Experimental Limits on Massive Neutrinos from e^+e^- Annihilations at 29 GeV

Received by OSTI

MAR 20 1989

C. Akerlof, J. Chapman, D. Errede, M.T. Ken,
D.I. Meyer, H. Neal, D. Nitz, R. Thun, R. Tschirhart
University of Michigan, Ann Arbor, MI 48109

M. Derrick, P. Kooijman, J.S. Loos ^[1],
B. Musgrave, L.E. Price, J. Repond, K. Sugano
Argonne National Laboratory, Argonne, IL 60439

D. Blockus, B. Brabson, J.M. Brom,
C. Jung ^[2] H. Ogren, H.W. Paik, D.R. Rust
Indiana University, Bloomington, IN 47405

S. Abachi, P. Baringer, B.G. Bylsma, R. DeBonte,
D. Koltick, E.H. Low, R.L. McIlwain, D.H. Miller,
C.R. Ng, L.K. Rangan ^[3], E.I. Shibata
Purdue University, W. Lafayette, IN 47907

B. Cork
Lawrence Berkeley Laboratory, Berkeley, CA 94720

ABSTRACT

A search was made in 29 GeV e^+e^- annihilations for massive neutrinos decaying to $e^\pm X^\mp(\nu)$ where X is a muon or meson. A 300pb^{-1} data sample yielded just one candidate event with a mass $m_{ex} > 1.8$ GeV. Significant limits are found for new neutrinos with masses from 1.8 to 6.7 GeV and with mixing parameters in the range $10^{-6} < |U|^2 < 1$.

[1] Present address: Bell Laboratories, Naperville, IL 60566.

[2] Present address: Stanford Linear Accelerator Center, Stanford, CA 94305.

[3] Present address: Lockheed Missiles and Space Co., Sunnyvale, CA 94086.

DISCLAIMER

This report was prepared as an account of work sponsored by an agency of the United States Government. Neither the United States Government nor any agency thereof, nor any of their employees, makes any warranty, express or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service by trade name, trademark, manufacturer, or otherwise does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or any agency thereof. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or any agency thereof.

MASTER

The replication of fermion families, as exemplified by the electron, muon and tau lepton, remains a deep puzzle of particle physics. The mass limits for neutrinos are small compared to the masses of the associated charged leptons,^[1] leading to the interesting possibility that evidence for additional particle families might first appear in the neutrino sector. Indeed, massive neutrinos could provide spectacular experimental signatures in existing detectors.^[2] Electron-positron annihilation is a particularly sensitive reaction for neutrino searches since the process $e^+e^- \rightarrow L^0\bar{L}^0$, mediated by a virtual Z^0 , is flavor conserving and is therefore not constrained by small mixing parameters connecting neutrinos of different families.

We report here final results from a neutrino search carried out with the High Resolution Spectrometers (HRS) at the PEP electron-positron collider at SLAC. The data were obtained at a c.m. energy of 29 GeV and correspond to an integrated luminosity of 300 pb^{-1} . Details of the HRS are given elsewhere.^[3] The spectrometer provided charged-particle tracking and electromagnetic calorimetry over 90% of the full solid angle. The detector did not identify muons. No attempt was made in this search to identify pions, kaons and protons although a limited capability to do so existed within certain momentum intervals.

A search for massive neutrinos can be interpreted only in terms of a specific model for their production and decay. For the production we take the process $e^+e^- \rightarrow Z^0 \rightarrow L^0\bar{L}^0$ where the neutrinos are assumed to have the usual V-A Standard-Model couplings to the virtual Z^0 . The differential production cross section is then given by:^[4]

$$\frac{d\sigma}{d\Omega} = \frac{G_F^2 M_Z^4 s \beta \{ [(1 - 4\sin^2\theta_w)^2 + 1](1 + \beta^2 \cos^2\theta) + 4(1 - 4\sin^2\theta_w)(\beta \cos\theta) \}}{256\pi^2 [(s - M_Z^2)^2 + M_Z^2 \Gamma_Z^2]} \quad (1)$$

where s is the square of the c.m. energy, θ is the neutrino production angle, and βc is the c.m. neutrino speed. Using the values $G_F = 1.166 \times 10^{-5} \text{ GeV}^{-2}$, $M_Z = 92.6 \text{ GeV}$, $\Gamma_Z = 2.8 \text{ GeV}$, $\sin^2\theta_w = 0.226$, and integrating over θ we obtain a total production cross section at $\sqrt{s} = 29 \text{ GeV}$ of $0.36\beta(3 + \beta^2)/4 \text{ pb}$. This

value includes a correction from initial-state radiation which lowers the cross section by 3%.^[5]

The massive neutrinos are assumed to decay, in analogy with quarks, via Standard-Model charged currents because of unitary mixing with the known, light neutrinos. The decay properties are then completely specified by the neutrino mass and the elements of the unitary matrix U that relates the mass and weak eigenstates of the neutrinos. We assume that the mixing is dominated by one particular light neutrino and consider separately the cases $L \rightarrow eW$, $L \rightarrow \mu W$ and $L \rightarrow \tau W$ where W is the virtual, charged vector boson which couples to fermion pairs in the usual manner. The partial decay widths of the massive neutrino to various final states such as $L \rightarrow \ell e \nu$, $\ell \mu \nu$, $\ell \tau \nu$, $\ell \pi$, $\ell \rho$, ℓA , and ℓ hadron continuum, where $\ell = e, \mu$, or τ , have been estimated with the same techniques first used in the evaluation of tau lepton decay modes.^[6] The neutrino lifetime is then given in terms of its mass M and mixing parameter U_ℓ by:

$$t = \frac{192\pi^3}{G_F^2 M^5 |U_\ell|^2} f(M) \quad (2)$$

where $f(M)$ is a calculated factor which summarizes the couplings and effects of phase space associated with all possible decay channels. Lifetimes for $|U_\ell|^2 \equiv 1$ are given in Fig. 1 as a function of the L^0 neutrino mass.

The experimental signature for heavy neutrino production chosen for this study is an event configuration with a pair of oppositely charged tracks, isolated in one hemisphere, where one track is identified in the calorimeter as an electron and the other as a non-showering particle X , so that X is not an electron. This $e^\pm X^\mp$ pair, accompanied by possible light, unobservable neutrinos is a potential signature for the weak decay of a neutral particle. No strong constraint is placed on the recoiling jet of particles opposite the $e^\pm X^\mp$ pair so that the search is for the process $e^+e^- \rightarrow (L^0 \rightarrow e^\pm X^\mp(\nu))(\bar{L}^0 \rightarrow \text{anything})$ and its charge conjugate. No attempt is made to identify the non-showering particle $X^\mp$ because of the limited

capabilities of the HRS. Preliminary results from this search, based on the initial third of the data, were reported previously.^[7]

The following decay modes contribute to final states with the above signature:

a.) $L \rightarrow eW$

$$L^0 \rightarrow e^- \mu^+ \nu$$

$$L^0 \rightarrow e^- (\pi^+ \text{ or } K^+)$$

$$L^0 \rightarrow e^- (\tau^+ \rightarrow \mu^+ \nu \nu) \nu$$

$$L^0 \rightarrow e^- (\tau^+ \rightarrow \pi^+ \nu \text{ or } K^+ \nu) \nu$$

b.) $L \rightarrow \mu W$

$$L^0 \rightarrow \mu^- e^+ \nu$$

$$L^0 \rightarrow \mu^- (\tau^+ \rightarrow e^+ \nu \nu) \nu$$

c.) $L \rightarrow \tau W$

$$L^0 \rightarrow (\tau^- \rightarrow e^- \nu \nu) \mu^+ \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow e^- \nu \nu) (\pi^+ \text{ or } K^+)$$

$$L^0 \rightarrow (\tau^- \rightarrow \mu^- \nu \nu) e^+ \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow \pi^- \nu \text{ or } K^- \nu) e^+ \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow e^- \nu \nu) (\tau^+ \rightarrow \mu^+ \nu \nu) \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow e^- \nu \nu) (\tau^+ \rightarrow \pi^+ \nu \text{ or } K^+ \nu) \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow \mu^- \nu \nu) (\tau^+ \rightarrow e^+ \nu \nu) \nu$$

$$L^0 \rightarrow (\tau^- \rightarrow \pi^- \nu \text{ or } K^- \nu) (\tau^+ \rightarrow e^+ \nu \nu) \nu$$

The total branching ratios for $L^0 \rightarrow \ell W \rightarrow e^\pm X^\mp(\nu)$, where $\ell = e, \mu$, or τ , are shown in Fig. 2 as a function of L^0 mass.

Event selection criteria were chosen to provide good efficiency for detecting possible $e^+e^- \rightarrow L^0 \bar{L}^0$ events while rejecting most known annihilation and two-photon reactions. We list the most important of these criteria with some brief comments:

Charged multiplicity: $n_{CH} \geq 4$ consistent with the assumption that the neutrinos decay via charged currents.

Kinematics: For the $e^\pm X^\mp$ pair, the magnitudes of the momenta satisfy $p_e > 1.0$, $p_x > 1.0$ and $(p_e + p_x) > 4.0$ GeV/c. This requirement was imposed for good electron identification which becomes less discriminating below 1 GeV/c. The opening angle of the pair satisfies $\cos\theta_{ex} > 0.2$ to reject random combinations from two separate jets.

Isolation of $e^\pm X^\mp$ pair: No other charged particles were allowed in the hemisphere centered on the total $e^\pm X^\mp$ pair momentum vector. No photons with $E > 0.1$ GeV were allowed within a cone of half-angle 60° centered on the pair momentum vector. A cone rather than a hemisphere requirement was made on photons since occasionally charged tracks from the recoiling jet would curl back in the 1.6 tesla field of the HRS towards the $e^\pm X^\mp$ hemisphere and give calorimeter signals near the boundary of the hemisphere that are difficult to interpret.

Particle identification: Electrons were identified by demanding that a calorimeter signal with energy E be spatially coincident with a charged track of momentum p such that $0.7 < E/p < 1.3$. Non-showering particles were identified by demanding that they deposit less than 0.5 GeV in the calorimeter where 0.2 GeV is a typical value for a minimum-ionizing particle.

Twenty-seven events satisfied these analysis requirements. Four of these events were kinematically consistent with four-prong QED processes. Since none of the four events had combinations of charged-particle pair masses consistent with the hypothesis $e^+e^- \rightarrow (L^0 \rightarrow e^-\pi^+)(\bar{L}^0 \rightarrow e^+\pi^-)$, they were rejected, leaving 23 candidate events for the neutrino search. For each event, the effective mass of the $e^\pm X^\mp$ pair (m_{ex}) and that of the recoiling jet of particles (m_j) is plotted in Fig. 3. The effective masses were calculated assuming $m_e = m_x = 0$.

The candidate events are characterized by $e^\pm X^\mp$ pairs with low effective mass, with only one pair having a mass above 1.73 GeV. None of the events have an anomalous decay vertex that would be the sign of the decay of a mas-

sive, long-lived neutrino. A number of background sources are expected to yield low-mass $e^\pm X^\mp$ pairs. These include semi-leptonic decays of isolated charmed particles, and low-multiplicity jets where either charged pions and photons overlap to yield fake electrons or where a charged pion interacts in the calorimeter and mimics an electron. In our previous study we found that the number of observed $e^\pm X^\mp$ low mass pairs is consistent with such backgrounds.^[4] We have not made a detailed investigation of possible physical backgrounds for events with high-mass $e^\pm X^\mp$ pairs. Processes that could contribute include $e^+e^- \rightarrow e^+e^-$ hadrons, $e^+e^- \rightarrow \tau^+\tau^-$ hadrons; and production of b-quarks.

We do not make a background subtraction but, rather, set limits on the production of massive neutrinos based on the following consideration: The number of events with $m_{ex} \leq 1.73$ GeV is consistent with background. Let us assume that these are background events. The simple mass contour drawn in Fig. 3 therefore represents an empirical boundary of the background. We then set limits on neutrino production by only searching the mass region to the right of this empirical mass contour. This search region contains either zero or one event, depending on neutrino mass. The validity of this approach to setting limits on neutrinos with mass above ~ 2 GeV is based on the following three points. First, the number of observed low-mass events is consistent with background estimates. Second, the distribution of $e^\pm X^\mp$ pair masses cuts off sharply at the mass of charmed mesons, as expected. Third, the region in the m_{ex} vs. m_j space open to the neutrino search represents a major fraction of the total phase space, so that small variations in the empirical mass contour in Fig. 3 do not have a significant impact on the limits that we find, especially for neutrino masses above ~ 3 GeV.

The one high-mass candidate event is characterized by an e^+X^- pair with $p_e = 11.2$ GeV/c, $p_x = 2.5$ GeV/c, $\cos \theta_{ex} = 0.59$ and $m_{ex} = 4.75$ GeV. If we assume that there is a missing light neutrino with energy of 0.8 GeV, given by the difference of the beam energy and that of the e^+X^- pair, then the effective mass of this neutrino and the e^+X^- pair is kinematically constrained to lie in the range from 4.9 to 8.1 GeV. The measured effective mass of the jet opposite the e^+X^-

pair is 6.39 GeV. This value includes contributions from both charged tracks and from neutral particles detected in the calorimeter. Since the reconstruction of the momentum vectors of neutral particles is not always reliable with a calorimeter of limited segmentation as in HRS, we have also calculated a lower limit on the jet mass using only the charged particle momenta, which are well measured. The charged particles in the jet have an effective mass of 4.9 GeV and a total energy of 8.7 GeV. Assuming that the neutral particles have an energy of $14.5 - 8.7 = 5.8$ GeV, one obtains a lower bound on the total jet mass of 6.43 GeV, which is close to the measured total jet mass. In setting cross section limits for massive neutrinos, we shall take the number of candidate events to be zero below 6.4 GeV and equal to one above this value.

Cross section limits at a given confidence level, CL, and as a function of neutrino mass M and mixing parameter $|U|^2$ are given by:

$$\sigma_{CL}(M, |U|^2) = \frac{N_{CL}(M)}{P_{ex}(M) \epsilon(M, |U|^2) \int L dt} \quad (3)$$

where $N_{90}(1.8 < M < 6.4 \text{ GeV}) = 2.30$ based on zero observed events, and $N_{90}(M > 6.4 \text{ GeV}) = 3.89$ for one candidate. The integrated luminosity $\int L dt$ is 300 inverse picobarn. The probability, $P_{ex}(M)$, for an event of the type $e^+e^- \rightarrow L^0 \bar{L}^0$ to yield at least one $e^\pm X^\mp$ pair is $2BR - (BR)^2$ where BR is the branching ratio for $L^0 \rightarrow e^\pm X^\mp$ shown in Fig. 2. Finally, $\epsilon(M, |U|^2)$ is the detection efficiency for $e^+e^- \rightarrow L^0 \bar{L}^0$ events that satisfy all selection criteria including the mass requirements on m_{ex} and m_J described above. The detection efficiencies depend on both M and $|U|^2$ since the neutrino lifetime is a function of both of these parameters.

Event detection efficiencies were calculated as a function of M and $|U|^2$ using a Monte Carlo program which generated fake events of the type $e^+e^- \rightarrow (L^0 \rightarrow e^\pm X^\mp(\nu)) (\bar{L}^0 \rightarrow \text{anything})$ and its charge conjugate according to the physics assumption stated above. The program generated neutrino decays to three-lepton and pion-lepton final states with the appropriate weak decay matrix elements.^[9]

Other decays, involving hadrons, were generated with an algorithm based on the Lund Monte Carlo program.^[10] These fake events were then converted to fake data using a detailed HRS simulation program. Finally, the fake data were analyzed with the same programs and selection criteria as were the real data.

We obtain limits on neutrinos as a function of both mass M and mixing $|U|^2$ by inserting the calculated detection efficiencies into Eq. 3 above. For prompt decays ($|U|^2 = 1$), event reconstruction efficiencies reach a maximum of about 0.23 at $M = 4$ GeV for $L \rightarrow eW$ and $L \rightarrow \mu W$ and a maximum of about 0.14 at $M = 5$ GeV for $L \rightarrow \tau W$. The $L \rightarrow \tau W$ efficiency is lower since many of the decay modes involve multiple light neutrinos and yield $e^\pm X^\mp$ pair masses that fall inside the mass contour described above. As $|U|^2$ decreases, the proper lifetime of the neutrino increases and is given by the values shown in Fig. 1 divided by $|U|^2$. The detection efficiency decreases with increasing lifetime since the efficiency of the charged-particle tracking algorithm decreases for long decay paths. The efficiency falls roughly linearly with mean decay path length λ and, for neutrino masses around 4 GeV, it is reduced by approximately a factor of two when $\lambda = 0.20\text{m}$. To obtain limits on neutrino production at a particular mass, $|U|^2$ was varied until the experimental cross section limit, which depends on $|U|^2$ by way of the detection efficiency, was equal to the theoretically expected value.

The results are given in Fig. 4 which shows the regions of neutrino mass M and mixing parameter $|U|^2$ excluded by this study. The 90% C.L. contours for $L \rightarrow eW$ and $L \rightarrow \mu W$ are almost identical and have been plotted as one curve in Fig. 4. The 90% C.L. cross section limits for $L \rightarrow \tau W$ are below the theoretical value over such a narrow mass range that the corresponding contour in Fig. 4 would not be very meaningful. We have therefore relaxed the confidence level to 80% for the $L \rightarrow \tau W$ limits and the corresponding results are shown in Fig. 4.

In conclusion, we have searched for massive neutrinos with Standard-Model couplings to the $W^\pm$ and Z^0 bosons and which mix with the known, light neutrinos. We obtain significant limits on the existence of such neutrinos in the mass

range from 1.8 to 6.7 GeV with mixing parameters $|U|^2$ as low as 3×10^{-6} . Other searches in electron-positron annihilations based on identifiable, secondary decay vertices have found significant limits over a similar range of mass and mixing parameters.[11] A general review on limits obtained from other types of experiments is given elsewhere.[12]

This work was supported by the U.S. Department of Energy under contracts W-31-109-ENG-38, DE-AC02-76ER01112, DE-AC03-76SF0098, DE-AC02-76ER01428, and DE-AC02-84ER40125. We express our appreciation to the PEP staff for their help in the experiment. We are particularly grateful to Y.S. Tsai for assisting us in understanding decay properties of heavy neutrinos.

REFERENCES

1. Particle Data Group, *Phys. Lett.* 170B (1986) and references therein.
2. R.E. Shrock, *Phys. Rev.* D24 (1981) 1232.
 R. Thun, *Phys. Lett.* 134B (1984) 459.
 M. Gronau, C.N. Leung, J. Rosner, *Phys. Rev.* D29 (1984) 2539.
 M.L. Perl, SLAC-PUB-3526 (1984).
3. D. Bender et al., *Phys. Rev.* D30 (1984) 515.
4. F.M. Renard, *Basics of Electron Positron Collisions*, (Editions Frontieres, Dreux, 1981).
5. J.D. Jackson and D.L. Scharre, *Nuclear Inst. Meth.* 128 (1975) 13.
6. Y.S. Tsai, *Phys. Rev.* D4 (1971) 2821.
 H.B. Thacker and J.J. Sakurai, *Phys. Lett.* 36B (1971) 103.
 Y.S. Tsai, SLAC-PUB-2450 (1979).
 F.J. Gilman and S.H. Rhie, *Phys. Rev.* D31 (1985) 1066.
7. D. Errede et al., *Phys. Lett.* 149B (1984) 519.
8. In Ref. 7 we estimated 5.5 ± 2.2 background events for an integrated luminosity of 106 pb^{-1} . This extrapolates to 16 ± 6 events for 300 inverse picobarns consistent with the 22 low-mass events observed in the data.
9. F. Bletzacker and H.T. Nieh, *Phys. Rev.* D16 (1977) 2115.
 Equations A2 and A6 of this reference have a sign error in the spin term.
 B. Humpert and P.N. Scharbach, *Phys. Rev.* D16 (1977) 2754.
10. B.A. Anderson et al. *Phys. Rep.* 97 (1983) 33.
11. R. Aleksan, in '86 Massive Neutrinos in Astrophysics and Particle Physics: Proceedings of the 6th Moriond Workshop, ed. O. Fackler and J. Tran Thanh Van (Editions Frontieres, Gif-sur-Yvette, France, 1986) p. 241.
 C. Wendt et al., SLAC-PUB-4177 (1987).
12. F. Gilman, *Comm. Nucl. Part. Phys.* 16 (1986) 231.

FIGURES

- 1.) Expected lifetime of massive neutrinos in the limit of complete mixing, $|U|^2 = 1$. The W in $L \rightarrow eW$, etc., is a virtual $W^\pm$ -boson.
- 2.) Expected branching ratio for the decay of massive neutrinos into $e^\pm X^\mp$ plus possible light neutrinos, as a function of mass. The particle $X^\mp$ is a muon, pion or kaon.
- 3.) Measured mass distribution for events with an isolated $e^\pm X^\mp$ pair recoiling against a jet of particles in the opposite hemisphere. The dashed line represents an empirical boundary used in setting limits on the production of massive neutrinos.
- 4.) Experimental limits on the production of massive neutrinos as a function of mass and mixing parameter $|U|^2$. The region surrounded by the curves is allowed, that outside the boundary is excluded.

FIG 1

FIG 2

FIG 3

FIG 4