=-i W_{1} \frac{Q_{1}}{\sqrt{t} \sqrt{t-t_{1}} \sqrt{t-t_{2}}(t-1)} \quad \text { and } \quad Q_{1}=\frac{\sqrt{1-t_{1}} \sqrt{1-t_{2}}}{\pi} .
$$

We determine $t_{1}$ and $t_{2}$ from

$$
\frac{h_{0}}{W_{1}}=Q_{1} \int_{0}^{t_{1}} \frac{d t}{\sqrt{t} \sqrt{t_{1}-t} \sqrt{t_{2}-t}(1-t)}
$$

$$
\frac{W_{0}}{W_{1}}=Q_{1} \int_{t_{1}}^{t_{2}} \frac{d t}{\sqrt{t} \sqrt{t-t_{1}} \sqrt{t_{2}-t}(1-t)}
$$

To evaluate the integrals, we use

$$
\int_{t_{1}}^{t_{2}} \frac{f(t) d t}{\sqrt{t-t_{1}} \sqrt{t_{2}-t}}=3 \int_{-1}^{1} \frac{f(t)}{\sqrt{4-x^{2}}} d x
$$

$$
t=\frac{2\left(t_{2}+t_{1}\right)+\left(t_{2}-t_{1}\right) x\left(3-x^{2}\right)}{4}
$$

We use Gaussian integration with segmented intervals for testing and accuracy purposes. We use a " 2 D " secant equation solver to determine $t_{1}$ and $t_{2}$ from the above integrals.
We describe the complex potentials for fluxes, fields:


Figure 3.

$$
\dot{F}=-\frac{Q_{2} V_{0}}{\sqrt{t} \sqrt{t-t_{1}}(t-1)} \quad \text { and } \quad Q_{2}=\frac{\sqrt{1-t_{1}}}{\pi} .
$$

We therefore have

$$
F=-Q_{2} V_{0} \int \frac{d t}{t^{2} \sqrt{1-t_{1} / t}(1-1 / t)}=-Q_{2} \int \frac{2 d u}{t_{1}\left(1-\left(1-u^{2}\right) / t_{1}\right)} V_{0}
$$

where

$$
\begin{gathered}
\sqrt{1-\frac{t_{1}}{t}}=u, \quad \frac{t_{1}}{t}=1-u^{2}, \quad \text { and thus } \quad \frac{d t}{t^{2}}=\frac{2 u d u}{t_{1}} \\
1-t_{1}=t_{3}^{2}
\end{gathered}
$$

Thus,

$$
\begin{aligned}
\frac{1}{V_{0}} F & =-Q_{2} \int \frac{2 d u}{u^{2}-t_{3}^{2}} \\
& =-Q_{2} \int\left(\frac{1}{u-t_{3}}-\frac{1}{u+t_{3}}\right) \frac{d u}{t_{3}} \\
& =\frac{Q_{2}}{t_{3}} \ln \frac{u+t_{3}}{u-t_{3}} \\
& =\frac{Q_{2}}{t_{3}} \ln \frac{\sqrt{1-t_{1} / t}+t_{3}}{\sqrt{1-t_{1} / t}-t_{3}} .
\end{aligned}
$$

Flux into pole / $V_{0}$ :

$$
\begin{aligned}
E_{p} & =\left.\frac{Q_{2}}{t_{3}} \ln \left|\frac{\sqrt{1+t_{1} / \tau}+t_{3}}{\sqrt{1+t_{1} / \tau}-t_{3}}\right|\right|_{0} ^{\infty} \\
& =\frac{Q_{2}}{t_{3}} \ln \frac{1+t_{3}}{1-t_{3}} .
\end{aligned}
$$

Flux into midplane (for $K$ ):

$$
\begin{aligned}
E_{M} & =\left(F\left(t_{2}\right)-F\left(t_{1}\right)\right) / V_{0} \\
& =\frac{Q_{2}}{t_{3}} \ln \frac{t_{3}+\sqrt{1-t_{1} / t_{2}}}{t_{3}-\sqrt{1-t_{1} / t_{2}}}
\end{aligned}
$$

with $K_{1}=2 \pi V_{0} E_{M}\left(1 / A_{e}\right)$ where $\left(1 / A_{e}\right)=.934 \cdot 10^{2} \mathrm{MKS}$.
We calculate the excess flux coefficient for the side of the pole $\left(V_{0}=1\right)$ :

$$
\left(A(t)-A(\infty)=\frac{y(t)-y(\infty)}{W_{1}}+E_{s}\right)_{t \rightarrow 1}
$$

$$
E_{s}=\int_{1}^{\infty}\left(-\dot{F}+\frac{\dot{z}}{i}\right) d t=\int_{1}^{\infty} \frac{1}{\sqrt{t} \sqrt{t-t_{1}}(t-1)} \underbrace{\left(Q_{2}-\frac{Q_{1}}{\sqrt{t-t_{2}}}\right)}_{G_{1}} d t
$$

where

$$
\begin{aligned}
G_{1} & =Q_{2}\left(1-\frac{\sqrt{1-t_{2}}}{\sqrt{t-t_{2}}}\right) \\
& =\frac{Q_{2}}{\sqrt{t-t_{2}}}\left(\sqrt{t-t_{2}}-\sqrt{1-t_{2}}\right) \\
& =\frac{Q_{2}(t-1)}{\sqrt{t-t_{2}}\left(\sqrt{t-t_{2}}+\sqrt{1-t_{2}}\right)}
\end{aligned}
$$

and thus

$$
\begin{aligned}
E_{s} & =Q_{2} \int_{1}^{\infty} \frac{d t}{\sqrt{t} \sqrt{t-t_{1}} \sqrt{t-t_{2}}\left(\sqrt{t-t_{2}}+\sqrt{1-t_{2}}\right)} \\
& =Q_{2} \int_{0}^{1} \frac{d t}{\sqrt{1-t_{1} t} \sqrt{1-t_{2} t}\left(\sqrt{t-t_{2} t}+\sqrt{t} \sqrt{1-t_{2}}\right)} .
\end{aligned}
$$

The field $B_{0}$ at $t=t_{1}$ is $(\dot{F} / \dot{z})$. With $V_{0}$ on pole,

$$
B_{0}=\frac{V_{0}}{W_{1}} \frac{Q_{2}}{Q_{1}} \sqrt{t_{2}-t_{1}}=\frac{V_{0}}{W_{1}} \frac{\sqrt{t_{2}-t_{1}}}{\sqrt{1-t_{2}}}=\frac{V_{0}}{D_{4}}
$$

where $D_{4}$ is an old notation and

$$
D_{4}=W_{1} \frac{\sqrt{1-t_{2}}}{\sqrt{t_{2}-t_{1}}}
$$

For the second definition of $K$,

$$
K_{2}=2 \pi \frac{1}{A_{e}} \sqrt{\frac{2 \int_{0}^{\lambda_{u} / 4} A^{2} d z}{\lambda_{u} / 4}}
$$

We need $\int F^{2} d z$, thus,

$$
\int A^{2} d z=\int_{t_{1}}^{t_{2}} F^{2} \dot{z} d t=G_{2}
$$

Therefore, we have that

$$
G_{2}=\frac{Q_{1} Q_{2}^{2}}{t_{3}^{2}} V_{0}^{2} W_{1} \int_{t_{1}}^{t_{2}}\left(\ln \frac{t_{3}+\sqrt{1-t_{1} / t}}{t_{3}-\sqrt{1-t_{1} / t}}\right)^{2} \frac{d t}{\sqrt{t} \sqrt{t-t_{1}} \sqrt{t_{2}-t}(1-t)}
$$

$$
\frac{G_{2}}{\lambda_{u} / 4}=\frac{G_{2}}{W_{0}}=V_{0}^{2} \frac{W_{1}}{W_{0}} \frac{Q_{1} Q_{2}^{2}}{t_{3}^{2}} J
$$

where

$$
J=\int_{t_{1}}^{t_{2}}\left(\ln \frac{t_{3}+\sqrt{1-t_{1} / t}}{t_{3}-\sqrt{1-t_{1} / t}}\right)^{2} \frac{d t}{\sqrt{t} \sqrt{t-t_{1}} \sqrt{t_{2}-t}(1-t)},
$$

and thus we may summarize

$$
K_{2}=2 \pi V_{0} \frac{1}{A_{e}} \sqrt{\frac{2 W_{1} Q_{2}^{2} Q_{1}}{t_{3}^{2} W_{0}} J} .
$$

## Program NPOLE1.BAS

```
PRINT:PRINT DATE$;" ";TIME$;". NPOLE1"
'GOTO BYPASS
PRINT "Determines parameter values and evaluates flux into midplane (Em) and"
PRINT "pole (Ep) of ID, and excess flux coefficient for side of pole (Es)."
PRINT "K1,K2,K3 are obtained by multiplying the printed values by the scalar"
PRINT "potential of pole in Tcm. K1 is for maximum deflection angle, K2 for"
PRINT "trajectory length effect, and K3 for B0*period. D4=VO/BO for V0=1."
REM--List of P1() elements:0>W01~(-2),1>T1,2>T2,3>T3,5>Q1,6>Q2,9>Function ID
BYPASS:
DEFINT J:DEFDBL A-Z
PI=4*ATN(1):A1$="##.###"~-~ ":TAP=0
A2\$=" Em Ep Ki K2 K3
A3$=" HO=##.## WO=##.## W1=##.##"
DIM P1(0:9),GX(1:4),GW(1:4)
SHARED PI,GX(),GW(),P1(),A1$
REM--GX,GH=(normalized) abscissas; P1=parameters for Gauss integrator
DATA .1834346425,.3626837834,.5255324099,.3137066459
DATA .7966664774,.2223810345,.9602898565,.1012285363
FOR J1=1 TO 4:READ GX(J1),GW(J1):NEXT J1:REH--Abscissas, weights for Gauss
REM
'C10=1:C20=1
PRINT:PRINT TAB(TAP);:PRINT A2$
DO
AGAIN:
INPUT;"HO, WO, \(>\) H1=", HOO,W00,W10:REM--INPUT unnormalized 1/2gap, period/4, IF HOO>O THEN HO=HOO:REM--pole to symmetry line distance, stored temporarily
IF WOO>O THEN WO=WOO:REM--in HOO,WOO,W1O, then in HO,WO,W1(=not-normalized).
IF W10>0 THEN W1=W10:REM--H01,W01=normalized with W1, used in program.
IF HOO=0 AND WOO=0 AND W1O=0 THEN END
IF WO<W1 THEN PRINT TAB(20);:PRINT "WO must be larger than W1!":GOTO AGAIN
PRINT TAB(24);:PRINT USING A3$;HO;WO;W1
н01=H0/W1:W01=WO/W1:P1(0)=1/(WO1*W01):DC1=.1:DC2=.1
GOSUB SOLVIT
PRINT TAB(TAP);:PRINT USING A1$;EM;EP;ES;K1;K2;K3;D4
LOOP
SOLVIT:
C11=C10 + DC1:C21=C20:C12=C10:C22=C20 + DC2
CALL EVAL (C10, C20, S10,S20):S10=S10-H01:S20=S20-w01:S00=ABS (S10) +ABS (S20)
CALL EVAL(C11,C21,S11,S21):S11=S11-H01:S21=S21-स01:S01=ABS(S11)+ABS(S21)
CALL EVAL (C12,C22,S12,S22) :S12=S12-H01:S22=S22-H01:S02=ABS(S12)+ABS(S22)
```

DO

REM--REARR puts "rorst" set into last column, to be discarded later GOSUB REARR

N1=1/((S11-S10)*(S22-S20)-(S12-S10)*(S21-S20)): REH--Start of 2D secant $\mathrm{D} 1=\mathrm{N} 1 *(\mathrm{~S} 20 * \mathrm{~S} 12-\mathrm{S} 10 * \mathrm{~S} 22): \mathrm{D} 2=\mathrm{N} 1 *(\mathrm{~S} 10 * \mathrm{~S} 21-\mathrm{S} 20 * \mathrm{~S} 11):$ REM--equation solver DC1 $=($ C11-C10) *D1 $+(C 12-C 10) * D 2: D C 2=(C 21-C 20) * D 1+(C 22-C 20) * D 2$

C12=C11: C22=C21:C11=C10:C21=C20:S12=S11:S22=S21:S11=S10:S21=S20
S02=S01:S01=S00:C10=C10+DC1:C20=C20+DC2:REM--Recommended new parameters
CALL EVAL (C10, C20,S10,S20):S10=S10-H01:S20=S20-W01:S00=ABS (S10) +ABS (S20)
LDOP UNTIL SOO<. 001
$\mathrm{P} 1(9)=3$ : $\operatorname{CALL} \operatorname{SGAUSSINT8}(0,1, \mathrm{G} 2,-.001): \mathrm{Q} 2=\mathrm{P} 1(6): \mathrm{T} 1=\mathrm{P} 1(1): \mathrm{T} 2=\mathrm{P} 1(2): \mathrm{T} 3=\mathrm{P} 1(3)$
$\mathrm{Q} 1=\mathrm{P} 1(5): \mathrm{ES}=\mathrm{Q} 2 * \mathrm{G} 2: \mathrm{EP}=\mathrm{Q} 2 / \mathrm{T} 3 * \mathrm{LOG}((1+\mathrm{T} 3) /(1-\mathrm{T} 3)): \mathrm{D} 4=\mathrm{W} 1 * \mathrm{SQR}((1-\mathrm{T} 2) /(\mathrm{T} 2-\mathrm{T} 1))$
$\mathrm{EM}=\mathrm{Q} 2 / \mathrm{T} 3 * \mathrm{LOG}(2 /(1-\mathrm{SQR}(1-\mathrm{T} 1 / \mathrm{T} 2) / \mathrm{T} 3)-1): \mathrm{K} 1=2 * \mathrm{PI} * \mathrm{EM} * .934: \mathrm{K} 3=4 * \mathrm{HO} / \mathrm{D} 4 * .934$
P1 (9) $=4$ : CALL SGAUSSINT8 ( $-1,1, \mathrm{~K} 2,-.001$ ) : K2 $=2 * \mathrm{PI} * \mathrm{Q} 2 / \mathrm{P} 1(3) * \mathrm{SQR}(2 * \mathrm{Q} 1 * 3 * \mathrm{~K} 2 / \mathrm{W01}) * .934$
RETURN

REARR:
IF S00>S01 THEN GOSUB SW01
IF S01>S02 THEN GOSUB S $\$ 12$
RETURN

SWO1:
SWAP S00,S01:SWAP S10,S11:SWAP S20,S21:SWAP C10,C11:SWAP C20,C21:RETURN SW12:
SWAP S01,S02:SWAP S11,S12:SWAP S21,S22:SWAP C11,C12:SWAP C21,C22:RETURN

SUB EVAL(C1,C2,S1,S2):REM--Calculates H01, W01 for set of parameters C1,C2>T1,T2
T1=P1 (0)*RANG (C1):T2=P1 (0) +(1-P1 (0))*RANG(C2)
$\mathrm{T} 3=\mathrm{SQR}(1-\mathrm{T} 1): \mathrm{P} 1(1)=\mathrm{T} 1: \mathrm{P} 1(2)=\mathrm{T} 2: \mathrm{P} 1(3)=\mathrm{T} 3$
$\mathrm{Q} 2=\mathrm{T} 3 / \mathrm{PI}: \mathrm{Q} 1=\mathrm{Q} 2 * \mathrm{SQR}(1-\mathrm{T} 2): \mathrm{P} 1(5)=\mathrm{Q} 1: \mathrm{P} 1(6)=\mathrm{Q} 2$
P1 (9) $=1$ : CALL SGAUSSINT8( $-1,1, \mathrm{G} 2,-.001$ ): S1=3*Q1*G2
P1 (9) $=2$ : CALL SGAUSSINT8 ( $-1,1, \mathrm{G} 2,-.001$ ) : S2=3*Q1*G2
END SUB

SUB SGAUSSINT8(XO,X3,G2,DG):REM-Gauss integrator, aith interval segmentation IF DG>0 THEN E1=DG:E2=0 ELSE E1=0:E2=-DG:REM-FFor DG>/<0,DG=absol./reI. error CALL GAUSSINT8(X0,X3,G2):J1=1:J4=16:REM--J4=largest \# subdiv.
DO
G1=G2:G2=0:J1=2*J1:DX=(X3-X0)/J1:REM--G1/G2=last/next computed integral
IF J1>J4 THEN PRINT " Not converged":END
FOR J2=0 TO J1-1
CALL GAUSSINT8(XO+J2*DX,X0+(J2+1)*DX,G3)

NEXT J2
LOOP UNTIL ABS(G2-G1)<E1+E2*ABS(G2) OR J1>J4
END SUB

```
SUB GAUSSINT8(X1,X2,G2):REM---m---m----Integrator; G2=value of integral
X0=.5*(X2+X1):X3=X0-X1:G2=0
ON P1(9) GOTO INTEGRAND1,INTEGRAND2,INTEGRAND3,INTEGRAND4
INTEGRAND1:
FOR J1=1 TO 4
    DX=GX(J1)*X3:G2=G2+GH(J1)*(GCT1(X0+DX)+GCT1(X0-DX))
NEXT J1:G2=G2*X3
EXIT SUB
INTEGRAND2:
FOR J1=1 TO 4
    DX=GX(J1)*X3:G2=G2+GW(J1)*(GCT2(XO+DX) +GCT2(XO-DX))
NEXT J1:G2=G2*X3
EXIT SUB
INTEGRAND3:
FOR J1=1 TO 4
    DK=GX(J1)*X3:G2=G2+GW(J1)*(GCT3(X0+DX)+GCT3(XO-DX))
NEXT J1:G2=G2*X3
EXIT SUB
INTEGRAND4:
FOR J1=1 TO 4
    DX=GX(J1)*X3:G2=G2+GW(J1)*(GCT4(XO+DX)+GCT4(XO-DX))
NEXT J1:G2=G2*X3
END SUB
FUNCTION GCT1(X):REM-------------_First of functions to be integrated.
TT=P1(1)*(2+X*(3-X*X))/4:GCT1=1/SQR((P1 (2)-TT)*(4-X*X))/(1-TT)
END FUNCTION
FUNCTION GCT2(X)
TT=((P1(2)+P1(1))*2+(P1(2)-P1(1))*X*(3-X*X))/4
GCT2=1/SQR(TT*(4-X*X))/(1-TT)
END FUNCTION
FUNCTION GCT3(X)
S1=SQR(1-P1(2)*X):GCT3=1/SQR(1-P1(1)*X)/(S1*(S1+SQR(X*(1-P1(2)))))
END FUNCTION
FUNCTION GCT4(X)
TT=((P1(2)+P1(1))*2+(P1(2)-P1(1))*X*(3-X*X))/4:T4=SQR(1-P1(1)/TT)/P1(3)
GCT4=(LOG((1+T4)/(1-T4)))}-2/(1-TT)/SQR(TT*(4-X*X))
END FUNCTIOR
```

```
FUNCTION RANG(X):REH-_-__-_-_-_-_-_-_-_-_-_-_-_-_
RANG =1/(1+EXP(X))
END FUNCTION
DEF FNPOLE (X)=((X+1)*LOG(X+1)-(X-1)*LOG(X-1))/PI
```


## Program Results

```
06-28-1993 10:14:02 NPOLE1
```

Determines parameter values and evaluates flux into midplane (Em) and pole (Ep) of ID, and excess flux coefficient for side of pole (Es). K1,K2,K3 are obtained by multiplying the printed values by the scalar potential of pole in Tcm. K1 is for maximum deflection angle, K2 for trajectory length effect, and $K 3$ for BO*period. $D 4=V O / B O$ for $V 0=1$.


## Error of Flux Calculation for Finite Pole Width with Excess Flux Coefficient



Figure 1.

$$
\dot{z}=-\frac{\sqrt{t^{2}-1}}{{\sqrt{t^{2}-a^{2}}}^{3}}
$$

We have

$$
h_{0}=\int_{0}^{\infty} \frac{\sqrt{1+t^{2}}}{{\sqrt{a^{2}+t^{2}}}^{3}} d t . \quad \text { and } \quad h_{1}=\int_{1}^{\infty} \frac{\sqrt{1-t^{2}}}{{\sqrt{t^{2}-a^{2}}}^{3}} d t
$$

We introduce

$$
\begin{gathered}
\sqrt{a^{2}+t^{2}}=\frac{1}{u}, \quad t=\sqrt{\frac{1}{u^{2}}-a^{2}}=\frac{\sqrt{1-a^{2} u^{2}}}{u} \\
d t=-\frac{\frac{a^{2} u^{2}}{\sqrt{1-a^{2} u^{2}}}+\sqrt{1-a^{2} u^{2}}}{u^{2}} d u=-\frac{1}{u^{2} \sqrt{1-a^{2} u^{2}}} d u \\
1+t^{2}=1-a^{2}+\frac{1}{u^{2}}=b^{2}+\frac{1}{u^{2}}, \quad b^{2}=1-a^{2}
\end{gathered}
$$

we therefore have

$$
h_{0}=-\frac{1}{a} \int_{0}^{1 / a} \frac{\sqrt{1+b^{2} u^{2}}}{\sqrt{1-a^{2} u^{2}}} d u
$$

And given $a u=\sin \varphi$, and $d u=\frac{\cos \varphi d \varphi}{a}$,

$$
\begin{aligned}
h_{0} & =-\frac{1}{a} \int_{0}^{\pi / 2} \sqrt{1+\frac{b^{2}}{a^{2}} \sin ^{2} \varphi} d \varphi \\
& =-\frac{1}{a} \int_{0}^{\pi / 2} \sqrt{1+\frac{b^{2}}{a^{2}}-\frac{b^{2}}{a^{2}} \cos ^{2} \varphi} d \varphi \\
& =-\frac{1}{a^{2}} \int_{0}^{\pi / 2} \sqrt{1-b^{2} \cos ^{2} \varphi} d \varphi \\
& =-\frac{1}{a^{2}} E\left(b^{2}\right) .
\end{aligned}
$$

For $t=\frac{1}{u}, d t=-\frac{d u}{u^{2}}$, and $u=\sin \alpha$

$$
\begin{aligned}
h_{1} & =-\int_{0}^{1} \frac{\sqrt{1 / u^{2}-1}}{\sqrt{1 / u^{2}-a^{2}}} \frac{d u}{u^{2}} \\
& =-\int_{0}^{1} \frac{\sqrt{1-u^{2}}}{{\sqrt{1-a^{2} u^{2}}}^{3}} d u \\
& =-\int_{0}^{\pi / 2} \frac{\cos ^{2} \alpha d \alpha}{\sqrt{1-a^{2} \sin ^{2} \alpha}}
\end{aligned}
$$

From Jahnke and Emde ${ }^{1}$ :

$$
h_{1}=-\frac{K\left(a^{2}\right)-E\left(a^{2}\right)}{a^{2}}
$$

and therefore

$$
\frac{h_{1}}{h_{0}}=\frac{K\left(a^{2}\right)-E\left(a^{2}\right)}{E\left(1-a^{2}\right)}
$$

[^8]

Figure 2.

$$
\begin{gathered}
\dot{F}=-\frac{2 a}{t^{2}-a^{2}}=\frac{1}{t+a}-\frac{1}{t-a} \\
\pi F=\ln \frac{t+a}{t-a}=\ln \frac{1+a / t}{1-a / t}
\end{gathered}
$$

The flux into the poleface is

$$
A(1)-A(\infty)=A_{\text {ideal }}=\frac{1}{\pi} \ln \frac{1+a}{1-a}
$$

Comparing this flux to the homogeneous flux and the excess flux for the end of a semi-infinite pole with half-gap $=h_{0}$, we have

$$
\begin{aligned}
A_{\text {approx }} & =\frac{h_{1}}{h_{0}}+\frac{1}{\pi}(2-\ln 4) \\
& =\frac{K\left(a^{2}\right)-E\left(a^{2}\right)}{E\left(1-a^{2}\right)}+.195 .
\end{aligned}
$$

Therefore we have

$$
G\left(\frac{h_{1}}{h_{0}}\right)=\frac{A_{\text {approx }}-A_{\text {ideal }}}{A_{\text {ideal }}}=\frac{A_{\text {approx }}}{A_{\text {ideal }}}-1=\frac{\frac{h_{1}}{h_{0}}+\frac{1}{\pi}(2-\ln 4)}{\frac{1}{\pi} \ln \frac{1+a}{1-a}}-1 .
$$

## Program EXCFLTST.BAS

CLS
DEFDBL A-Z
PRINT DATE \$ ;" ";TIME\$ ;" EXCFLTST"
REM--Error of flux calculation for finite width pole with excess flux
REM--coefficient. IMPUT parameter = 1/2-width of pole / 1/2-gap.
PI=4*ATN(1):A1\$="dA=\#\#.\#\#\#-~-- $\mathrm{dA} / \mathrm{A}=\# \# . \# \# \#^{-\cdots-\cdots} \mathrm{dX} / \mathrm{H} 1=\# \# . \# \# \#^{-\cdots-n}$
$E 1=(2-L O G(4)) / P I: X 2=1: D Y=1 E-6$
DIM P1 (0:2)
DO
INPUT;"H1/HO=", $\mathbf{H}$
$\mathrm{X} 1=.9 * \mathrm{X} 2: \mathrm{P} 1(0)=\mathrm{H} 0:$ REM-G1=2/(1+EXP $(\mathrm{PI} * \mathrm{HO} 0+2)): E 1=2 / \mathrm{PI} *(1-\mathrm{G} 1-L O G(2-\mathrm{G} 1))$
CALL SECANTS (X1,X2,DY,Y2,P1())
$A 1=1 / \operatorname{SQR}(1+\operatorname{EXP}(X 2)): A 0=L O G((1+A 1) /(1-A 1)) / P I: A E=H 0+E 1$
PRINT TAB(15);:PRINT USING A1\$;AE-AO;AE/AO-1; (AE-AO)/H0
LOOP

SUB SECANTS (X1,X2,DY,Y2,P1())
CALL FCTY( $\mathrm{X} 1, \mathrm{Y} 1, \mathrm{P} 1())$ : CALL $\operatorname{FCTY}(\mathrm{X} 2, Y 2, \mathrm{P} 1())$
IF ABS (Y1)<ABS (Y2) THEN SHAP Y1,Y2:SHAP X1,X2
J1\%=0
DO
$D X=Y 2 *(X 1-X 2) /(Y 2-Y 1)$
$\mathrm{X} 1=\mathrm{X} 2: \mathrm{Y} 1=\mathrm{Y} 2: \mathrm{X} 2=\mathrm{X} 1+\mathrm{DX}: J 1 \%=\mathrm{J} 1 \%+1$
CALL FCTY(X2,Y2,P1())
LOOP UNTIL ABS (Y2)<DY OR J1\%=15
IF $\mathrm{J} 1 \%=15$ THEN PRINT "NOT COAVERGED"
END SUB

SUB FCTY(X1,Y1,P1())
$A 2=1 /(1+E X P(X 1))$
Y1=(ELLK (A2)-ELLE(A2))/ELLE(1-A2)-P1(0)
END SUB

FUNCTION ELLK (X1)
$\mathrm{X}=1-\mathrm{X} 1: \mathrm{S} 1=.01451196212 * \mathrm{X}+.03742563713: \mathrm{S} 1=\mathrm{S} 1 * \mathrm{~K}+.03590092383$
$\mathrm{S} 1=\mathrm{S} 1 * \mathrm{X}+.09666344259: \mathrm{S} 1=\mathrm{S} 1 * \mathrm{X}+1.38629436112: \mathrm{S} 2=.00441787012 * \mathrm{X}+.03328355346$
S2=S2*X+.06880248576:S2=S2*X+.12498593597:S2=S2*X+.5
ELLK=S1-S2*LOG(X)
END FUNCTION

## FUNCTION ELLE(X1)

$\mathrm{X}=1-\mathrm{X} 1: \mathrm{S} 1=.01736506451 * \mathrm{X}+.04757383546: \mathrm{S} 1=\mathrm{S} 1 * \mathrm{X}+.0626060122$
S1=S1*X+. $44325141463: S 2=X * .00526449639+.04069697526$
S2=S2*X+. 09200180037:S2=S2*X+. 2499836831
ELLE $=X * S 1+1-X * S 2 * L O G(X)$
END FUNCTION

## Program Results

06-16-1993 09:09:24 EXCFLTST
$\mathrm{H} 1 / \mathrm{H} 2=.1 \mathrm{dA}=4.832 \mathrm{E}-02 \mathrm{dA} / \mathrm{A}=1.956 \mathrm{E}-01 \mathrm{dX} / \mathrm{H} 1=4.832 \mathrm{E}-01$
$\mathrm{H} 1 / \mathrm{H} 2=.2 \mathrm{dA}=2.240 \mathrm{E}-02 \mathrm{dA} / \mathrm{A}=6.022 \mathrm{E}-02 \mathrm{dX} / \mathrm{H} 1=1.123 \mathrm{E}-01$
$\mathrm{H} 1 / \mathrm{H} 2=.3 \mathrm{dA}=1.129 \mathrm{E}-02 \mathrm{dA} / \mathrm{A}=2.331 \mathrm{E}-02 \mathrm{dX} / \mathrm{H} 1=3.762 \mathrm{E}-02$
$\mathrm{H} 1 / \mathrm{H} 2=.4 \mathrm{dA}=5.848 \mathrm{E}-03 \mathrm{dA} / \mathrm{A}=9.920 \mathrm{E}-03 \mathrm{dX} / \mathrm{H} 1=1.462 \mathrm{E}-02$
H1/H2=.5 $\mathrm{dA}=3.074 \mathrm{E}-03 \mathrm{dA} / \mathrm{A}=4.441 \mathrm{E}-03 \mathrm{dX} / \mathrm{H} 1=6.149 \mathrm{E}-03$
H1/H2=.6 $\quad \mathrm{dA}=1.627 \mathrm{E}-03 \mathrm{dA} / \mathrm{A}=2.050 \mathrm{E}-03 \quad \mathrm{dX} / \mathrm{H} 1=2.712 \mathrm{E}-03$
$\mathrm{H} 1 / \mathrm{H} 2=.7 \quad \mathrm{dA}=8.645 \mathrm{E}-04 \quad \mathrm{dA} / \mathrm{A}=9.665 \mathrm{E}-04 \mathrm{dX} / \mathrm{H} 1=1.235 \mathrm{E}-03$
$\mathrm{H} 1 / \mathrm{H} 2=.8 \mathrm{dA}=4.602 \mathrm{E}-04 \mathrm{dA} / \mathrm{A}=4.626 \mathrm{E}-04 \mathrm{dX} / \mathrm{H} 1=5.752 \mathrm{E}-04$
Н1/H2=.9 $\mathrm{dA}=2.452 \mathrm{E}-04 \quad \mathrm{dA} / \mathrm{A}=2.239 \mathrm{E}-04 \mathrm{dX} / \mathrm{H} 1=2.725 \mathrm{E}-04$
H1/H2=1 $\mathrm{dA}=1.307 \mathrm{E}-04 \mathrm{dA} / \mathrm{A}=1.094 \mathrm{E}-04 \mathrm{dX} / \mathrm{H} 1=1.307 \mathrm{E}-04$
H1/H2 $=2 \quad \mathrm{dA}=3.008 \mathrm{E}-07 \mathrm{dA} / \mathrm{A}=1.370 \mathrm{E}-07 \mathrm{dX} / \mathrm{H} 1=1.504 \mathrm{E}-07$

## Excess Flux Into Pole and Flux Into Side of Gm40



Figure 1.

The conformal map is described by

$$
\pi \dot{z}=\frac{\sqrt{t}(a-1)^{2}}{(t-1)(t-a)^{2}}
$$

To determine the value $a$ that produces the desired $D$, we use $t=a+\tau,|\tau| \ll a$. Expanding in $\tau$ gives

$$
\pi \frac{d z}{d \tau}=\frac{(a-1)^{2} \sqrt{a}\left(1+\frac{\tau}{2 a}\right)}{(a-1)\left(1+\frac{\tau}{a-1}\right)} \cdot \frac{1}{\tau^{2}}
$$

Expanding more, and then integrating over the half-circle around $t=a$, we get

$$
\begin{aligned}
D & =-(a-1) \sqrt{a}\left(\frac{1}{2 a}-\frac{1}{a-1}\right) \\
& =\sqrt{a} \cdot \frac{a+1}{2 a} \\
& =\frac{1}{2}\left(\sqrt{a}+\frac{1}{\sqrt{a}}\right) .
\end{aligned}
$$

By substitution and integration, we have

$$
\frac{\pi z}{(a-1)^{2}}=\frac{\partial}{\partial a} \int \frac{\sqrt{t} d t}{(t-1)(t-a)}=\frac{\partial}{\partial a} J
$$

April, 1993. Note 0131u-w.
where, for $t=W^{2}$ and $d t=2 W d W$,

$$
\begin{aligned}
J & =\int \frac{2 t d W}{(t-1)(t-a)} \\
& =\int \frac{2}{a-1}\left(\frac{a}{t-a}-\frac{1}{t-1}\right) d W \\
& =\frac{1}{a-1} \int\left(\sqrt{a}\left(\frac{1}{W-\sqrt{a}}-\frac{1}{W+\sqrt{a}}\right)-\left(\frac{1}{W-1}-\frac{1}{W+1}\right)\right) d W \\
& =\frac{1}{a-1}\left(\sqrt{a} \ln \frac{\sqrt{a}-W}{\sqrt{a}+W}+\ln \frac{1+W}{1-W}\right):
\end{aligned}
$$

Further, we have that

$$
\begin{gathered}
\sqrt{a} \frac{\partial}{\partial a} \ln \frac{\sqrt{a}-W}{\sqrt{a}+W}=\frac{1}{2}\left(\frac{1}{\sqrt{a}-W}-\frac{1}{\sqrt{a}+W}\right)=\frac{W}{a-t} \\
\left(\frac{1}{\sqrt{a}-1 / \sqrt{a}}\right)^{\prime}=-\frac{1}{2 a} \cdot \frac{\sqrt{a}+1 / \sqrt{a}}{(\sqrt{a}-1 / \sqrt{a})^{2}}=-\frac{D}{(a-1)^{2}} \\
\left(\frac{1}{a-1}\right)^{\prime}=-\frac{1}{(a-1)^{2}}
\end{gathered}
$$

Thus,

$$
J^{\prime}=-\frac{1}{(a-1)^{2}}\left(\ln \frac{1+W}{1-W}+D \ln \frac{\sqrt{a}-W}{\sqrt{a}+W}\right)+\left(\frac{1}{a-1} \cdot \frac{W}{a-t}\right)
$$

and therefore,

$$
\pi z=(a-1) \frac{W}{a-t}+D \ln \frac{\sqrt{a}+W}{\sqrt{a}-W}-\ln \frac{1+W}{1-W}
$$

Further, for

$$
\begin{gathered}
\pi \dot{F}=-\frac{a-1}{(t-1)(t-a)}=\frac{1}{t-1}-\frac{1}{t-a} \\
\pi F=\ln \frac{1-t}{1-t / a} .
\end{gathered}
$$

The flux into the side of the pole, for $-\infty \leq t \leq 0$, is

$$
A_{S}=\frac{1}{\pi} \ln (a)
$$

We describe the excess flux into the poleface by

$$
\begin{gathered}
\Delta A_{P}=F(0)-F(1-\varepsilon)-(z(0)-z(1-\varepsilon)) \text { follows } \varepsilon \rightarrow 0 \\
\pi \Delta A_{P}=\ln \frac{1-1 / a}{\varepsilon}+\left(1+D \ln \frac{\sqrt{a}+1}{\sqrt{a}-1}-\ln \frac{2}{\varepsilon / 2}\right) \\
\frac{a-1}{\varepsilon a} \frac{\varepsilon}{4}=\frac{a-1}{4 a} \\
\Delta A_{P}=\frac{1}{\pi}\left(1+\ln \frac{a-1}{4 a}+D \ln \frac{\sqrt{a}+1}{\sqrt{a}-1}\right) .
\end{gathered}
$$

The definition of $\Delta A_{P}$ means that the flux into the pole surface is the same as the uniform flux into a pole whose width is increased, on both sides, by the product of the half-gap and the expression for $\Delta A_{P}$. The definition of $A_{S}$ means that the total flux into each side of the pole equals the product of the scalar potential of the pole and the expression for $A_{S}$.
From our expression for D , and $a-2 D \sqrt{a}+1=0$, we have

$$
\sqrt{a}=D+\sqrt{D^{2}-1}
$$

We may now eliminate $a$ from $A_{S}$ and $\Delta A_{P}$. Thus,

$$
A_{S}=\frac{2}{\pi} \ln \left(D+\sqrt{D^{2}-1}\right)
$$

and further,

$$
\begin{gathered}
1 / \sqrt{a}=D-\sqrt{D^{2}-1} \\
\frac{\sqrt{a}+1}{\sqrt{a}-1}=\frac{D+1+\sqrt{D^{2}-1}}{D-1+\sqrt{D^{2}-1}}=\frac{\sqrt{D+1}}{\sqrt{D-1}} \cdot \frac{\sqrt{D+1}+\sqrt{D-1}}{\sqrt{D-1}+\sqrt{D+1}}=\sqrt{\frac{D+1}{D-1}} \\
\frac{a-1}{4 a}=\frac{\sqrt{a}-\sqrt{1 / a}}{4 \sqrt{a}}=\frac{\sqrt{D^{2}-1}}{2\left(D+\sqrt{D^{2}-1}\right)}=\frac{1}{2\left(1+1 / \sqrt{1-1 / D^{2}}\right)} \\
\triangle A_{P}=\frac{1}{\pi}\left(1+\frac{D}{2} \ln \frac{D+1}{D-1}-\ln \left(2\left(1+\frac{D}{\sqrt{D^{2}-1}}\right)\right)\right) .
\end{gathered}
$$



Figure 1.

Status characterized by status vector $v=\binom{V}{\Phi}$, where $V$ is the scalar potential with respect to the midplane, and $\Phi$ is the flux transported to the right. Going "downstream", $V$ and $\Phi$ change because of the $\int H d s$ "loss" in iron (and due to small gaps), and because of flux going to the midplane. Over a short distance,

$$
\begin{equation*}
\frac{d \Phi}{d x}=-V \cdot \varepsilon \tag{1}
\end{equation*}
$$

with $\varepsilon$ to 0 th approximation (detailed later in this note) is given by

$$
\begin{equation*}
\varepsilon=\frac{W}{h} \tag{2}
\end{equation*}
$$

with $h$ having the value of the half-gap, and $W$ being the width over which the flux "escapes" to the midplane.

$$
\begin{equation*}
\frac{d V}{d x}=-\Phi \cdot k_{2} \tag{3}
\end{equation*}
$$

April, 1993. Note 0129u-w.
with $k_{2}$ in 0 th approximation (also detailed later) given by

$$
\begin{equation*}
k_{2}=\frac{1}{a \mu}=\gamma / a \tag{4}
\end{equation*}
$$

where $a$ is the cross-section area of the flux "duct", and $\mu$ is the permeability. The voltage drop due to small gaps perpendicular to the flux flow will be added later. Within the section with constants $k_{2}$ and $\varepsilon$, we get

$$
\begin{equation*}
V^{\prime \prime}=-k_{2} \Phi^{\prime}=V k^{2} \quad \text { with } \quad k^{2}=\varepsilon k_{2} . \tag{5}
\end{equation*}
$$

The solution within the uniform section of length $x$ is

$$
\begin{gather*}
V=\alpha C+\beta S \text { with } C=\cosh k x \text { and } S=\sinh k x \\
\Phi=-V^{\prime} / k_{2}=-(\alpha S+\beta C) k / k_{2} \\
v(x)=\left(\begin{array}{cc}
C & S \\
-S k / k_{2} & -C k / k_{2}
\end{array}\right)\binom{\alpha}{\beta} \text { with }\binom{\alpha}{\beta}=\left(\begin{array}{cc}
1 & 0 \\
0 & -k_{2} / k
\end{array}\right) v(0), \\
v(x)=M \cdot v(0), \quad v=\binom{V}{\Phi} \text { and } M=\left(\begin{array}{cc}
C & -S k_{2} / k \\
-S k / k_{2} & C
\end{array}\right) \tag{6}
\end{gather*}
$$

By reversing the direction arrow of $\Phi$, i.e., by re-defining the sign of $\Phi$, the off-diagonal minus signs disappear. The sequence of sections with different properties are taken into account by multiplying their matrices. $v$ remains unchanged when crossing the interface from one section to the next unless there is a (steering) coil, or a local field clamp, thus introducing an additive $\Delta v$ when going through that interface.

It is clear that $1 / k$ is the important scaling distance that describes how transported flux decays.

## Structure of Solution to Simple Problem.



Figure 2.

There are field clamps at each end, i.e. at point 0 and point 5. There are $\pm \Delta V$ coils at the interfaces between points 1 and 2, and between points 3 and 4. The status vectors $v_{0}=\binom{0}{\Phi_{0}}$ and $v_{5}=\binom{0}{\Phi_{5}}$ describe that the points 0 and 5 are located in the midplane, and that they contain the to-be-determined values $\Phi_{0}$ and $\Phi_{5}$ which represent the fluxes going to the midplane through the field clamps. Of similar interest are the $\Phi$-components of $v_{2}$ and $v_{4}$.
Given $\Delta v_{0}=\binom{\Delta v_{0}}{0}$, we describe the coil(s) by

$$
\begin{gather*}
v_{2}=M_{01} v_{0}+\Delta v_{0}  \tag{7.1}\\
v_{4}=M_{23} v_{2}-\Delta v_{0}=M_{23} M_{01} v_{0}+\left(M_{23}-I\right) \Delta v_{0} \tag{7.2}
\end{gather*}
$$

where $I$ is the unit matrix.

$$
\begin{equation*}
v_{5}=M_{45} v_{4}=\underbrace{M_{05}}_{a_{i k}} v_{0}+\underbrace{M_{45}\left(M_{23}-I\right)}_{b_{i k}} \Delta v_{0} \tag{7.3}
\end{equation*}
$$

where $a_{i k}$ and $b_{i k}$ are elements of these matrices, and thus

$$
\begin{gather*}
v_{5}=\Phi_{0}\binom{a_{12}}{a_{22}}+\binom{b_{11}}{b_{21}} \Delta V_{0}=\binom{0}{\Phi_{5}} . \\
\Phi_{0}=-\Delta V_{0} b_{11} / a_{12}  \tag{7.4}\\
\Phi_{5}=\Phi_{0} a_{22}+\Delta V_{0} b_{21}=\Delta V_{0}\left(b_{21}-a_{22} b_{11} / a_{12}\right) \\
\Phi_{5}=\Delta V_{0}\left(a_{12} b_{21}-a_{22} b_{11}\right) / a_{12} \tag{7.5}
\end{gather*}
$$

With $\Phi_{0}$ and $\Phi_{5}$ now known, (7.1) and (7.2) give the flux produced by the coils in the section delimited by points 2 and 3 .

## Details of $k_{2}$.

One has to be careful to use the correct value for $\mu$. If the field associated with this flux is parallel to the pre-existing field, one has to use $\mu=\frac{d B}{d H}$. If it is perpendicular to the pre-existing flux, one must use $\mu=\frac{B}{H}$ which is the "normal $\mu$ ".

Now we must look at the effect of a thin gap over a large area. $\Delta V$ across that gap from flux $\Phi$ is gotten from $\Phi=\int \frac{\Delta V d a}{g}$ and thus $\Delta V=\frac{\Phi}{\int \frac{d a}{g}}$. If a gap-less length $L$ of $\mu$ is associated with this gap, the total $\Delta V$ is given by

$$
\begin{gather*}
\Delta V=\Phi\left(\frac{1}{\int \frac{d a}{g}}+\gamma \frac{L}{A}\right)=\Phi L\left(\frac{\gamma}{a}+\frac{1 / L}{\int \frac{d g}{a}}\right), \\
k_{2}=\frac{\gamma}{a}+\frac{1 / L}{\int \frac{d a}{g}} . \tag{8}
\end{gather*}
$$

## Details of $\varepsilon$.

Only the general approach and the results derived in a separate note are given here. There are three contributions to $\varepsilon$ : flux from the top, from the sides, and from the poles facing the midplane.


Figures 3(a) and 3(b).

(c)

(d)

Figures 3(c) and 3(d).

To get the flux into the top per unit length in direction perpendicular to the paper plane, we use as a model the solid block that touches the midplane of Figure 3(b). For the flux into each side, we use the geometry of Figure 3(c) and calculate the flux into the side. If the side has "pole structure" we take it into account with an excess voltage drop coefficient approximation (if necessary). For flux from the poles to the midplane, we calculate the flux for the geometry of Figure 3(d), and we use
the excess flux coefficient for a solid block of Figure 3(c) to correct the width $2 W_{0}$ of the cross-section shown in Figure 3(a).

Results for the Geometry of Figure 3(d).

$$
\begin{equation*}
\varepsilon_{P}=\frac{\Phi(\lambda / 4)}{V_{0}} \frac{4}{\lambda}\left(W_{0}+\Delta W_{0}\right) 2 \tag{9.1}
\end{equation*}
$$

with $\frac{\Phi(\lambda / 4)}{V_{0}}$ calculated by POISSON or an analytical program.
We calculate $\Delta W_{0}$ from the geometry of Figure 3(c), with

$$
\begin{equation*}
\Delta W_{0}=h_{0} \frac{1}{\pi}\left(1+\frac{D}{2} \ln \left(\frac{D+1}{D-1}\right)-\ln \left(2\left(1+\frac{D}{\sqrt{D^{2}-1}}\right)\right)\right) \tag{9.2}
\end{equation*}
$$

The contribution from the flux into the top is

$$
\begin{equation*}
\varepsilon_{T}=\frac{2}{\pi} \ln \frac{1+a_{1}}{1-a_{1}} \tag{10.1}
\end{equation*}
$$

where $a_{1}$ is determined from

$$
\begin{equation*}
\frac{h_{1}}{W_{0}}=\frac{E(b)-a_{1}^{2} K(b)}{E\left(a_{1}\right)-b^{2} K\left(a_{1}\right)} \tag{10.2}
\end{equation*}
$$

with $b^{2}=1-a_{1}^{2}$, and

$$
\begin{equation*}
E\left(a_{1}\right)=\int_{0}^{\pi / 2} \sqrt{1-a_{1}^{2} \sin ^{2} \varphi} d \varphi \quad \text { and } \quad K\left(a_{1}\right)=\int_{0}^{\pi / 2} \frac{d \varphi}{\sqrt{1-a_{1}^{2} \sin ^{2} \varphi}} \tag{10.3}
\end{equation*}
$$

The flux into the sides contributes

$$
\varepsilon_{S}=\frac{4}{\pi} \ln \left(D+\sqrt{D^{2}-1}\right)
$$

with $D$ given by (9.2).
(11.1) assumes smooth sides, i.e., the excess potential drop is ignored. It should be noted that the area $a$ in (8) is smaller than the cross-section shown in Figure 3(a), the latter includes the poles, while the former does not.

We make here further clarifications on units. If we were to deal with a uniform field over a width $W$ of a flat pole, at distance $h_{0}$ from the midplane, $\varepsilon$ would be exactly $\varepsilon=\frac{W}{H}$. That is, $\frac{d \Phi}{d x}$ and $V$ have the same dimensions, meaning that either $\mu_{0}=4 \pi \cdot 10^{-7}$ is incorporated in the vector potential $V$, or $\mu_{0}$ is left out of the definition of $\Phi$. The meaning of $\varepsilon$ is the flux per unit length in the axial direction of the structure on potential $V$, divided by $V$.
$\varepsilon_{S}$ with excess potential drop is given by

$$
\begin{equation*}
\varepsilon_{S}=\frac{4}{\pi} \ln \left(D_{1}+\sqrt{D_{1}^{2}-1}\right) \tag{11.2}
\end{equation*}
$$

with

$$
\begin{equation*}
D_{1}=\frac{h_{1}+\frac{2}{\pi} \Delta L}{h_{0}+\frac{2}{\pi} \Delta L} \tag{11.3}
\end{equation*}
$$



Figure 4.

$$
\begin{equation*}
\alpha=\frac{\lambda / 4}{h_{3}} \quad \text { and } \quad \Delta L=\frac{h_{3}}{\pi}((\alpha+1) \ln (1+1 / \alpha)+(\alpha-1) \ln (1-1 / \alpha)) \tag{11.4}
\end{equation*}
$$

The effect of $\Delta L$ will be very small under most circumstances. The excess flux
potential drop is too small to be of concern for $\varepsilon_{T}$.

## 3D Scalar Potential for Saturation-Caused Fields in the Insertion Device

This entails the same approach as for the case of $\mu=\infty$, except that the condition $\partial V / \partial x=0$ at $y=h$ is to be dropped:

$$
\begin{aligned}
V & =\sum \cos n k_{z} z \cdot g_{n}(x, y) \\
\nabla^{2} V & =0 \Longrightarrow \nabla^{2} g_{n}=n^{2} k_{z}^{2} g_{n}
\end{aligned}
$$

We introduce $n k_{z} x=u$, and $n k_{z} y=v$ :

$$
\begin{equation*}
\nabla_{u, v}^{2} g=g . \tag{1}
\end{equation*}
$$

We construct $g(u, v)$ that has the following properties: odd in $y, g(-v)=-g(v)$, and gives field approximating $\cosh \varepsilon u-1$ for $y=0 . \varepsilon$ is arbitrary, real or imaginary, and the field equals 0 for $u=0$ when letting $\varepsilon \rightarrow 0$ at end.
We try $g=\cosh \varepsilon u \sinh a v$. To satisfy (1):

$$
\varepsilon^{2}+a^{2}=1 \text { and thus } a=\sqrt{1-\varepsilon^{2}}
$$

has to hold. We add a function of $v$, such that $g_{v}^{\prime}$ is proportional to $\cosh \varepsilon u-1$ for $v=0$. The only odd function of $v$ that will satisfy this requirement and also satisfy (1) is $-a \sinh v$, thus

$$
\begin{equation*}
g=\cosh \varepsilon u \sinh a v-a \sinh v \tag{2}
\end{equation*}
$$

One can use the superposition of such functions with different $\varepsilon$, but this would probably not be practical.
The expansion for $\varepsilon \rightarrow 0$ is

$$
\begin{equation*}
g=\frac{\varepsilon^{2}}{2}\left(u^{2} \sinh v-v \cosh v+\sinh v\right) \tag{3}
\end{equation*}
$$

For $v=0$, we obtain the expected sextupole field:

$$
\begin{equation*}
g_{v}^{\prime}(u, 0)=\frac{\varepsilon^{2}}{2} u^{2} \tag{4}
\end{equation*}
$$

At the pole, where $v_{h}=n k_{z} h$,

$$
\begin{equation*}
g_{u}^{\prime}\left(u, v_{h}\right)=\frac{\varepsilon^{2}}{2} 2 u \sinh v_{h} \tag{5}
\end{equation*}
$$

It is this field in the $x$-direction that is responsible for the sextupole field in the midplane.

February, 1992. Note 0125u-w.
(5) allows us to make an estimate of the saturation effects in the midplane during the design phase. Thus,

$$
\begin{equation*}
\frac{g_{v}^{\prime}(u, 0)}{g_{u}^{\prime}\left(u, v_{h}\right)}=\frac{u}{2 \sinh v_{h}}=\frac{n k_{z} x}{2 \sinh n k_{z} h}=\frac{x}{2 h} \cdot \frac{n k_{z} h}{\sinh n k_{z} h} . \tag{6}
\end{equation*}
$$

It is interesting to note that every additional expansion of (2) in $\varepsilon^{2}$ leads to a new solution to (1) describing the fields in the midplane to the highest orders $\sim x^{4}, x^{6}$, etc.
To check on (3), its expansion in $k_{z}$ up to the 3 rd order terms in $\{u, v\}$ gives, as expected,

$$
\begin{align*}
g & =\frac{\varepsilon^{2}}{2}\left(u^{2} v-v\left(1+\frac{v^{2}}{2}-\left(1+\frac{v^{2}}{6}\right)\right)\right) \\
& =\frac{\varepsilon^{2}}{2}\left(u^{2} v-\frac{v^{3}}{3}\right) \\
& =\frac{\varepsilon^{2}}{6} \Im(u+i v)^{3} . \tag{7}
\end{align*}
$$

## Scalar Potential for 3D Insertion Device Fields

In the 2 D case,

$$
\begin{equation*}
V=\sum_{n=\text { odd }} b_{n} \cos n k_{z} z \cdot \sinh n k_{z} y \quad \text { with } \quad k_{z}=\frac{2 \pi}{\lambda} \tag{1}
\end{equation*}
$$

In order to simplify matters, we drop the sum, and re-introduce it at the end.
The effects of lateral ends are equally periodic in $z$, thus $n k_{z} \Rightarrow k_{z}$, and

$$
\begin{equation*}
V_{L E}=\cos k_{z} z \cdot g(x, y) \tag{2}
\end{equation*}
$$

Where $g(x, y)$ is valid only in the vacuum region of the $\{x, y\}$ space.
Further, we have that

$$
\begin{equation*}
\nabla^{2} V=0 \quad \Longrightarrow \quad \nabla^{2} g=k_{z}^{2} g \tag{3}
\end{equation*}
$$

where $g$ is the Fourier expansion coefficient as a function of $x, y$,
At the pole surface, for integer $\mu, z=\mu \lambda / 2$ and $y=h=$ half gap, $B_{x}=B_{y}=0$. We expand $g$ in a Fourier series in $y$. We have that $g \sim \sin m k_{y} y$ for $k_{y}=\pi / h$, and

$$
g=\sum a_{m} \sin m k_{y} y \cdot \cosh k_{m} x
$$

with $k_{m}^{2}=k_{z}^{2}+m^{2} k_{y}^{2}$. We use $a_{m}=b_{0} b_{m} / \cosh k_{m} W$, where W is half the pole width, and we expect $b_{m}$ to be only weakly dependent on $W$.

$$
\begin{cases}V=\sum_{n=\mathrm{odd}} b_{n_{0}} \cos n k_{z} z \cdot\left(\sinh n k_{z} y+g_{n}\right),  \tag{4}\\ g_{n}=\sum_{m=1} \frac{b_{n m} \sin m k_{y} y \cdot \cosh k_{n m} x}{\cosh k_{n m} W}, & \text { with } k_{z}=2 \pi / \lambda, k_{y}=\pi / h \\ k_{n m}^{2}=n^{2} k_{z}^{2}+m^{2} k_{y}^{2}, & \text { with } \frac{k_{z}}{k_{y}}=\frac{2 h}{\lambda}=\frac{\text { gap }}{\lambda} \\ k_{n m}^{2}=m^{2} k_{y}^{2}\left(1+\frac{n^{2}}{m^{2}}\left(\frac{k_{z}}{k_{y}}\right)^{2}\right), & \end{cases}
$$

Under most circumstances, $\frac{k_{z}}{k_{y}} \leq .5$. For $n=1, k_{n m} \approx m k_{y}$.

February, 1992. Note 0124u-w.

In the region of interest, only the case of $m=1$ is of importance. That is, the dominant term is

$$
\begin{equation*}
g_{1}=\frac{b_{11} \sin k_{y} y \cdot \cosh k_{11} x}{\cosh k_{11} W} \tag{5}
\end{equation*}
$$

We may now proceed to conclude that

$$
\begin{equation*}
-H_{y}=k_{z} \sum n b_{n_{0}} \cos n k_{z} z\left(\cosh n k_{z} y+\sum b_{n m} \frac{m k_{y}}{n k_{z}} \cdot \frac{\cos m k_{y} y \cdot \cosh k_{n m} x}{\cosh k_{n m} W}\right) \tag{6}
\end{equation*}
$$

From (6), we expect $b_{n m}<0$, and $\frac{m k_{y}}{n k_{z}}\left|b_{n m}\right|$ to be in the order of 1 , but probably less than 1.

## Suggestions for Magnetic Measurements.

Make all measurements as function of $z$, filter out random errors, and then do the harmonic analysis by measuring the quantities derived from $\sinh n k_{z} y+g_{n}$. To measure field components, measure $B_{y}$ at $y=0$ for a number of values of $x$ close enough to the lateral edge to get values of $b_{n 1}$ and $b_{n 2}$. Then measurements of $B_{x}$ close to the lateral ends are made, at $y \simeq h / 2$, to check the validity of $V(x, y, z)$. If agreement is reached, an investigatation of whether $b_{n m}$ are more easily obtained from $B_{x}$ measurements is to be done. To verify the model, compare the measurements at individual points, without the harmonic analysis, to the model calculations.
After sufficient measurements, make a table that lists the $b_{n m}$ coefficients as functions of two dimensionless products (i.e. $h / \lambda$ and $W / h$ ), and possibly find a practical formula to represent the data. A possible complication may result from saturation in the iron which may dominate the behavior of the field as a function of $x$.
Examination of experimental data shows that decay of field errors as one moves away from the lateral edge of the insertion device can be much slower than this description indicates. A possible cause of this may be $H_{x}$ at pole surface caused by saturation.
(11.1) assumes smooth sides, i.e., the excess potential drop is ignored. It should be noted that the area $a$ in (8) is smaller than the cross-section shown in Figure 3(a), the latter includes the poles, while the former does not.

We make here further clarifications on units. If we were to deal with a uniform field over a width $W$ of a flat pole, at distance $h_{0}$ from the midplane, $\varepsilon$ would be exactly $\varepsilon=\frac{W}{H}$. That is, $\frac{d \Phi}{d x}$ and $V$ have the same dimensions, meaning that either $\mu_{0}=4 \pi \cdot 10^{-7}$ is incorporated in the vector potential $V$, or $\mu_{0}$ is left out of the definition of $\Phi$. The meaning of $\varepsilon$ is the flux per unit length in the axial direction of the structure on potential $V$, divided by $V$.
$\varepsilon_{S}$ with excess potential drop is given by

$$
\begin{equation*}
\varepsilon_{S}=\frac{4}{\pi} \ln \left(D_{1}+\sqrt{D_{1}^{2}-1}\right) \tag{11.2}
\end{equation*}
$$

with

$$
\begin{equation*}
D_{1}=\frac{h_{1}+\frac{2}{\pi} \Delta L}{h_{0}+\frac{2}{\pi} \Delta L} \tag{11.3}
\end{equation*}
$$



Figure 4.

$$
\begin{equation*}
\alpha=\frac{\lambda / 4}{h_{3}} \quad \text { and } \quad \Delta L=\frac{h_{3}}{\pi}((\alpha+1) \ln (1+1 / \alpha)+(\alpha-1) \ln (1-1 / \alpha)) \tag{11.4}
\end{equation*}
$$

The effect of $\Delta L$ will be very small under most circumstances. The excess flux
potential drop is too small to be of concern for $\varepsilon_{T}$.

## 3D Scalar Potential for Saturation-Caused Fields in the Insertion Device

This entails the same approach as for the case of $\mu=\infty$, except that the condition $\partial V / \partial x=0$ at $y=h$ is to be dropped:

$$
\begin{gathered}
V=\sum \cos n k_{z} z \cdot g_{n}(x, y) \\
\nabla^{2} V=0 \quad \Longrightarrow \quad \nabla^{2} g_{n}=n^{2} k_{z}^{2} g_{n}
\end{gathered}
$$

We introduce $n k_{z} x=u$, and $n k_{z} y=v$ :

$$
\begin{equation*}
\nabla_{u, v}^{2} g=g \tag{1}
\end{equation*}
$$

We construct $g(u, v)$ that has the following properties: odd in $y, g(-v)=-g(v)$, and gives field approximating $\cosh \varepsilon u-1$ for $y=0 . \varepsilon$ is arbitrary, real or imaginary, and the field equals 0 for $u=0$ when letting $\varepsilon \rightarrow 0$ at end.
We try $g=\cosh \varepsilon u \sinh a v$. To satisfy (1):

$$
\varepsilon^{2}+a^{2}=1 \text { and thus } a=\sqrt{1-\varepsilon^{2}}
$$

has to hold. We add a function of $v$, such that $g_{v}^{\prime}$ is proportional to $\cosh \varepsilon u-1$ for $v=0$. The only odd function of $v$ that will satisfy this requirement and also satisfy (1) is $-a \sinh v$, thus

$$
\begin{equation*}
g=\cosh \varepsilon u \sinh a v-a \sinh v . \tag{2}
\end{equation*}
$$

One can use the superposition of such functions with different $\varepsilon$, but this would probably not be practical.
The expansion for $\varepsilon \rightarrow 0$ is

$$
\begin{equation*}
g=\frac{\varepsilon^{2}}{2}\left(u^{2} \sinh v-v \cosh v+\sinh v\right) . \tag{3}
\end{equation*}
$$

For $v=0$, we obtain the expected sextupole field:

$$
\begin{equation*}
g_{v}^{\prime}(u, 0)=\frac{\varepsilon^{2}}{2} u^{2} \tag{4}
\end{equation*}
$$

At the pole, where $v_{h}=n k_{z} h$,

$$
\begin{equation*}
g_{u}^{\prime}\left(u, v_{h}\right)=\frac{\varepsilon^{2}}{2} 2 u \sinh v_{h}, \tag{5}
\end{equation*}
$$

It is this field in the $x$-direction that is responsible for the sextupole field in the midplane.

February, 1992. Note 0125u-w.
(5) allows us to make an estimate of the saturation effects in the midplane during the design phase. Thus,

$$
\begin{equation*}
\frac{g_{v}^{\prime}(u, 0)}{g_{u}^{\prime}\left(u, v_{h}\right)}=\frac{u}{2 \sinh v_{h}}=\frac{n k_{z} x}{2 \sinh n k_{z} h}=\frac{x}{2 h} \cdot \frac{n k_{z} h}{\sinh n k_{z} h} \tag{6}
\end{equation*}
$$

It is interesting to note that every additional expansion of (2) in $\varepsilon^{2}$ leads to a new solution to (1) describing the fields in the midplane to the highest orders $\sim x^{4}, x^{6}$, etc.
To check on (3), its expansion in $k_{z}$ up to the 3rd order terms in $\{u, v\}$ gives, as expected,

$$
\begin{align*}
g & =\frac{\varepsilon^{2}}{2}\left(u^{2} v-v\left(1+\frac{v^{2}}{2}-\left(1+\frac{v^{2}}{6}\right)\right)\right) \\
& =\frac{\varepsilon^{2}}{2}\left(u^{2} v-\frac{v^{3}}{3}\right) \\
& =\frac{\varepsilon^{2}}{6} \Im(u+i v)^{3} . \tag{7}
\end{align*}
$$

## Scalar Potential for 3D Insertion Device Fields

In the 2D case,

$$
\begin{equation*}
V=\sum_{n=\text { odd }} b_{n} \cos n k_{z} z \cdot \sinh n k_{z} y \quad \text { with } \quad k_{z}=\frac{2 \pi}{\lambda} \tag{1}
\end{equation*}
$$

In order to simplify matters, we drop the sum, and re-introduce it at the end.
The effects of lateral ends are equally periodic in $z$, thus $n k_{z} \Rightarrow k_{z}$, and

$$
\begin{equation*}
V_{L E}=\cos k_{z} z \cdot g(x, y) \tag{2}
\end{equation*}
$$

Where $g(x, y)$ is valid only in the vacuum region of the $\{x, y\}$ space.
Further, we have that

$$
\begin{equation*}
\nabla^{2} V=0 \quad \Longrightarrow \quad \nabla^{2} g=k_{z}^{2} g \tag{3}
\end{equation*}
$$

where $g$ is the Fourier expansion coefficient as a function of $x, y$,
At the pole surface, for integer $\mu, z=\mu \lambda / 2$ and $y=h=$ half gap, $B_{x}=B_{y}=0$. We expand $g$ in a Fourier series in $y$. We have that $g \sim \sin m k_{y} y$ for $k_{y}=\pi / h$, and

$$
g=\sum a_{m} \sin m k_{y} y \cdot \cosh k_{m} x
$$

with $k_{m}^{2}=k_{z}^{2}+m^{2} k_{y}^{2}$. We use $a_{m}=b_{0} b_{m} / \cosh k_{m} W$, where W is half the pole width, and we expect $b_{m}$ to be only weakly dependent on $W$.

$$
\begin{cases}V=\sum_{n=\text { odd }} b_{n_{0}} \cos n k_{z} z \cdot\left(\sinh n k_{z} y+g_{n}\right),  \tag{4}\\ g_{n}=\sum_{m=1} \frac{b_{n m} \sin m k_{y} y \cdot \cosh k_{n m} x}{\cosh k_{n m} W}, & \text { with } k_{z}=2 \pi / \lambda, k_{y}=\pi / h \\ k_{n m}^{2}=n^{2} k_{z}^{2}+m^{2} k_{y}^{2}, & \text { with } \frac{k_{z}}{k_{y}}=\frac{2 h}{\lambda}=\frac{\text { gap }}{\lambda} \\ k_{n m}^{2}=m^{2} k_{y}^{2}\left(1+\frac{n^{2}}{m^{2}}\left(\frac{k_{z}}{k_{y}}\right)^{2}\right), & \end{cases}
$$

Under most circumstances, $\frac{k_{z}}{k_{y}} \leq .5$. For $n=1, k_{n m} \approx m k_{y}$.

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In the region of interest, only the case of $m=1$ is of importance. That is, the dominant term is

$$
\begin{equation*}
g_{1}=\frac{b_{11} \sin k_{y} y \cdot \cosh k_{11} x}{\cosh k_{11} W} \tag{5}
\end{equation*}
$$

We may now proceed to conclude that

$$
\begin{equation*}
-H_{y}=k_{z} \sum n b_{n_{0}} \cos n k_{z} z\left(\cosh n k_{z} y+\sum b_{n m} \frac{m k_{y}}{n k_{z}} \cdot \frac{\cos m k_{y} y \cdot \cosh k_{n m} x}{\cosh k_{n m} W}\right) \tag{6}
\end{equation*}
$$

From (6), we expect $b_{n m}<0$, and $\frac{m k_{y}}{n k_{z}}\left|b_{n m}\right|$ to be in the order of 1 , but probably less than 1.

## Suggestions for Magnetic Measurements.

Make all measurements as function of $z$, filter out random errors, and then do the harmonic analysis by measuring the quantities derived from $\sinh n k_{z} y+g_{n}$. To measure field components, measure $B_{y}$ at $y=0$ for a number of values of $x$ close enough to the lateral edge to get values of $b_{n 1}$ and $b_{n 2}$. Then measurements of $B_{x}$ close to the lateral ends are made, at $y \simeq h / 2$, to check the validity of $V(x, y, z)$. If agreement is reached, an investigatation of whether $b_{n m}$ are more easily obtained from $B_{x}$ measurements is to be done. To verify the model, compare the measurements at individual points, without the harmonic analysis, to the model calculations.
After sufficient measurements, make a table that lists the $b_{n m}$ coefficients as functions of two dimensionless products (i.e. $h / \lambda$ and $W / h$ ), and possibly find a practical formula to represent the data. A possible complication may result from saturation in the iron which may dominate the behavior of the field as a function of $x$.

Examination of experimental data shows that decay of field errors as one moves away from the lateral edge of the insertion device can be much slower than this description indicates. A possible cause of this may be $H_{x}$ at pole surface caused by saturation.

## Gradient Measurement in Insertion Device

The beam is in the $z$ direction. The midplane is the $\{x, z\}$ plane. We use a vibrating coil to measure $\partial B_{y} / \partial x$.

As a general mechanical design principle, make the wanted resonance frequency and its harmonics different from the resonance frequencies of other vibrating modes.

We want to measure $\partial B_{y} / \partial x$. We move a $B_{y}$-coil in the $x$ direction that is "long" in $z$ and short in $x$. The problem arises that this motion may excite vibration in the $y$ direction, adding a $\partial B_{y} / \partial y$ signal. A better way to collect the same information is to measure $\partial B_{x} / \partial y$, by vibrating a $B_{x}$-coil in the $y$ direction such that it is "long" in $z$ and short in $y$. Possible contamination due to $\partial B_{x} / \partial z$ drops out in the Fourier analysis in $z$.

## Undulator Trajectory and Radiation

We begin with the following definitions:

$$
\begin{gathered}
\dot{\mathbf{r}} \gamma m=-e\left(\mathbf{e}_{x} \dot{x}+\mathbf{e}_{y} \dot{y}+\mathbf{e}_{z} \dot{z}\right) \times \mathrm{e}_{y} B, \\
\bar{x} \gamma m=e \dot{z} B=e \dot{z} A_{z}^{\prime} \\
\dot{x} \gamma m=\epsilon A(z), \\
d t \beta c=d s=d z \sqrt{1+x^{\prime 2}}, \quad \dot{x}=x^{\prime} \frac{\beta c}{\sqrt{1+x^{\prime 2}}}, \\
A=B_{0} \int \cos k z d z=\frac{B_{0}}{k} \sin k z \\
\frac{x^{\prime}}{\sqrt{1+x^{\prime 2}}}= \\
\varepsilon \sin k z, \quad \varepsilon=\frac{B_{0} / k}{\beta \gamma m c / e}=\frac{K}{\gamma} \\
\frac{1}{x^{\prime 2}}+1=\frac{1}{\varepsilon^{2} \sin ^{2} k z}, \quad x^{\prime}=\frac{e s}{\sqrt{1-\varepsilon^{2} \sin ^{2} k z}}
\end{gathered}
$$

Thus,

$$
J=\int \dot{x} e^{i \varphi} d t
$$

where

$$
\begin{gathered}
\varphi=\omega\left(t-\frac{z}{c}\right)=\frac{\omega}{\beta c}(\beta c t-\beta z)=\frac{\omega}{\beta c} \int\left(\sqrt{1+x^{\prime 2}}-\beta\right) d z \\
1+x^{\prime 2}=\frac{1}{1-\varepsilon^{2} \sin ^{2} k z^{2}} \\
\beta=\sqrt{1-\frac{1}{\gamma^{2}}} \simeq 1-\frac{1}{2 \gamma^{2}}
\end{gathered}
$$

and furthermore,

$$
\sqrt{1+x^{\prime 2}}-\beta=1+\frac{\varepsilon^{2} \sin ^{2} k z}{2}-1+\frac{1}{2 \gamma^{2}}=\frac{1}{2 \gamma^{2}}\left(1+\varepsilon^{2} \gamma^{2} \sin ^{2} k z\right)
$$

and

$$
\varphi=\frac{\omega}{2 \beta c \gamma^{2}} \int\left(1+K^{2} \sin ^{2} k z\right) d z=\frac{\omega}{2 \beta c \gamma^{2}} \int\left(1+\frac{K^{2}}{2}-\frac{K^{2}}{2} \cos 2 k z\right) d z
$$

where, for $\beta \approx 1$,

$$
\varphi=\frac{\omega}{2 c \gamma^{2}}\left(1+\frac{K^{2}}{2}\right) z-\frac{\omega}{2 c \gamma^{2}} \frac{K^{2}}{2} \frac{\sin 2 k z}{2 k}
$$

Therefore,

$$
\begin{gathered}
J=\int x^{\prime} d z \cdot e^{i \varphi}, \quad \text { with } \frac{\omega}{c}=k_{L} \\
J=\frac{\varepsilon}{2 i} \int e^{i\left(\frac{k_{J}}{2 \gamma^{2}}\left(1+\frac{K^{2}}{2}\right)-k\right) z-i\left(\frac{k_{L}}{2 \gamma^{2}} \frac{K^{2}}{2} \frac{\sin 2 k z}{2 k}\right)} d z
\end{gathered}
$$

where

$$
\begin{gathered}
\Delta k=\left(\frac{k_{L}}{2 \gamma^{2}}\left(1+\frac{K^{2}}{2}\right)-k\right) \\
e^{i u \sin x}=\sum J_{n}(u) e^{i n x}, \quad \text { with } \quad u=\frac{k_{L} K^{2}}{8 \gamma^{2} k} \quad \text { and } \quad x=k z
\end{gathered}
$$

Thus,

$$
J=\frac{\varepsilon}{2 \dot{i}} \int e^{i(\Delta k z-2 n k z)} J_{n}(u) d z
$$

Further, from

$$
\frac{k_{L}\left(1+K^{2} / 2\right)}{2 \gamma^{2}}-(2 n+1) k=0
$$

and solving for $\frac{k_{L}}{k}$, we have

$$
u=\frac{(n+1 / 2) K^{2} / 2}{1+K^{2} / 2}
$$

## Mathematical Representation of Undulator and Wiggler Fields

Undulator and wiggler fields that are not uniform in the transverse direction are usually derived from

$$
\begin{equation*}
V=V \cosh k_{1} x \sinh k_{2} y \cos z \quad \text { with } \quad k_{1}^{2}+k_{2}^{2}=k^{2} \tag{1}
\end{equation*}
$$

Starting with $\nabla^{2} V$ in cylindrical co-ordinates, we have

$$
r^{2} \nabla^{2} V=\left(r^{2} \frac{\partial^{2}}{\partial r^{2}}+r \frac{\partial}{\partial r}+\frac{\partial^{2}}{\partial \varphi^{2}}+r^{2} \frac{\partial^{2}}{\partial z^{2}}\right) V=0
$$

Assuming, without loss of generality, midplane symmetry, we write

$$
\begin{equation*}
V=\sum F_{n} \sin n \varphi \cos k z \tag{2.1}
\end{equation*}
$$

thus getting

$$
\begin{equation*}
\left(r^{2} \frac{\partial^{2}}{\partial r^{2}}+r \frac{\partial}{\partial r}-n^{2}-k^{2} r^{2}\right) F_{n}=0 \tag{3}
\end{equation*}
$$

and therefore,

$$
\begin{equation*}
F_{n}=a_{n} I_{n}(k r) \tag{2.2}
\end{equation*}
$$

An interesting consequence is that whether one uses (1) or (2), one would get the same fields and pole shapes for a sufficiently small $k r$.

$$
\begin{equation*}
\sum a_{n} I_{n}(k r) \sin n \varphi=V_{0} \cosh \left(k_{1} r \cos \varphi\right) \sinh \left(k_{2} r \sin \varphi\right) \tag{4}
\end{equation*}
$$

and, in particular, this means that

$$
\begin{equation*}
a_{n} I_{n}(k r) \pi=V_{0} \int_{0}^{2 \pi} \cosh \left(k_{1} r \cos \varphi\right) \sinh \left(k_{2} r \sin \varphi\right) \sin n \varphi d \varphi \tag{5}
\end{equation*}
$$

This must hold in particular for $k r \ll 1$, i.e. by comparing the lowest order term in $r$ and executing the trivial integrations, one gets $a_{n}$ which then leads to an extremely interesting integral representation of $I_{n}(k r)$.

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[^0]:    December, 1986. Note 0022bpm.

    * For subscript 2, the quantities are not averaged over time but are time independent, as are all time integrated quantities.

[^1]:    May, 1993. Note 0335 csem.

[^2]:    June, 1981. Note 0076csem.
    $\dagger$ REC can be considered this way if one assumes differential permeability $\equiv 1$.

[^3]:    January, 1994. Note 0611thry.

[^4]:    December, 1993. Note 0609thry.
    $\dagger$ Spencer, J. E. and Enge, H. A., Nuclear Instruments and Methods 49, 181 (1967).

[^5]:    January, 1993. Note 0606thry.

[^6]:    $\dagger$ In the case under discussion here, $B_{z}=0$ at both ends even though $B_{z} \neq 0$ in the fringe field region.

[^7]:    August, 1993. Note 0139u-w.
    $\dagger$ See document $0138 \mathrm{u}-\mathrm{w}$ for the origins of this equation.
    $\ddagger$ See document 0140 u -w for derivations of $\varepsilon_{1}$ and $\varepsilon_{2}$.

[^8]:    1 Table of Functions with Formulae and Curves, Dover Publications, 1945: p. 56.

