ANL/FPP/TM--241

DE89 014375

ANL/FPP/TM-241

ARGONNE NATIONAL LABORATORY
9700 South Cass Avenue
Argonne, Illinois 60439-4801

KAOS/LIB-V: A LIBRARY OF NUCLEAR RESPONSE FUNCTIONS
GENERATED BY KAOS-V CODE
FROM ENDF/B-V AND OTHER DATA FILES

by

Y. Farawila*, Y. Gohar, and C. Maynard**

April 1989

Work supported by

Office of Fusion Energy
U.S. Department of Energy
Under Contract W-31-109-Eng-38

DISCLAIMER

This report was prepared as an account of work sponsored by an agency of the United States Government. Neither the United States Government nor any agency thereof, nor any of their employees, makes any warranty, express or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service by trade name, trademark, manufacturer, or otherwise does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or any agency thereof. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or any agency thereof.

* Nuclear Engineering and Health Physics Programs, Georgia Institute of Technology, Atlanta, Georgia 30332.

** Nuclear Engineering and Engineering Physics Department, University of Wisconsin-Madison, Madison, Wisconsin 53706.

MASTER

DISTRIBUTION OF THIS DOCUMENT IS UNLIMITED *ep*

TABLE OF CONTENTS

	<u>Page</u>
ABSTRACT	1
I. INTRODUCTION.....	2
I.1 General Background.....	2
I.2 Kerma Factor Evaluation Procedure.....	4
II. DESCRIPTION OF THE KAOS/LIB-V LIBRARY CONTENTS.....	7
II.1 List of Materials in KAOS/LIB-V.....	8
II.2 Nuclear Response.....	11
III. SUMMARY OF KERMA FACTOR EVALUATIONS.....	15
III.1 Single-Isotope Materials.....	16
III.2 Kerma Factor Evaluation Summary for Natural Elements.....	30
III.3 Fissionable Materials.....	50
IV. GRAPHICAL REPRESENTATION OF KAOS/LIB-V RESPONSE FUNCTIONS.....	51
REFERENCES	400
APPENDIX A DECAY ENERGY TABLES.....	404
APPENDIX B VITAMIN-E 174 GROUP STRUCTURE.....	411

KAOS/LIB-V: A LIBRARY OF NUCLEAR RESPONSE FUNCTIONS
GENERATED BY KAOS-V CODE
FROM ENDF/B-V AND OTHER DATA FILES

by

Y. Farawila, Y. Gohar, and C. Maynard

ABSTRACT

KAOS/LIB-V: A library of processed nuclear responses for neutronics analyses of nuclear systems has been generated. The library was prepared using the KAOS-V code and nuclear data from ENDF/B-V. The library includes kerma (kinetic energy released in materials) factors and other nuclear response functions for all materials presently of interest in fusion and fission applications for 43 nonfissionable and 15 fissionable isotopes and elements. The nuclear response functions include gas production and tritium-breeding functions, and all important reaction cross sections. KAOS/LIB-V employs the VITAMIN-E weighting function and energy group structure of 174 neutron groups.

Auxiliary nuclear data bases, e.g., the Japanese evaluated nuclear data library JENDL-2 were used as a source of isotopic cross sections when these data are not provided in ENDF/B-V files for a natural element. These are needed mainly to estimate average quantities such as effective Q-values for the natural element. This analysis of local energy deposition was instrumental in detecting and understanding energy balance deficiencies and other problems in the ENDF/B-V data. Pertinent information about the library and a graphical display of the main nuclear response functions for all materials in the library are given.

I. INTRODUCTION

KAOS/LIB-V: a library of processed nuclear responses is described in this report. The library includes neutron kerma-factors, gas production, tritium breeding cross sections, and all important reaction cross sections, for a large number of materials of interest in fusion and fission systems. The library was generated with the KAOS-V code employing basic nuclear data from ENDF/B-V.

The kerma (kinetic energy released in materials) factors are the key nuclear responses in the library because of the importance of the nuclear heating. The kerma factor evaluation procedure, description of the KAOS/LIB-V contents and summary of kerma factor evaluations are given in this report. Also, a graphical display of the main nuclear response functions for materials in the library are shown to help in the nuclear design analysis.

I.1 General Background

Heating is an important parameter of any nuclear system. The energy multiplication in a fusion reactor blanket is a very important factor in the assessment of the economic performance of the power plant. More important are the severe design constraints due to the undesirable heating effects. This is particularly true for the plasma facing components where excessive heating is expected. Other fusion reactor components where heating effects are important are the superconducting magnets. The accurate calculation of heating in the magnet due to the leaking neutrons and photons is a key factor in the design of the cryogenic heat removal systems. Accurate neutron kerma factors are necessary for heating calculations and are also in increasing demand for biomedical applications. Direct experimental measurements of kerma factors are limited to a few isotopes and over the small range of high incident

neutron energies of particular importance to radiation therapy. The microscopic cross section data bases, especially the [ENDF/B files [1] provide the means for computing kerma factors for many materials and for incident neutron energies up to 20 MeV. The computed kerma factors from microscopic cross sections however, suffer from large uncertainties that cruelly manifest themselves sometimes as negative values [2,3]. It is generally believed - justifiably so - that the fundamental improvement in the computed kerma factors rests on improving the nuclear data files. The huge inertia which is involved in adjusting the massive nuclear data files and the fact that these are generally satisfactory for their main purpose - transport calculations - make it difficult to take this route.

Thus, there is a need for improved methods in calculating nuclear heating parameters from the existing data. This work is a response to this need, especially for the fusion community.

1.2 Kerma Factor Evaluation Procedure

The code KAOS-V (Kerma And Other Stuff) was written to be used as the main kerma factor evaluation tool. A brief code description along with a user's guide are given in reference [4]. The evaluation of the recommended neutron kerma factors for a material depends largely on the individual characteristics of that material and the particularities of its evaluated nuclear data. The kerma factor evaluation procedure for each of the 64 materials covered in this work is briefly described in section III. There is however a general procedure that applies to most of these evaluations. This general procedure is described as follows:

- * The ENDF/B-V data for the material under study is reviewed. The nuclear structure, excitation levels, and decay modes for the target and product nuclei are reviewed using the Table of Isotopes. The reaction Q-values are also checked. This review helps primarily in planning the next steps.
- * For materials where resonance parameters are given in ENDF/B-V file 2, an NJOY run [5,6] was made to reconstruct the cross sections and perform Doppler broadening at 300 degree Kelvin. The NJOY module HEATR is also run [6] to make comparisons with subsequent KAOS-V runs. The NJOY PENDF file is kept for these purposes.
- * A KAOS-V run is made in the Direct Energy Balance mode. The partial and total kerma factors and gamma energy production cross section are printed and plotted. The intermediate files in which these results are stored are saved for subsequent comparison tests. The total kerma factors produced in this stage are in good agreement with the NJOY (HEATR) results for most materials.
- * A KAOS-V run is made in the kinematics mode. Again, the partial and total kerma factors are printed, plotted, and saved for the subsequent comparisons.

- * The total kerma factors produced by Direct Energy Balance and Kinematics are compared. Good agreement is expected for materials for which the ENDF/B-V files include detailed gamma production for each reaction. On the other hand, if substantial disagreement was found, further testing and comparison is made in the partial kerma factor level. The interactive mode of the KAOS-V code is frequently used in this stage which facilitates detailed analysis of partial kerma factors and provides the means for performing non-standard tests.
- * Partial kerma factors for each reaction produced by the Direct Energy Balance and Kinematics methods are compared. The gamma energy production is also compared with the excitation energy of the residual nuclei, a common test to identify energy balance problems.
- * For natural elements, the energy-dependent Q-values are estimated using additional data sources as described earlier. The Direct Energy Balance and Kinematics runs are repeated using the variable Q-values. The sensitivity of the total kerma factors to the introduction of the variable Q-values is studied. Depending on this sensitivity, more effort was made in estimating the more important Q-values, and the use of variable Q-values was dropped altogether for some other reactions.
- * After making the necessary data corrections (if any), the information collected through the above testing and comparisons is used to make the final decision on the recommended set of prompt kerma factors. In general, at low incident neutron energies, kerma factors are dominated by radiative capture and elastic scattering. The methods for calculating kerma from these two processes are identical for both Direct Energy Balance and Kinematics approaches. At higher incident neutron energies, the inelastic scattering is

turned on. If the ENDF/B-V file 12 is used to represent inelastic gammas in a manner consistent with file 3 cross sections, the agreement between the two methods will be extended to cover the range of incident neutron energies below the thresholds of other nonelastic reactions. In this incident energy range, the total kerma factors using Direct Energy Balance for many materials experiences a minimum and becomes more sensitive to gamma energy imbalance or representation problems and is often negative, and the Kinematics method is therefore used. For incident neutron energies high enough to produce excited residuals following the emission of charged particles, Direct Energy Balance is generally preferred. The reasons for this preference are the fact that the Kinematics method is not equipped to account for residual excitation and generally overestimates kerma in this range, and also because the magnitude of kerma factors is higher and the sensitivity to gamma errors is therefore reduced. To preserve the continuity of the final recommended set of kerma factors, the connection between the results from the two methods is made at an incident neutron energy at which the two results are in agreement.

* The decay of the unstable products is then studied. The decay heat is included to produce additional kerma response functions as detailed in later sections.

II. DESCRIPTION OF THE KAOS/LIB-V LIBRARY CONTENTS

The nuclear response library KAOS/LIB-V (Kerma And Other Stuff) is generated by using the code KAOS-V [4] as an evaluation tool. The evaluated nuclear data files ENDF/B-V are the chief source of data used in these nuclear response evaluations. Doppler broadening at 300°K was performed for materials where resonance parameters are provided. The effect of Doppler broadening on the group values was very small (< 2.0% in most cases). The library contains 22 response functions for 64 materials. The responses are presented in Vitamin-E 174 group structure, where the first group refers to the lowest neutron energy. The nuclear responses are presented in the units of (eV.barn) for kerma factor responses and gamma energy production, and (barn) for all the other responses. The library is saved in a readable text file. The materials are ordered such that those with lower atomic number are listed first, and lighter isotopes are also presented before the heavier isotopes of the same element. For each material, 22 titled responses are given. FIDO format [7] was used to fill the zero groups before thresholds, and also to zero the responses that do not apply or are not provided for a certain material. This helped reduce the storage size needed for the library. An interactive program, RETRIEVE, was written to retrieve selected responses and perform group collapsing from the Vitamin-E [8] 174 group values.

II.1 List of Materials in KAOS/LIB-V

No.	Material Name	ENDF/B-V MAT number	KAOS/LIB-V Symbol
1	Hydrogen	1301	h-1
2	Deuterium	1302	h-2
3	Helium	1270	he-4
4	Lithium-6	1303	li-6
5	Lithium-7	1397	li-7
6	Beryllium	1304	be-9
7	Boron-10	1305	b-10
8	Boron-11	1160	b-11
9	Carbon	1306	c-12
10	Nitrogen	1275	n-14
11	Oxygen	1276	o-16
12	Fluorine	1309	f-19
13	Sodium	1311	na-23
14	Magnesium	1312	mg-nat
15	Aluminum	1313	al-27
16	Silicon	1314	si-nat
17	Phosphorus	1315	p-31
18	Sulfur	1347	s-nat
19	Chlorine	1149	cl-nat
20	Potassium	1150	k-nat
21	Calcium	1320	ca-nat
22	Titanium	1322	ti-nat
23	Vanadium	1323	v-nat
24	Chromium	1324	cr-nat
25	Manganese	1325	mn-55
26	Iron	1326	fe-nat
27	Cobalt	1327	co-59
28	Nickel	1328	ni-nat
29	Copper	1329	cu-nat

Nc.	Material Name	ENDF/B-V MAT number	KAOS/LIB-V Symbol
30	Zirconium-90	1385	zr-90
31	Zirconium-91	1386	zr-91
32	Zirconium-92	1387	zr-92
33	Zirconium-94	1388	zr-94
34	Zirconium-96	1389	zr-96
35	Niobium	1189	nb-93
36	Molybdenum	1321	mo-nat
37	Hafnium-174	1374	hf-174
38	Hafnium-176	1376	hf-176
39	Hafnium-177	1377	hf-177
40	Hafnium-178	1378	hf-178
41	Hafnium-179	1383	hf-179
42	Hafnium-180	1384	hf-180
43	Tantalum	1285	ta-181
44	Tungsten-182	1475	w-182
45	Tungsten-183	1476	w-183
46	Tungsten-184	1477	w-184
47	Tungsten-186	1478	w-186
48	Lead	1382	pb-nat
49	Bismuth	1375	bi-209
50	Thorium-232	1390	th-232
51	Protactinium-233	1391	pa-233
52	Uranium-233	1393	u-233
53	Uranium-234	1394	u-234
54	Uranium-235	1395	u-235
55	Uranium-236	1396	u-236
56	Uranium-238	1398	u-238
57	Neptunium-237	1337	np-237
58	Plutonium-238	1338	pu-238

No.	Material Name	ENDF/B-V MAT number	KAOS/LIB-V Symbol
59	Plutonium-239	1399	pu-239
60	Plutonium-240	1380	pu-240
61	Plutonium-241	1381	pu-241
62	Plutonium-242	1342	pu-242
63	Americium-241	1361	am-241
64	Americium-243	1363	am-243

II.2 Nuclear Responses

There are 22 nuclear response function tables for each material in the library. These responses are outlined below:

- 1- Recommended Prompt Kerma Factor: The recommended prompt kerma factor can be either calculated by direct energy balance, kinematics, or a combination of these methods in different energy ranges according to the description in the summary provided in section III. This response is provided for all the materials included in this library. The response is given in units of (eV.barn).
- 2- Recommended Kerma Factor with Charged Particle Decay Heat: In this response, the decay heat deposited locally via the emitted charged particles are added to the recommended set of prompt kerma factors. This response is provided for all materials. In case there is no charged particle decay heat (within the half-life cutoff of 1 day), this response will be identical to the recommended prompt kerma.
- 3- Recommended Kerma Factor Plus Total Decay heat: In this response, the total decay heat including energy carried by charged particles and gamma rays are added to the recommended prompt kerma factor response.
- 4- Direct Energy Balance Prompt Kerma Factor: The prompt kerma factors produced by the direct energy balance method in units of (eV.barn) are listed. Notice that in the case gamma files are not provided in ENDF/B-V, the Direct Energy Balance Prompt kerma factor includes the gamma energy as deposited locally. It is possible to find negative kerma factors in this list. These can be used for comparison purposes and in special cases to account for a global energy balance in large systems. This response is presented for most but not all materials, where zeros are entered instead.

- 5- Direct Energy Balance Kerma Factor with Charged Particle Decay Heat: Charged Particle decay heat is added to the direct energy balance prompt kerma in the same way described for the recommended set. If the prompt direct energy balance was zeroed out, this response will be also zeroed out.
- 6- Direct energy Balance Kerma Factor Plus Total Decay Heat: The total decay heat is added to the direct energy balance kerma in the same way described for the recommended set.
- 7- Kinematics Prompt Kerma Factor: The prompt kerma factor produced by the kinematics method is listed. This response is provided for most, but not all materials.
- 8- Kinematics Kerma Factor with Charged Particle Decay Heat: The charged particle decay heat is added to the prompt kinematic kerma factor in the same way described for the recommended set. This response will be zeroed out in the case the prompt kinematic kerma factor is not provided.
- 9- Kinematics Kerma Plus Total Decay Heat: The total decay heat is added to the kinematics prompt kerma in the same way described for the third response. Again, this response will be zeroed out if the prompt kinematic set is not provided, and it will be identical to the prompt set if decay heat (within the 1 day half-life cutoff) is zero.
- 10- Prompt Gamma Energy Production: The total prompt gamma energy production is given in the units of (eV.barn). With the exception of Zirconium isotopes, where gamma files are not provided in the ENDF/B- V, this response is not calculated from kinematics, but rather by direct processing of ENDF/B-V files. In a few cases, interpolation and other minor corrections were made.

- 11- Hydrogen-1 Production Cross Section: In this response the H-1 gas production cross section is given in barns. The response is computed by summing up all proton-producing reaction cross sections multiplied by the respective number of emitted protons. For gas production responses, the delayed component of gas production due to radioactive decay is also added.
- 12- Deuterium Production Cross Section: The production of H-2 gas is presented in the same way as in response 11.
- 13- Tritium Production Cross Section: The production of H-3 is presented in this response in the same way as response 11.
- 14- Helium-3 Production Cross Section: See response 11.
- 15- Helium-4 Production Cross Section: See response 11.
- 16- Total Cross Section: The total cross section is computed from the ENDF/B-V file 2 resonance parameters (if any) and MT=1 file 3 data.
- 17- Elastic Scattering Cross Section: Computed from ENDF/B-V resonance file 2 (if any) and MT=2 file 3 data.
- 18- Total Inelastic Cross Section: This response is equivalent to the data presented in ENDF/B-V file 3 MT=4. The (n,nx) and (n,xn) reactions, where x represents a charged particle e.g. proton or alpha particle, are not included if they are not contained in MT=4 (and equivalently MT=51-91 series).
- 19- Radiative Capture Cross Section: Computed from ENDF/B-V file 2 resonance parameters (if any) and file 3 MT=102 data.
- 20- Total $(n,2n)$ Cross Section: All non-redundant reactions producing two neutrons are added to form this response. These include sequential representation (MT=6-9), direct reaction (MT=16), $(n,2n\alpha)$ reaction (MT=24), and $(n,2n2\alpha)$ reaction (MT=30).

21- Total (n,3n) Cross Section: Reactions producing 3 neutrons are added to form this response. These include (n,3n) and (n,3n α) reactions (MT=17, MT=25).

22- Number of Fission Neutrons * Cross Section: In this response, the total number of fission neutrons (prompt plus delayed) is given. It is computed by multiplying the total number of neutrons produced per fission obtained from ENDF/B-V file 1 by the fission cross section. The fission cross section is computed from file 2 resonance parameters and file 3 MT=18 data.

III. SUMMARY OF KERMA FACTOR EVALUATIONS

In this chapter, the kerma factor evaluation procedure will be summarized for the processed 64 materials included in the KAOS/LIB-V library. The degree of details will vary from one material to another according to the complexity of the analysis that was needed and the amount of data substitution from sources other than ENDF/B-V general files. Since this kind of analysis depends on the critical review of the nuclear data, comments about the data problems will be embedded in the kerma factor evaluation description. The kerma factor evaluation summary of the single-isotope materials will be presented first, followed by that for natural elements, and finally the kerma factor evaluation for fissionable materials.

The abbreviations KIN and DEB are used throughout this summary to stand for KINEMATICS and DIRECT ENERGY BALANCE respectively. Reactions producing charged particles, e.g. (n,p) and (n,α) reactions are referred to as (n,x) reactions. The kerma factors are grouped in Vitamin-E structure, and are ordered such that the first group refers to the group with the lowest neutron energy.

III.1 Single-Isotope Materials

1-H-1

This hydrogen isotope has only two reactions, elastic scattering and radiative capture which simplifies the nuclear response calculation. The calculated kerma factors using the direct energy balance and kinematics are equivalent.

1-H-2

There are only elastic scattering, radiative capture, and (n,2n) reactions for this material. The elastic and capture reactions receive the same treatment in both kinematics and direct energy balance procedures. However, the different treatment of (n,2n) reaction resulted in a KIN total kerma factor greater by a maximum of 34% than the DEB kerma (at 20 MeV incident neutron energy). The DEB result was recommended and adopted for the KAOS/LIB-V library. The KIN result is less reliable because of the indistinguishability error between the first and second secondary neutron energies in ENDF/B-V file 5. This error is proportional to $1/A$ which is particularly large for deuterium where the nuclear mass ratio $A=2$ is small.

2-He-4

Elastic scattering is the only energy deposition mechanism below 20 MeV. The DEB and KIN results are therefore equivalent.

3-Li-6

Gamma production is limited to radiative capture photons, and the 3.562 MeV inelastic photon. The inelastic gamma production is represented in

ENDF/B-V file 12, and the KIN kerma factor is therefore equivalent to that produced by DEB.

The nuclear data for Li-6 was evaluated by I. Bondarenko and Eh. Petrov [9] for the purpose of calculating kerma factors. They published their kerma factor result as a log-scale graph, which made the comparison rather approximate. Their kerma factors were found to be larger than those computed from ENDF/B-V by roughly 20% above the incident neutron energy of 3 MeV.

3-Li-7

Gamma production is limited to radiative capture photons, and the 0.4776 MeV inelastic photon. The inelastic photon is represented in ENDF/B-V file 12 which preserves energy balance. The DEB and KIN kerma results are therefore equivalent.

The nuclear data of Li-7 was evaluated by I. Bondarenko and Eh. Petrov [9] for the purpose of calculating kerma factors. Comparison was made between our results and theirs which were published as a log-scale graph in the incident neutron energy range 0.001-16. MeV. In the keV region, the ENDF/B-V kerma factors including decay heat was found to be in good agreement with that of Bondarenko and Petrov. Above 4 MeV however, their kerma factors were found to be higher than ours by roughly 20%.

4-Be-9

Special care was taken to compute the (n,2n) reaction kerma factor which contributes a significant portion of the total energy deposition. The reaction is represented in ENDF/B-V as a sequence of two events resulting in the formation of the unstable Be-8. The problem lies in the absence of any

information about the possible excitation energy of the product Be-8 before it splits into two alpha particles. The computed KIN kerma factors ignoring the excitation of Be-8 should be underestimated and was actually found to be less than the DEB result at high incident neutron energies. The second neutron's energy obtained from ENDF/B-V file 5 for low incident neutron energy was found to be larger than the kinematic prediction assuming isotropic emission. The reason can be attributed to the anisotropic angular distribution of the second neutron and its correlation to the first one.

DEB method, on the other hand, requires neither intermediate nucleus excitation energy nor angular distribution information. Furthermore, the qualitative analysis of KIN results indicates the superiority and justifies the use of DEB. Although the ENDF/B-V representation of the $(n,2n)$ reaction considers Be-8 as the reaction product, the immediate decay of the unstable Be-8 to two alpha particles was taken to contribute to the prompt kerma.

5-B-10

The DEB results for prompt kerma factors are equal to those for KIN to 0.1% accuracy. This is largely due to the level (n,x) reaction representation in ENDF/B-V which is ideal for consistent gamma file representation. However there is still a need to make some comments about the data representation:

1. For (n,α_0) and (n,α_1) reactions (MT= 780 and 781 respectively) the ENDF/B-V Q-values were found to be zeros. These were corrected to 2.79 and 2.312 MeV respectively.
2. The tenth inelastic level (MT = 60) is flagged to decay by alpha emission (LR = 22). There is however a 6.029 MeV photon in file 13 which indicates a

decay of the same level by gamma emissions. The cross section for this photon's production is too small to worry about the overall accuracy of kerma factors and gas production response functions, but the need to account for branching in the data representation is recognized. The resulting error is significant for other materials, e.g., F-19.

3. The two 0.511 MeV gammas produced due to the decay of the 4th level (n, p_4) by pair production ($LR = 40$) were not accounted for in the gamma files. The resulting error in kerma factors was found to be negligible.

5-B-11

There are no gamma files in ENDF/B-V for this material and kinematics method is the only option. It was not necessary to use an accurate value for η -factor to account for the capture gamma spectrum on the recoil energy. The reason is that the capture kerma is dominated by the decay energy of B-12 with the short half life of 0.02 sec, and $\eta = 1$ was used.

6-C-12

The ENDF/B-V gamma files represent only the photons produced in the inelastic scattering and capture reactions. This makes KIN and DEB results equivalent and the ENDF/B-V is self-consistent. The alpha particle-producing reactions are particularly important at high incident energies. The (n, α_0)Be-9 reaction [18] leaves the Be-9 nucleus in the ground state, and was adequately evaluated by studying the reciprocal reaction Be-9(α, n)C-12. The contribution from (n, α_0) reaction calculated from ENDF/B-V was found to be in agreement with the experimental measurement at 14.1 MeV [11]. The 2.43 MeV excited state of Be-9 decays by disintegration to a neutron and Be-8 which is

unstable in the ground state and splits into two alpha particles. The secondary neutron average energy was calculated from the excitation level and mass difference assuming that all events are isotropic in the center-of-mass. The resulting secondary neutron average energy increases linearly with the incident neutron energy as:

$E_n' = 0.401 + 0.0375 E \text{ MeV}$, which is always larger than the 0.6 MeV computed from the incident energy-independent evaporation spectrum given in ENDF/B-V. The resulting effect on total kerma factor was too small to account for the discrepancies with experimental results. Because of its importance in health physics and medical applications, direct measurement of the carbon kerma factor were carried out above 14 MeV. Results by Barschall, DeLuca and others [10-13] are systematically lower than the computed kerma by 5-25%. These results reveal an overestimation [22] of the evaluated $(n,n'\alpha)$ cross section listed in ENDF/B-V file. The experimental kerma factors were not used in the KAOS/LIB-V library pending confirmation of the diagnosis of the problem and the subsequent modification of data files in the coming version ENDF/B-VI. This is consistent with the methodology adopted here for calculating kerma factors for heating purposes consistent with transport calculations. Other data sources, experimental or evaluated, are used only in the absence of corresponding ENDF/B-V information but are not used to replace them.

7-N-14

The primary capture photon of 10.835 MeV is not isotropically emitted. However, the anisotropy begins at a high incident neutron energy of 1 MeV at which the recoil energy is dominated by the incident neutron impact not the

recoiling photon. The isotropic photon assumption is therefore used without compromising accuracy. In addition, the low energy kerma for this material is dominated by the (n,p) reaction, not capture. Due to the use of ENDF/B-V reaction numbers MT=700's series to represent the (n,p) , (n,t) and (n,α) reactions, a good match (to 4%) between the KIN and DEB results was obtained. The KIN result was adopted because of the possible neglect of $(n,2n)$ gammas which will result in a slightly overestimated DEB kerma.

8-0-16

With the detailed gamma file, it was possible to obtain matching KIN and DEB kerma factors that differ by no more than 0.3%.

9-F-19

The prompt kerma factors at low incident neutron energy were found to be much smaller than those computed by the NJOY module HEATR. This discrepancy is due to the absence of an ENDF/B-V gamma file 12 for radiative capture, and HEATR was not able to perform the kinematic evaluation of capture kerma factors using the alternative file 13, MT=102. DEB and KIN methods results were calculated and compared. The kerma factor values using both methods are almost identical below an incident neutron energy of 1-5 MeV (groups 1 through 120). Between 5.0 and 6.0 MeV incident energy the oscillating error did not exceed a maximum of 20%. Above 6 MeV incident energy, the kerma factors computed by the two methods diverged considerably where DEB results were as large as three times the KIN values. A comparison between gamma energy production from file 13 and excitation energy from file 3 was made. It was found that excitation energy is much larger than gamma production

energy above 6 MeV incident neutron energy - a result cited earlier by MacFarlane [23]. The ENDF/B-V data evaluation was reviewed and the error was found. The inelastic de-excitation flag LR was set to zero for all levels and continuum which indicates decay by gamma emission. Levels of excitation energy of 4.683 MeV and above decay also by alpha and proton emission [24]. The adjustment of the LR flag to indicate particle rather than gamma emission for MT=64 through 91 [25] should solve the problem. The binding energy of the alpha is 4.01 MeV which is much lower than the 7.99 MeV proton binding energy. Also, up to the highest energy level in the Table of Isotopes [24] of 10.411 MeV, no proton emission was indicated. From the above argument, the decay of levels 4.683 MeV and above is dominantly by alpha emission. The flag LR is therefore adjusted to LR=22 for MT=64- 91. After making the LR and the accompanying Q-value changes, the gap between KIN and DEB kerma factor values were greatly narrowed. The new KIN results were found to be larger than the DEB kerma factors by a maximum of 27%. This difference is not large. It also carries the correct sign since KIN kerma factors are expected to be larger than DEB kerma factors at high incident neutron energy when detailed gamma production is not provided and KIN procedures ignore the possible emission of gammas following charged particle reactions. The new DEB kerma factors are therefore recommended for the KAOS/LIB-V library. The change of LR flag does not change the transport cross sections and is still consistent with the model's commitment in that regard. However, users of the KAOS/LIB-V library should be aware, that the helium gas production response function is affected.

11-Na-23

Below the incident neutron energy of 2 MeV, the kerma factors produced by KIN and DEB methods are identical. The difference between the two methods oscillates between 2 MeV and 6.5 MeV incident energies with an amplitude of nearly 30%. This error can be attributed to the representation of gamma production as a function of incident neutron energy using too few points, although the spectral data are adequately represented. Above the incident neutron energy of 6.5 MeV, the ratio between DEB to KIN kerma factors increases gradually until it approaches 2 near 20 MeV incident neutron energy. The problem is less severe but qualitatively similar to the F-19 case described earlier. The Q-value for (n,n'p) and (n,n' α) reactions are -8.79 MeV and -10.47 MeV respectively. Allowing 1-2 MeV for the Coulomb barrier, these reactions are energetically possible above an incident neutron energy of about 10 MeV. Unlike F-19, the above reactions are not considered in ENDF/B-V data for Na-23. Until the possibility for these reactions is quantitatively determined in the entire incident neutron energy range, the KIN results were chosen as the recommended kerma factors. The DEB kerma factors are also kept in the KAOS/LIB-V library. If the (n,n'x) reactions were considered, the resulting kerma should be between the present KIN and DEB values but closer to the KIN estimation. This was verified at the incident neutron energy of 14.6 MeV using the measurement of (n,n'p) cross section $\sigma = 270 \pm 30$ mb by Yu Kozy et al. [26]. The overestimated DEB kerma factor was therefore reduced by $\sigma |Q|$, i.e., reduced from 5.86×10^6 eV.barn to 3.49×10^6 eV.barn. The excitation energy of Na-23 prior to the proton emission was estimated to be 11 MeV. The KIN kerma factor should be increased by $\sigma.(E^* - |Q|)$, i.e., increased from 2.98×10^6 eV.barn to 3.53×10^6

eV.barn. The two new results are different by only 1% which is a surprisingly good agreement.

13-Al-27

Detailed gamma production for each reaction is provided in ENDF/B-V. The DEB and KIN methods are therefore equivalent for all the reactions except for the (n,2n) reaction. The (n,2n) reaction is far from dominating the energy deposition. The KIN and DEB evaluations of kerma factors are found to be close. The DEB result is accepted for KAOS/LIB-V library.

15-P-31

The relative difference between DEB and KIN kerma factors is as small as 1% for incident neutron energies below 1.5 MeV. Between 1.5 MeV and 2.3 MeV incident neutron energy, the DEB kerma factors are less than the KIN kerma factors by a maximum of 18%.

Above the incident neutron energy of 2.3 MeV, the relative difference between the two methods increases to a maximum of 57% indicating underestimated gamma production or an overlooked inelastically excited level decay by particle emission. The inelastic scattering is represented only by a continuum reaction due to the scarcity of experimental data [27]. The KIN results were adopted for prompt kerma factors.

25-Mn-55

The KIN evaluation of kerma factors is adopted for the KAOS/LIB-V library. The KIN evaluation was judged to be superior to the DEB evaluation because of the problems with the gamma files. The capture gamma energy

production was 36-53% too high compared to the available energy. Although gamma heating will be significantly overestimated for low incident neutron energies, the neutron heating is not affected as much. This is particularly valid for this material because capture kerma is dominated by the decay heat of the residual Mn-56. For groups 75 through 105 ($E = 0.13 - 0.67$ MeV), the DEB values are much larger than the KIN ones. This underestimation of gamma production between the incident neutron energy of 0.13 and 0.2 MeV is purely due to the neglect of the 1.258×10^5 eV first level inelastic photon. Between the incident energy of 0.2 MeV and 0.67 MeV, the gamma file which represents capture photons does not contain the inelastic photon because the spectrum is truncated at 0.3 MeV gamma energy. Gamma production is overestimated between the incident neutron energies of 0.57 MeV and 10.0 MeV resulting not only in DEB kerma factors much lower than the corresponding KIN results, but also negative DEB kerma in 40 out of 55 groups in this incident neutron energy range. Above the incident neutron energy of 10 MeV, the DEB kerma factors are found to be larger than the KIN equivalent indicating gamma production underestimation. The evaluated gamma production deviates considerably from the experimental data used for the evaluation [28] and was not helpful in evaluating kerma factors.

27-Co-59

The kerma factors produced by DEB for groups 116 through 120 and 167 through 173 (incident neutron energy between 1.2 - 1.5 MeV, and 14 - 17 MeV) were found to be negative. The observed negative group kerma factors are due to overestimated gamma production. DEB grouped kerma factors much larger than the KIN counterparts were also observed. The overall gamma files are

not as terrible as it may sound, for the absolute gamma production error is only ~20%. Due to the kerma sensitivity to gamma production errors and the inevitable inconsistency when gamma production is evaluated independently from the neutron files, the KIN method was used in the entire incident neutron energy range.

40-Zr-90,91,92,94, and 96

There are no gamma files for any of the zirconium isotopes in ENDF/B-V evaluations. The KIN method had to be used to compute the neutron kerma factors recommended for the KAOS/LIB-V library. Thermal capture gamma spectra and multiplicity for natural zirconium are obtained from the ENDL-2 files. The η -factor of 2.327 was computed from the above data and used in computing capture kerma factors for each zirconium isotope. The resulting capture kerma is better than that obtained using the single-photon approximation even though the individual isotopic η -factors were not available.

Depending on the particular heating application, the DEB kerma factors with gamma energy assumed to be deposited locally could be also used. It was necessary to correct capture Q-values for each isotope, since a single Q-value of 7.98 MeV was given for all the isotopes in ENDF/B-V. The new Q-values for Zr-90, Zr-91, Zr-92, Zr-94 and Zr-96 are 7.1973, 8.635, 6.732, 6.4709 and 5.5809 MeV respectively.

41-Nb-93

Between the inelastic scattering threshold of 0.0293 MeV and 0.1 MeV, the 0.029 MeV photon is not represented in the gamma files. The resulting

error in DEB kerma was confined to 0.4% because energy deposition by elastic scattering is dominant in this energy range. Above the incident neutron energy of 0.1 MeV and up to 1 MeV however, all the 43 group DEB kerma factors are negative. Gamma production in that incident neutron energy range is obtained by interpolating spectral and production cross section data given at 0.1 and 1 MeV. This poorly-represented gamma production would be further overestimated if the provided interpolation schemes were used directly due to the parabolic interpolation problem. Between the incident neutron energy of 1.0 - 4.0 MeV, the error oscillates due to having too few gamma points. The 4 gamma points in file 13 at 1. MeV intervals are not enough to produce the details required for 28 groups. The KIN method was therefore recommended for incident neutron energies up to 4.0 MeV. Above 4 MeV however, 12 gamma points at intervals of 1 or 2 MeV were provided in files 13 and 15 and provide better details for the remaining 33 groups. The DEB method kerma factors were found to be less than the KIN counterparts by an average of 30% above 4 MeV incident neutron energy. This could be explained as due to the subtraction of the (n,x) gamma not accounted for in the KIN method. The recommended kerma factors are computed by the KIN method for the first 141 groups, and by the DEB method for groups 142 through 174.

72-Hf-174,176,177,178,179, and 180

There are no gamma files for any of the Hafnium isotopes in ENDF/B-V evaluations. The KIN method had to be used to compute the neutron kerma factors recommended for the KAOS/LIB-V library. Thermal capture gamma spectra and multiplicity for natural Hafnium are obtained from the ENDL-2 files. The η -factor of 2.446 was computed from the above data and used in

computing capture kerma factors for each Hafnium isotope. The resulting capture kerma factor is better than that obtained using the single-photon approximation even though the individual isotopic η -factors were not available.

Depending on the particular heating application, the DEB kerma factors with gamma energy assumed to be deposited locally could be also used. It was necessary to correct capture Q-values for each isotope, since a single Q-value of 7.18 MeV was given for all the isotopes in ENDF/B-V. The new Q-values for Hf-174, Hf-176, Hf-177, Hf-178, Hf-179, and Hf-180 are 6.789, 6.383, 7.627, 6.099, 7.388, and 5.695 MeV respectively.

73-Ta-181

The KIN method is accepted in the entire incident neutron energy range. The DEB results suffer from the gamma inaccuracy that resulted - when overestimated - in 34 negative kerma groups. Other positive DEB kerma factor group values are generally overestimated by an order of magnitude.

74-W-182, 183, 184, and 186

The ENDF/B-V data file (MAT=1474) for natural tungsten is produced by combining isotopic data files using the NJOY module MIXER. The DEB method was used to calculate kerma factors for natural tungsten (MAT=1474). The data files (MAT=1475, 1476, 1477 and 1478) were used to calculate the individual isotopic kerma factors. The kerma factors for natural tungsten are then obtained from the isotopic kerma by abundance-weighted averaging. Kerma factors obtained for MAT=1474 were found to be negative for groups 110 through 112 and 120 through 123. Isotopic kerma factors for these groups are

all positive (with the exception of W-184 group 122) and the weighted natural tungsten kerma factors are all positive. This discrepancy is rather odd because natural element kerma factors calculated with a single Q -value per reaction tends to be over-estimated. The error - if at all explainable - can be due to the reduced availability of level inelastic series where 40 levels (MT=51-90) were provided for the natural file, compared to a total of 60 levels for the 4 isotopes. The individual isotope approach was therefore chosen. Detailed gamma production is provided for each isotope for inelastic, capture, $(n,2n)$, and $(n,3n)$ reactions. The DEB kerma factor groups 149-151 for W-183 and group 122 for W-184 are negative. No (n,p) or (n,α) gammas were considered which leaves the DEB method without merit relative to the KIN method. The KIN method was therefore used to calculate kerma factors for all the 4 isotopes in the entire incident neutron energy range.

83-Bi-209

Many negative DEB kerma factor groups were found. The problems with gamma files - spotted by MacFarlane [23] are generally confirmed here. With almost 50% chance of producing the isomeric state of Bi-210 in the radiative capture reaction, MacFarlane suggestion of increasing capture gamma multiplicity by 10% needs to be slightly reduced to only 6.5%. The KIN method was used to produce the recommended kerma factors in the entire incident neutron energy range.

III.2 Kerma Factor Evaluation Summary For Natural Elements

12-Mg-nat

The energy-dependent Q-values are evaluated as follows:

(n,2n) Q-value

Below an incident neutron energy of 12 MeV, the only participating isotope is Mg-25, therefore

$$Q(E) = -7.33 \text{ MeV}, \quad \text{Threshold} < E < 12.0 \text{ MeV}.$$

Between 12 and 17 MeV, abundance weighting of Q-values of Mg-25 and Mg-26 results in:

$$Q(E) = -9.2 \text{ MeV}, \quad \text{for } 12.0 < E < 17.0 \text{ MeV}.$$

Above an incident neutron energy of 17 MeV, the cross section σ_{24} for the more abundant isotope Mg-24 is taken from Reference [29] and used to get:

$$Q(E) = (\sigma_{24} a_{24} Q_{24} + (\sigma_{\text{nat}} - \sigma_{24} a_{24}) Q^*) / \sigma_{\text{nat}}$$

where $a_{24} = 0.786$ is the abundance of Mg-24, and Q_{24} is its (n,2n) Q-value. $Q^* = -9.2 \text{ MeV}$ is the Q-value of the Mg-25 and Mg-26 isotopes weighted by their respective abundances. σ_{nat} is the (n,2n) reaction cross section for natural Magnesium provided in ENDF/B-V.

(n,p) Q-value

Threshold energies for the (n,p) reactions for Mg-24 and Mg-26 are estimated from:

$$E_{\text{th}} = |Q| (A + 1) / A + 0.1 z Z \quad \text{MeV}$$

These threshold energies are used to divide the incident neutron energy range where one, two, or three isotopes undergo the (n,p) reaction. In each region, the Q-value is obtained by abundance weighting. The resulting Q-value is tabulated as follows:

E (MeV)	Q (MeV)
4.95	-3.049
6.7	-3.3
7.0	-4.54
9.8	-4.54
10.0	-4.9
20.0	-4.9

(n, α) Q-value

The energy dependence of the (n, α) Q-value is obtained by the same technique used for (n,p) reaction. The resulting Q-value is tabulated as follows:

E (MeV)	Q (MeV)
2.11	0.4818
5.0	0.4818
6.0	-2.16
8.0	-2.16
9.0	-2.5
20.0	-2.5

(n,np) and (n, $n\alpha$) Q-values

The most abundant isotope Mg-24 has the largest Q-value and therefore dominates these reactions. Other isotopes' Q-values are not much smaller than the Mg-24 Q-value. The ENDF/B-V Q-value is used directly where the slight overestimation of the resulting kerma is expected to be minimal.

Decay Heat

Abundance-weighted cross sections are used to calculate the decay kerma factor. The proper adjustments for threshold energies are also made. The (n,2n) cross section for Mg-24 used previously to compute the Q-value is again used for decay calculations.

Recommended Prompt Kerma Factors

The DEB and KIN calculations were made with and without the energy dependent Q-values. Two negative DEB kerma factor groups (121 and 126) were found for both constant and variable Q-values. The use of energy-dependent Q-values resulted in lowering kerma by a maximum of 25%. The KIN evaluation with variable Q-values is recommended for the entire incident neutron energy range.

¹⁴Si-nat

Reaction Q-values provided in ENDF/B-V are equal to those for the most abundant isotope, Si-28 (92%), except for the (n,2n) reaction where the Q-value for Si-29 with 4.7% abundance was used. Because of the high threshold energy for Si-28 (n,2n) reaction of nearly 18 MeV, the effect of the less abundant isotopes on the Q-value needs to be considered for the DEB calculations. Between the threshold group 157 and group 165, the (n,2n) DEB kerma factors using the ENDF/B-V Q-value were found to be negative, while the last 4 groups were found to be much larger than the KIN counterparts. The use of a variable Q-value should bring these overestimated groups closer to the KIN value, but the problem with the negative groups remains unchanged. The KIN kerma factors are formally independent of the Q-value and much less

sensitive to errors in the secondary neutron energy, and is therefore preferred. Energy distributions of secondary proton and alpha particles from $(n,n'p)$ and $(n,n'\alpha)$ reactions are provided in the ENDF/B-V file 5, MT=719 and 799 respectively. One should expect these data to appear in file 25. This is the only material however, that a KIN treatment was possible for charged particle-producing reactions. A good agreement between DEB and KIN results was observed, especially at lower energies. The KIN kerma is larger than DEB by a maximum of 28% for the $(n,n'\alpha)$ reaction and by 15% for the $(n,n'p)$ reaction. Errors in inelastic gamma production result in negative DEB inelastic kerma for groups 140 and 170 through 174. Between groups 142 and 165, the underestimated gamma production results in overestimated DEB inelastic kerma by a maximum of over 100%. The (n,p) reaction was treated by both DEB and KIN methods. The KIN treatment was possible for this material because the level and continuum cross sections are given and the continuum proton energy is given in file 5 MT = 718. The maximum difference between the two results did not exceed 14%.

The (n,α) reaction was examined in a similar manner to the (n,p) reaction. It was found that the KIN kerma factors exceed the DEB kerma factors by ~80% for the first 4 groups (group 136 threshold to 140), but the absolute value of this partial kerma factor is small in that range of higher incident neutron energies. However, the maximum difference does not exceed 23%. The prompt kerma factors produced by the KIN method are recommended.

Decay energy production was computed only for the most abundant isotope Si-28. The decay of Si-27 formed by the $(n,2n)$ reaction is ignored due to the high reaction threshold of ~18 MeV. Only the β^- decay of Al-28 produced by the (n,p) reaction is considered.

16-S-nat

In addition to the natural element evaluation provided in the general ENDF/B-V files under material number MAT=1347, a separate ENDF/B-V evaluation is provided in the dosimetry files for S-32, the most abundant sulfur isotope (95%). These data were used to separate the 'less abundant isotopes' cross sections for the purpose of estimating the variation of the elemental Q-values with incident neutron energy.

(n,2n) Q-values

Between the threshold and 11.5 MeV incident neutron energy, the (n,2n) reaction is energetically possible only with the isotope S-33 and its Q-value of -8.642 is used. In the incident neutron energy range of 11.5 - 15.1 MeV, the abundance-weighted Q-value of S-33 and S-34, $Q^* = -11.0$ MeV is taken. Above 15.1 MeV, the cross section of S-32 and the combined cross section of S-33 and S-34 and their abundances are used to obtain the energy-dependent Q-value as:

$$Q(E) = (\sigma_{32} a_{32} Q_{32} + (\sigma_{\text{nat}} - \sigma_{32} a_{32}) Q^*) / \sigma_{\text{nat}}$$

where $a_{32}=0.9502$ is the abundance of S-32, σ_{32} and σ_{nat} are the (n,2n) microscopic cross sections for S-32 and natural sulfur respectively, $Q_{32}=-15.0424$ MeV is the Q-value for S-32, and $Q^*=-11.0$ MeV is the Q-value of the S-33 and S-34 isotopes obtained by abundance weighting.

(n,n' α) Q-value

The reaction is not included in the ENDF/B-V (MAT = 1316) evaluation for S-32. The Q-value of the most abundant S-32 is the largest in addition to

being close in magnitude to the other isotopes' Q-values. No energy-dependent Q-value was used for this reaction.

(n,np) Q-value

The combined abundance-weighted cross section of S-33 and S-34 was found to oscillate between small positive and negative values when computed from the ENDF/B-V data for natural sulfur and S-32, MAT=1347 and MAT=1316 respectively. The minor isotopes' participation is therefore minimal and the single Q-value for S-32 was kept for kerma factor evaluation.

(n,p) Q-value

Below the S-32 threshold of 1.6 MeV, the S-33 Q-value of 0.53 MeV was used. The Q-value then linearly decreases with incident energy up to 2.0 MeV above which the -0.928 MeV Q-value for S-32 was used.

(n,d) Q-value

The reaction is not considered for S-32 evaluation. The Q-value provided for the elemental evaluation is equal to that of the abundant S-32 and therefore accepted.

(n,t) Q-value

The S-33 Q-value was used for incident neutron energies below the 13.09 MeV threshold of S-32. The S-32 Q-value is taken above 14 MeV with a linear transition allowed between the threshold and 14 MeV.

(n, α) Q-value

Below 1.4 MeV incident energy, the reaction is energetically possible only in S-32 and S-33 isotopes. With the exception of the low energy end where the S-nat evaluation gives a zero cross section while the S-32 evaluation provides a positive value, the two evaluations provided for a formally complete determination of the elemental Q-value by abundance and cross section weighting. Above 1.5 MeV, the Q-value for the most abundant isotope S-32 was used because the two ENDF/B-V evaluations resulted in a combined S-33 and S-34 cross section with many negative points.

(n,2p) Q-value

The reaction is not represented in the S-32 evaluation. The S-32 threshold of 9.9 MeV marked the transition between the S-33 Q-value of -9.00 MeV below the threshold and the S-32 Q-value of -9.57 MeV above the threshold.

Decay Heat

The S-32 evaluation was used to compute decay heat due to S-32 (n,2n) S-31 and S-32 (n,t) P-30 reactions. The S-nat (MAT 1347) evaluation was used with abundance and threshold corrections to compute the decay heat from S-34 (n,p) P-34, S-34 (n, α) Si-31, and S-32 (n,2p) Si-31 reactions.

Prompt Kerma Factors

A correction to ENDF/B-V (MAT = 1347, MF = 12, MT = 102) was made by adding a point (8.99×10^5 eV, 2.82297) to the capture gamma multiplicity table. This corrected the obviously unintended linear decline of the

multiplicity to zero at the 0.9 MeV cutoff. This correction resulted in improving the gamma production but less significant improvement to the capture kerma. The introduction of energy-dependent Q-values resulted in the reduction of the DEB prompt kerma factors by a maximum of 40%. Gamma production is found to be underestimated between 0.9 and 10 MeV incident energy. The KIN prompt kerma factors are accepted for incident neutron energies up to 10 MeV. Above 10 MeV, the DEB method was used to account for possible (n,x) gammas.

17-Cl-nat

Variable Q-values were not needed for this material. The maximum error in the total prompt kerma due to using a single Q-value for the $(n,2n)$ reaction does not exceed 1%. Cross sections for the Cl-35 (n,p) and Cl-35 (n,α) reactions [29] are about an order of magnitude larger than those for Cl-37. This justified keeping ENDF/B-V Q-values which are equal to those of the most abundant isotope Cl-35 for these two reactions. For the $(n,n'p)$ and $(n,n'\alpha)$ reactions, the Q-values of the two isotopes are different by 24% and 11% respectively, but the combined error in total kerma due to using the single Cl-35 Q-value provided in ENDF/B-V is less than 1%. The KIN prompt kerma factors were found to agree with the DEB kerma factors with a maximum error of 15% below the incident neutron energy of 6.4 MeV. Above 6.4 MeV, the DEB kerma factors become increasingly larger than the KIN counterparts up to a factor of 3 near the incident neutron energy of 20 MeV. Only inelastic and capture photons are represented in the gamma files. The underestimation of gamma production for incident neutron energies above 6 MeV could be due to neglecting other nonelastic gammas. The KIN method is used to produce the

recommended kerma factors. The abundance-weighted isotopic cross sections for decay heat calculations are obtained by adjusting the ENDF/B-V cross sections for the isotopic thresholds.

19-K-nat

The utilization of variable Q-values was not necessary because of the high abundance of K-39 (93.3%). An attempt to separate the K-39 (n,p) cross section was made by subtracting the abundance weighted cross section for the K-41 (n,p) reaction found in the partial evaluation ENDF/B-V MAT=7191. The resulting K-39 cross section contained many negative points and was discarded. Level inelastic gamma production between the 2.59 MeV threshold of K-39 and 6.8 MeV matches the inelastic excitations computed from neutron files. Between the incident neutron energy of 1.01 and 6.8 MeV, the continuum inelastic scattering cross section represents the reaction in the less abundant isotope K-41. No inelastic gamma production was represented in that range. Above 6.8 MeV incident energy, the inelastic gamma was underestimated. It seems like the inelastic contribution of K-41 was entirely ignored resulting in overestimating the DEB kerma factors by an average of 10% and a maximum of 39%. The KIN method was therefore used to produce recommended prompt kerma factors. For decay heat calculations, abundance ratios were used to obtain isotopic cross sections from elemental (n,2n), (n, γ) and (n, α) cross sections. The (n, α) cross section was trimmed at the threshold of K-41. The partial evaluation MAT = 7191 was used to compute decay heat resulting from the K-41 (n,p) Ar-41 reaction.

20-Ca-nat

The most abundant isotope Ca-40 with abundance 96.9% dominates the ENDF/B-V elemental evaluation. Its Q-values are used for all reactions

except for the (n,2n) reaction where the higher Q-value of Ca-48 was used to include a small contribution from this isotope. This overestimated Q-value did not affect the accuracy of the total kerma. With the detailed gamma production provided in ENDF/B-V file 12, the relative difference between the DEB and KIN results did not exceed 1% for most groups, and the DEB kerma factors are accepted to fill the recommended kerma factor field in the KAOS/LIB-V library. The small decay heat contribution from Ca-40 (n,2n) Ca-39 and Ca-40 (n,t) K-38 reactions was computed. The threshold energy for the (n,2n) reaction was restored to 16.0 MeV for the purpose of the decay heat calculations.

22-Ti-nat

In addition to the natural element evaluation in ENDF/B-V, MAT=1322, there are partial evaluations for Ti-46, Ti-47, Ti-48, and Ti-50 represented by material numbers MAT = 7226, 7227, 7228, and 7220 respectively. With no dominating isotope with large abundance, the estimation of the variation of the elemental Q-values becomes necessary. These are evaluated as follows:

(n,2n) Q-value

Prompt DEB kerma factors were found to be sensitive to changes in the (n,2n) Q-values at high incident neutron energy. An initial estimation of the energy dependence of the Q-value was made. This estimation depended on approximating the weighting isotopic cross sections with magnitudes proportional to their abundances as a ramp function that starts at their respective threshold energies. A better estimation of the energy-dependent Q-value was made by taking the elemental cross section function weighted by

abundance and translating it to begin at each isotope's threshold. This latter estimation of the Q-value variation with energy was found to be in good agreement with that obtained by using the evaluated isotopic cross sections from reference [30]. The (n, γ) reaction Q-value used in this evaluation is tabulated below:

E (MeV)	Q (MeV)
threshold	-8.14
11.0	-8.4
12.0	-8.6
13.0	-9.6
16.0	-10.5
20.0	-11.0

(n,p) Q-value

The ENDF/B-V zero Q-value was replaced. The isotopic cross sections are available only for Ti-46, Ti-47, and Ti-48 from ENDF/B-V partial evaluations. Since the Ti-49 Q-value is close to that of Ti-46, both abundances were combined and the modified abundances along with the three available cross sections are used to compute the reaction Q-value before the threshold of Ti-50. The effect of Ti-50 on the Q-value was assumed to be small and the previous evaluation was extended to the incident neutron energy of 20 MeV.

(n, α) Q-value

The ENDF/B-V zero Q-value was replaced. The new energy-dependent Q-value was constructed using threshold energies and abundance weighting. The resulting Q-value is tabulated below:

E (MeV)	Q (MeV)
1.0E-11	1.357
3.0	1.357
4.0	0.79
5.0	0.7
6.0	0.0
7.0	-1.0
7.5	-1.2
10.0	-1.3
20.0	-1.35

Other reaction Q-values were not changed. The reason is that the reaction cross section is small or the variation of isotopic Q-values is sufficiently small.

Isotopic cross sections for producing radioactive decay heat were obtained as follows: the Ti-46 (n,2n) cross section is obtained from reference [31]. Threshold-adjusted, abundance-weighted elemental cross sections were used to compute decay heat from the other reactions as needed. The DEB and KIN methods were used to compute prompt kerma factors with and without the variable Q-values. There is an oscillating DEB/KIN error for groups 114 through 148. There are also five DEB negative groups (129 through 133). The KIN result was accepted for groups 1 through 148 i.e., up to the incident neutron energy of 6.0 MeV. Above 6.0 MeV, DEB results were taken to account for the (n,x) gamma production.

23-V-NAT

The only stable isotope is V-51. Although the abundance of the long lived isotope V-50 is only 0.25%, its Q-values for (n,p) and (n, α) reactions are larger than those for the principal isotope V-51 [32] and are listed in the ENDF/B-V evaluation. The adjusted Q-values are taken such that the V-50

Q-values are considered only below the V- 51 threshold. The elemental Q-value for (n,p) reaction used in this evaluation is tabulated below:

E (MeV)	Q (MeV)
1.0E-11	2.995
1.717	2.995
1.8	-1.6837
20.0	-1.6837

The (n, α) Q-value is similarly tabulated as follows:

E (MeV)	Q (MeV)
1.0E-11	0.7577
2.096	0.7577
2.1	-2.0546
20.0	-2.0546

The comparison between the DEB and KIN prompt kerma factor results shows good agreement for groups 1 through 137, i. e., up to the incident neutron energy of 3.0 MeV. The removal of a parabolic interpolation overestimation error in gamma production reduced the number of negative DEB kerma factor groups from 5 to 4 (groups 139- 142). The difference between both calculations fluctuates to a maximum of $\pm 50\%$ for groups 144 through 162. The KIN results were therefore accepted up to the incident neutron energy of 12.0 MeV (groups 1 through 163). Above 12.0 MeV, the DEB kerma factors are less than the KIN counterparts by an average of 30% reflecting the consideration of (n,x) gamma production, and is therefore accepted in that range. For decay heat calculation, the (n,p) reaction cross section was set to zero before the threshold energy of V-51.

24-Cr-nat

The gamma production cross section in ENDF/B-V file 13, MT=3 is zero at the incident neutron energy of 20 MeV. The situation was corrected by

inserting the value of 2.4267 barn taken from reference [33]. Isotopic evaluations for the four chromium isotopes were found in the JENDL-2 library. These data are used to evaluate cross section and abundance-weighted Q-values for the (n,p) and (n, α) reactions. Threshold adjustments were needed to fit the principal ENDF/B-V data source. Other reaction cross sections found in ENDF/B-V elemental evaluations were not accounted for in some or all JENDL isotopic evaluations. The overestimation of the DEB kerma factors due to use of the single ENDF/B-V Q-values for these reactions is expected to be small. The use of energy-dependent Q-values affect the KIN kerma factors above group 145. The largest reduction amounts to 24% in group 153. Large gamma production errors result in large errors in the DEB kerma factors with 33 negative groups. The KIN method was therefore used to produce the recommended prompt kerma factors in the entire incident neutron energy range.

Abundance weighting is used for decay heat computations. This was done because not all the radioactivity-producing reactions are accounted for in the JENDL-2 evaluations, and also because of the large discrepancies between ENDF/B-V and the combined elemental cross sections from JENDL-2.

26-Fe-nat

In addition to the elemental ENDF/B-V.2 evaluation, other data sources were considered for the kerma factor evaluation. Among those are the isotopic ENDF/B-V dosimetry files and INDL-V partial evaluations, JENDL-2 isotopic evaluations, and the Los Alamos evaluation [34] of Fe-54 and Fe-56.

Variable Q-values

The isotopic evaluations of JENDL-2 were used to construct the capture Q-value as a function of incident neutron energy. On the other hand, an

abundance-weighted single Q-value was listed in the ENDF/B-V.2 evaluation. The maximum relative difference between the energy-dependent Q-value and the single ENDF/B-V.2 value was 2.65% in the 0.8 - 2.0 MeV incident neutron energy range. This small difference does not account for the energy imbalance problems in that range contrary to the explanation provided by Larson (reference [35] page 25). The ENDF/B-V.2 capture Q-value was used. The (n,p) Q-value listed in ENDF/B-V.2 is equal to that of the Fe-54 isotope with only a 5.8% abundance. The elemental (n,p) cross section was constructed from JENDL-2, INDL-V, partial ENDF/B-V MAT = 6430 and MAT = 6431 evaluations, and the Los Alamos Fe-54 and Fe-56 evaluations. All the constructed elemental cross section are found to be in good agreement with the ENDF/B-V.2 value. Because JENDL-2 is the most complete of all data sources where all isotopic data were found, these data were used to evaluate the variable (n,p) Q-value. The maximum reduction in the total prompt kerma factor due to using a variable (n,p) Q-value amounts to 19% at the incident neutron energy of 10 MeV. The (n, α) Q-value in ENDF/B-V.2 is also equal to that of Fe-54. Construction of the elemental (n, α) cross section from the available isotopic data sources was done. The INDL-V data were not sufficient, while the JENDL-2 reconstructed elemental (n, α) cross section was much larger than that of ENDF/B-V.2. The constructed elemental (n, α) cross section from Los Alamos data is also larger than that of ENDF/B-V.2 but with better agreement. The variable (n,d) Q-value evaluated from the Los Alamos evaluation data was found to affect the total prompt kerma by no more than 1% and was subsequently dropped. Other reaction Q-values were also studied and the listed ENDF/B-V.2 values were found adequate. Only the (n,p) reaction variable Q-values needed to be considered. Gamma production files in the revision 2 of ENDF/B-V.2 data were modified in response to MacFarlane's work

on energy imbalance [23,35]. The new gamma production was modified above 2 MeV to "reproduce" the estimated heating, which makes these gamma files particularly suspicious for the purpose of evaluating kerma factors. A thorough review of gamma production was therefore necessary. Apart from flaws irrelevant to energy balance, e.g., the branching ratios for MT = 57 which are taken opposite to those of the Table of Isotopes, other energy balance problems were found. Inelastic gamma represented in file 12 are not only resulting from the excitation of the respective levels by inelastic scattering of the corresponding file 3 data, but also the excitation of these levels due to gamma decay of higher levels. Although this practice reduces data volume a little, the energy balance is no longer as strict as expected from file 12 representation. A comparison between excitation energy production and inelastic gamma production shows that the former was larger by a maximum of 9%. In the incident neutron energy range 0.85 - 2.122 MeV, the gamma production in the ENDF/B-V files MF=13, MF=15 represents capture photons. As mentioned earlier, the energy imbalance that produced negative kerma factors in that range can not be due to elemental Q-value problems. The elemental (n,γ) cross section constructed from JENDL-2 data is in better qualitative agreement with the ENDF/B-V.2 capture gamma production than the ENDF/B-V.2 (n,γ) cross section. The possibly underestimated ENDF/B-V.2 (n,γ) cross section needs to be reviewed in this range, and more gamma resolution than the provided 4 point histogram is needed.

The difference between the DEB and KIN results alternates between positive and negative values for groups 110-146. The DEB kerma factor of group 136 is negative. Although the DEB problem below the incident neutron energy of 2.122 MeV is due to a mismatch between neutron and gamma files, the

coarse histogram gamma production representation is the source of the DEB error between the incident energy of 2.122 and 5.5 MeV. The KIN prompt kerma factors are therefore recommended below 5.5 MeV. Above 5.5 MeV, the DEB kerma factors are accepted to account for the (n,x) gamma production.

For decay heat calculations, the isotopic cross sections from the JENDL-2 files were normalized to reproduce the ENDF/B-V.2 cross sections. No radioactivity-producing reaction in Fe-57 was found in JENDL-2 and considering its small abundance, its decay heat contribution was ignored. Fe-58 (n,p) and (n, α) reactions are the only radioactivity-producing reactions found for this isotope in JENDL-2 data and their small contribution was considered.

28-Ni-nat

In addition to the ENDF/B-V elemental evaluation, JENDL-2 isotopic evaluations for the 5 Nickel isotopes were used. Variable Q- values for (n,p) and (n, α) reactions were computed from JENDL-2 data. The maximum reduction of total kerma factors due to using these variable Q- values amounts to only 5%. The effect of the other reactions' variable Q-values was assumed to be small and was not considered. The DEB and KIN results are in good agreement up to group 113 where the error does not exceed 3.8%. Between group 114 and 141, the difference oscillates between positive and negative values with 3 negative DEB kerma factor groups 121, 122, and 127. The reason for this DEB inaccuracy is the coarse histogram gamma production representation. In that range, there are only 6 histogram intervals which do not provide the resolution needed for the 28 groups. Therefore for the first 141 groups, up to an incident neutron energy of 4.0 MeV, the KIN results are recommended. Above 4.0 MeV, there are 14 histogram intervals for gamma

production which provide better resolution for the remaining 33 groups. To account for (n,x) gamma production, the DEB kerma factors are recommended above 4.0 MeV.

Decay heat production is produced by reactions with the low- abundance isotopes Ni-61, Ni-62, and Ni-64. The JENDL-2 cross sections for these isotopes were therefore used for decay heat production calculations directly without normalization to elemental ENDF/B-V values.

29-Cu-nat

There are only two copper natural isotopes Cu-63 and Cu-65 with abundances of 69.2% and 30.8% respectively. This made the unique determination of isotopic data possible if data for either isotope is available. ENDF/B-V activation files provide for partial isotopic evaluations which are prepared by the same evaluator, C. Fu, and therefore consistent with the elemental evaluation. The Cu-63 (n, γ), (n,p) and (n,d) cross sections were obtained from ENDF/B-V MAT = 7293, while the Cu-65 (n,2n) cross section was obtained from ENDF/B-V, MAT = 7295. The (n,n' α) cross sections for both isotopes are taken from JENDL-2 and normalized to ENDF/B-V elemental values. The Cu-65 (n,n'p) cross section is taken from reference [36] and is tabulated below:

Neutron Energy E (MeV)	Cu-65 (n,n'p) cross section (barn)
8.0	0.0
9.5	1.0E-3
10.5	6.0E-3
12.0	7.0E-3
13.5	1.7E-2
14.5	1.8E-2
16.0	3.0E-2
20.0	6.0E-2

The isotopic cross section data mentioned above were used to compute variable Q-values for $(n,2n)$, $(n,n'\alpha)$, $(n,n'p)$, (n,p) , and (n,α) reactions. Variable Q-values for (n,d) and $(n,He3)$ reactions are not considered because their maximum effect can not exceed 7.0×10^3 and 6.0×10^3 eV.barn respectively. Away from the sensitive groups near the negative DEB kerma factor groups 169 and 170, the maximum reduction in DEB kerma factors due to variable Q-values amounts to 75% at the last group 174. The DEB kerma factor is negative in the incident neutron energy range of 1.25 through 1.65 MeV i. e. groups 118 through 122. Kerma factor groups 169 and 170 are also negative. Comparison with KIN results shows that the difference between kerma factors computed by the two methods oscillates between positive and negative values in the entire incident neutron energy range. The gamma production represents only capture, inelastic, and $(n,2n)$ photons, i. e. there is no representation for possible (n,x) gammas. The KIN method is therefore used to produce the recommended kerma factors. The isotopic cross sections are used to compute decay heat production.

42-Mo-nat

In the elemental ENDF/B-V evaluation, the only reactions represented are the elastic, continuum inelastic, $(n,2n)$, $(n,3n)$ and capture. While the (n,p) and (n,α) reactions are ignored in the ENDF/B-V evaluation, non-trivial cross sections of the order of 25 mb and 10 mb respectively are to be found in JENDL-2 evaluations. DEB kerma factors with constant Q-values are negative for 97 out of 174 groups. The use of variable Q-values would introduce more negative groups. The DEB kerma is particularly sensitive to gamma production errors because there are no (n,x) reactions in the

representation. The accuracy of gamma files can not be judged before detailed level inelastic, (n,p) , and (n,α) reactions are introduced to the neutron files. Since the KIN method is well suited to treat all the listed reactions in ENDF/B-V without the need of any Q-value information, this method was used to produce the recommended kerma factors.

For decay heat calculations, the Mo-92 $(n,2n)$ cross section was taken from JENDL-2 and normalized to ENDF/B-V values. Abundance weighting was used to compute heating from the Mo-100 (n,γ) reaction.

82-Pb-nat

In the elemental ENDF/B-V evaluation, the only reactions represented are elastic and inelastic scattering, $(n,2n)$, $(n,3n)$, and radiative capture. Small $(n,n'p)$, $(n,n'\alpha)$, (n,p) , and (n,α) reaction cross sections are found in the JENDL-2 evaluations. With constant Q-values, the DEB kerma factors were found to be negative for 26 groups. With variable Q-values generated from JENDL-2 for $(n,2n)$ and $(n,3n)$ reactions, 4 more groups became negative. The negative kerma error can be attributed to the DEB kerma factor sensitivity to gamma production because there are no charged particle-producing reactions in the representation. Improvements in that area is expected to have a good impact on DEB kerma factor accuracy. On the other hand, the KIN treatment of the represented reactions is complete and independent of Q-values. The KIN kerma factors are therefore recommended.

For decay heat computations, JENDL-2 isotopic cross sections normalized to elemental ENDF/B-V are used.

III.3 Fissionable Materials

Kerma factors were evaluated for 15 fissionable materials listed in section II.1. The kerma factors for these materials are dominated by the contribution of the fission reaction. This is particularly true for fissile materials where thermal neutrons are able to induce fission. The other reactions whose cross sections are given in ENDF/B-V are capture, elastic scattering, inelastic scattering, $(n,2n)$, and $(n,3n)$ reactions. For fertile materials, the fission effective threshold energy is lower than that for the $(n,2n)$ and $(n,3n)$ reactions. Below the fission threshold, the kerma factor will be dominated by capture and scattering interactions. Since capture and scattering treatment is equivalent for both DEB and KIN general approaches, and the fission kerma factor computation method is not a kinematic one, it was appropriate to apply the DEB method to all fissionable materials. It should be noted however, that gamma files are not provided in the ENDF/B-V data for some materials. These materials include Pa-233, U-234, U-236, Np-237, and Pu-238. In this case, the computed DEB kerma factors are larger than the KIN ones because they include the unwanted component of gamma energy from fission and other reactions. The KIN results are therefore recommended for the KAOS/LIB-V library for these materials. Both DEB and KIN kerma factors are saved for all the fissionable materials in the KAOS/LIB-V library.

The radioactive decay contribution to the local energy deposition was calculated from the delayed beta component of the fission energy release given in ENDF/B-V file 1. The charged particle decay energy from radioactive nuclei formed by non-fission reactions was also added. Similar treatment was given to gamma decay heat component.

IV. GRAPHICAL REPRESENTATION OF KAOS/LIB-V RESPONSE FUNCTIONS

In this section, graphs of all nuclear response functions are presented for all the materials included in the KAOS/LIB-V library.

Only the recommended set of prompt and total kerma factors are plotted. The log-scale graphs show the low incident neutron energy part in more detail. The linear scale graphs show the high incident neutron energy part more clearly.

H-1 RECOMMENDED PROMPT KERMA FACTOR

H-1 RECOMMENDED PROMPT KERMA FACTOR

H-1 PROMPT GAMMA ENERGY PRODUCTION

H-1 TOTAL CROSS SECTION

H-1 ELASTIC CROSS SECTION

H-1 RADIATIVE CAPTURE CROSS SECTION

H-2 RECOMMENDED PROMPT KERMA FACTOR

H-2 RECOMMENDED PROMPT KERMA FACTOR

H-2 PROMPT GAMMA ENERGY PRODUCTION

H-2 HYDROGEN-1 PRODUCTION CROSS SECTION

H-2 TRITIUM PRODUCTION CROSS SECTION

H-2 TOTAL CROSS SECTION

H-2 ELASTIC CROSS SECTION

H-2 RADIATIVE CAPTURE CROSS SECTION

H-2 TOTAL (n,2n) REACTION CROSS SECTION

HE-4 RECOMMENDED PROMPT KERMA FACTOR

HE-4 RECOMMENDED PROMPT KERMA FACTOR

HE-4 TOTAL CROSS SECTION

HE-4 ELASTIC CROSS SECTION

LI-6 PROMPT GAMMA ENERGY PRODUCTION

LI-6 HYDROGEN-1 PRODUCTION CROSS SECTION

LI-6 DEUTERIUM PRODUCTION CROSS SECTION

LI-6 TRITIUM PRODUCTION CROSS SECTION

LI-6 HELIUM-4 PRODUCTION CROSS SECTION

LI-6 TOTAL CROSS SECTION

LI-6 ELASTIC CROSS SECTION

LI-6 TOTAL INELASTIC CROSS SECTION

LI-6 RADIATIVE CAPTURE CROSS SECTION

LI-6 TOTAL (n,2n) REACTION CROSS SECTION

LI-7 PROMPT GAMMA ENERGY PRODUCTION

LI-7 HYDROGEN-1 PRODUCTION CROSS SECTION

LI-7 DEUTERIUM PRODUCTION CROSS SECTION

LI-7 TRITIUM PRODUCTION CROSS SECTION

LI-7 HELIUM-4 PRODUCTION CROSS SECTION

LI-7 TOTAL CROSS SECTION

LI-7 ELASTIC CROSS SECTION

LI-7 TOTAL INELASTIC CROSS SECTION

LI-7 RADIATIVE CAPTURE CROSS SECTION

LI-7 TOTAL (n,2n) REACTION CROSS SECTION

LI-7 TOTAL (n,3n) CROSS SECTION

BE-9 PROMPT GAMMA ENERGY PRODUCTION

BE-9 HYDROGEN-1 PRODUCTION CROSS SECTION

BE-9 HELIUM-4 PRODUCTION CROSS SECTION

BE-9 TOTAL CROSS SECTION

BE-9 ELASTIC CROSS SECTION

BE-9 RADIATIVE CAPTURE CROSS SECTION

BE-9 TOTAL (n,2n) REACTION CROSS SECTION

B-10 RECOMMENDED PROMPT KERMA FACTOR

B-10 RECOMMENDED PROMPT KERMA FACTOR

B-10 PROMPT GAMMA ENERGY PRODUCTION

B-10 HYDROGEN-1 PRODUCTION CROSS SECTION

B-10 DEUTERIUM PRODUCTION CROSS SECTION

B-10 TRITIUM PRODUCTION CROSS SECTION

B-10 HELIUM-4 PRODUCTION CROSS SECTION

B-10 TOTAL CROSS SECTION

B-10 ELASTIC CROSS SECTION

B-10 TOTAL INELASTIC CROSS SECTION

B-10 RADIATIVE CAPTURE CROSS SECTION

B-11 HYDROGEN-1 PRODUCTION CROSS SECTION

B-11 TRITIUM PRODUCTION CROSS SECTION

B-11 HELIUM-4 PRODUCTION CROSS SECTION

B-11 TOTAL CROSS SECTION

B-11 ELASTIC CROSS SECTION

B-11 TOTAL INELASTIC CROSS SECTION

B-11 RADIATIVE CAPTURE CROSS SECTION

B-11 TOTAL (n,2n) REACTION CROSS SECTION

C-12 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)

C-12 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)

C-12 PROMPT GAMMA ENERGY PRODUCTION

C-12 HYDROGEN-1 PRODUCTION CROSS SECTION

C-12 DEUTERIUM PRODUCTION CROSS SECTION

C-12 HELIUM-4 PRODUCTION CROSS SECTION

C-12 TOTAL CROSS SECTION

C-12 ELASTIC CROSS SECTION

C-12 TOTAL INELASTIC CROSS SECTION

C-12 RADIATIVE CAPTURE CROSS SECTION

N-14 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

N-14 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

N-14 PROMPT GAMMA ENERGY PRODUCTION

N-14 HYDROGEN-1 PRODUCTION CROSS SECTION

N-14 DEUTERIUM PRODUCTION CROSS SECTION

N-14 TRITIUM PRODUCTION CROSS SECTION

N-14 HELIUM-4 PRODUCTION CROSS SECTION

N-14 TOTAL CROSS SECTION

N-14 ELASTIC CROSS SECTION

N-14 TOTAL INELASTIC CROSS SECTION

N-14 RADIATIVE CAPTURE CROSS SECTION

N-14 TOTAL (n,2n) REACTION CROSS SECTION

0-16 PROMPT GAMMA ENERGY PRODUCTION

0-16 HYDROGEN-1 PRODUCTION CROSS SECTION

O-16 DEUTERIUM PRODUCTION CROSS SECTION

O-16 HELIUM-4 PRODUCTION CROSS SECTION

0-16 TOTAL CROSS SECTION

0-16 ELASTIC CROSS SECTION

O-16 TOTAL INELASTIC CROSS SECTION

O-16 RADIATIVE CAPTURE CROSS SECTION

F-19 PROMPT GAMMA ENERGY PRODUCTION

F-19 HYDROGEN-1 PRODUCTION CROSS SECTION

F-19 DEUTERIUM PRODUCTION CROSS SECTION

F-19 TRITIUM PRODUCTION CROSS SECTION

F-19 HELIUM-4 PRODUCTION CROSS SECTION

F-19 TOTAL CROSS SECTION

F-19 ELASTIC CROSS SECTION

F-19 TOTAL INELASTIC CROSS SECTION

F-19 RADIATIVE CAPTURE CROSS SECTION

F-19 TOTAL (n,2n) REACTION CROSS SECTION

NA-23 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

NA-23 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

NA-23 PROMPT GAMMA ENERGY PRODUCTION

NA-23 HYDROGEN-1 PRODUCTION CROSS SECTION

NA-23 HELIUM-4 PRODUCTION CROSS SECTION

NA-23 TOTAL CROSS SECTION

NA-23 ELASTIC CROSS SECTION

NA-23 TOTAL INELASTIC CROSS SECTION

NA-23 RADIATIVE CAPTURE CROSS SECTION

NA-23 TOTAL (n,2n) REACTION CROSS SECTION

MG-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

MG-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

MG-NAT PROMPT GAMMA ENERGY PRODUCTION

MG-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

MG-NAT HELIUM-4 PRODUCTION CROSS SECTION

MG-NAT TOTAL CROSS SECTION

MG-NAT ELASTIC CROSS SECTION

MG-NAT TOTAL INELASTIC CROSS SECTION

MG-NAT RADIATIVE CAPTURE CROSS SECTION

MG-NAT TOTAL (n,2n) REACTION CROSS SECTION

AL-27 PROMPT GAMMA ENERGY PRODUCTION

AL-27 HYDROGEN-1 PRODUCTION CROSS SECTION

AL-27 DEUTERIUM PRODUCTION CROSS SECTION

AL-27 TRITIUM PRODUCTION CROSS SECTION

AL-27 HELIUM-4 PRODUCTION CROSS SECTION

AL-27 TOTAL CROSS SECTION

AL-27 ELASTIC CROSS SECTION

AL-27 TOTAL INELASTIC CROSS SECTION

AL-27 RADIATIVE CAPTURE CROSS SECTION

AL-27 TOTAL (n,2n) REACTION CROSS SECTION

SI-NAT PROMPT GAMMA ENERGY PRODUCTION

SI-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

SI-NAT DEUTERIUM PRODUCTION CROSS SECTION

SI-NAT HELIUM-4 PRODUCTION CROSS SECTION

SI-NAT TOTAL CROSS SECTION

SI-NAT ELASTIC CROSS SECTION

SI-NAT TOTAL INELASTIC CROSS SECTION

SI-NAT RADIATIVE CAPTURE CROSS SECTION

SI-NAT TOTAL (n,2n) REACTION CROSS SECTION

P-31 PROMPT GAMMA ENERGY PRODUCTION

P-31 HYDROGEN-1 PRODUCTION CROSS SECTION

P-31 HELIUM-4 PRODUCTION CROSS SECTION

P-31 TOTAL CROSS SECTION

P-31 ELASTIC CROSS SECTION

P-31 TOTAL INELASTIC CROSS SECTION

P-31 RADIATIVE CAPTURE CROSS SECTION

P-31 TOTAL (n,2n) REACTION CROSS SECTION

S-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

S-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

S-NAT PROMPT GAMMA ENERGY PRODUCTION

S-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

S-NAT DEUTERIUM PRODUCTION CROSS SECTION

S-NAT TRITIUM PRODUCTION CROSS SECTION

S-NAT HELIUM-4 PRODUCTION CROSS SECTION:

S-NAT TOTAL CROSS SECTION

S-NAT ELASTIC CROSS SECTION

S-NAT TOTAL INELASTIC CROSS SECTION

S-NAT RADIATIVE CAPTURE CROSS SECTION

S-NAT TOTAL (n,2n) REACTION CROSS SECTION

CL-NAT PROMPT GAMMA ENERGY PRODUCTION

CL-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

CL-NAT HELIUM-4 PRODUCTION CROSS SECTION

CL-NAT TOTAL CROSS SECTION

CL-NAT ELASTIC CROSS SECTION

CL-NAT TOTAL INELASTIC CROSS SECTION

CL-NAT RADIATIVE CAPTURE CROSS SECTION

CL-NAT TOTAL (n,2n) REACTION CROSS SECTION

K-NAT PROMPT GAMMA ENERGY PRODUCTION

K-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

K-NAT HELIUM-4 PRODUCTION CROSS SECTION

K-NAT TOTAL CROSS SECTION

K-NAT ELASTIC CROSS SECTION

K-NAT TOTAL INELASTIC CROSS SECTION

K-NAT RADIATIVE CAPTURE CROSS SECTION

K-NAT TOTAL (n,2n) REACTION CROSS SECTION

CA-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

CA-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

CA-NAT PROMPT GAMMA ENERGY PRODUCTION

CA-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

CA-NAT DEUTERIUM PRODUCTION CROSS SECTION

CA-NAT TRITIUM PRODUCTION CROSS SECTION

CA-NAT TOTAL CROSS SECTION

CA-NAT ELASTIC CROSS SECTION

CA-NAT TOTAL INELASTIC CROSS SECTION

CA-NAT RADIATIVE CAPTURE CROSS SECTION

CA-NAT TOTAL (n,2n) REACTION CROSS SECTION

TI-NAT PROMPT GAMMA ENERGY PRODUCTION

TI-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

TI-NAT DEUTERIUM PRODUCTION CROSS SECTION

TI-NAT TRITIUM PRODUCTION CROSS SECTION

TI-NAT HELIUM-3 PRODUCTION CROSS SECTION

TI-NAT HELIUM-4 PRODUCTION CROSS SECTION

TI-NAT TOTAL CROSS SECTION

TI-NAT ELASTIC CROSS SECTION

TI-NAT TOTAL INELASTIC CROSS SECTION

TI-NAT RADIATIVE CAPTURE CROSS SECTION

TI-NAT TOTAL (n,2n) REACTION CROSS SECTION

TI-NAT TOTAL (n,3n) CROSS SECTION

V-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

V-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

V-NAT PROMPT GAMMA ENERGY PRODUCTION

V-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

V-NAT DEUTERIUM PRODUCTION CROSS SECTION

V-NAT TRITIUM PRODUCTION CROSS SECTION

V-NAT HELIUM-4 PRODUCTION CROSS SECTION

V-NAT TOTAL CROSS SECTION

V-NAT ELASTIC CROSS SECTION

V-NAT TOTAL INELASTIC CROSS SECTION

V-NAT RADIATIVE CAPTURE CROSS SECTION

V-NAT TOTAL (n,2n) REACTION CROSS SECTION

CR-NAT PROMPT GAMMA ENERGY PRODUCTION

CR-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

CR-NAT DEUTERIUM PRODUCTION CROSS SECTION

CR-NAT TRITIUM PRODUCTION CROSS SECTION

CR-NAT HELIUM-3 PRODUCTION CROSS SECTION

CR-NAT HELIUM-4 PRODUCTION CROSS SECTION

CR-NAT TOTAL CROSS SECTION

CR-NAT ELASTIC CROSS SECTION

CR-NAT TOTAL INELASTIC CROSS SECTION

CR-NAT RADIATIVE CAPTURE CROSS SECTION

CR-NAT TOTAL (n,2n) REACTION CROSS SECTION

CR-NAT TOTAL (n,3n) CROSS SECTION

MN-55 PROMPT GAMMA ENERGY PRODUCTION

MN-55 HYDROGEN-1 PRODUCTION CROSS SECTION

MN-55 DEUTERIUM PRODUCTION CROSS SECTION

MN-55 HELIUM-3 PRODUCTION CROSS SECTION

MN-55 HELIUM-4 PRODUCTION CROSS SECTION

MN-55 TOTAL CROSS SECTION

MN-55 ELASTIC CROSS SECTION

MN-55 TOTAL INELASTIC CROSS SECTION

MN-55 RADIATIVE CAPTURE CROSS SECTION

MN-55 TOTAL (n,2n) REACTION CROSS SECTION

FE-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

FE-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

FE-NAT PROMPT GAMMA ENERGY PRODUCTION

FE-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

FE-NAT DEUTERIUM PRODUCTION CROSS SECTION

FE-NAT TRITIUM PRODUCTION CROSS SECTION

FE-NAT HELIUM-3 PRODUCTION CROSS SECTION

FE-NAT HELIUM-4 PRODUCTION CROSS SECTION

FE-NAT TOTAL CROSS SECTION

FE-NAT ELASTIC CROSS SECTION

FE-NAT TOTAL INELASTIC CROSS SECTION

FE-NAT RADIATIVE CAPTURE CROSS SECTION

FE-NAT TOTAL (n,2n) REACTION CROSS SECTION

CD-59 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

CD-59 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

CO-59 PROMPT GAMMA ENERGY PRODUCTION

CO-59 HYDROGEN-1 PRODUCTION CROSS SECTION

CO-59 DEUTERIUM PRODUCTION CROSS SECTION

CO-59 TRITIUM PRODUCTION CROSS SECTION

CO-59 HELIUM-3 PRODUCTION CROSS SECTION

CO-59 HELIUM-4 PRODUCTION CROSS SECTION

CO-59 TOTAL CROSS SECTION

CO-59 ELASTIC CROSS SECTION

CO-59 TOTAL INELASTIC CROSS SECTION

CO-59 RADIATIVE CAPTURE CROSS SECTION

CO-59 TOTAL (n,2n) REACTION CROSS SECTION

NI-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

NI-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

NI-NAT PROMPT GAMMA ENERGY PRODUCTION

NI-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

NI-NAT DEUTERIUM PRODUCTION CROSS SECTION

NI-NAT HELIUM-4 PRODUCTION CROSS SECTION

NI-NAT TOTAL CROSS SECTION

NI-NAT ELASTIC CROSS SECTION

NI-NAT TOTAL INELASTIC CROSS SECTION

NI-NAT RADIATIVE CAPTURE CROSS SECTION

NI-NAT TOTAL (n,2n) REACTION CROSS SECTION

CU-NAT PROMPT GAMMA ENERGY PRODUCTION

CU-NAT HYDROGEN-1 PRODUCTION CROSS SECTION

CU-NAT DEUTERIUM PRODUCTION CROSS SECTION

CU-NAT HELIUM-3 PRODUCTION CROSS SECTION

CU-NAT HELIUM-4 PRODUCTION CROSS SECTION

CU-NAT TOTAL CROSS SECTION

CU-NAT ELASTIC CROSS SECTION

CU-NAT TOTAL INELASTIC CROSS SECTION

CU-NAT RADIATIVE CAPTURE CROSS SECTION

CU-NAT TOTAL (n,2n) REACTION CROSS SECTION

ZR-90 RECOMMENDED PROMPT KERMA FACTOR

ZR-90 RECOMMENDED PROMPT KERMA FACTOR

ZR-90 PROMPT GAMMA ENERGY PRODUCTION

ZR-90 HYDROGEN-1 PRODUCTION CROSS SECTION

ZR-90 ELASTIC CROSS SECTION

ZR-90 TOTAL INELASTIC CROSS SECTION

ZR-90 RADIATIVE CAPTURE CROSS SECTION

ZR-90 TOTAL (n,2n) REACTION CROSS SECTION

ZR-91 RECOMMENDED PROMPT KERMA FACTOR

ZR-91 RECOMMENDED PROMPT KERMA FACTOR

ZR-91 PROMPT GAMMA ENERGY PRODUCTION

ZR-91 HYDROGEN-1 PRODUCTION CROSS SECTION

ZR-91 HELIUM-4 PRODUCTION CROSS SECTION

ZR-91 TOTAL CROSS SECTION

ZR-91 ELASTIC CROSS SECTION

ZR-91 TOTAL INELASTIC CROSS SECTION

ZR-91 RADIATIVE CAPTURE CROSS SECTION

ZR-91 TOTAL (n,2n) REACTION CROSS SECTION

ZR-92 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-92 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-92 PROMPT GAMMA ENERGY PRODUCTION

ZR-92 HYDROGEN-1 PRODUCTION CROSS SECTION

ZR-92 HELIUM-4 PRODUCTION CROSS SECTION

ZR-92 TOTAL CROSS SECTION

ZR-92 ELASTIC CROSS SECTION

ZR-92 TOTAL INELASTIC CROSS SECTION

ZR-92 RADIATIVE CAPTURE CROSS SECTION

ZR-92 TOTAL (n,2n) REACTION CROSS SECTION

ZR-94 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-94 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-94 PROMPT GAMMA ENERGY PRODUCTION

ZR-94 HYDROGEN-1 PRODUCTION CROSS SECTION

ZR-94 HELIUM-4 PRODUCTION CROSS SECTION

ZR-94 TOTAL CROSS SECTION

ZR-94 ELASTIC CROSS SECTION

ZR-94 TOTAL INELASTIC CROSS SECTION

ZR-94 RADIATIVE CAPTURE CROSS SECTION

ZR-94 TOTAL (n,2n) REACTION CROSS SECTION

ZR-96 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-96 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

ZR-96 PROMPT GAMMA ENERGY PRODUCTION

ZR-96 HELIUM-4 PRODUCTION CROSS SECTION

ZR-96 TOTAL CROSS SECTION

ZR-96 ELASTIC CROSS SECTION

ZR-96 TOTAL INELASTIC CROSS SECTION

ZR-96 RADIATIVE CAPTURE CROSS SECTION

ZR-96 TOTAL (n,2n) REACTION CROSS SECTION

NB-93 RECOMMENDED PROMPT KERMA FACTOR

NB-93 RECOMMENDED PROMPT KERMA FACTOR

NB-93 PROMPT GAMMA ENERGY PRODUCTION

NB-93 HYDROGEN-1 PRODUCTION CROSS SECTION

NB-93 PROMPT GAMMA ENERGY PRODUCTION

NB-93 HYDROGEN-1 PRODUCTION CROSS SECTION

NB-93 HELIUM-4 PRODUCTION CROSS SECTION

NB-93 TOTAL CROSS SECTION

NB-93 ELASTIC CROSS SECTION

NB-93 TOTAL INELASTIC CROSS SECTION

NB-93 RADIATIVE CAPTURE CROSS SECTION

NB-93 TOTAL (n,2n) REACTION CROSS SECTION

NB-93 TOTAL (n,3n) CROSS SECTION

MO-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

MO-NAT RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

MO-NAT PROMPT GAMMA ENERGY PRODUCTION

MO-NAT TOTAL CROSS SECTION

MO-NAT ELASTIC CROSS SECTION

MO-NAT TOTAL INELASTIC CROSS SECTION

MO-NAT RADIATIVE CAPTURE CROSS SECTION

MO-NAT TOTAL (n,2n) REACTION CROSS SECTION

HF-174 RECOMMENDED PROMPT KERMA FACTOR

HF-174 RECOMMENDED PROMPT KERMA FACTOR

HF-174 PROMPT GAMMA ENERGY PRODUCTION

HF-174 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-174 TOTAL CROSS SECTION

HF-174 ELASTIC CROSS SECTION

HF-174 TOTAL INELASTIC CROSS SECTION

HF-174 RADIATIVE CAPTURE CROSS SECTION

HF-174 TOTAL (n,2n) REACTION CROSS SECTION

HF-176 RECOMMENDED PROMPT KERMA FACTOR

HF-176 RECOMMENDED PROMPT KERMA FACTOR

HF-176 PROMPT GAMMA ENERGY PRODUCTION

HF-176 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-176 TOTAL CROSS SECTION

HF-176 ELASTIC CROSS SECTION

HF-176 TOTAL INELASTIC CROSS SECTION

HF-176 RADIATIVE CAPTURE CROSS SECTION

HF-176 TOTAL (n,2n) REACTION CROSS SECTION

HF-177 RECOMMENDED PROMPT KERMA FACTOR

HF-177 RECOMMENDED PROMPT KERMA FACTOR

HF-177 PROMPT GAMMA ENERGY PRODUCTION

HF-177 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-177 TOTAL CROSS SECTION

HF-177 ELASTIC CROSS SECTION

HF-177 TOTAL INELASTIC CROSS SECTION

HF-177 RADIATIVE CAPTURE CROSS SECTION

HF-177 TOTAL (n,2n) REACTION CROSS SECTION

HF-178 PROMPT GAMMA ENERGY PRODUCTION

HF-178 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-178 TOTAL CROSS SECTION

HF-178 ELASTIC CROSS SECTION

HF-178 TOTAL INELASTIC CROSS SECTION

HF-178 RADIATIVE CAPTURE CROSS SECTION

HF-178 TOTAL (n,2n) REACTION CROSS SECTION

HF-179 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

HF-179 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

HF-179 PROMPT GAMMA ENERGY PRODUCTION

HF-179 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-179 TOTAL CROSS SECTION

HF-179 ELASTIC CROSS SECTION

HF-179 TOTAL INELASTIC CROSS SECTION

HF-179 RADIATIVE CAPTURE CROSS SECTION

HF-179 TOTAL (n,2n) REACTION CROSS SECTION

HF-180 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

HF-180 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

HF-180 PROMPT GAMMA ENERGY PRODUCTION

HF-180 HYDROGEN-1 PRODUCTION CROSS SECTION

HF-180 TOTAL CROSS SECTION

HF-180 ELASTIC CROSS SECTION

HF-180 TOTAL INELASTIC CROSS SECTION

HF-180 RADIATIVE CAPTURE CROSS SECTION

HF-180 TOTAL (n,2n) REACTION CROSS SECTION

TA-181 RECOMMENDED PROMPT KERMA FACTOR

TA-181 RECOMMENDED PROMPT KERMA FACTOR

TA-181 PROMPT GAMMA ENERGY PRODUCTION

TA-181 HYDROGEN-1 PRODUCTION CROSS SECTION

TA-181 TOTAL CROSS SECTION

TA-181 ELASTIC CROSS SECTION

TA-181 TOTAL INELASTIC CROSS SECTION

TA-181 RADIATIVE CAPTURE CROSS SECTION

TA-181 TOTAL (n,2n) REACTION CROSS SECTION

TA-181 TOTAL (n,3n) CROSS SECTION

W-182 RECOMMENDED PROMPT KERMA FACTOR

W-182 RECOMMENDED PROMPT KERMA FACTOR

W-182 PROMPT GAMMA ENERGY PRODUCTION

W-182 HYDROGEN-1 PRODUCTION CROSS SECTION

W-182 HELIUM-4 PRODUCTION CROSS SECTION

W-182 TOTAL CROSS SECTION

W-182 ELASTIC CROSS SECTION

W-182 TOTAL INELASTIC CROSS SECTION

W-182 RADIATIVE CAPTURE CROSS SECTION

W-182 TOTAL (n,2n) REACTION CROSS SECTION

W-182 TOTAL (n,3n) CROSS SECTION

W-183 RECOMMENDED PROMPT KERMA FACTOR

W-183 RECOMMENDED PROMPT KERMA FACTOR

W-183 PROMPT GAMMA ENERGY PRODUCTION

W-183 HYDROGEN-1 PRODUCTION CROSS SECTION

W-183 HELIUM-4 PRODUCTION CROSS SECTION

W-183 TOTAL CROSS SECTION

W-183 ELASTIC CROSS SECTION

W-183 TOTAL INELASTIC CROSS SECTION

W-183 RADIATIVE CAPTURE CROSS SECTION

W-183 TOTAL (n,2n) REACTION CROSS SECTION

W-183 TOTAL (n,3n) CROSS SECTION

W-184 PROMPT GAMMA ENERGY PRODUCTION

W-184 HYDROGEN-1 PRODUCTION CROSS SECTION

W-184 HELIUM-4 PRODUCTION CROSS SECTION

W-184 TOTAL CROSS SECTION

W-184 ELASTIC CROSS SECTION

W-184 TOTAL INELASTIC CROSS SECTION

W-184 RADIATIVE CAPTURE CROSS SECTION

W-184 TOTAL (n,2n) REACTION CROSS SECTION

W-184 TOTAL (n,3n) CROSS SECTION

W-186 PROMPT GAMMA ENERGY PRODUCTION

W-186 HYDROGEN-1 PRODUCTION CROSS SECTION

W-186 ELASTIC CROSS SECTION

W-186 TOTAL INELASTIC CROSS SECTION

W-186 RADIATIVE CAPTURE CROSS SECTION

W-186 TOTAL (n,2n) REACTION CROSS SECTION

W-186 TOTAL (n,3n) CROSS SECTION

PB-NAT PROMPT GAMMA ENERGY PRODUCTION

PB-NAT TOTAL CROSS SECTION

PB-NAT ELASTIC CROSS SECTION

PB-NAT TOTAL INELASTIC CROSS SECTION

PB-NAT RADIATIVE CAPTURE CROSS SECTION

PB-NAT TOTAL (n,2n) REACTION CROSS SECTION

PB-NAT TOTAL (n,3n) CROSS SECTION

BI-209 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)

BI-209 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)

BI-209 PROMPT GAMMA ENERGY PRODUCTION

BI-209 HYDROGEN-1 PRODUCTION CROSS SECTION

BI-209 HELIUM-4 PRODUCTION CROSS SECTION

BI-209 TOTAL CROSS SECTION

BI-209 ELASTIC CROSS SECTION

BI-209 TOTAL INELASTIC CROSS SECTION

BI-209 RADIATIVE CAPTURE CROSS SECTION

BI-209 TOTAL (n,2n) REACTION CROSS SECTION

BI-209 TOTAL (n,3n) CROSS SECTION

TH-232 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

TH-232 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

TH-232 PROMPT GAMMA ENERGY PRODUCTION

TH-232 TOTAL CROSS SECTION

TH-232 ELASTIC CROSS SECTION

TH-232 TOTAL INELASTIC CROSS SECTION

TH-232 RADIATIVE CAPTURE CROSS SECTION

TH-232 TOTAL (n,2n) REACTION CROSS SECTION

TH-232 TOTAL (n,3n) CROSS SECTION

TH-232 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

PA-233 TOTAL CROSS SECTION

PA-233 ELASTIC CROSS SECTION

PA-233 TOTAL INELASTIC CROSS SECTION

PA-233 RADIATIVE CAPTURE CROSS SECTION

PA-233 TOTAL (n,2n) REACTION CROSS SECTION

PA-233 TOTAL (n,3n) CROSS SECTION

PA-233 NUMBER OF FISSION NEUTRONS * CROSS SECTION

U-233 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

U-233 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

U-233 PROMPT GAMMA ENERGY PRODUCTION

U-233 TOTAL CROSS SECTION

U-233 ELASTIC CROSS SECTION

U-233 TOTAL INELASTIC CROSS SECTION

U-233 RADIATIVE CAPTURE CROSS SECTION

U-233 TOTAL (n,2n) REACTION CROSS SECTION

U-233 TOTAL (n,3n) CROSS SECTION

U-233 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

U-234 TOTAL CROSS SECTION

U-234 ELASTIC CROSS SECTION

U-234 TOTAL INELASTIC CROSS SECTION

U-234 RADIATIVE CAPTURE CROSS SECTION

U-234 TOTAL (n,2n) REACTION CROSS SECTION

U-234 TOTAL (n,3n) CROSS SECTION

U-234 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

U-235 PROMPT GAMMA ENERGY PRODUCTION

U-235 TOTAL CROSS SECTION

U-235 ELASTIC CROSS SECTION

U-235 TOTAL INELASTIC CROSS SECTION

U-235 RADIATIVE CAPTURE CROSS SECTION

U-235 TOTAL (n,2n) REACTION CROSS SECTION

U-235 TOTAL (n,3n) CROSS SECTION

U-235 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

U-236 TOTAL CROSS SECTION

U-236 ELASTIC CROSS SECTION

U-236 TOTAL INELASTIC CROSS SECTION

U-236 RADIATIVE CAPTURE CROSS SECTION

U-236 TOTAL (n,2n) REACTION CROSS SECTION

U-236 TOTAL (n,3n) CROSS SECTION

U-236 NUMBER OF FISSION NEUTRONS * CROSS SECTION

U-238 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

U-238 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

U-238 PROMPT GAMMA ENERGY PRODUCTION

U-238 TOTAL CROSS SECTION

U-238 ELASTIC CROSS SECTION

U-238 TOTAL INELASTIC CROSS SECTION

U-238 RADIATIVE CAPTURE CROSS SECTION

U-238 TOTAL (n,2n) REACTION CROSS SECTION

U-238 TOTAL (n,3n) CROSS SECTION

U-238 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

NP-237 TOTAL CROSS SECTION

NP-237 ELASTIC CROSS SECTION

NP-237 TOTAL INELASTIC CROSS SECTION

NP-237 RADIATIVE CAPTURE CROSS SECTION

NP-237 TOTAL (n,2n) REACTION CROSS SECTION

NP-237 TOTAL (n,3n) CROSS SECTION

NP-237 NUMBER OF FISSION NEUTRONS * CROSS SECTION

PU-238 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-238 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-238 TOTAL CROSS SECTION

PU-238 ELASTIC CROSS SECTION

PU-238 TOTAL INELASTIC CROSS SECTION

PU-238 RADIATIVE CAPTURE CROSS SECTION

PU-238 TOTAL (n,2n) REACTION CROSS SECTION

PU-238 TOTAL (n,3n) CROSS SECTION

PU-238 NUMBER OF FISSION NEUTRONS * CROSS SECTION

PU-239 PROMPT GAMMA ENERGY PRODUCTION

PU-239 TOTAL CROSS SECTION

PU-239 ELASTIC CROSS SECTION

PU-239 TOTAL INELASTIC CROSS SECTION

PU-239 RADIATIVE CAPTURE CROSS SECTION

PU-239 TOTAL (n,2n) REACTION CROSS SECTION

PU-239 TOTAL (n,3n) CROSS SECTION

PU-239 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

PU-240 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-240 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-240 PROMPT GAMMA ENERGY PRODUCTION

PU-240 TOTAL CROSS SECTION

PU-240 ELASTIC CROSS SECTION

PU-240 TOTAL INELASTIC CROSS SECTION

PU-240 RADIATIVE CAPTURE CROSS SECTION

PU-240 TOTAL (n,2n) REACTION CROSS SECTION

PU-240 TOTAL (n,3n) CROSS SECTION

PU-240 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

PU-241 PROMPT GAMMA ENERGY PRODUCTION

PU-241 TOTAL CROSS SECTION

PU-241 ELASTIC CROSS SECTION

PU-241 TOTAL INELASTIC CROSS SECTION

PU-241 RADIATIVE CAPTURE CROSS SECTION

PU-241 TOTAL (n,2n) REACTION CROSS SECTION

PU-241 TOTAL (n,3n) CROSS SECTION

PU-241 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

PU-242 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-242 RECOMMENDED PROMPT KERMA FACTOR (solid line)
 REC. K. FACTOR WITH CH. PARTICLE DECAY HEAT (dashed line)
 REC. K. FACTOR PLUS TOTAL DECAY HEAT (chain dotted line)

PU-242 PROMPT GAMMA ENERGY PRODUCTION

PU-242 TOTAL CROSS SECTION

PU-242 ELASTIC CROSS SECTION

PU-242 TOTAL INELASTIC CROSS SECTION

PU-242 RADIATIVE CAPTURE CROSS SECTION

PU-242 TOTAL (n,2n) REACTION CROSS SECTION

PU-242 TOTAL (n,3n) CROSS SECTION

PU-242 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

AM-241 PROMPT GAMMA ENERGY PRODUCTION

AM-241 TOTAL CROSS SECTION

AM-241 ELASTIC CROSS SECTION

AM-241 TOTAL INELASTIC CROSS SECTION

AM-241 RADIATIVE CAPTURE CROSS SECTION

AM-241 TOTAL (n,2n) REACTION CROSS SECTION

AM-241 TOTAL (n,3n) CROSS SECTION

AM-241 NUMBER OF FISSION NEUTRONS $\times$ CROSS SECTION

AM-243 PROMPT GAMMA ENERGY PRODUCTION

AM-243 TOTAL CROSS SECTION

AM-243 ELASTIC CROSS SECTION

AM-243 TOTAL INELASTIC CROSS SECTION

AM-243 RADIATIVE CAPTURE CROSS SECTION

AM-243 TOTAL (n,2n) REACTION CROSS SECTION

AM-243 TOTAL (n,3n) CROSS SECTION

AM-243 NUMBER OF FISSION NEUTRONS × CROSS SECTION

REFERENCES:

- 1- R. Kinsey, "ENDF-102 Data Formats and Procedures for the Evaluated Nuclear Data File, ENDF" BNL-NCS-50496 (ENDF 102) (Oct. 1979), Revised by B. Magurno (Nov. 1983).
- 2- R. Little, and R. Seamon, "Negative Heating Numbers", Los Alamos National Laboratory Memorandum to P. Young, and E. Arthur, CSEWG Evaluations Committee. (June 17, 1981).
- 3- R. Little, and R. Seamon; Los Alamos National Laboratory, Radiation Transport Group, A Letter to J. Huang, University of Wisconsin-Madison (Nov. 1982).
- 4- Y. Farawila et al., "KAOS-V: A Program to Evaluate Neutron Kerma Factors and other Nuclear Responses" Argonne National Laboratory Report, ANL/FPP/TM-240 (1989).
- 5- R. MacFarlane, D. Muir, and R. Boicourt, "The NJOY Nuclear Data Processing System, Vol. I: User's manual" LA-9303-M, Vol. I (ENDF- 324).
- 6- R. MacFarlane, D. Muir, and R. Boicourt, "The NJOY Nuclear Data Processing System, Vol. II: The NJOY, RECONR, BROADR, HEATER, and THERMR Modules" LA-9303-M, Vol. II (ENDF-324).
- 7- W. W. Engle, Jr., "A User Manual for ANISN: A One Dimensional Discrete Ordinates Transport Code with Anisotropic Scattering" Oak Ridge National Laboratory Report K-1693 (1973).
- 8- R. W. Roussin et al., "Vitamin-E ^{174}N , ^{38}G General Purpose Cross-Section Library" Proceedings of the International Conference on Nuclear Data for Basic and Applied Science, Santa Fe, New Mexico (May 1985).
- 9- I. M. Bondarenko and Eh. E. Petrov, "Evaluation of Neutron- ^6Li and- ^7Li Interaction Cross-Sections for Calculating Kerma Factors" INDC(CCP)-265/L, (Dec. 1986).

- 10- P. DeLuca, Jr., H. Barschall, M. Burhoe, and R. Haight, "Carbon Kerma Factor for 18- and 20-MeV Neutrons" Nuclear Science and Engineering: 94,192-198 (1986).
- 11- P. DeLuca, Jr., H. Barschall, R. Haight, and J. McDonald, "Kerma Factor of Carbon for 14.1-MeV Neutrons" Radiation Research 100, 78-86 (1984).
- 12- R. Haight, S. Grimes, R. Johnson, and H. Barschall, "The $^{12}\text{C}(n,\alpha)$ Reaction and the Kerma Factor for Carbon at $E_n=14.1$ MeV" Nuclear Science and Engineering: 87,41-47 (1984).
- 13- P. DeLuca, Jr., H. Barschall, R. Haight, and J. McDonald, "Measured Neutron Carbon Kerma Factors from 14.1 MeV to 18 MeV" 1984, Proc. 5th Symp. Neutron Dosimetry Munich/Neuherberg, FRG, EUR-9762, p.193, Commission of the European Communities, Luxembourg (1985).
- 14- B. Antolokovic and J. Hudomalj, " 3α Decay of the 10.84 and 11.83 MeV States of ^{12}C " Nuclear Physics A237 (1975) 253-259.
- 15- G. M. Frye, Jr., L. Rosen, and L. Stewart, "Disintegration of Carbon into Three Alpha Particles by 12-20 Mev Neutrons" Physical Review Vol. 99, No. 5 (Sept. 1955).
- 16- B. Antolkovic and Z. Dolenec, "The Neutron-Induced $^{12}\text{C}(n,n')3\alpha$ Reaction at 14.4 MeV in a Kinematically Complete Experiment" Nuclear Physics A237 (1975) 235-252
- 17- E. A. Davis et al., "Disintegration of O^{16} and C^{12} by Fast Neutrons" Nuclear Physics 48 (1963) 169-175.
- 18- R. A. Al-Kital and R. A. Peck, Jr., " $\text{C}^{12}(n,\alpha)\text{Be}^9$ Reaction Induced by 14-MeV Neutrons" Physical Review Vol. 130, No. 4 (May 1963) 1500- 1503.
- 19- A. W. Obst, T.B. Grandy, and J.L. Weil, "Reaction $^9\text{Be}(\alpha,n)^{12}\text{C}$ from 1.7 to 6.4 MeV" Physical Review C, Vol.5 No.3 (March 1972).

- 20- L. Van der Zwan and K.W. Geiger, "The $^9\text{Be}(\alpha, n)^{12}\text{C}$ Cross Section between 1.5 and 7.8 MeV" Nuclear Physics A152 (1970) 481-494.
- 21- V. Verbinski et al., "Neutrons from $^9\text{Be}(\alpha, n)$ Reaction for E_α between 6 and 10 MeV" Physical Review Vol.170, No.4 (June 1968) 916- 923.
- 22- H. Schumacher, H. Menzel, and J. Coyne, "Critical Evaluation of Neutron Kerma Factors Using Theoretical and Experimental Ionization Yield Spectra" 1984, Proc. 5th Symp. Neutron Dosimetry Munich/Neuherberg, FRG, EUR-9762, p. 213, Commission of the European Communities, Luxembourg (1985).
- 23- R. MacFarlane, "Energy Balance of ENDF/B-V.2" Prepared for the meeting of CSEWG at BNL (May 1984).
- 24- C. Lederer, and V. Shirley, Eds "Table of Isotopes", 7th ed. (Wiley, New York, 1978).
- 25- C. Fu., Oak Ridge National Laboratory, Private Communications.
- 26- Yu. E. Kozy and G. A. Prokopets, "The Reactions $(n, n'\gamma)$, (n, np) , and (n, pn) in ^{23}Na , ^{27}Al , and ^{32}S at a Neutron Energy 14.6 MeV" Sov. J. Nucl. Phys. 28(1), (July 1978).
- 27- R. J. Howerton and M. H. MacGregor, "The LLL Evaluated Nuclear Data Library (ENDL): Descriptions of Individual Evaluations for $Z=0-98$ " UCRL-50400 Vol.15 Part D Rev.1 (May 1978).
- 28- G. L. Morgan, " $\text{Mn}(n, x\gamma)$ Reaction Cross Section for Incident Neutron Energies between 0.2 and 20.0 MeV" ORNL/TM-5531 (Aug. 1976).
- 29- D. E. Cullen et al., "Supplemental Neutron-Induced Interactions ($Z \leq 35$): Graphical, Experimental Data" UCRL-50400 Vol. 8 Rev. 1 Part A (July 1976).

- 30- C. Philis et al., "Titanium-II: An Evaluated Nuclear Data File" ANL/NDM 28, (June 1977).
- 31- P. Guenther et al., "Fast Neutron Cross Sections of Vanadium and an Evaluated Neutronic File" ANL/NDM-24, (May 1977).
- 32- G. L. Morgan and E. Newman, "The $\text{Cr}(n,x\gamma)$ Reaction Cross Section for Incident Neutron Energies Between 0.2 and 20.0 MeV" ORNL-TM-5098 (ENDF-222), (Jan. 1976).
- 33- E. Arthur, and P. Young, "Evaluated Neutron-Induced Cross Sections for $^{54,56}\text{Fe}$ to 40 MeV" Los Alamos National Laboratory report LA-8626- MS (ENDF-304) (Dec. 1980).
- 34- B. Magurno, and P. Young, "ENDF-201 Supplement I - ENDF/B-V.2 Summary Documentations" BNL-NCS-17541 (ENDF-201) (Jan. 1985).
- 35- D. M. Hetrick, C. Y. Fu, and D. C. Larson, "Calculated Neutron- Induced Cross Sections for $^{63,65}\text{Cu}$ from 1 to 20 MeV and Comparisons with Experiments" ORNL/TM-9083 (ENDF-337) (Aug. 1984).

Appendix A
DECAY ENERGY TABLES

The decay energies for all the materials in the KAOS/LIB-V library are listed below. The materials that do not produce decay heat with half-life within the 1 day cutoff are not listed. For single-isotope materials, charged particles decay energy $E_d(\text{ch})$ and gamma decay energy $E_d(\gamma)$ are listed. The producing reactions are identified using the ENDF/B-V MT number. For natural elements, the radioactive reaction products are also identified.

Single-Isotope Materials

	MT	$E_d(\text{ch})$	$E_d(\gamma)$
3-Li-6			
	103	1.568	0.0
3-Li-7			
	102	9.316	0.0
	104	1.568	0.0
4-Be-9			
	103	6.177	0.0
	104	9.316	0.0
	107	1.568	0.0

	MT	$E_d(\text{ch})$	$E_d(\gamma)$
5-B-11			
	102	6.3405	0.05771
	103	4.825	1.351
	107	9.3031	0.0
6-C-12			
	103	6.4	0.0
7-N-14			
	16	0.4794	1.022
8-O-16			
	103	2.664	4.605
9-F-19			
	16	0.2456	0.9903
	102	2.481	1.6326
	103	1.701	1.008
	107	2.664	4.605
11-Na-23			
	102	0.5626	4.1228
	103	1.9011	0.14664
	107	2.4808	1.6326
13-Al-27			
	102	1.2315	1.7789
	103	0.6884	0.8842
	107	0.56265	4.1228

	MT	$E_d(\text{ch})$	$E_d(\gamma)$
15-P-31			
	16	1.44	1.022
	103	0.589	0.0
	107	1.2315	1.7789
25-Mn-55			
	102	0.833	2.047
	103	1.093	0.0
27-Co-59			
	106	1.1016	0.0655
	107	0.833	2.047
40-Zr-92			
	103	1.48	0.246
40-Zr-94			
	103	1.9	0.6
	107	0.615	1.124
40-Zr-96			
	102	1.184	1.532
	107	1.74	1.19
72-Hf-178			
	103	0.8584	0.0
72-Hf-179			
	103	0.4877	0.214
72-Hf-180			
	103	0.5652	1.62

	MT	$E_d(\text{ch})$	$E_d(\gamma)$
74-W-184			
	103	0.385	2.481
74-W-186			
	28	0.677	0.298
	102	0.521	0.2727
	103	1.71	0.885
	107	0.4684	0.722
83-Bi-209			
	107	0.5845	0.0

Natural Elements Decay Data

	MT	Product	$E_d(\text{ch})$	$E_d(\gamma)$
12-Mg-nat				
	16	Mg-23	1.334	1.06
	28	Na-24	0.5626	4.1228
		Na-25	1.494	0.4244
	102	Mg-27	0.6884	0.8842
	103	Na-24	0.5626	4.1228
		Na-25	1.494	0.4244
		Na-26	3.3565	2.176
	107	Ne-23	1.9011	0.14664
14-Si-nat				
	103	Al-28	1.255	1.779

	MT	Ed(ch)	Ed(γ)	
16-S-nat				
	16	S-31	2.00	1.022
	103	P-34	2.3	0.319
	105	P-30	1.44	1.022
	107	Si-31	0.589	0.0
	111	Si-31	0.589	0.0
17-Cl-nat				
	16	Cl-34	1.524	1.022
	102	Cl-38	1.51	1.419
	103	S-37	0.780	2.9295
	107	P-34	2.30	0.319
19-K-nat				
	16	K-38	1.2115	3.1896
	102	K-42	1.4185	0.2744
	103	Ar-41	0.4492	1.294
	107	Cl-38	1.51	1.419
20-Ca-nat				
	16	Ca-39	2.58	1.022
	105	K-38	1.2115	3.1896
22-Ti-nat				
	16	Ti-45	0.3785	0.879
	28	Sc-49	0.8292	0.0
	102	Ti-51	0.8678	0.3688
	103	Sc-49	0.8292	0.0
		Sc-50	1.658	3.125
	104	Sc-49	0.8292	0.0
	105	Sc-44	0.6029	2.07

MT	Product	$E_d(\text{ch})$	$E_d(\gamma)$
23-V-nat			
102	V-52	1.0931	1.434
103	Ti-51	0.8673	0.3687
24-Cr-nat			
16	Cr-49	0.645	1.127
28	V-52	1.093	1.434
	V-53	0.998	1.037
102	Cr-55	1.12	0.0
103	V-52	1.093	1.434
	V-53	0.998	1.037
	V-54	1.313	3.944
104	V-52	1.093	1.434
	V-53	0.998	1.037
105	V-52	1.093	1.434
106	Ti-51	0.867	0.369
	Ti-52	0.72	0.142
107	Ti-51	0.867	0.369
26-Fe-nat			
16	Fe-53	1.153	1.1808
103	Mn-56	0.833	2.047
	Mn-58	2.907	0.0
107	Cr-55	1.094	0.0
28-Ni-nat			
28	Co-61	0.4522	0.09163
102	Ni-65	0.6364	0.5427
103	Co-61	0.4522	0.09163
	Co-62	1.664	1.512
	Co-64	3.329	0.181
107	Fe-61	1.581	1.129

	MT	Product	$E_d(\text{ch})$	$E_d(\gamma)$
29-Cu-nat				
	16	Cu-62	1.288	0.9995
		Cu-64	0.1297	0.2053
	22	Co-61	0.5816	0.09771
	102	Cu-64	0.1297	0.2053
		Cu-66	1.071	0.08315
	103	Ni-65	0.6364	0.5413
	106	Co-61	0.5816	0.09771
		Co-63	1.549	0.1362
	107	Co-62	1.664	1.512
42-Mo-nat				
	16	Mo-91	1.47	0.9576
	102	Mo-101	0.505	1.5
82-Pb-nat				
	16	Pb-203	0.0	0.3
	102	Pb-209	0.2095	0.0

Decay Energy for Non-fission Reactions in Fissionable Materials

Material	MT	$E_d(\text{ch})$	$E_d(\gamma)$
Th-232	102	0.603	0.0
Pa-233	102	0.291	1.646
U-238	102	0.6174	0.0
Pu-242	102	0.29	0.0
Am-241	102	0.2734	0.0
Am-243	102	0.72	0.0

Appendix B

Vitamin-E 174 Group Structure

group	Top Energy eV	group	Top Energy eV
1	1.00001E-01	31	5.82947E+02
2	4.13994E-01	32	7.48518E+02
3	5.31578E-01	33	9.61116E+02
4	6.82560E-01	34	1.23410E+03
5	8.76425E-01	35	1.58461E+03
6	1.12535E+00	36	2.03468E+03
7	1.44498E+00	37	2.24867E+03
8	1.85539E+00	38	2.48517E+03
9	2.38237E+00	39	2.61259E+03
10	3.05902E+00	40	2.74654E+03
11	3.92786E+00	41	3.03539E+03
12	5.04348E+00	42	3.35463E+03
13	6.47595E+00	43	3.70744E+03
14	8.31529E+00	44	4.30742E+03
15	1.06770E+01	45	5.53084E+03
16	1.37096E+01	46	7.10174E+03
17	1.76035E+01	47	9.11882E+03
18	2.26033E+01	48	1.05946E+04
19	2.90232E+01	49	1.17088E+04
20	3.72665E+01	50	1.50344E+04
21	4.78512E+01	51	1.93045E+04
22	6.14421E+01	52	2.18749E+04
23	7.88932E+01	53	2.35786E+04
24	1.01301E+02	54	2.41755E+04
25	1.30073E+02	55	2.47875E+04
26	1.67017E+02	56	2.60584E+04
27	2.14454E+02	57	2.70001E+04
28	2.75364E+02	58	2.85011E+04
29	3.53575E+02	59	3.18278E+04
30	4.53999E+02	60	3.43067E+04

Group	Top Energy eV	Group	Top Energy eV
61	4.08677E+04	91	2.97211E+05
62	4.63092E+04	92	2.98491E+05
63	5.24752E+04	93	3.01974E+05
64	5.65622E+04	94	3.33733E+05
65	6.73794E+04	95	3.68832E+05
66	7.19981E+04	96	3.87742E+05
67	7.94987E+04	97	4.07622E+05
68	8.25034E+04	98	4.50492E+05
69	8.65169E+04	99	4.97871E+05
70	9.80366E+04	100	5.23397E+05
71	1.11090E+05	101	5.50232E+05
72	1.16786E+05	102	5.78444E+05
73	1.22773E+05	103	6.08101E+05
74	1.29068E+05	104	6.39279E+05
75	1.35686E+05	105	6.72055E+05
76	1.42642E+05	106	7.06512E+05
77	1.49956E+05	107	7.42736E+05
78	1.57644E+05	108	7.80817E+05
79	1.65727E+05	109	8.20850E+05
80	1.74224E+05	110	8.62936E+05
81	1.83156E+05	111	9.07180E+05
82	1.92547E+05	112	9.61640E+05
83	2.02419E+05	113	1.00259E+06
84	2.12797E+05	114	1.10803E+06
85	2.23708E+05	115	1.16484E+06
86	2.35177E+05	116	1.22456E+06
87	2.47235E+05	117	1.28735E+06
88	2.73237E+05	118	1.35335E+06
89	2.87247E+05	119	1.42274E+06
90	2.94518E+05	120	1.49569E+06

Group	Top Energy eV	Group	Top Energy eV
121	1.57237E+06	151	6.70320E+06
122	1.65299E+06	152	7.04688E+06
123	1.73774E+06	153	7.40818E+06
124	1.82684E+06	154	7.78801E+06
125	1.92050E+06	155	8.18731E+06
126	2.01897E+06	156	8.60708E+06
127	2.12248E+06	157	9.04837E+06
128	2.23130E+06	158	9.51229E+06
129	2.30686E+06	159	1.00000E+07
130	2.34570E+06	160	1.05127E+07
131	2.36525E+06	161	1.10517E+07
132	2.38521E+06	162	1.16183E+07
133	2.46597E+06	163	1.22140E+07
134	2.59240E+06	164	1.25232E+07
135	2.72532E+06	165	1.34986E+07
136	2.86505E+06	166	1.38403E+07
137	3.01194E+06	167	1.41907E+07
138	3.16637E+06	168	1.45499E+07
139	3.32871E+06	169	1.49183E+07
140	3.67879E+06	170	1.56831E+07
141	4.06570E+06	171	1.64872E+07
142	4.49329E+06	172	1.69046E+07
143	4.72367E+06	173	1.73325E+07
144	4.96585E+06	174	1.96403E+07
145	5.22046E+06		
146	5.48812E+06		
147	5.76950E+06		
148	6.06531E+06		
149	6.37628E+06		
150	6.59238E+06		

DISTRIBUTION LIST FOR ANL/FPP/TM-241

Internal:

C. Adams	D. Gruen	C. Reed
M. Billone	P. Guenther	A. Smith
R. Blomquist	A. Hassanein	D. Smith
J. Brooks	T. Hua	D. Sze
Y. Chang	C. Johnson	C. Till
R. Clemmer	A. Krauss	L. Turner
D. Ehst	R. Lawson	S. Yang
K. Evans	L. LeSage	FPP Files (25)
P. Finn	B. Loomis	ANL Contract File
E. Fujita	S. Majumdar	ANL Libraries
E. Gelbard	R. Mattas	ANL Patent Dept.
Y. Gohar (10)	B. Picologlou	TIS Files (3)
L. Greenwood	K. Porges	

External:

DOE-OSTI, for distribution per UC-420 (37)
Manager, Chicago Operations Office

U.S. Department of Energy, Office of Fusion Energy

S. Berk
M. Cohen
R.J. Dowling
G. Haas
G.R. Nardella
A.L. Opdenaker
R. Price
T. Reuther
S. Staten

U.S. Department of Energy, Division of Nuclear Physics

S.L. Whetstone

Battelle Pacific Northwest Laboratory

L. Carter
D.G. Doran
G. Hollenberg

Brookhaven National Laboratory

S. Baron
M.R. Bhat
M. Divadeenam
C.L. Dunford
N.E. Holden
R.R. Kinsey
V. McLane
S.F. Mughabghab
S. Pearlstein
P.F. Rose
J.R. Tuli

Central Institute of Physics, Bucharest
V. Zoita

Columbia University
R.A. Gross

Commission of the European Communities
G. Casini
J. Darvis
C. Ponti
D. Wattecamps

Culham Laboratory
G.J. Butterworth

Duke University
R. Walter

EG&G Idaho, Inc.
J. Crocker
D. Holland
S. Piet
C.W. Reich
J. Smith

Fusion Engineering Design Center
C. Flanagan
T. Shannon

Fusion Power Associates
S. Dean

GA Technologies, Inc.
E. Cheng
K. Schultz

Georgia Institute of Technology
Y. Farawila (10)
W. Stacey, Jr.

Grumman Space Systems
L. Deutsch
J. Rathke
M. Stauber
A. Tobin

I.V. Kurchatov Institute of Atomic Energy
D.V. Markovskiy
O. Schipakin
G. Shatalov
L.V. Tocheniy
S. Zimin

Japan Atomic Energy Research Institute

T. Kuroda
K. Maki
T. Nakamura
Y. Seki
T. Tomabechi

Kernforschungszentrum Karlsruhe

J.E. Vetter

Lawrence Berkeley Laboratory

B. Cooper
W. Lindquist

Lawrence Livermore National Laboratory

J. Doggett
J. Gilleland
M.W. Guinan
R. Haight
L. Hansen
C. Henning
R. Howerton
S. Iwasaki
J.D. Lee
G. Logan
J.R. Miller
L.J. Perkins
R. Moir
P.D. Soran
K.I. Thomassen

Los Alamos National Laboratory

R.E. Alcouffe
J. Anderson
E. Arthur
L. Booth
J. Briesmeister
F.W. Brinkley
R. Brown
F. Clinard
J.W. Davidson
D. Dudziak
J. Hendricks
N. Jarmie
R.J. Juzaitis
R. Krakowski
P. Lisowski
R.C. Little
R.J. Livak
R. MacFarlane
R. Miller
D. Muir
R.D. O'Dell
R.E. Seamon

W.L. Thompson
J.T. West III
P. Young

Massachusetts Institute of Technology

L. Bromberg
D. Cohn
M. Kazimi
D. Montgomery

Max-Planck Institute fur Plasmaphysik

J.L. Boutard
M. Chazalon
W. Daenner
H. Kranse
A. Knoblock
E. Salpietro
R. Toschi
K. Verschuur
G. Vieider

McDonnell Douglas Astronautics Company

J. Davis
D.A. DeFreece

National Bureau of Standards

R. Reed
O. Wasson

North Carolina State University

C. Gould

Oak Ridge National Laboratory

R.G. Alsmiller
C.C. Baker
V.C. Baker
L.A. Berry
B.L. Bishop
J.A. Blair
E.E. Bloom
D. Campbell
J. Cannon
C. Fu
T. Gabriel
G. Haste
D. Hetrick
D. Larson
R.A. Lillie
M.J. Martin
H. McCurdy
R. Roussin, RSIC
R.T. Santoro
J. Scott
J. Stout
D.W. Swain

Ohio University

S.M. Grimes
R.O. Lane

Ontario Hydro, Canada

B. Stasko

Osaka University

K. Sumita
A. Takahashi

Princeton Plasma Physics Laboratory

H. Furth
D. Jassby
L.P. Ku
R. Little
P.H. Rutherford
J. Schmidt

Purdue University

P. Lykoudis

Rensselaer Polytechnic Institute

M.J. Embrechts
D. Steiner

Sandia National Laboratories

W. Bauer
M.A. Sweeney

SIN, Villigen, Switzerland

C. Marinucci

TRW, Inc.

J. Garner
J. Gordon
J.A. Maniscalco

University of California-Los Angeles

M. Abdou
R. Conn
N. Ghoneim
Y. Ikeda
A. Kumar
F. Najmabadi
M.Z. Youssef

University of Illinois-Urbana

G.H. Miley
M.H. Ragheb

University of Michigan-Ann Arbor

G. Knoll

University of Missouri-Columbia

A. Bolon

University of Oregon

A. Klein

University of Tennessee-Knoxville

T.W. Kerlin

P. Stevens

University of Tokyo, Japan

Y. Oka

University of Washington-Seattle

F.L. Ribe

G.L. Woodruff

University of Wisconsin-Madison

G. Kulcinski

C. Maynard (10)

M. Sawan

W. Vogelsang

Westinghouse Hanford Company

F.M. Mann

R.J. Puigh